

Masahito Yamazaki (Kavli IPMU, University of Tokyo)

March 30, 2022

[X00] 総括班 村山 (KIPMU)	[A01]軽いDM 高橋 (東北大)	[A02]重いDM 村瀬 (PSU)	[A03]マクロDM 柳 (名古屋大)	
[B01] レーザー干渉計 道村 (東大)	axion, dilaton ( <b>円偏光</b> )	背景重力波 (相転移など)	背景重力波 (inflationなど)	[C02
[B02] すばる分光 高田 (KIPMU)	fuzzy DM, SIDM 3D DM <b>地図</b>	<b>矮小銀河内の対消滅</b> 3D DM <b>地図</b>	PBH, UCMH, DM subhalo, 3D DM地図	2]宇宙構:
[B03] イメージング 宮崎 (NAOJ)	DM subhalo DM <b>地図</b>	DM subhalo DM地図	PBH, UCMH (重カマイクロレンズ)	造形成理
[B04] X線 山崎(典) (ISAS)	sterile neutrino moduli (輝線、連続光)	ダークマター崩壊 (輝線、連続光)	PBH蒸発 (X線背景放射)	論安藤
[B05] <i>e<sup>+</sup>e<sup>-</sup></i> 加速器 西田(KEK)	dark photon SIMP	高エネルギーの間接検証 (余剰次元、Higgs)	高エネルギーの間接検証 (余剰次元、Higgs)	(アムステリ
[B06] CMB 小松(MPA)	axion (CMB <b>偏光</b> )	宇宙初期の対消滅 <sub>N-c</sub>	PBH ( $ au$ )	レダム大)
[C01]量子重力理論 山崎(雅) (KIPMU)				
🤇 🦯 quantum gravity??				









## The Team









Yasunori Nomura (Berkeley)

Ryo Saito (Yamaguchi) Satoshi Shirai (IPMU) Syuhei Iguro (IPMU -> Karlruhe)

## Activity of Hired PD

Syuhei Iguro (2021/Apr - 2021/Sep, move to Karlsruhe )

Expert of flavor and collider physics:





# Unified approach to secondary CMB B-mode polarization

T. Namikawa, A. Naruko, **RS**, A. Taruya, and D. Yamauchi: JCAP 10 (2021) 029 (1) [B06] [A02] [C01] [C02]

## **Cosmic Microwave Background (CMB) anisotropies**



- Anisotropies of order 10<sup>-5</sup>: a source of rich information on the early universe, calculable by cosmological perturbation theory
- Recent precise measurements → Small, nonlinear effects

## Curve-of-Sight (CoS) approach :

\_ensed

A new approach to compute all secondary (non-primordial) nonlinear effects to CMB anisotropies in a single framework.

- A solution to the Boltzmann equation of CMB photons at nonlinear orders (An extension of the Line-of-Sight approach [Seljak & Zaldarriaga 96] at the linear order.)
- Weak gravitational lensing can be integrated to the Boltzmann equation.



## Application : an accuracy test of the remapping approach

The standard remapping approach is reliable to subtract the lensing B mode to detect the primordial GWs in upcoming CMB experiments.



- Estimation of all nonlinear effects not in the remapping approach. Redshift, time delay,…
- The correction is O(0.01)%
  - ~ Inflationary B mode with the tensor-to-scalar ratio  $r = O(10^{-5})$

#### Yasunori Nomura

![](_page_10_Picture_1.jpeg)

![](_page_10_Picture_2.jpeg)

What ensures the stability of dark matter?

![](_page_10_Picture_4.jpeg)

 → Any linearly realized global symmetry must be explicitly broken with O(1) strength at the string scale (the cutoff scale of low energy field theory).

Y.N., Phys. Rev. **D101** (2020) 066024

- $\cdots$  suggests specific candidates for dark matter
  - string axion
  - particle whose stability is ensured by an accidental symmetry ... occurs naturally for composite dark matter

#### Yasunori Nomura

#### Chiral Dark Sector

![](_page_11_Picture_2.jpeg)

A very **simple**, perhaps the simplest, model of composite dark matter with the properties **consistent with the black hole physics**.

![](_page_11_Figure_4.jpeg)

the **most general** Lagrangian consistent with **gauge** symmetry

→ stable dark matter with the correct abundance (dark pion, dark nucleon)

K. Harigaya and Y.N., Phys. Rev. D94 (2016) 035013
R. T. Co, K. Harigaya and Y.N., *Phys. Rev. Lett.* 118 (2017) 101801

··· rich phenomenology

possibility of two-component dark matter (dark pion and nucleon), dark radiation for  $a = 0, \cdots$ 

#### Plan

- detailed study of the model
  - $\cdots$  latest constraints, prospect for future observations,  $\cdots$
- analysis of general features of the class of similar models

 $\cdots$  effect of U(1) gauge symmetry on the confining phase transition in the early universe,  $\cdots$ 

## Examples of MY's Recent Research

## Holography for ensemble averages of CFTs

(as suggested by "wormholes" in 2d gravity / black hole information paradox)

M. Ashwinkumar, M. Dodelson, A. Kidambi, J. Leedom, MY (JHEP '21., and to appear )

## Revisiting Θ-angles/axions in Yang-Mills

R. Kitano, R. Matsudo, N. Yamada + MY ('21) and in progress Implications? Cf. Y. Nomura + T. Watari + MY ('17), M. Ibe + T.T. Yanagida + MY ('18) [e.g. A01?]

## Refining/quantifying some swampland constraints?

Ongoing discussions with J. Leedom, T. Rudelius and others

### More constraints from UV completion / positivity, causality, unitarity $\cdots$

Discussion with T. Noumi (part of this scheme), K. Aoki, J. Tokuda and C01 group

![](_page_12_Picture_10.jpeg)

![](_page_13_Picture_0.jpeg)

quantum dark sector shower, leading to characteristic signal (e.g. lepton jets with 4, 6, 8,  $\cdots \mu$ 's) [Many papers recently; cf. Tanaka-san's talk]

## **Classical Monte Carlo:**

## insufficient because of quantum interference effects

(only approximate, e.g. large Nc limit, small off-diagonal flavor coupling,  $\cdots$ )

![](_page_14_Figure_3.jpeg)

Important quantum interference effects between different flavors

Quantum processes are better simulated by quantum computers!

cf. [Bauer, de Jong, Nachman, Provasol ('19)] as a toy model for QCD

![](_page_14_Picture_7.jpeg)

So Chigusa (UC Berkeley)

MY

#### From [Chigusa-MY], to appear

![](_page_15_Figure_1.jpeg)

Sometimes dramatic enhancement for n-dark-photon events for large n (e.g. 2n  $\mu$  events); implications for DM collider search @ e.g. FASER? [cf. Otono-san's talk] Dark Matter + Quantum Computers? Quantum Sensors? new ideas / questions / collaborations welcome!

![](_page_16_Picture_1.jpeg)

2022 ~ MY as member (課題推進者)

![](_page_16_Picture_3.jpeg)

ibm\_Kawasaki @ Utokyo 27 qubit machine