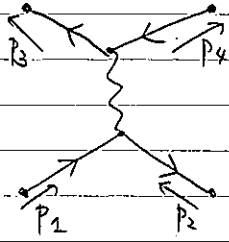


# § 4. Ultraviolet Divergence and Regularization

## § 4.1. An example: Self-Energy

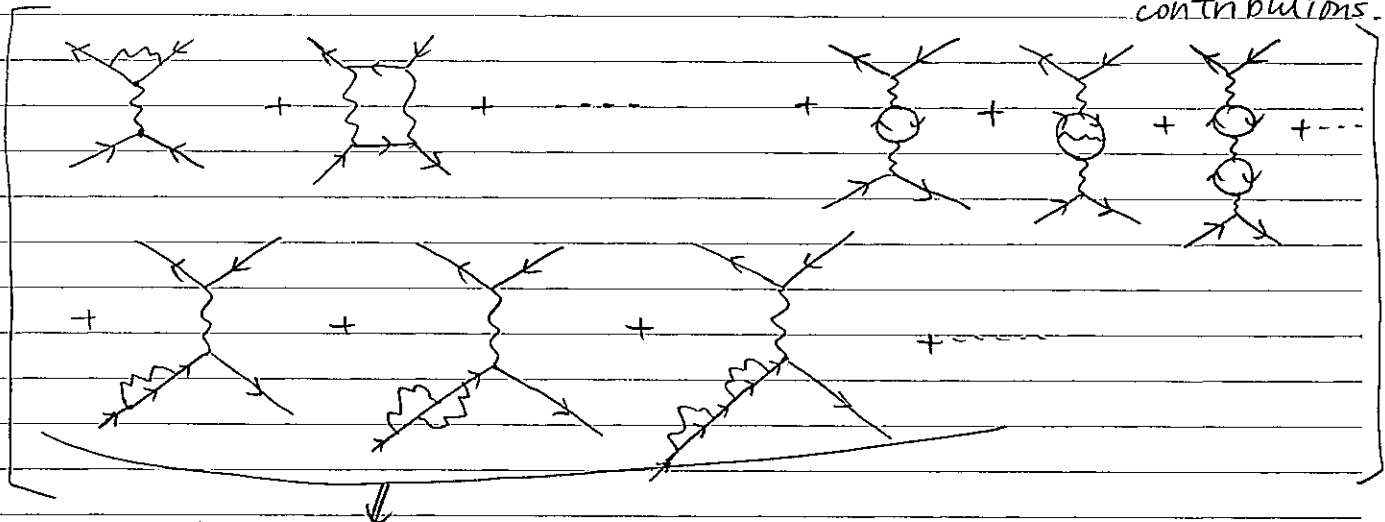
Consider the  $e^- + e^+ \rightarrow \mu^- + \mu^+$  process in QED.



$$= \frac{i[\not{p}_3 + M]}{(p_3^2 - M^2 + i\epsilon)} (-ieQ_{\mu}) \gamma^\mu \frac{i[\not{p}_2 + M]}{(p_2^2 - M^2 + i\epsilon)} (-ieQ_e) \gamma^\nu \frac{i[\not{p}_1 + m]}{(p_1^2 - m^2 + i\epsilon)} (-ieQ_e) \gamma^\nu \frac{i[\not{p}_4 + m]}{(p_4^2 - m^2 + i\epsilon)} (-ieQ_{\mu}) \gamma^\mu$$

$$\times \frac{(-i\eta_{\mu\nu})}{(p_1 + p_2)^2 + i\epsilon} (2\pi)^4 \delta^4(p_3 + p_4 - p_1 - p_2).$$

This is the leading order ( $\mathcal{O}(e^2)$ ) contribution to the 4-point time-ordered correlation function, but there are other contributions.



An observation: those contributions have the following structure:

$$(\text{a common factor}) \times \left[ \frac{i(\not{p}_2 + m)}{(p_1^2 - m^2 + i\epsilon)} + \frac{i(\not{p}_2 + m)}{(p_1^2 - m^2 + i\epsilon)} \times (M_1)_\lambda \frac{\not{p}_2 + m}{(p_2^2 - m^2 + i\epsilon)} + \frac{i(\not{p}_2 + m)}{(p_1^2 - m^2 + i\epsilon)} \times (M_1')_\lambda \frac{i(\not{p}_1 + m)}{(p_1^2 - m^2 + i\epsilon)} \right]$$

$$+ \frac{i(\not{p}_1 + m)}{(p_1^2 - m^2 + i\epsilon)} \times M_1 \times \frac{i(\not{p}_1 + m)}{(p_1^2 - m^2 + i\epsilon)} \times M_2 \frac{i(\not{p}_1 + m)}{(p_1^2 - m^2 + i\epsilon)} + \dots$$

$$\cong (\text{the common factor}) \times \frac{i}{(\not{p}_1 - m + i\epsilon) - i(M_1 + M_1')}$$

(after integrating loop momenta)

↓

4x4 matrix valued function of  $p_1^\mu$  and  $m$ .

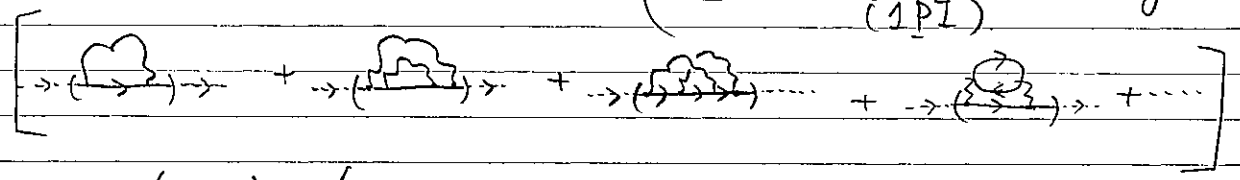
So, it is better to deal with perturbative corrections

- at the core of a particle-particle interaction

- the 1st unseparable corrections on an external state
  - the 2nd unseparable =
  - the 3rd unseparable =
  - ⋮
- can be organized into a geometric series.

⇒ just enough to compute single unseparable corrections.

(1 particle irreducible diagrams)  
(1PI)



$$= \sum_{\text{sum}} (\mathcal{U}'\text{'s}) \left( \begin{array}{l} \text{including the vertex factors at the both ends} \\ \text{but not including } (2\pi)^4 \delta^4(p_2^{\text{in}} - p_1^{\text{out}}) \\ \text{not including } \frac{i(p_1+m)}{(p_1^2-m^2+i\epsilon)} \text{ at the both ends} \end{array} \right)$$

$$= -i \left[ \frac{1}{2} A(p_1^2, m^2) + B(p_1^2, m^2) \right]$$

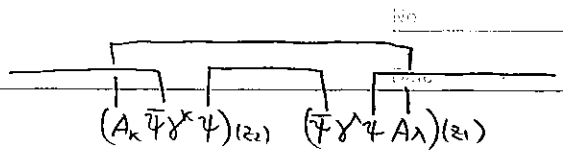
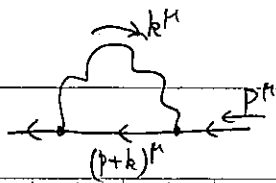
So, the 4-point time-ordered correlation function is in the form of

$$\text{(the common factor)} \times \left[ \frac{i}{(1-A(p_1^2, m^2)) p_1^2 - (m+B) + i\epsilon} \right]$$

$$\frac{i [(1-A) p_1^2 + (m+B)]}{[(1-A)^2 p_1^2 - (m+B)^2 + i\epsilon]}$$

A pole in  $(p_2)^2 \Leftrightarrow$  an on-shell particle in the Hilbert space  
 The residue  $\Leftrightarrow$  relative normalization between fields and states.

- The particle mass (where the pole is located) is shifted from the one we started with because of the interaction.
- What we see as a particle is NOT an elementary particle but a quantum mechanical cloud of interacting fields presenting themselves as a pole in the momentum space.
- Don't we need consistency in the calculation??



$$-i [A(p,m) \not{p} + B(p,m)] \stackrel{1\text{-loop}}{=} \int \frac{d^4 k}{(2\pi)^4} (-ieQ \gamma^k) \frac{i \not{p+k} + m}{(p+k)^2 - m^2 + i\epsilon} (-ieQ \gamma^k) \frac{(-i \eta_{\kappa\lambda})}{(k^2 + i\epsilon)}$$

§4.2 Evaluate this 1-loop self-energy.

$$[\not{p}A + B] = -i(eQ)^2 \int \frac{d^4 k}{(2\pi)^4} \frac{\left\{ \gamma^k \not{p+k} + m \right\} \gamma_{\kappa} = -2(p+k) + 4m}{[(p+k)^2 - m^2 + i\epsilon][k^2 + i\epsilon]} \left\{ \begin{array}{l} \gamma^{\kappa} \gamma^{\nu} \gamma_{\kappa} = -2\gamma^{\nu} \\ \gamma^{\mu} \gamma_{\mu} = 4 \mathbb{1}_{4 \times 4} \end{array} \right. \leftarrow$$

Calculation trick ①

$$\left\{ \begin{array}{l} \frac{1}{\alpha\beta} = \int_0^1 dx \frac{1}{[x\alpha + (1-x)\beta]^2} \quad \left( \text{RHS} = \int_0^1 dx \frac{1}{(\alpha-\beta)^2 \left(x + \frac{\beta}{\alpha-\beta}\right)^2} = \frac{1}{(\alpha-\beta)^2} \left( \frac{\alpha-\beta}{\beta} - \frac{\alpha-\beta}{\alpha} \right) = \text{LHS} \right) \\ \frac{1}{\alpha\beta\gamma} = \int_0^1 \int_0^1 dx dy \frac{1}{[x\alpha + y\beta + (1-x-y)\gamma]^3} \quad \left( \text{proof = similar} \right) \\ \frac{1}{\alpha_1 \alpha_2 \dots \alpha_n} = \int_{0 \leq x_i} d^n x \frac{\delta(x, x_i - 1) (n-1)!}{[x_1 \alpha_1 + x_2 \alpha_2 + \dots + x_n \alpha_n]^n} \quad \left( \text{OFT II, week 15} \right) \end{array} \right.$$

$$\int_0^1 [\not{p}A + B] = -i(eQ)^2 \int \frac{d^4 k}{(2\pi)^4} \int_0^1 dx \frac{[-2(p+k) + 4m]}{[x(p+k)^2 - x m^2 + (1-x)k^2 + i\epsilon]^2}$$

denominator:  $k^2 + 2xk \cdot p + x p^2 - x m^2 = \underbrace{(k + xp)^2}_{\equiv (k')^2} + (1-x^2)p^2 - x m^2$

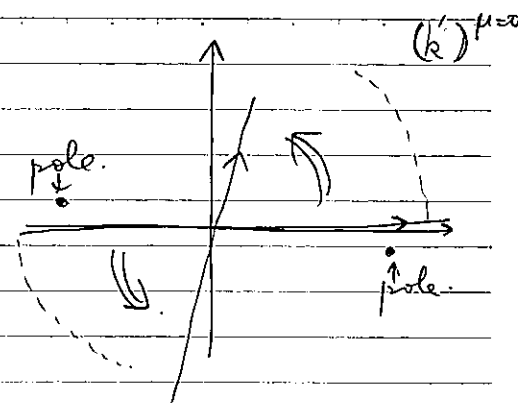
$$= -i(eQ)^2 \int dx \int \frac{d^4 k'}{(2\pi)^4} \frac{[-2k' - 2(1-x)\not{p} + 4m]}{[(k')^2 + x(1-x)p^2 - m^2 + i\epsilon]^2}$$

Because the denominator is even under  $k'^{\mu} \rightarrow -k'^{\mu}$ ,  
just drop the  $(-2k')$  term in the numerator.  
(odd)

( $x$ 's introduced here: Feynman parameters)

Calculation trick ②

$$\int_{-\infty}^{+\infty} \frac{d(k^0)}{2\pi} \frac{1}{\left[ (k^0)^2 - |\vec{k}'|^2 + (\text{something}) + i\epsilon \right]^\nu}$$



rotate the integration contour  
(wick rotation)

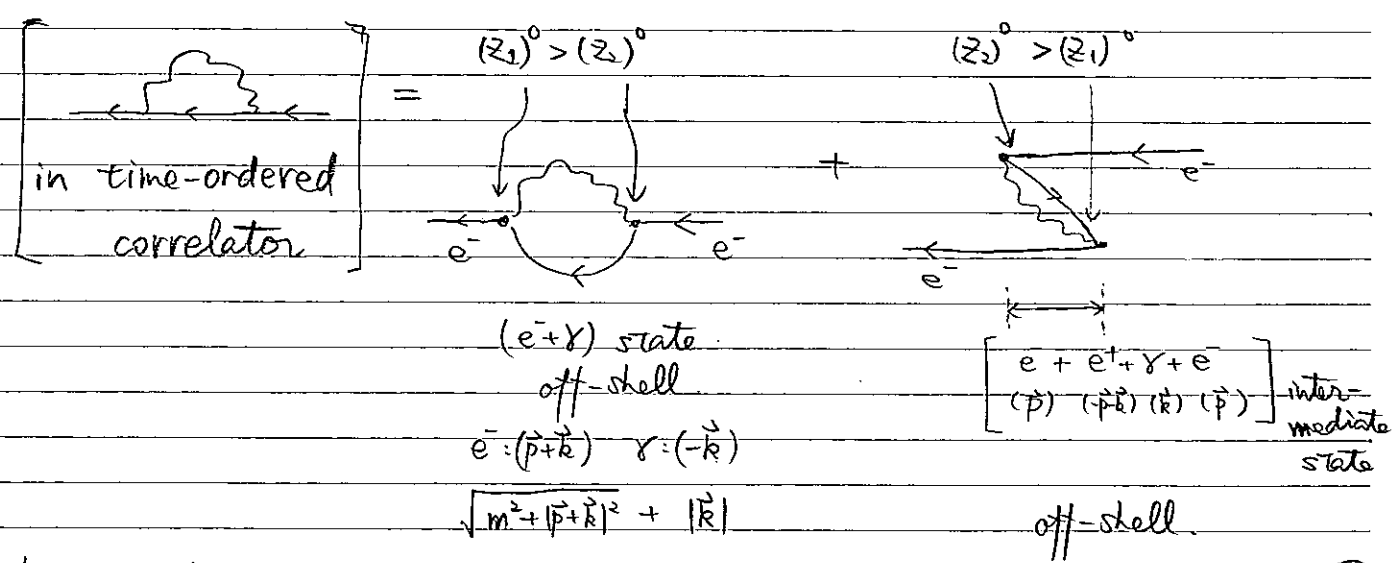
$$\int_{-\infty}^{+\infty} d(k^0) \Rightarrow \int_{-i\infty}^{+i\infty} d(k^0) = i \int_{-\infty}^{+\infty} dk^E \quad \text{Euclidean signature}$$

$$[\not{x} A + B] = \underbrace{[-(eQ)^2]}_{+(eQ)^2} (+i) \int_0^1 dx \int \frac{d^4 k_E}{(2\pi)^4} \frac{[-2\not{x}(1-x) + \not{4}]}{[-k_E^2 - 2m^2 + x(1-x)p^2]^2}$$

From the region  $|k_E|^2 \gg m^2, |p^2|$ , this integration diverges.

$$e^2 \int d|k_E| \left( \frac{k_E^3}{(k_E)^2} \right) \sim \text{log divergence}$$

Remember that



Ultraviolet

(Divergence): there are just too many intermediate states for the correlator to be finite. labeled by  $\vec{k} \in \mathbb{R}^3$  PLUS

### § 4.3 Regularization

UV divergence is from <sup>(many)</sup> high-energy states.

But we do not know for sure whether they are really "there".

- the Standard model : not sure beyond TeV  
certainly not above  $M_{pl} \sim 2 \times 10^{19}$  GeV.
- QED : not above 80 GeV ( $m_W$ )
- $\pi^\pm, N$  : not above 700 GeV ( $p^\pm, p^0$ )
- condensed matter system : not  $|\vec{k}| > (1/a_B)$

Brillouin zone boundary  
(lattice spacing  
not continuous)

Do we need to know the right UV modification to those theories??

No! regularization + renormalization

Modify QED Lagrangian at high-energy

$$\text{eg } \mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4\Lambda^2} F_{\mu\nu} D^\kappa D_\kappa F^{\mu\nu} + \bar{\Psi} (i\gamma^\mu D_\mu - m) \Psi$$

- higher covariant derivative
  - Pauli-Villars
  - string theory .....
- their combination .....

Now

$$\begin{aligned} [\not{A} + B] &\Rightarrow -i(eQ)^2 \int \frac{d^4k}{(2\pi)^4} \frac{\gamma^\kappa (\not{p} + \not{k} + m) \gamma_\kappa}{[(p+k)^2 - m^2 + i\epsilon] [k^2 - \frac{b^2}{\Lambda^2} + i\epsilon]} \rightarrow (k^2)(k^2 - \Lambda^2)/\Lambda^2 \\ &= -i(eQ)^2 \int dx dy \int \frac{d^4k}{(2\pi)^4} \frac{[-2(\not{p} + \not{k}) + 4m] (-2\Lambda^2)}{[k^2 + 2xk \cdot p - x^2 m^2 - y\Lambda^2 + x p^2 + i\epsilon]^3} \\ &= -i(eQ)^2 \int dx dy \int \frac{d^4k}{(2\pi)^4} \frac{[-2(1-x)\not{p} + 4m] (-2\Lambda^2)}{[(k')^2 - x^2 m^2 - y\Lambda^2 + x(1-x)p^2 + i\epsilon]^3} \end{aligned}$$

Wick rotation: then.

$$\begin{aligned}
 [A+B] &= -i(eQ)^2 \int d^3x d^3y \frac{i \int d^4k_E (-2\Lambda^2) [-2(1-x)\not{p} + \not{x}m]}{(2\pi)^4 (-) [k_E^2 + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} \\
 &= \frac{(eQ)^2}{(2\pi)^4} [2\pi^3 = \text{vol}(S^3)] \frac{1}{2} \int_0^{+\infty} d(k_E^2) (k_E^2) (2\Lambda^2) \frac{[-2(1-x)\not{p} + \not{x}m]}{[k_E^2 + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} \\
 &= \frac{(eQ)^2}{16\pi^2} \int_0^{+\infty} dk \frac{k}{[k + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} (2\Lambda^2) [-2(1-x)\not{p} + \not{x}m] \\
 &\left( \text{Now } \int_0^{+\infty} dk \frac{k}{[k + (xxx)]^3} = \frac{1}{(xxx)} \int_0^{+\infty} dk \frac{k+1-1}{(k+1)^3} = \frac{1}{(xxx)} \left(1 - \frac{1}{2}\right) = \frac{1}{2(xxx)} \right. \\
 &= \frac{\alpha Q^2}{4\pi} \int_0^1 dx \int_0^1 dy \frac{\Lambda^2 [-2(1-x)\not{p} + \not{x}m]}{[xm^2 + y\Lambda^2 - x(1-x)p^2]^3} \quad \text{converges}
 \end{aligned}$$

$y$ -integral then

$$= \frac{\alpha Q^2}{4\pi} \int_0^1 dx [-2(1-x)\not{p} + \not{x}m] \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right)$$

We found that

$$\begin{aligned}
 A(p^2, m^2) &= \frac{\alpha Q^2}{4\pi} \int_0^1 dx -2(1-x) \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right) \\
 B(p^2, m^2) &= \frac{\alpha Q^2}{4\pi} \int_0^1 dx \not{x}m \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right)
 \end{aligned}$$

A and B remain finite. so long as  $\Lambda$  is finite  
(not infinite)

The UV divergence is under control.  
(regularization)