

§ 2.3. Feynman rules for time-ordered product

How to compute $\langle 0 | T \{ \phi_I^1(x_1) \phi_I^2(x_2) \dots \phi_I^n(x_n) \exp[-i \int dt V_I(t)] \} | 0 \rangle$
 etc. full. com. amp. ?
 $\langle 0 | T \{ \phi_I^1(x) \phi_I^2(y) \} | 0 \rangle_{\text{PI}}$

"Elementary" fields are in the form of

$$\begin{cases} \psi_I(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(u_s(\vec{p}) a_{\vec{p},s} e^{-ip \cdot x} + v_s(\vec{p}) b_{\vec{p},s}^\dagger e^{ip \cdot x} \right) \\ \bar{\psi}_I(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(\bar{u}_s(\vec{p}) a_{\vec{p},s}^\dagger e^{ip \cdot x} + \bar{v}_s(\vec{p}) b_{\vec{p},s} e^{-ip \cdot x} \right) \end{cases}$$

Dirac field.
= means onshell

A complex ^{scalar} boson: just drop the 4-component spinors $u_s(\vec{p})$ etc.
 relativistic

A real relativistic boson: drop distinction between $a_{\vec{p}}^\dagger$ & $b_{\vec{p}}^\dagger$.
 scalar

e.g.

$$\begin{aligned} \langle 0 | T \{ \psi_I(x) \bar{\psi}_I(y) \} | 0 \rangle &= \int \frac{d^3p}{(2\pi)^3} \left(\frac{\theta(x^0 - y^0)}{2E_p} u_s(\vec{p}) \bar{u}_s(\vec{p}) e^{-ip \cdot (x-y)} + \frac{\theta(y^0 - x^0)}{2E_p} v_s(\vec{p}) \bar{v}_s(\vec{p}) e^{ip \cdot (x-y)} \right) \\ &= \int \frac{d^4p}{(2\pi)^4} \frac{i(\not{p} + m)}{(p^2 - m^2 + i\epsilon)} e^{-ip \cdot (x-y)} =: D(x-y) \end{aligned}$$

the principle: bring ann. operators to the right
 creat. ops' to the left.

ann. ops } become a part of [,] or { } or otherwise act on } $|0\rangle$
 creat. ops } to be C-valued. } $\langle 0|$

exploit all possible combinatorics. (contraction patterns).

Supplementary notes on Dirac fermion

We use $\mathcal{L} = \bar{\psi}(i\gamma^\mu D_\mu - m)\psi$ as the Lagrangian.

The gamma matrices satisfy the relation $\{\gamma^\mu, \gamma^\nu\} = 2\eta^{\mu\nu}$.
 $\eta^{\mu\nu} = \text{diag}(+, -, -, -)$.

To the equation of motion $(i\gamma^\mu D_\mu - m)\psi = 0$, there are 2 solutions of the form $\psi(x) = u_s(\vec{p}) e^{-ip \cdot x} = u_s(\vec{p}) e^{-iE_p t} e^{i\vec{p} \cdot \vec{x}}$, ($s=1,2$).

In a chiral basis, where $\gamma^0 = \begin{pmatrix} 0 & 1_{2 \times 2} \\ 1_{2 \times 2} & 0 \end{pmatrix}$, $\gamma^i = \begin{pmatrix} 0 & \tau^i \\ -\tau^i & 0 \end{pmatrix}$,

$$u_s(\vec{p}) = \begin{pmatrix} \sqrt{p \cdot \sigma} \xi_s \\ \sqrt{p \cdot \bar{\sigma}} \xi_s \end{pmatrix}, \quad \begin{pmatrix} (\sigma^\mu)_{2 \times 2} = (1, \vec{\tau}) \\ (\bar{\sigma}^\mu)_{2 \times 2} = (1, -\vec{\tau}) \end{pmatrix}$$

Two other solutions to the equation of motion are given by

$$\psi(x) = v_s(\vec{p}) e^{ip \cdot x} = v_s(\vec{p}) e^{iE_p t} e^{-i\vec{p} \cdot \vec{x}} \quad (s=1,2),$$

with $v_s(\vec{p}) = \begin{pmatrix} \sqrt{p \cdot \sigma} \xi_s \\ -\sqrt{p \cdot \bar{\sigma}} \xi_s \end{pmatrix}$ in the chiral basis above.

We usually take the 2-component spinors ξ_r to be orthonormal, $\xi_r^\dagger \xi_s = \delta_{rs}$.

Q One can verify that the following relations hold: [see Peskin-Schroeder §3]

$$\begin{aligned} \sum_{s=1,2} u_s(\vec{p}) \bar{u}_s(\vec{p}) &= \not{p} + m, & \sum_{s=1,2} v_s(\vec{p}) \bar{v}_s(\vec{p}) &= \not{p} - m, \\ u_r^\dagger(\vec{p}) u_s(\vec{p}) &= 2E_p \xi_r^\dagger \xi_s = v_r^\dagger(\vec{p}) v_s(\vec{p}) \\ \bar{u}_r(\vec{p}) u_s(\vec{p}) &= 2m \xi_r^\dagger \xi_s = -\bar{v}_r(\vec{p}) v_s(\vec{p}). \end{aligned}$$

In the non-relativistic limit, $\frac{1}{\sqrt{2E_p}} u_s(\vec{p})$ becomes $\frac{1}{\sqrt{2}} \begin{pmatrix} \xi_s \\ \xi_s \end{pmatrix}$ in the chiral basis.

[$\begin{pmatrix} \xi_s \\ 0 \end{pmatrix}$ in the non-rela. basis]

Therefore.

time-ordered correlation fns. are given by summing up contributions (amplitudes) from all possible contraction patterns. (Feynman diagram).

The amplitude of a given diagram is given by a product of.

[propagator] $D(x-y)$ for each one of contractions.
and. $([ann. creat.], \{ann, creat.\})$
[vertex] remaining coefficients. (eg. $-i \int dt' \int d^3x (\bar{\psi} \gamma^\mu e A_\mu \psi)$)
integrated over all the spacetime coordinates of \mathbb{I}_x .

Its Fourier transform version. is often more convenient.

$\psi(x) = \int \frac{d^4p}{(2\pi)^4} e^{-ip \cdot x} \tilde{\psi}(p)$. then...
for each vertex $\int d^4x \prod_i (e^{-ip_i \cdot x}) \times \text{coeff.} = (2\pi)^4 \delta^4(\sum_i p_i) \times \text{coeff.}$
for each propagator $\int \frac{d^4x d^4y}{(2\pi)^8} e^{-ip \cdot x} e^{iq \cdot y} D(x-y) = \int \frac{d^4p}{(2\pi)^4} e^{-ip \cdot x} e^{iq \cdot y} D(p) = (2\pi)^4 \delta^4(p+q) \times D(p)$.

Instead of $\#(\text{vertex})$ spacetime integrals, we are left with.

$\#(\text{propagator (internal lines)}) - \#(\text{vertex}) = -\chi(\text{graph}) = -h_0 + h_1 + \dots$
 $h_0 := \#(\text{connected components}) = 1$ (of fully connected)
 $h_2 := \#(\text{loops})$
actually $[h_2 = \#(\text{loops})]$ momentum integrals.

$h_1 = 0$ (tree level) leading order contributions.
 $h_2 = 1$ (1-loop) next-to-leading order contributions.
!
!

§2.4. Other quantities of interest in QFT

Time-ordered product correlation functions
(and scattering amplitudes derived from them)
are not the only class of observables in QFT's.

Two other classes of interest.

- In-In formalism / real time formalism in thermal field theory /
Schwinger-Keldysh formalism (known in many different names).

examples: $\langle \text{init. state} | \{ \mathcal{O}_1(x_1) \mathcal{O}_2(x_2) \dots \mathcal{O}_n(x_n) \} | \text{init. state} \rangle$

The initial state $|\text{init. state}\rangle$ is... often... the ground state
or thermal ensemble, and interactions are turned on later.
(backgrounds)

- ✓ conductance. (current-current correlation) in a time-dep.
background field
- ✓ decay / scattering of particles in a thermal plasma.
- ✓ growth (time evolution) of inflation field fluctuations
in the background of metric of expanding universe.

The time coordinate starts at $t=t_*$ for the initial state. It then goes up.
(operators are sorted out that way), and then comes back to t_*
in this class of observables.

- Out-of-time-ordered correlation functions.

example $\langle \Omega | [A(t, \vec{x}), B(0, \vec{y})]^2 | \Omega \rangle$ — (*) vanishes @ $t=0$ ($\vec{x} \neq \vec{y}$).
How fast does it grow?

If A and B are unitary operators

$$(*) = 2 - (\langle \Omega | B^\dagger A^\dagger B A | \Omega \rangle + \text{c.c.}).$$

There are experimental ways to measure such (*)'s.

§3 Scattering Processes at the Leading Order

§3.1 Vector field propagators.

gauge symmetry = redundant description.

$A_\mu(x,t)$ and $A'_\mu := A_\mu - (\partial_\mu \chi(x,t))$ are physically equivalent.

Choose a gauge

example: Coulomb gauge $A_0 = 0$ and $\vec{\nabla} \cdot \vec{A} = 0$.

when the photon is not coupled to matter.

Mode decomposition + canonical quantization.

$$\Rightarrow \begin{cases} \vec{A}(\vec{x}, t) = \int \frac{d^3\vec{p}}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \sum_{r=1,2} \left(\vec{E}_r(\vec{p}) a_{\vec{p},r} e^{-ip \cdot x} + \vec{E}_r(\vec{p}) a_{\vec{p},r}^\dagger e^{ip \cdot x} \right) \\ A_0(\vec{x}, t) = 0. \end{cases} \quad \vec{p} \cdot \vec{E}_r(\vec{p}) = 0 \quad (\text{transverse polarization})$$

This results in

$$\begin{aligned} G_{ij}(x,y) &:= \langle 0 | T \{ A_i(x) A_j(y) \} | 0 \rangle \\ &= \int \frac{d^4p}{(2\pi)^4} \frac{i e^{-ip \cdot (x-y)}}{p^2 + i\epsilon} \left[\sum_r \vec{E}_r(\vec{p}) \otimes \vec{E}_r(\vec{p}) \right]_{ij}. \end{aligned}$$

$\rightarrow \left[\frac{1}{p^2} - \frac{\vec{p}_i \otimes \vec{p}_j}{|\vec{p}|^2 p^2} \right]$

But this Green function does not satisfy the EOM.

$$\left[\partial^2 \delta_{\mu\kappa} - \partial_\mu \partial^\kappa \right] G_{\kappa\lambda} J^\lambda = J_\mu \dots \quad (*)$$

Modification A (temporal / axial gauge)

$$G_{\mu\nu} = \frac{-i}{p^2 + i\epsilon} \left[\eta_{\mu\nu} - \frac{(p_\mu n_\nu + n_\mu p_\nu)}{(p \cdot n)} + \frac{p_\mu p_\nu n^2}{(p \cdot n)^2} \right] \quad \text{for some } n_\mu.$$

For $n^\mu = (1, \vec{0})$ $n^\mu G_{\mu\nu} = 0$, $\nabla^\mu \partial^\nu G_{\mu\nu} = \frac{i p(p)}{p_0}$, (*)

are all satisfied. $\left[\partial^\mu J_\mu = 0 \right]$ (use)

The 3x3 part is $\left[-1 + \frac{\vec{p} \otimes \vec{p}}{(p^0)^2} \right]$

(Modification B (Coulomb gauge))

replace by $\left[\eta_{\mu\nu} - \frac{(p_\mu n_\nu + n_\mu p_\nu)^{(n \cdot p)}}{(p \cdot n)^2 - p^2 n^2} + \frac{p_\mu p_\nu n^2}{(p \cdot n)^2 - p^2 n^2} \right]$

The Coulomb gauge. (or axial gauge, temporal gauge) propagator is rarely use for practical computations in QFT.

More convenient choice is

$$G_{\mu\nu} = \int \frac{d^4 p}{(2\pi)^4} \frac{i e^{-i p \cdot (x-y)}}{p^2 + i\epsilon} \left[\eta_{\mu\nu} + (\xi - 1) \frac{p_\mu p_\nu}{p^2} \right].$$

• $\xi = 1$: Feynman gauge.

• $\xi = 0$: Landau gauge

• $\xi = \infty$: unitary gauge.

(For derivation, see Peskin-Schroeder §9.)

§ 3.2. $e^+e^- \rightarrow \mu^+\mu^-$ cross section

We use a Lagrangian $\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\psi} (\gamma^\mu D_\mu - m) \psi$.

$$D_\mu = (\partial_\mu - ieQA_\mu)$$

$$\left(\begin{array}{l} e > 0 \\ Q = \text{charge. } (Q_e = -1) \end{array} \right)$$

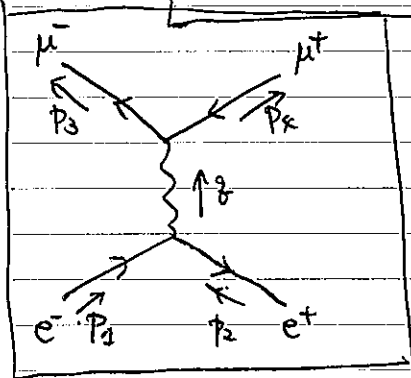
Consider.

$$(*) = \langle 0 | T \left\{ \bar{\psi}_{m,c}(-p_3) \bar{\psi}_{m,d}(-p_4) \psi_{e,b}(p_2) \psi_{e,a}(p_1) \int d^4x ie \left\{ Q_m (\bar{\psi}_m \gamma^\mu \psi_m) + Q_e (\bar{\psi}_e \gamma^\mu \psi_e) \right\} A_\mu \right. \right. \\ \left. \left. \int d^4y ie \left\{ Q_m (\bar{\psi}_m \gamma^\nu \psi_m) + Q_e (\bar{\psi}_e \gamma^\nu \psi_e) \right\} A_\nu \times \frac{1}{2i} \right\} | 0 \rangle.$$

$$(p_2^\mu, p_3^\mu, p_4^\mu, p_1^\mu > 0)$$

□ (contraction) : pairing of annihilation and creation operators

$$\Rightarrow \frac{2}{2} \times \frac{i(\not{p}_3 + m_m)_{sf}}{p_3^2 - m_m^2 + i\epsilon} ieQ_m [\gamma^\mu]_{fg} \frac{i(-\not{p}_4 + m_m)_{gd}}{p_4^2 - m_m^2 + i\epsilon} \\ \times \frac{i(\not{p}_2 + m_e)_{bh}}{p_2^2 - m_e^2 + i\epsilon} ieQ_e [\gamma^\nu]_{hg} \frac{i(\not{p}_1 + m_e)_{ia}}{p_1^2 - m_e^2 + i\epsilon} \times \left(\frac{-i\eta_{\mu\nu}}{q^2 + i\epsilon} \right) \\ \times (2\pi)^4 \delta^4(p_1 + p_2 - q) \times (2\pi)^4 \delta^4(q - p_3 - p_4) \frac{d^4q}{(2\pi)^4}$$



$$\text{Res} (*) = \left[\bar{u}_f(\vec{p}_3) \gamma^\mu v_s(\vec{p}_4) \right] \left[\bar{v}_{s'}(\vec{p}_2) \gamma^\nu u_r(\vec{p}_1) \right] \frac{\eta_{\mu\nu}}{(p_1 + p_2)^2 + i\epsilon} \\ \times i(2\pi)^4 \delta^4(p_1 + p_2 - p_3 - p_4) \times (e^2 Q_e Q_\mu)$$

$$= (2\pi)^4 \delta^4(p_1 + p_2 - p_3 - p_4) \cdot i\mathcal{M}(e^+e^- \rightarrow \mu^+\mu^-)$$

(LSZ formula)

Now, we hope to compute $|\mathcal{M}|^2$ (and then the cross section).

But the expression is a mess, when we retain m_e, m_μ etc.
r.s., r'.s'.

The expression of $|\mathcal{M}|^2$ simplifies when we think of

$\rightarrow \sum_{r,s} \left| \mathcal{M}(\bar{e}_r + e_s^+ \rightarrow \bar{\mu}_r + \mu_s^+) \right|^2$ ignore the spin of μ^+, μ^- in the final state.

$\rightarrow \frac{1}{2} \sum_{r'=1}^2 \frac{1}{2} \sum_{s'=1}^2 \sum_{r,s=1}^2 \left| \mathcal{M}(\bar{e}_{r'} + e_{s'}^+ \rightarrow \bar{\mu}_r + \mu_s^+) \right|^2$ initial state e^+, e^- are not polarized.

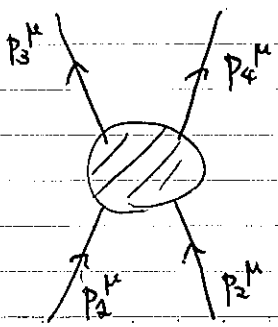
Now

$$\begin{aligned} \overline{|\mathcal{M}|^2} &= \frac{1}{4} \sum_{r,s,r',s'} \left(\frac{e^2 Q_e Q_\mu}{(p_1 + p_2)^2} \right)^2 \times \left[\bar{u}_r(\vec{p}_3) \gamma^\mu v_{s'}(\vec{p}_4) \right] \left[\bar{v}_{s'}(\vec{p}_4) \gamma^\kappa u_r(\vec{p}_3) \right] \\ &\quad \times \left[\bar{v}_{s'}(\vec{p}_2) \gamma_\mu u_{r'}(\vec{p}_1) \right] \left[\bar{u}_{r'}(\vec{p}_1) \gamma_\kappa v_{s'}(\vec{p}_2) \right] \\ &= \frac{1}{4} \left(\frac{e^2 Q_e Q_\mu}{(p_1 + p_2)^2} \right)^2 \times \text{Tr}_{4 \times 4} \left[\gamma^\mu (\not{p}_2 - m_\mu) \gamma^\kappa (\not{p}_3 + m_\mu) \right] \times \text{Tr}_{4 \times 4} \left[\gamma_\mu (\not{p}_1 + m_e) \gamma_\kappa (\not{p}_2 - m_e) \right] \\ &= \frac{1}{4} \left(\frac{e^2 Q_e Q_\mu}{(p_1 + p_2)^2} \right)^2 \times \left\{ -m_\mu^2 \eta^{\mu\kappa} - (p_3 \cdot p_4) \eta^{\mu\kappa} + p_3^\mu p_4^\kappa + p_4^\kappa p_3^\mu \right\} \times 4 \\ &\quad \times \left\{ -m_e^2 \eta_{\mu\kappa} - (p_1 \cdot p_2) \eta_{\mu\kappa} + p_{1\mu} p_{2,\kappa} + p_{1,\kappa} p_{2,\mu} \right\} \times 4 \\ &= 4 \left(\frac{e^2 Q_e Q_\mu}{(p_1 + p_2)^2} \right)^2 \times \left[\begin{aligned} &\left\{ (p_3 \cdot p_4) + m_\mu^2 \right\} \left\{ (p_1 \cdot p_2) + m_e^2 \right\} \times 4 \\ &+ \left\{ p_1 \cdot p_4 \right\} \left\{ p_2 \cdot p_3 \right\} + \left\{ p_4 \cdot p_2 \right\} \left\{ p_1 \cdot p_3 \right\} \right] \times 2 \\ &- \left[\left\{ (p_1 \cdot p_2) + m_e^2 \right\} \left\{ p_3 \cdot p_4 \right\} \right] \times 2 - \left[\left\{ (p_3 \cdot p_4) + m_\mu^2 \right\} \left\{ p_1 \cdot p_2 \right\} \right] \times 2 \end{aligned} \right] \end{aligned}$$

(see suppl. notes)

It is conventional to use the following variables (Mandelstam variables)

in 2-body \rightarrow 2-body scattering:



$s := (p_1 + p_2)^2$
 $t := (p_1 - p_3)^2$
 $u := (p_1 - p_4)^2$

Due to the momentum conservation $(p_1 + p_2 - p_3 - p_4)^\mu = 0$,

there is a relation $s + t + u = m_1^2 + m_2^2 + m_3^2 + m_4^2$.

Using these kinematical variables, the spin-sum/average MF can be rewritten as follows:

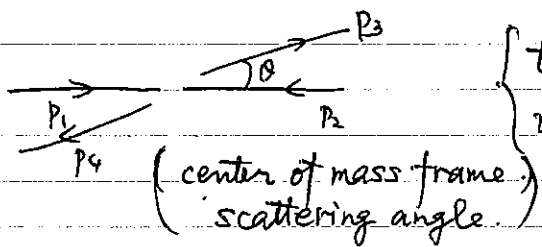
$$\frac{1}{4} \sum_{r,s} \sum_{r',s'} |M|^2 = 4 \left(\frac{e^2 Q_e Q_\mu}{s} \right)^2 \left[s^2 + \frac{1}{2} \left\{ (-u + m_e^2 + m_\mu^2)^2 + (-t + m_e^2 + m_\mu^2)^2 \right\} - s \left(\frac{s-2m_\mu^2}{2} \right) - s \left(\frac{s-2m_e^2}{2} \right) \right]$$

$$= 4 \left(\frac{e^2 Q_e Q_\mu}{s} \right)^2 \left[\frac{u^2 + t^2}{2} + (s-u-t)(m_e^2 + m_\mu^2) + (m_e^2 + m_\mu^2)^2 \right].$$

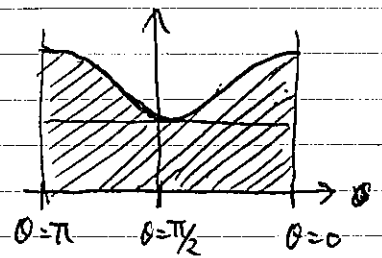
Here are two limits of interest.

• High energy limit : $(2m_\mu)^2 \ll s$.

$\Rightarrow \frac{1}{4} \sum_{r,s} \sum_{r',s'} |M|^2 \approx 2 (e^2 Q_e Q_\mu)^2 \frac{(u^2 + t^2)}{s^2} = (e^2 Q_e Q_\mu)^2 (1 + \cos^2 \theta)$



$\begin{cases} t \approx -s \cdot \sin^2(\frac{\theta}{2}) \\ u \approx -s \cos^2(\frac{\theta}{2}) \end{cases}$



Using [homework I-1]

$$d\sigma \approx \frac{d(\cos \theta)}{32\pi s} (\beta \approx 1) \left(\frac{1}{4} \sum_{r,s} \sum_{r',s'} |M|^2 \right) \approx \left[d(\cos \theta) (1 + \cos^2 \theta) \Rightarrow \frac{d}{3} \right] \times \frac{\pi \alpha_e^2}{2s} Q_\mu^2 \Rightarrow \frac{4\pi}{3} \frac{\alpha_e^2}{s} Q_\mu^2$$

• Just above the threshold : $s \sim (2m_\mu)^2 \Rightarrow t \sim u \sim -m_\mu^2$

$\Rightarrow \frac{1}{4} \sum_{r,s} \sum_{r',s'} |M|^2 \approx 2 (e^2 Q_e Q_\mu)^2 \quad (\theta\text{-indep.})$

Using [homework I-1]

$$\int d\sigma \approx \int_{-1}^{+1} d(\cos \theta) \frac{\beta}{32\pi s} \left(\frac{1}{4} \sum_{r,s} \sum_{r',s'} |M|^2 \right) \approx \pi \frac{\alpha_e^2}{s} Q_\mu^2 \times \beta$$

$\beta = \frac{|\vec{p}_\mu|}{E_\mu}$

$\alpha_e = \frac{e^2}{4\pi}$

Supplementary notes

$\{\gamma^\mu, \gamma^\nu\} = 2\eta^{\mu\nu}$: the definition of gamma matrices.

We can use this anti-commutation relation to derive the followings.

First,

$$\text{tr}_{4 \times 4} [\gamma^\mu \gamma^\nu] = \frac{1}{2} \left(\text{tr} [\gamma^\mu \gamma^\nu] + \text{tr} [\gamma^\nu \gamma^\mu] \right) = \frac{1}{2} \text{tr} [\{\gamma^\mu, \gamma^\nu\}] = \frac{2}{2} \text{tr} [\mathbb{1}_{4 \times 4}] \eta^{\mu\nu} = 4 \eta^{\mu\nu}$$

cyclic rotation within a trace.

Secondly,

$$\begin{aligned} \text{tr}_{4 \times 4} [\gamma^\mu \gamma^\nu \gamma^\kappa \gamma^\lambda] &= \text{tr}_{4 \times 4} [\{\gamma^\mu, \gamma^\nu\} \gamma^\kappa \gamma^\lambda] = \text{tr}_{4 \times 4} [\gamma^\nu \gamma^\mu \gamma^\kappa \gamma^\lambda] \\ &= 2\eta^{\mu\nu} \text{tr} [\gamma^\kappa \gamma^\lambda] - \text{tr} [\gamma^\nu \{\gamma^\mu, \gamma^\kappa\} \gamma^\lambda] + \text{tr} [\gamma^\nu \gamma^\kappa \gamma^\mu \gamma^\lambda] \\ &= 2\eta^{\mu\nu} \text{tr} [\gamma^\kappa \gamma^\lambda] - 2\eta^{\mu\kappa} \text{tr} [\gamma^\nu \gamma^\lambda] + \text{tr} [\gamma^\nu \gamma^\kappa \{\gamma^\mu, \gamma^\lambda\}] \\ &\quad - \text{tr} [\gamma^\nu \gamma^\kappa \gamma^\lambda \gamma^\mu] \\ &= 2 \times 4 \left(\eta^{\mu\nu} \eta^{\kappa\lambda} - \eta^{\mu\kappa} \eta^{\nu\lambda} + \eta^{\mu\lambda} \eta^{\nu\kappa} \right) - \text{tr} [\gamma^\mu \gamma^\nu \gamma^\kappa \gamma^\lambda]. \end{aligned}$$

Therefore $\text{tr}_{4 \times 4} [\gamma^\mu \gamma^\nu \gamma^\kappa \gamma^\lambda] = 4 \left(\eta^{\mu\nu} \eta^{\kappa\lambda} - \eta^{\mu\kappa} \eta^{\nu\lambda} + \eta^{\mu\lambda} \eta^{\nu\kappa} \right)$.