

June 6-8, 2018

Galaxy-Evolution Workshop 2018@Ehime Univ.

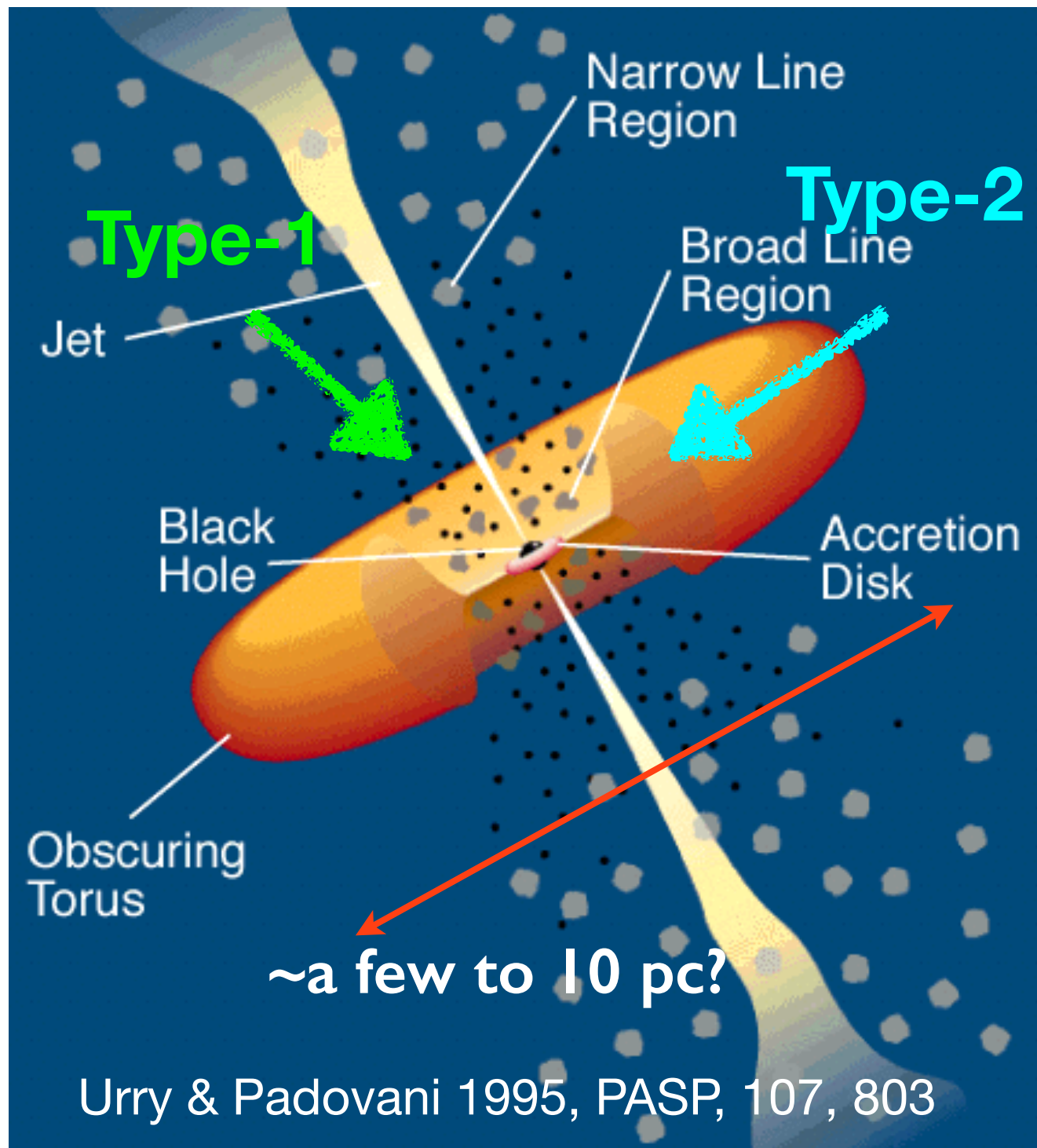
**The Circumnuclear *Multi-phase* Torus
in the Circinus Galaxy Revealed by ALMA**

→ Izumi et al. 2018c, to be submitted

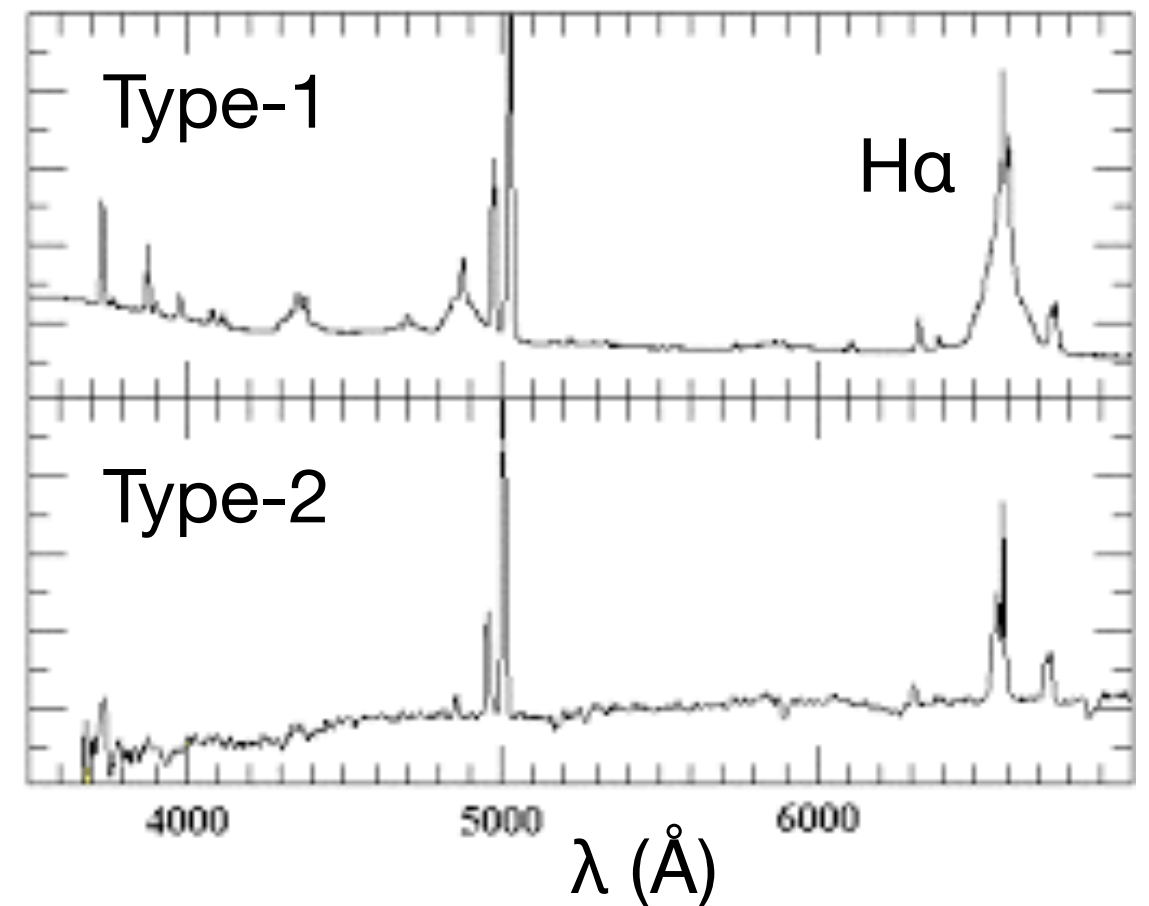
**Takuma Izumi(NAOJ Fellow),
K.Wada, R.Fukushige, S.Hamamura (Kagoshima Univ.),
K.Kohno (IoA/Univ. of Tokyo)**

SMBH obscuration: **Torus**

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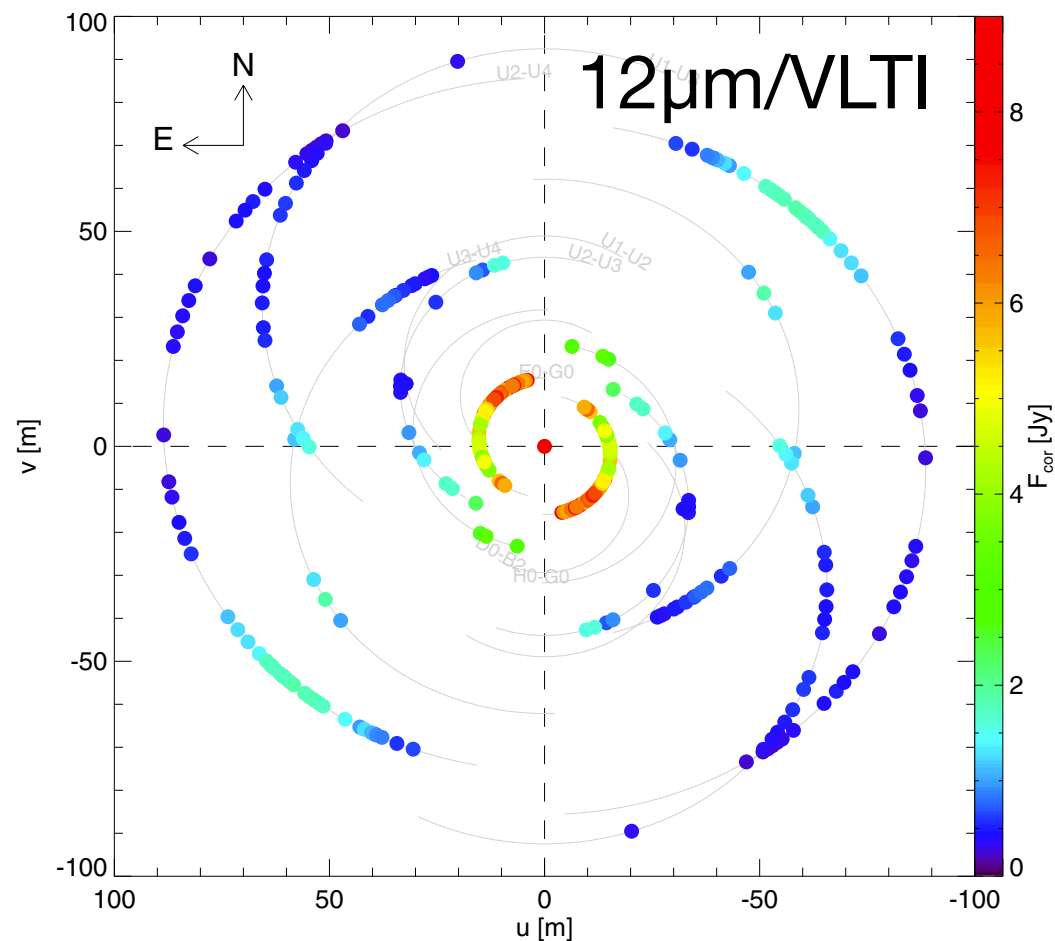
(Strict-) Unified scheme
(e.g., Antonucci 1993, ARA&A, 31, 473)



Q. Physical origin of the torus??

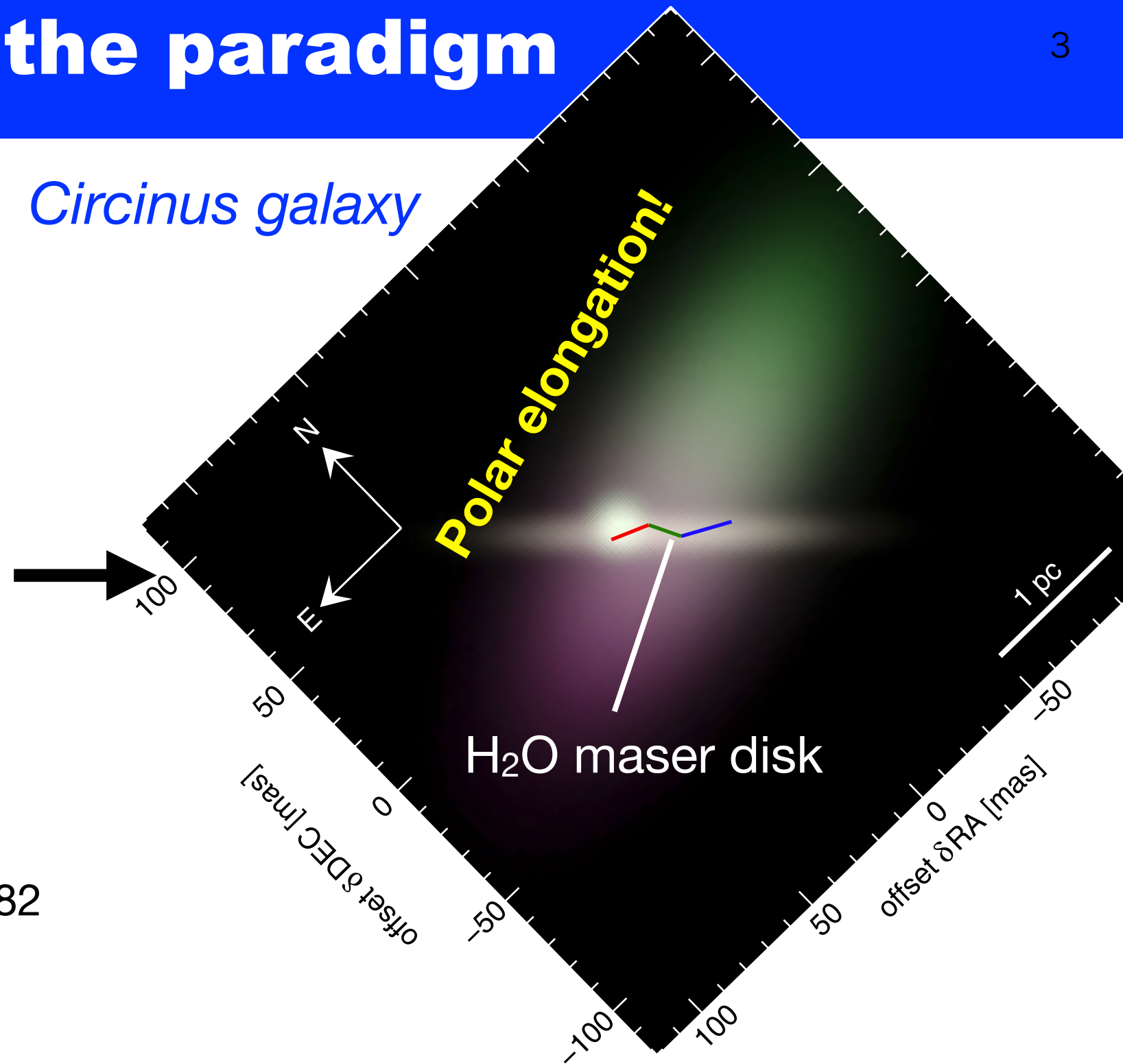
Big challenge to the paradigm

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Tristram et al. 2014, A&A, 563, A82

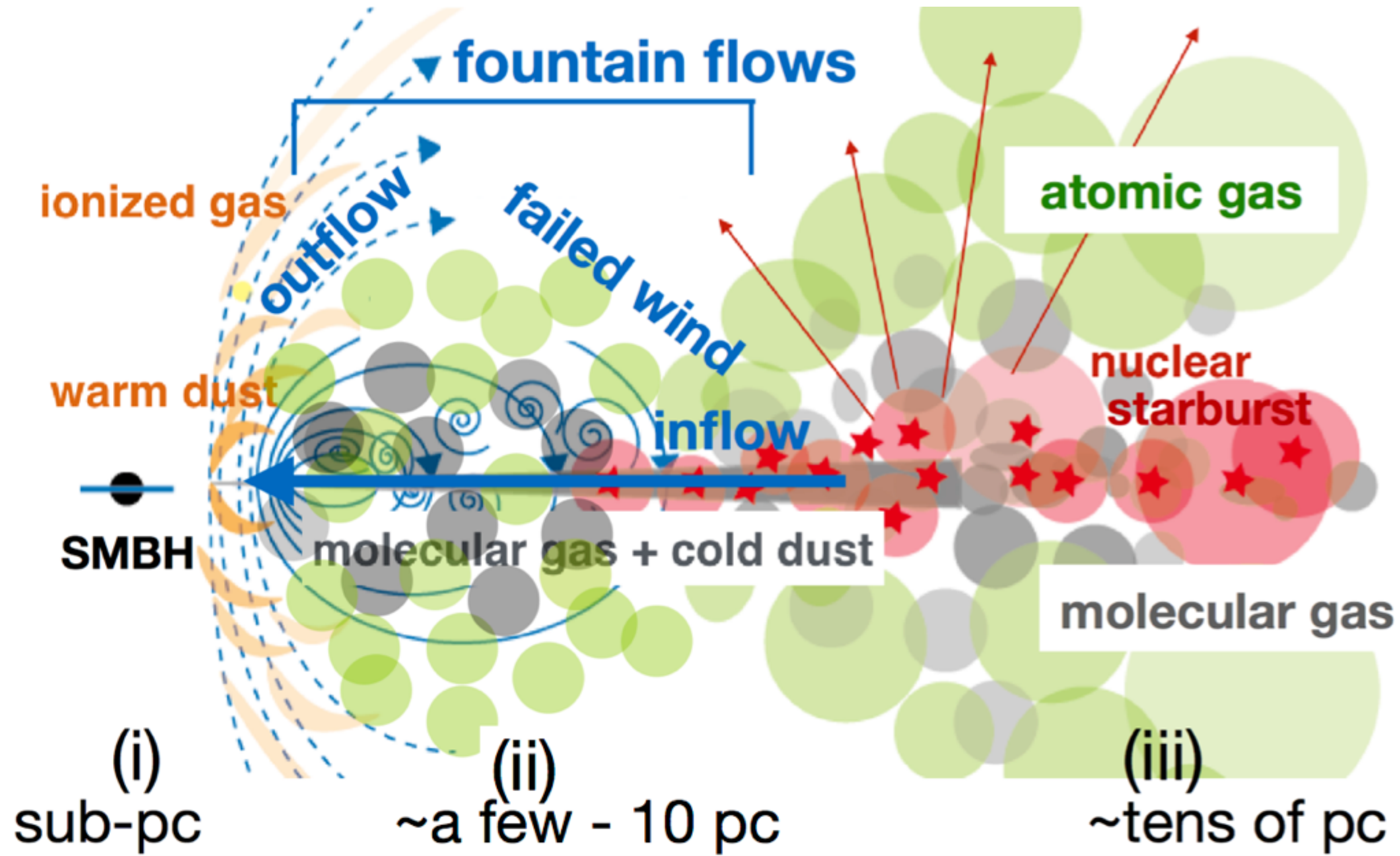
Circinus galaxy

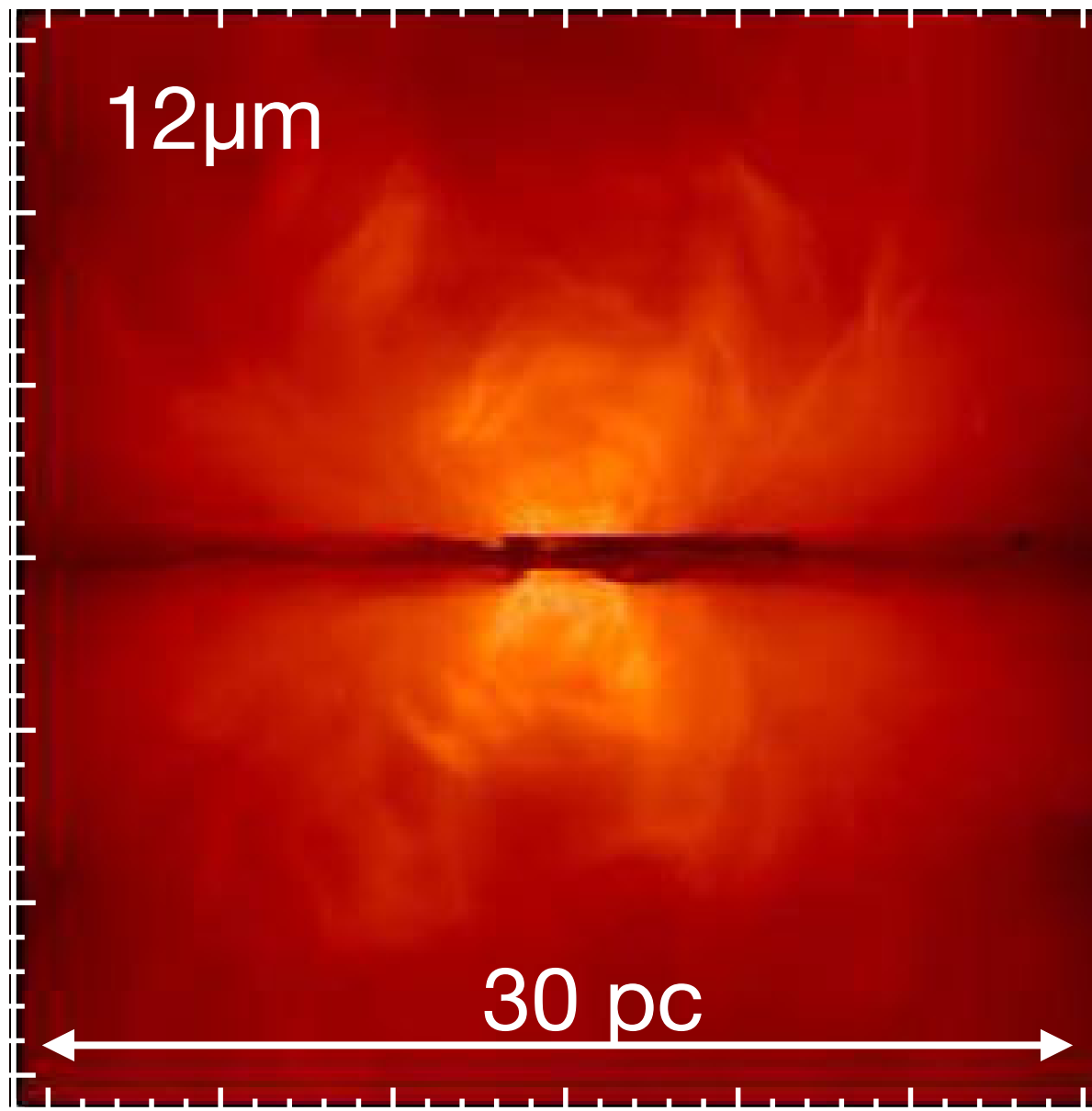


- Polar elongation in MIR continuum of the Circinus galaxy!?
- Statistical confirmation (e.g., Lopez-Gonzaga et al. 2016)

Multi-phase Dynamic Torus model

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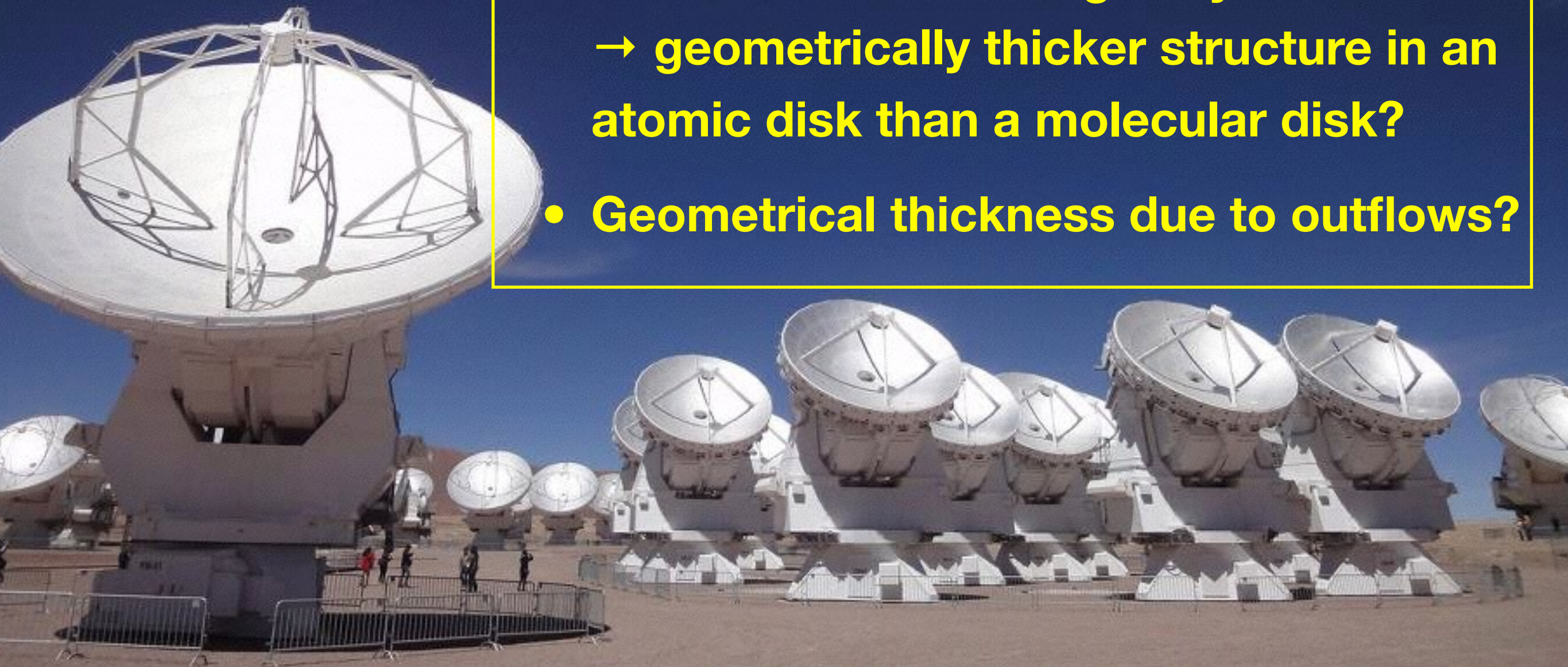


Schartmann et al. 2014,
MNRAS, 445, 3878

- Indeed, MIR polar elongation was reproduced
- IR SED is well-explained, too.
- Consistent with observations (morphological/photometric manner)

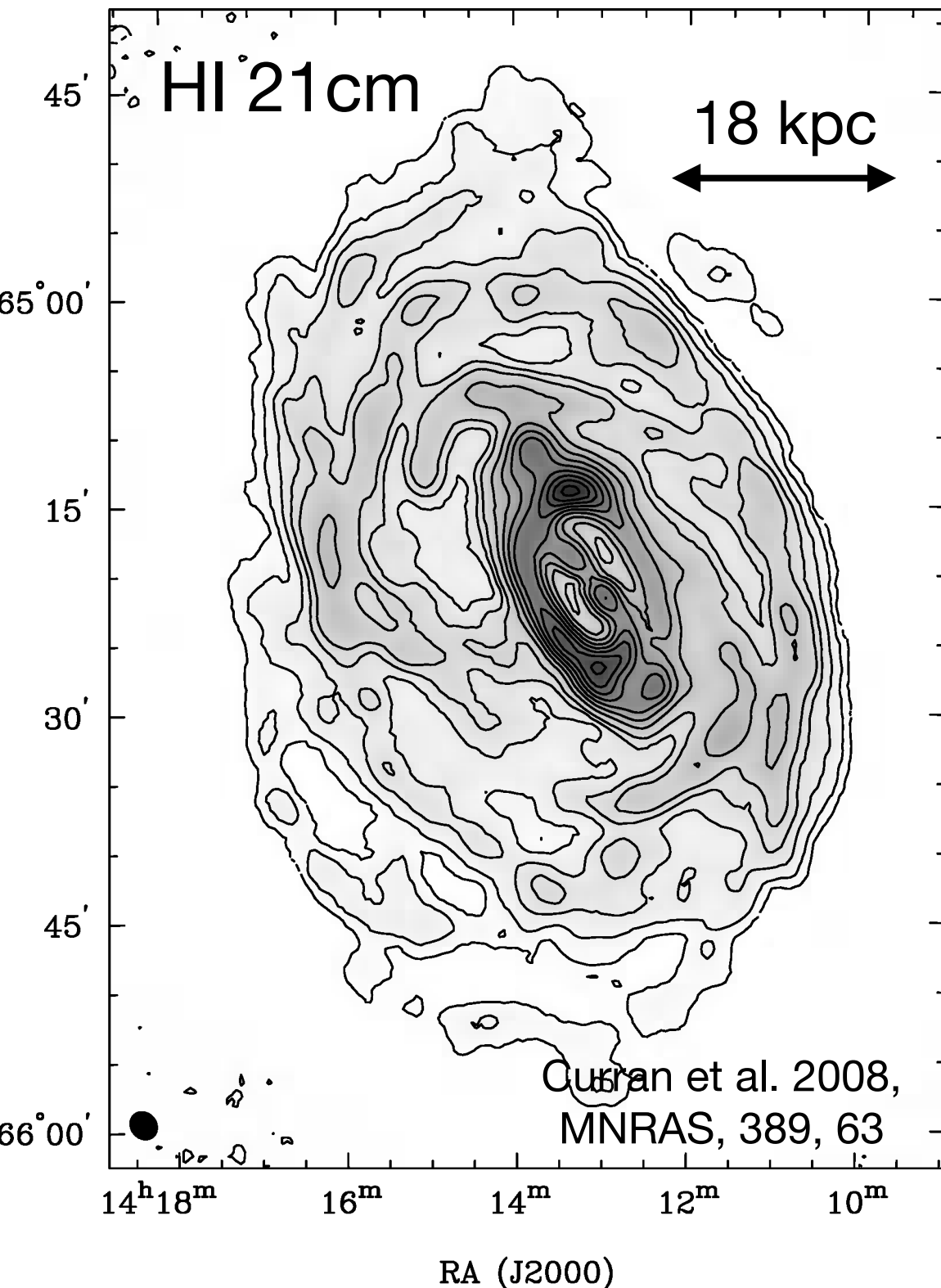
Out Study: Multi-phase circumnuclear obscuring structures studied with ALMA

- **Atomic vs molecular gas dynamics?**
→ geometrically thicker structure in an atomic disk than a molecular disk?
- **Geometrical thickness due to outflows?**



ALMA Cycle 4 Observations (Band 7 + 8) ⁷

The Circinus Galaxy (4.2 Mpc)

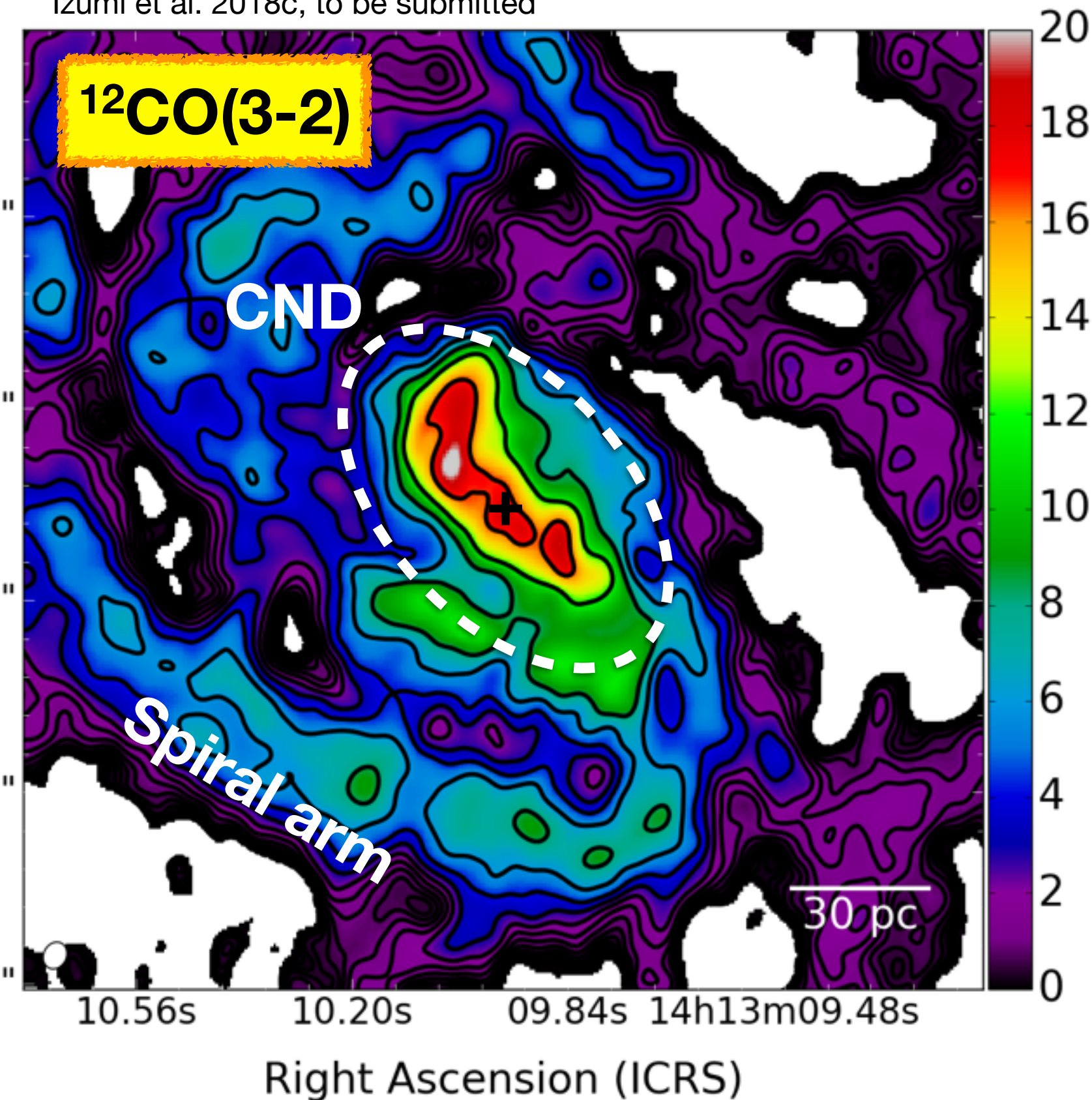


- Compton-thick AGN (Arevalo+2014)
 - $N_H \sim (6-10) \times 10^{24} \text{ cm}^{-2}$
 - $L_{2-10\text{keV}} \sim (2-5) \times 10^{42} \text{ erg/s}$
- Low SFR (i.e., weak SN-feedback)
→ **Test the fountain scheme**
- High resolution [CO\(3-2\)](#) + [\[CI\]\(1-0\)](#) in ALMA Cycle 4 (PI: T.Izumi)
→ molecular + atomic structures and their dynamical differences
- $\theta \sim 5 \text{ pc (CO) \& } \sim 15 \text{ pc (CI)}$

Molecular & Atomic gas distributions

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Izumi et al. 2018c, to be submitted

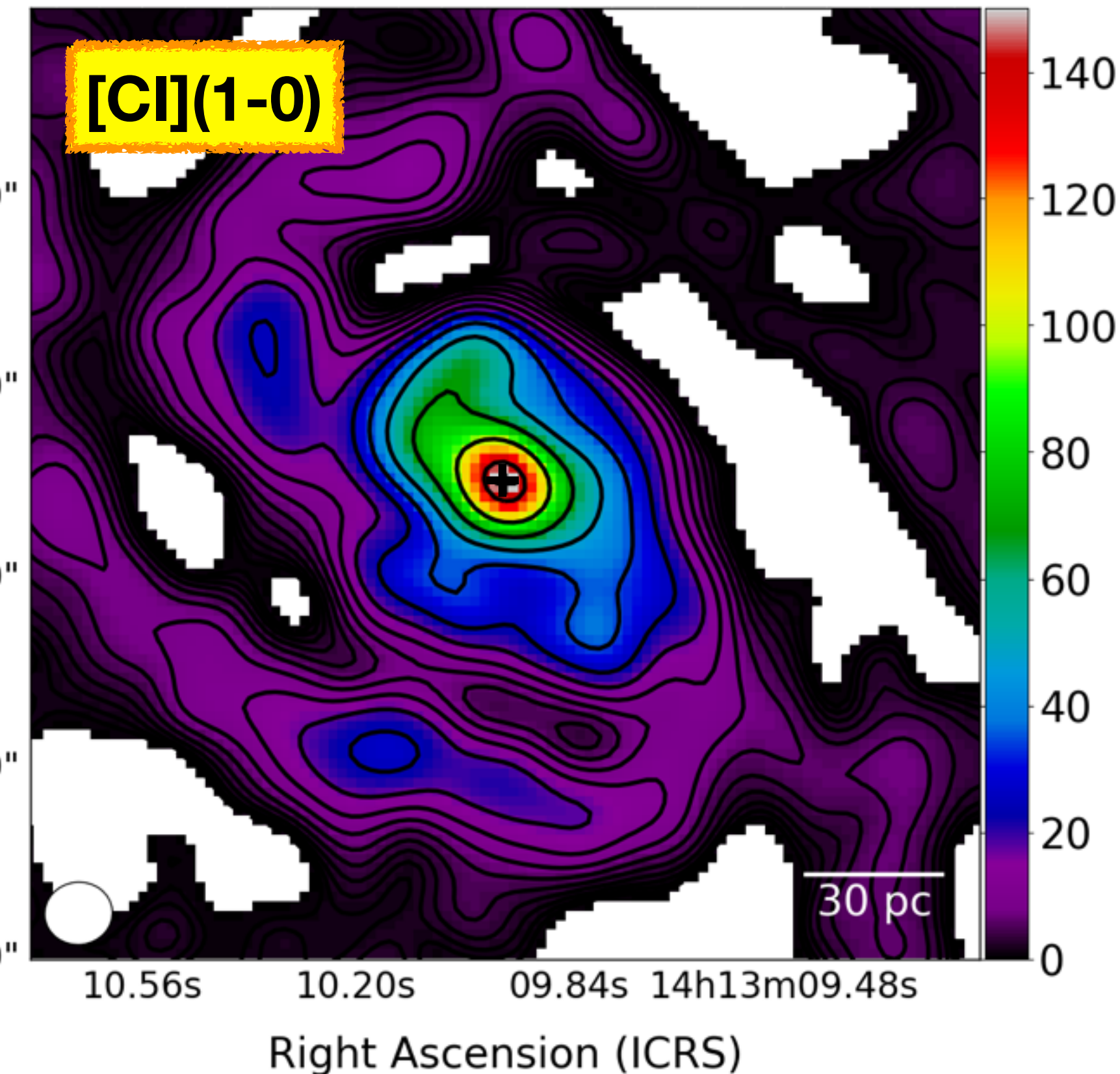


- CND ($D \sim 70$ pc) + Spirals
- **$M_{\text{H}_2} \sim 3 \times 10^6 M_{\text{sun}}$ (CND)**
 - c.f., $M_{\text{BH}} \sim 2 \times 10^6 M_{\text{sun}}$
- **$N_{\text{H}_2} \sim 5 \times 10^{23} \text{ cm}^{-2}$ (AGN)**
 - *beam-averaged*
 - N_{H} (Xray) $\sim 6 \times 10^{24} \text{ cm}^{-2}$
- **CND will be a significant nuclear obscurer!**

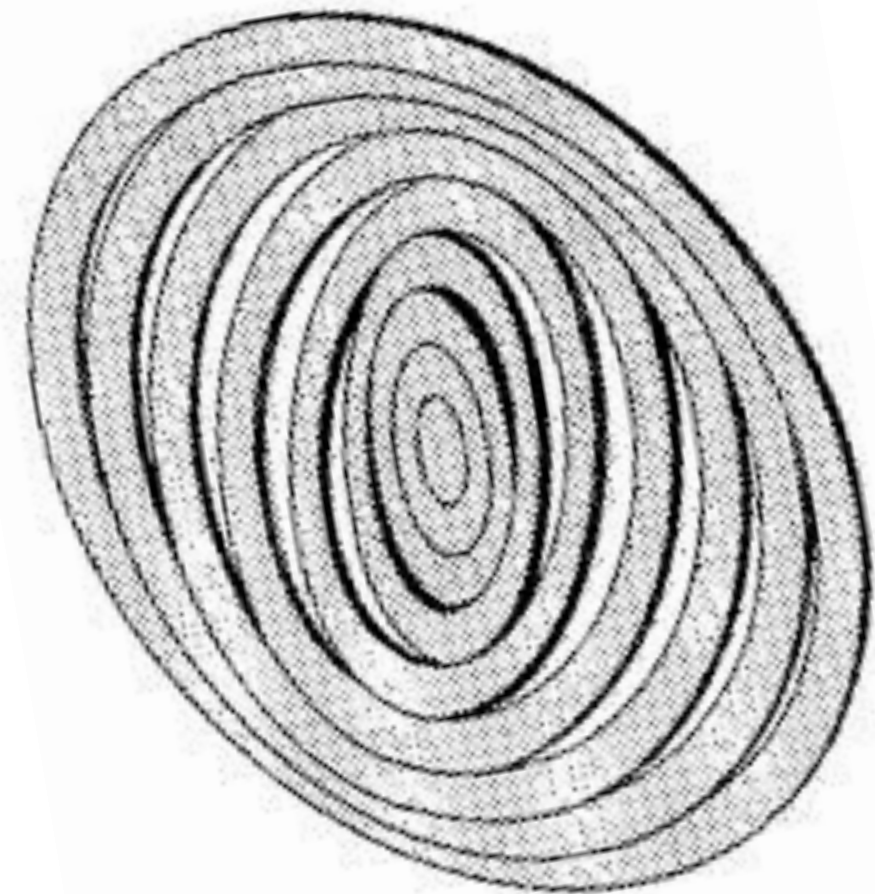
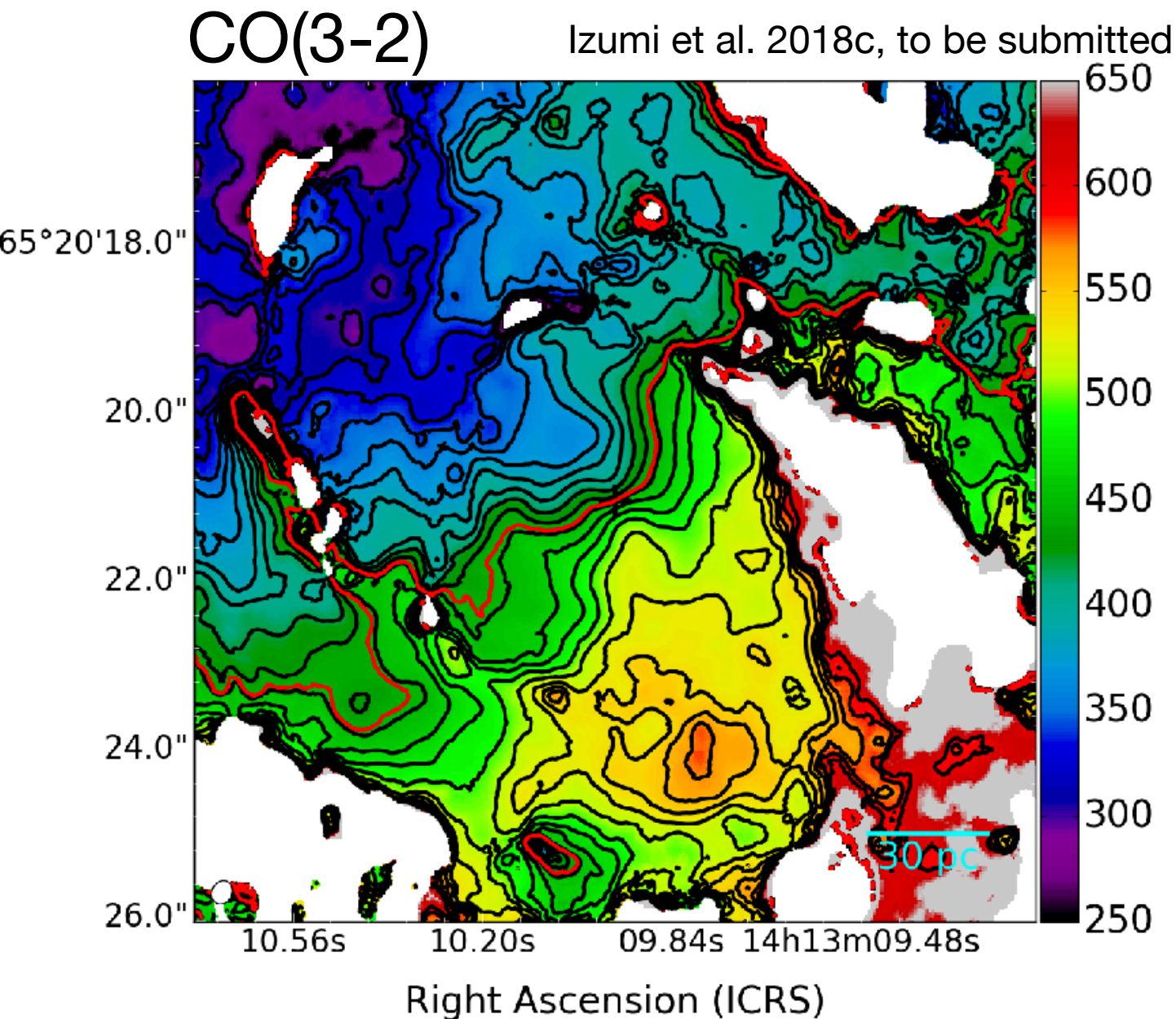
Molecular & Atomic gas distributions

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Izumi et al. 2018c, to be submitted



- *Similar 2D distribution to the CO(3-2)*
- Tracing the same volume???
- But we don't know 3D structures!

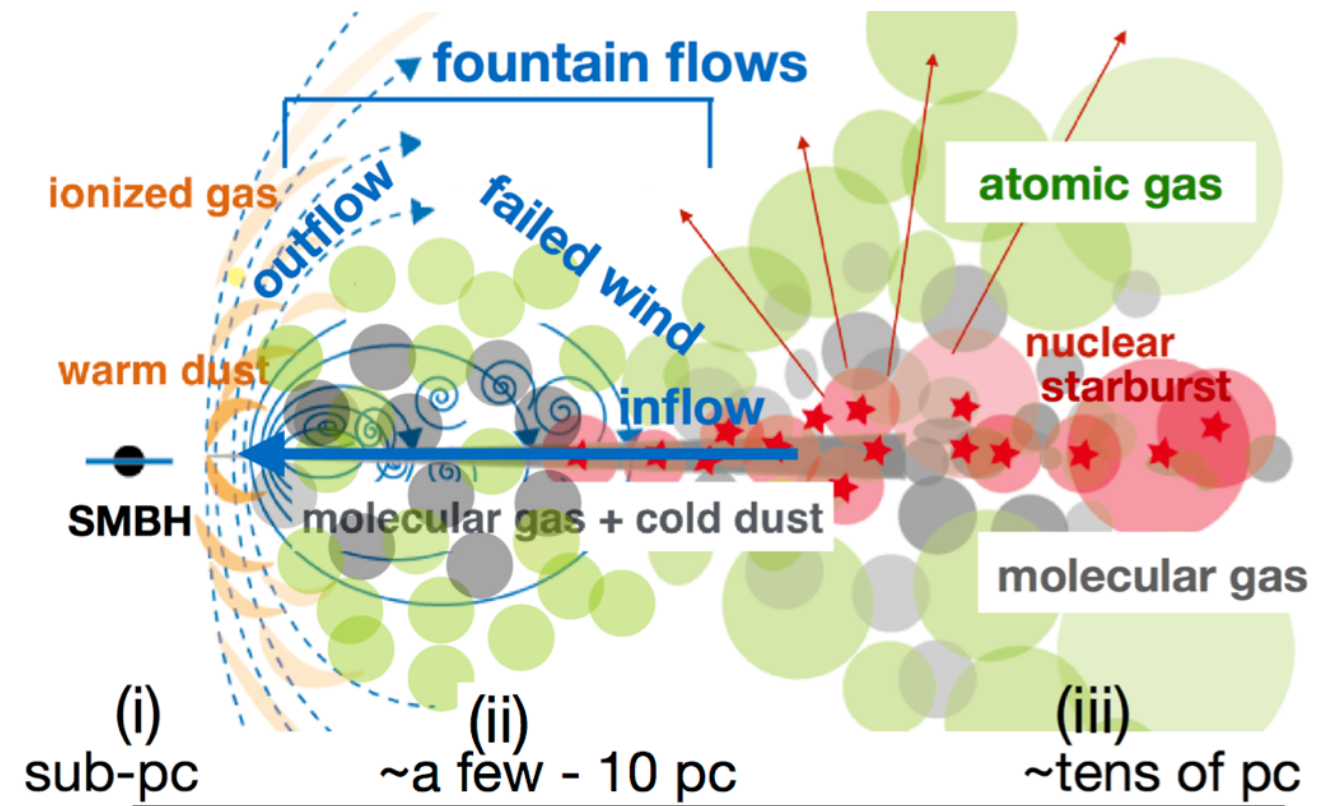
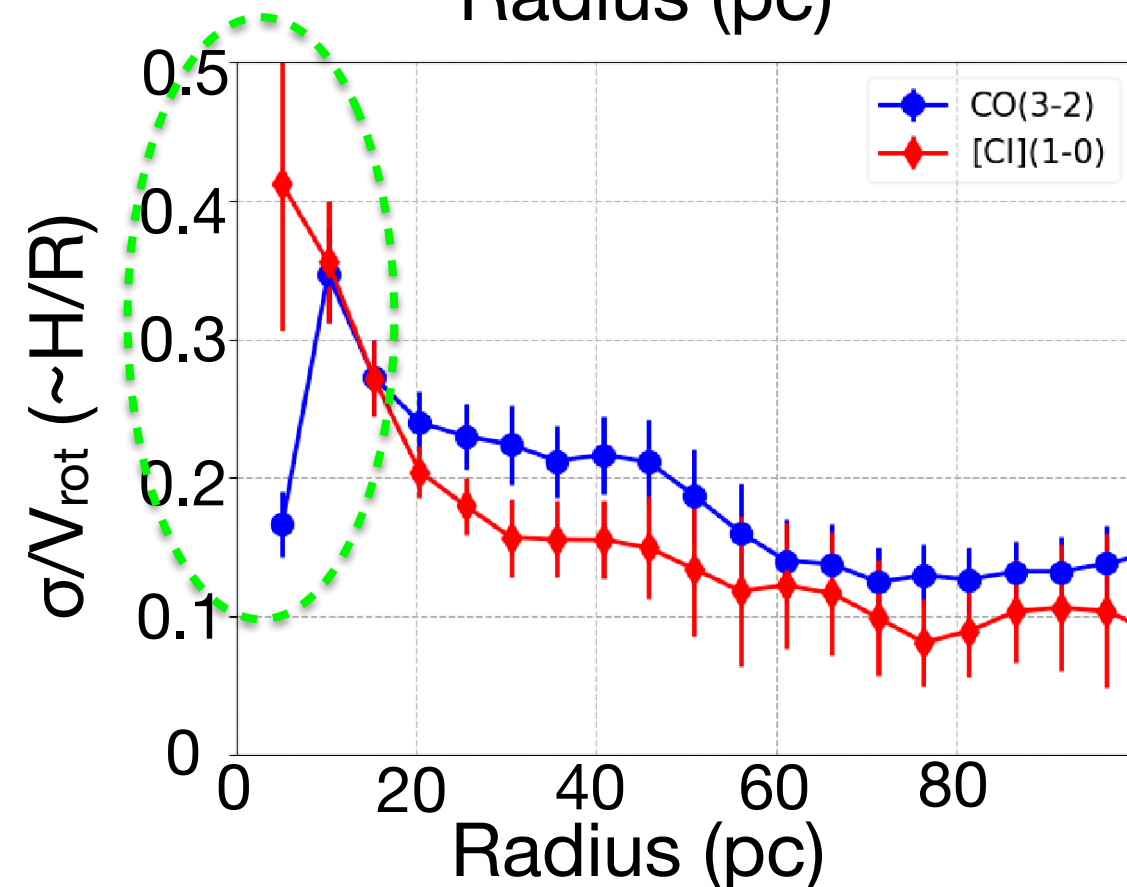
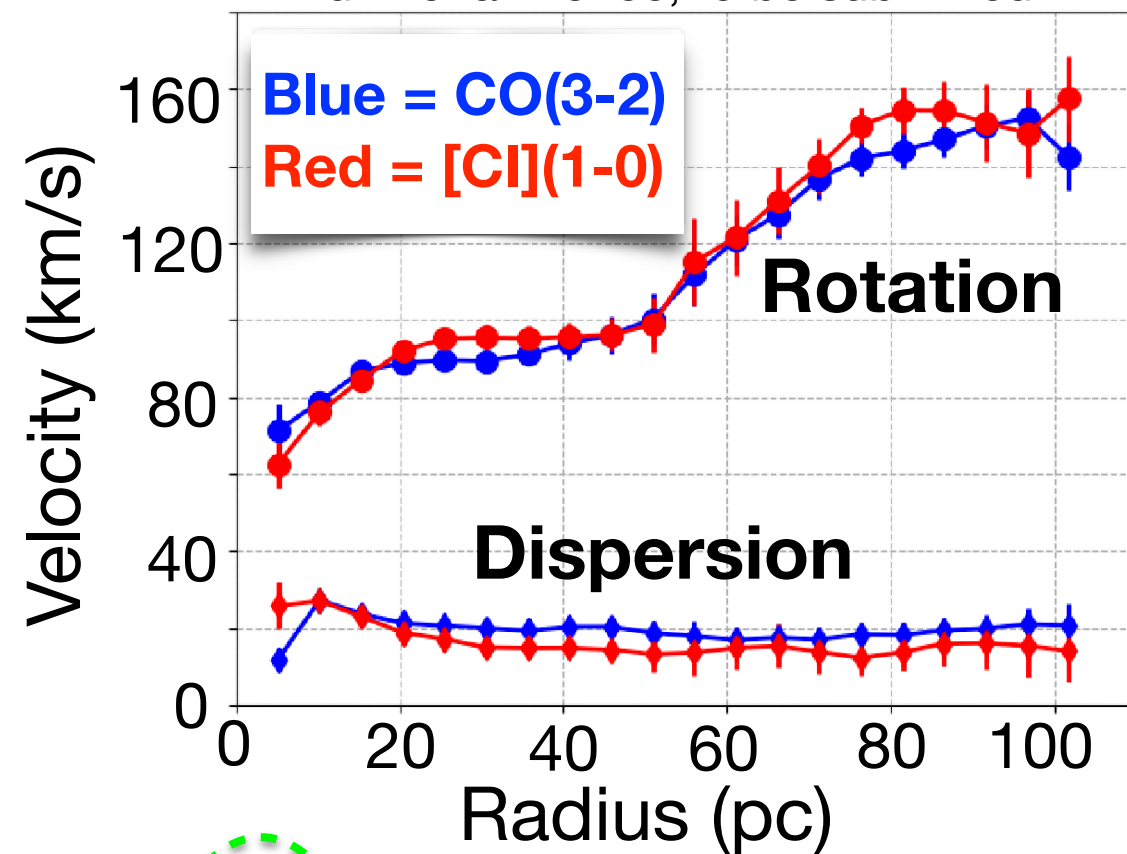


- Global motion is dominated by rotation
- We decomposed the dynamics with tilted-ring models
 - V_{rot} , σ , inclination, P.A.

CI vs CO: Multi-phase Torus Dynamics

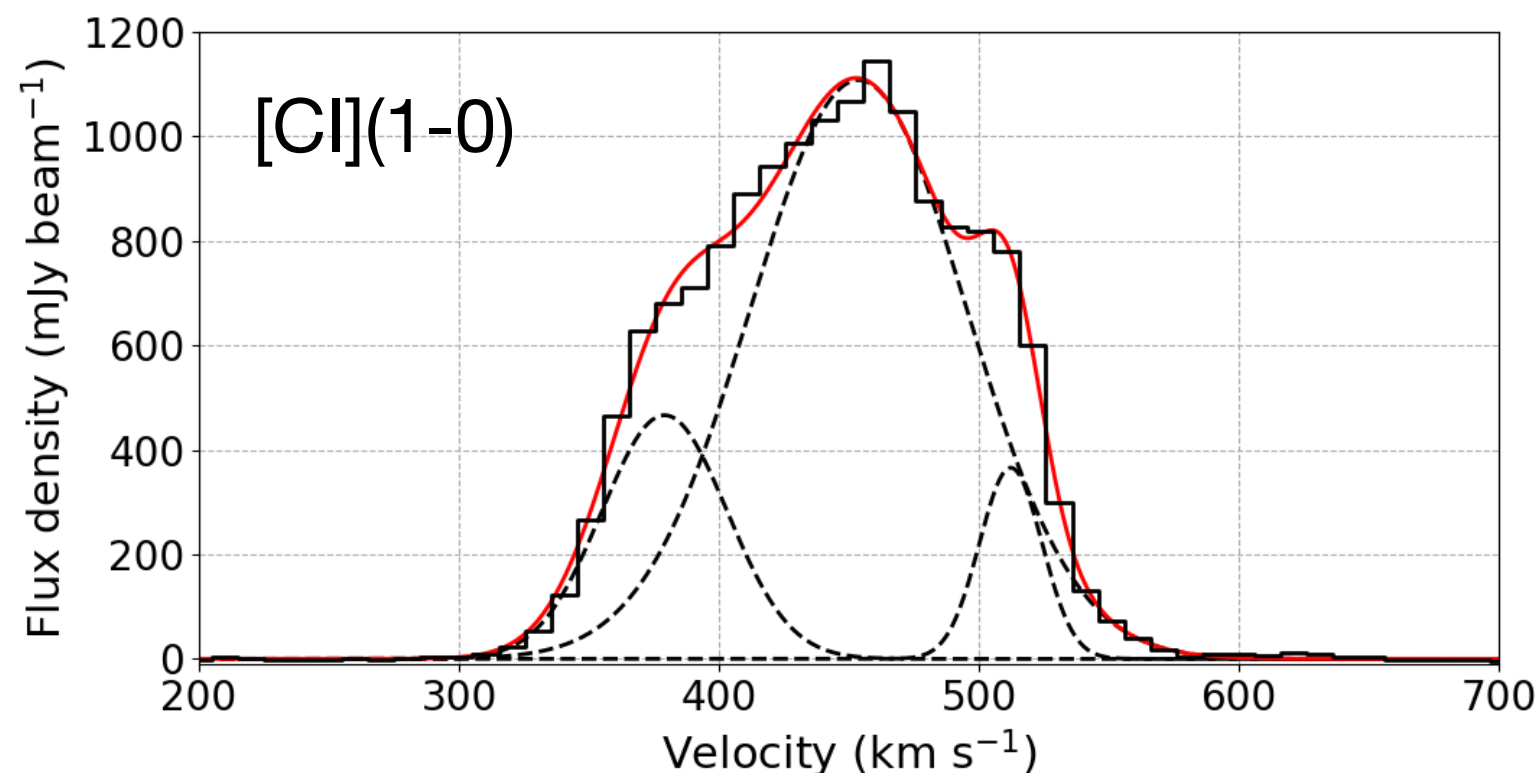
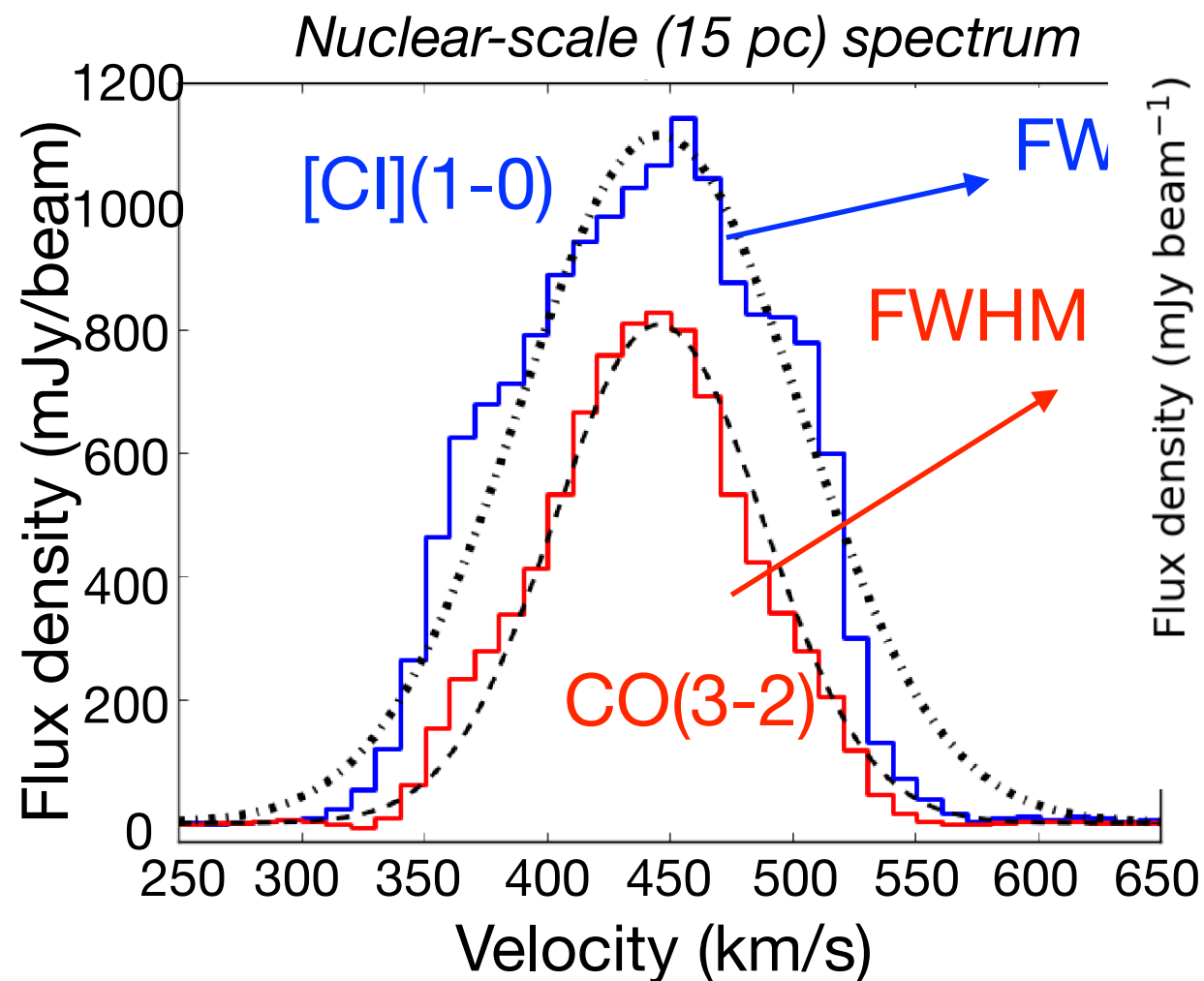
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Izumi et al. 2018c, to be submitted

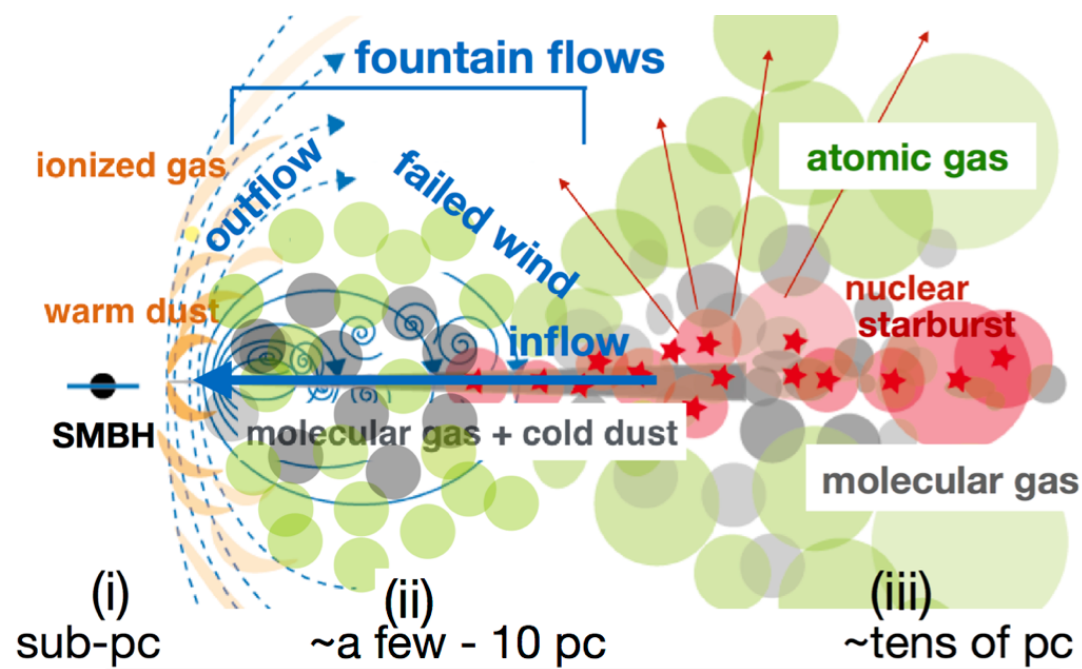


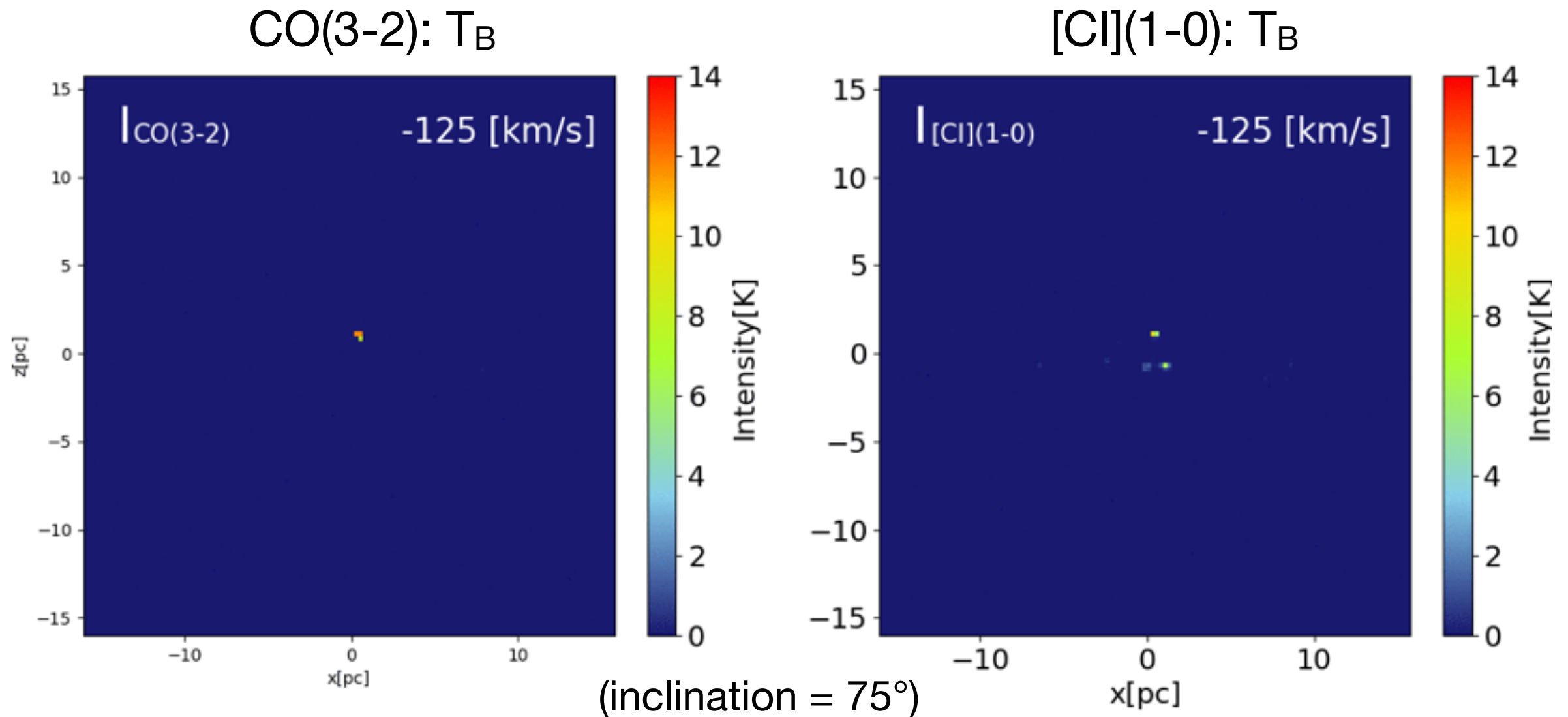
- Geometrically thin mol. & atomic disks at $r > 20$ pc (c.f., low-SFR in Circinus)
- **Geometrically thicker atomic disk than the dense mol. disk at $r < 10$ pc!**
→ **Multi-phase geometrical structures!**

Then, what makes the geometrical thickness? 12



- Nuclear [Cl](1-0) spectrum shows a larger FWHM than CO(3-2).
- Multi-Gaussians yielded a better fit
→ Almost symmetric emergence of the additional blue- and red-components
- **Coherent motion like outflows??**
 - consistent with the fountain model
 - will fall back to the disk (failed wind)

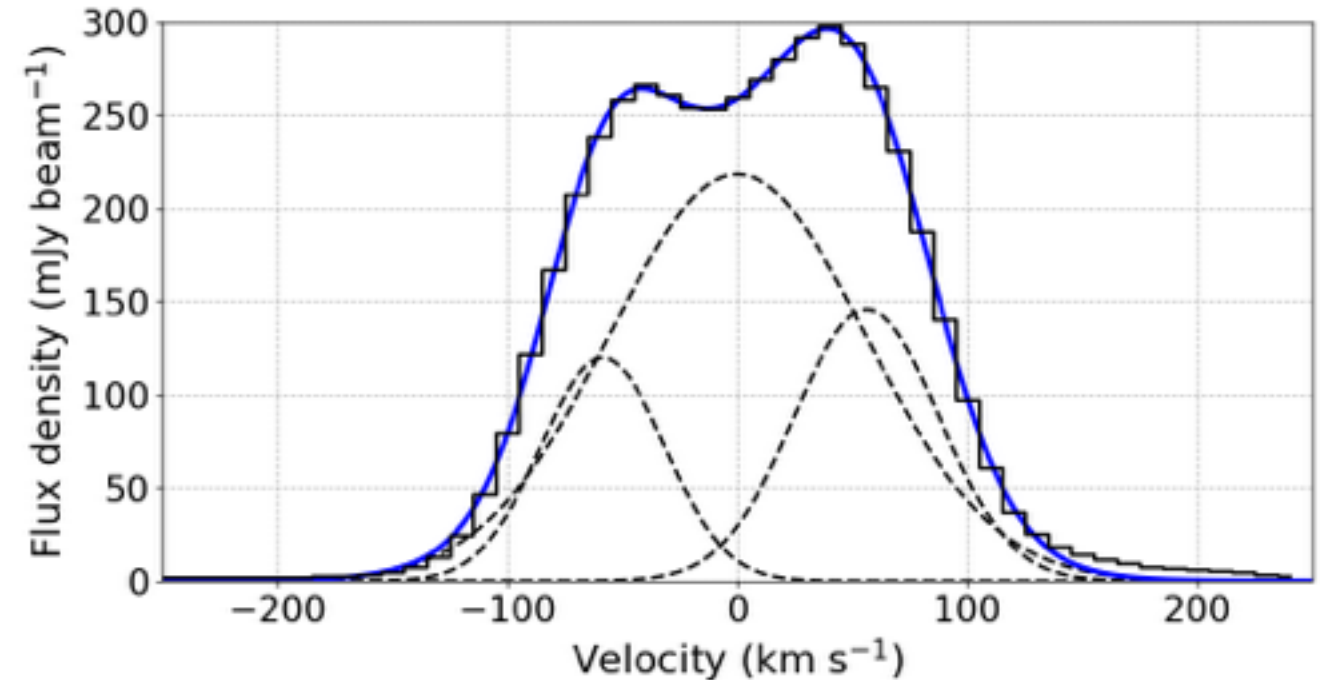
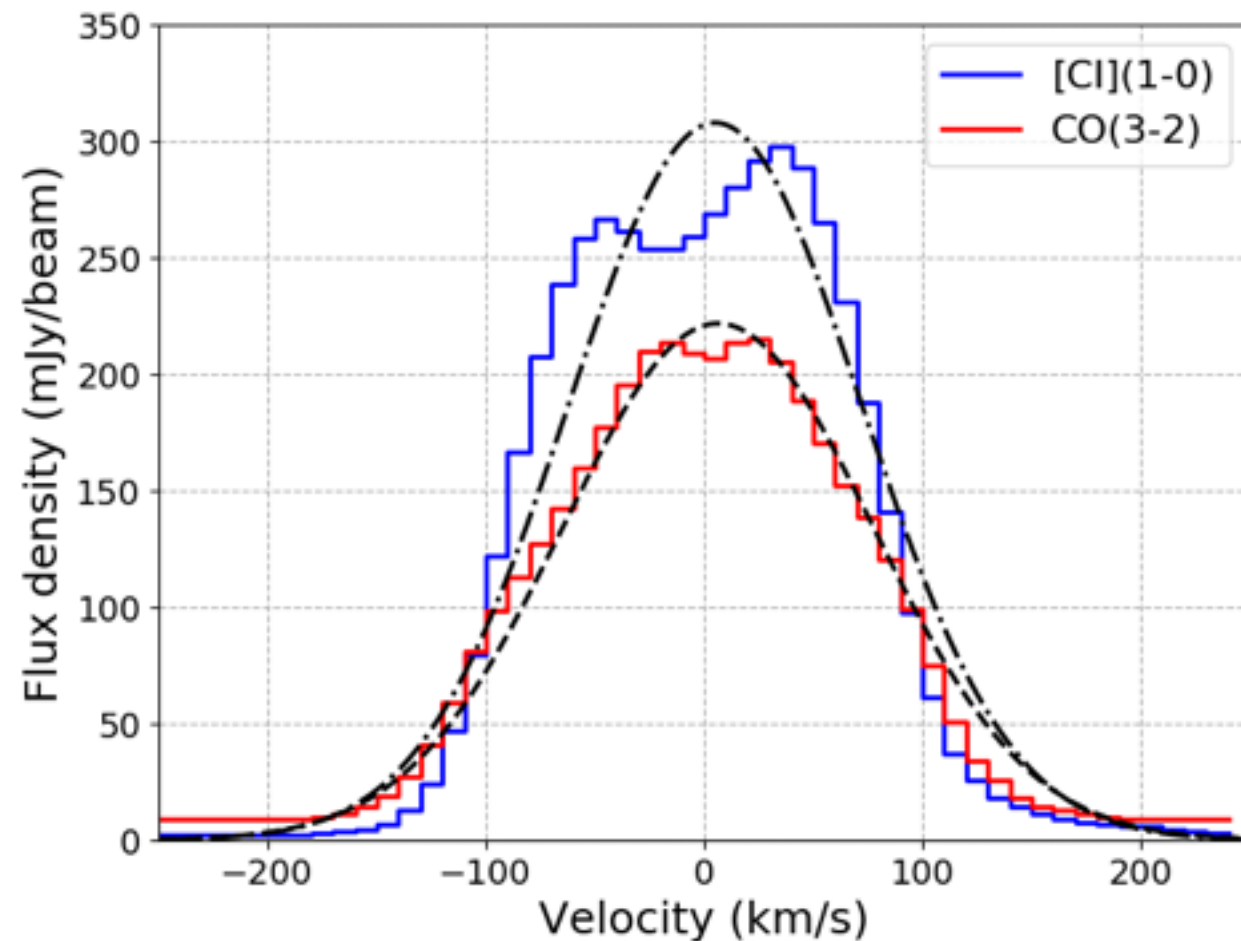




Izumi et al. 2018c, to be submitted

- M_{BH} , Eddington ratio, CND-scale M_{gas} : Circinus-like values
- **Hydrodynamic simulation + XDR chemistry (Wada+16) + rad. transfer**
- CO(3-2) → mid-plane of the CND/torus
- [Cl](1-0) → mid-plane + **puffed-up component due to outflows**

Izumi et al. 2018c, to be submitted



- Indeed, we found different line profiles for the simulated CO(3-2) and [Cl](1-0)
- *Triple-Gaussians* can well fit the profile → outflow components stand out
- Good consistency between ALMA obs. and our simulation
→ Support the fountain scheme, where atomic outflows play the key role in determining the “torus” geometry.

- Now we are testing the multi-phase dynamic torus model with ALMA: $\sim 5\text{-}15$ pc resolution CO(3-2) and [Cl](1-0) data are in hand.
- Gas (and dust) distributions = CND/torus + spiral arms.
- The CND can provide a significant nuclear obscuration (as a torus).
- Cl (atom) gas is in a geometrically thicker volume than the CO (molecule) gas at $r < 10$ pc.
→ multi-phase gas dynamics in the torus!
- The thickness of the atomic disk would be due to nuclear atomic outflows, as expected in our fountain model
→ **Physical origin of the torus!?**