

## § 1.3 Regularization

(many)

UV divergence is from high-energy states.

But we do not know for sure whether they are really "there".

- the Standard model : not sure beyond TeV  
certainly not above  $M_{pl} \sim 2 \times 10^{16}$  GeV.
- QED : not above 80 GeV ( $m_{W^\pm}$ )
- $\pi^\pm, N$  : not above 700 MeV ( $p^\pm, p^0$ )
- condensed matter system : not  $|\vec{k}| > (1/a_B)$

Brillouin zone boundary  
(lattice spacing  
not continuous)

Do we need to know the right UV modification to those theories??

No!

regularization + renormalization

Modify QED Lagrangian at high-energy

$$\text{eg. } \mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4\Lambda^2} F_{\mu\nu} D^\kappa D_\kappa F^{\mu\nu} + \bar{\psi} (i \gamma^\mu D_\mu - m) \psi$$

- higher covariant derivative
  - Pauli-Villars
  - string theory .....
- their combination .....

Now

$$\begin{aligned} \cancel{[A+B]} &\stackrel{(1)}{\Rightarrow} -i(eQ)^2 \int \frac{d^4 k}{(2\pi)^4} \frac{\gamma^\kappa (\cancel{p+k} + m) \gamma_\kappa}{(p+k)^2 - m^2 + i\epsilon} \frac{1}{\left[ k^2 - \frac{k^2}{\Lambda^2} + i\epsilon \right]} \rightarrow (k^2)(k^2 - \Lambda^2)/\Lambda^2 \\ &= -i(eQ)^2 \int \frac{d^4 k}{(2\pi)^4} \int_0^1 dx dy \frac{[2(\cancel{p+k} + m)] (-2\Lambda^2)}{[k^2 + 2xk \cdot p - xm^2 - y\Lambda^2 + xp^2 + i\epsilon]^3} \\ &= -i(eQ)^2 \int dx dy \int \frac{d^4 k'}{(2\pi)^4} \frac{[-2(1-x)\cancel{p} + 4m] (-2\Lambda^2)}{[(k')^2 - xm^2 - y\Lambda^2 + x(1-x)p^2 + i\epsilon]^3} \end{aligned}$$

Wick rotation: then.

$$\begin{aligned}
[\not{A} + B]^{(1)} &= -i(eQ)^2 \int d^4x d^4y \cdot i \frac{d^4k_E}{(2\pi)^4} \frac{(-2\Lambda^2) [-2(1-x)\not{p} + \not{q}m]}{[k_E^2 + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} \\
&= \frac{(eQ)^2}{(2\pi)^4} \left[ 2\pi^2 = \text{vol}(S^2) \right] \frac{1}{2} \int_0^{+\infty} d(k_E^2) (k_E^2) (2\Lambda^2) \frac{[-2(1-x)\not{p} + \not{q}m]}{[(k_E^2) + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} \\
&= \frac{(eQ)^2}{16\pi^2} \int_0^{+\infty} dk K \frac{K}{[K + xm^2 + y\Lambda^2 - x(1-x)p^2]^3} (2\Lambda^2) [-2(1-x)\not{p} + \not{q}m] \\
&\quad \left( \text{Now } \int_0^{+\infty} dk \frac{k}{[k + (xxx)]^3} = \frac{1}{(xxx)} \int_0^{+\infty} dk \frac{k+1-1}{(k+1)^3} = \frac{1}{(xxx)} \left( 1 - \frac{1}{2} \right) = \frac{1}{2(xxx)} \right) \\
&= \frac{\alpha Q^2}{4\pi} \int_0^1 dx \int_0^1 dy \frac{\Lambda^2 [-2(1-x)\not{p} + \not{q}m]}{[xm^2 + y\Lambda^2 - x(1-x)p^2]} \quad \text{converges}
\end{aligned}$$

$\not{q}$ -integral then

$$= \frac{\alpha Q^2}{4\pi} \int_0^1 dx [-2(1-x)\not{p} + \not{q}m] \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right)$$

We found that

$$\begin{aligned}
A(p^2, m^2) &= \frac{\alpha Q^2}{4\pi} \int_0^1 dx -2(1-x) \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right) \\
B(p^2, m^2) &= \frac{\alpha Q^2}{4\pi} \int_0^1 dx \not{q}m \ln \left( \frac{(1-x)\Lambda^2 + xm^2 - x(1-x)p^2}{xm^2 - x(1-x)p^2} \right)
\end{aligned}$$

A and B remain finite so long as  $\Lambda$  is finite (not infinite)

The UV divergence is under control.  
(regularization)

Regularization methods may be different from the true mechanism adopted by the nature in rendering amplitudes UV finite. This is not a problem (as we will see later), because the choice of regularization method does not matter after completing the procedure of renormalization.

## §2 Renormalized Perturbation Theory

### §2.1 Idea

We have seen that

$$\left\{ \begin{aligned} & \int \frac{d^4 p}{(2\pi)^4} \int \frac{d^4 q}{(2\pi)^4} e^{i p \cdot x} e^{-i q \cdot y} \langle 0 | T \{ \Psi_2(x) \bar{\Psi}_2(y) \exp \left[ i \int d^4 z \mathcal{L}_{int}(z) \right] \} | 0 \rangle_{\text{connected}} \\ &= (2\pi)^4 \delta^4(p-q) \frac{i}{\not{p} - m - (A \not{p} + B) + i\epsilon} = (2\pi)^4 \delta^4(p-q) \left( \frac{i \not{p} \cdot z_2}{p^2 - m_{\text{pole}}^2 + i\epsilon} + \dots \right) \\ & \Psi_2(x) = \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{1}{\sqrt{2E_{\vec{p}}}} \left( u(\vec{p})_s \underset{\text{(free)}}{e^{-i p \cdot x}} a_{\vec{p},s} + v(\vec{p})_r \underset{\text{(free)}}{e^{i p \cdot x}} b_{\vec{p},s}^\dagger \right) \end{aligned} \right.$$

$$\Rightarrow \text{Renormalized fields } \frac{1}{\sqrt{z_2}} \Psi_2(x) =: [\Psi_2]_r(x) \text{ and } \frac{1}{\sqrt{z_2}} \bar{\Psi}_2 =: [\bar{\Psi}_2]_r$$

have the right normalization.

also a theory parameter "m" should be chosen

$$\text{so that the "pole" } \left( \begin{aligned} & (1 - A(p^2, m^2)) p^2 - (m + B(p^2, m^2))^2 = 0 \\ & \text{has a solution } p^2 = m_{\text{pole}}^2 \end{aligned} \right)$$

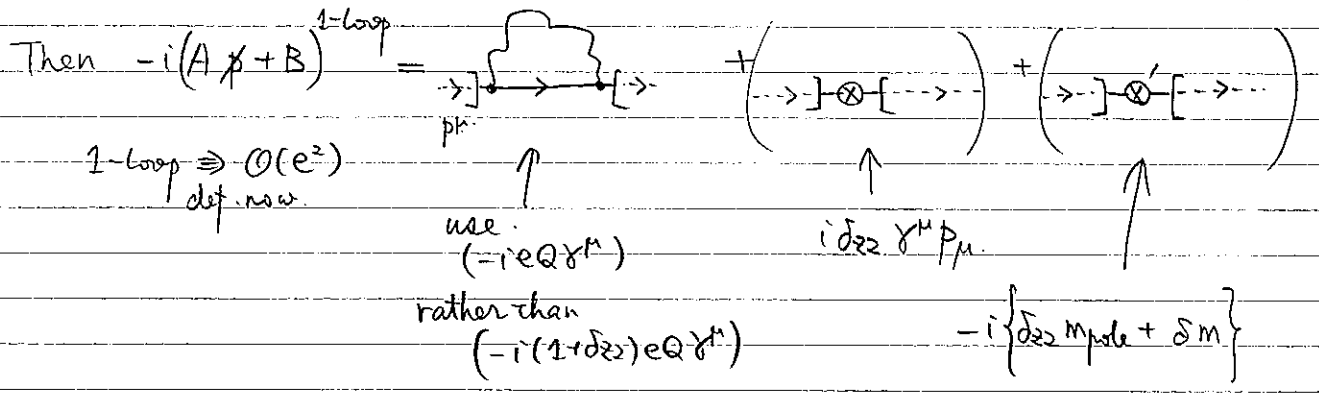
has the right value.

QED action

$$\mathcal{L} = \bar{\Psi} \left[ i \gamma^\mu (\partial_\mu + i e Q A_\mu) - m \right] \Psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

(For now, keep  $e$  and  $A_\mu$  untouched  $\Rightarrow$  allowed in QED)

$$\begin{aligned}
 \mathcal{L}_{\text{free}} &= \bar{\Psi} [i\gamma^\mu \partial_\mu - m] \Psi \\
 &= \bar{\Psi}_r [i\gamma^\mu \partial_\mu - m] \Psi_r \\
 &= (1 + \delta_{32}) \bar{\Psi}_r [i\gamma^\mu \partial_\mu - (m_{\text{pole}} + \delta m)] \Psi_r \\
 &= \underbrace{\bar{\Psi}_r [i\gamma^\mu \partial_\mu - m_{\text{pole}}] \Psi_r}_{\mathcal{L}_0} + \underbrace{\delta_{32} \bar{\Psi}_r i\gamma^\mu \partial_\mu \Psi_r - \{\delta_{32} m_{\text{pole}} + (1 + \delta_{32}) \delta m\} \bar{\Psi}_r \Psi_r}_{\text{treat as a part of } \mathcal{L}_{\text{int}} \text{ (along with } (1 + \delta_{32}) \bar{\Psi}_r \gamma^\mu \Psi_r A_\mu(-eQ)}
 \end{aligned}$$



We know already that  $\delta_{32} \sim \mathcal{O}(e^2)$

To be more explicit,

$$\begin{cases} A^{(1)}(p^2, m^2) = \frac{\alpha Q^2}{4\pi} \int_0^1 dx -2(1-x) \ln \left( \frac{(1-x)\Lambda^2 + x m^2 - x(1-x)p^2}{+x m^2 - x(1-x)p^2} \right) \\ B^{(1)}(p^2, m^2) = \frac{\alpha Q^2}{4\pi} \int_0^1 dx \quad \times m \ln \left( \frac{(1-x)\Lambda^2 + x m^2 - x(1-x)p^2}{+x m^2 - x(1-x)p^2} \right) \end{cases}$$

★  $p^2 = m_{\text{pole}}^2$  must be a solution to

$$(1 - A^{(1)}(p^2, m^2))^2 p^2 = (m + B^{(1)}(p^2, m^2))^2$$

Keeping terms of  $\mathcal{O}(1)$  and  $\mathcal{O}(\alpha)$  but not  $\mathcal{O}(\alpha^2)$ ,

$$m_{\text{pole}}^2 = m^2 + \left\{ 2m B^{(1)}(m^2, m^2) + 2m^2 A^{(1)}(m^2, m^2) \right\} + \mathcal{O}(\alpha^2)$$

$$m^2 = m_{\text{pole}}^2 - \left\{ 2m_{\text{pole}} B^{(1)}(m_{\text{pole}}^2, m_{\text{pole}}^2) + 2m_{\text{pole}}^2 A^{(1)}(m_{\text{pole}}^2, m_{\text{pole}}^2) \right\} + \mathcal{O}(\alpha^2)$$

- When we keep track of the difference  $(m^2 - m_{\text{pole}}^2)$  of order  $\alpha$ , we can use  $m^2$  and  $m_{\text{pole}}^2$  interchangeably for the arguments of  $A^{(1)}$  and  $B^{(1)}$ .
- $\left. \begin{array}{l} m_{\text{pole}}^2 \text{ is determined by } m^2, \Lambda^2 \text{ and } e^2. \\ m^2 \text{ is determined by } m_{\text{pole}}^2, \Lambda^2 \text{ and } e^2. \end{array} \right\}$

$$\Rightarrow m = m_{\text{pole}} + (\delta m) \quad \text{expressed in terms of } m_{\text{pole}}^2, \Lambda^2, e^2.$$

★ The residue at the pole is determined by

$$\frac{(1 - A(p^2, m^2)) \cancel{p^2} + \dots}{(1 - A(p^2, m^2))^2 p^2 - (m + B(p^2, m^2))^2} \Rightarrow \text{residue } (1 - A^{(1)}) \left| \frac{1 - A^{(1)}}{(1 - A^{(1)})^2} - 2 \frac{\partial A^{(1)}}{\partial p^2} \Big|_{(m_{\text{pole}}^2)} - 2m \frac{\partial B^{(1)}}{\partial p^2} \right|$$

$$Z_2^{(1)} = (1 + \delta Z_2^{(1)}) = 1 + A^{(1)}(m_{\text{pole}}^2, m_{\text{pole}}^2) + 2m_{\text{pole}}^2 \left( \frac{\partial A^{(1)}}{\partial p^2} \right) (m_{\text{pole}}^2, m_{\text{pole}}^2) + 2m_{\text{pole}} \frac{\partial B^{(1)}}{\partial p^2} (m_{\text{pole}}^2, m_{\text{pole}}^2)$$

We can use  $m^2$  and  $m_{\text{pole}}^2$  interchangeably in the arguments of  $A^{(1)}$ ,  $B^{(1)}$  etc., because we ignore  $\mathcal{O}(\alpha^2)$  terms in  $\delta Z_2$  in this calculation.

Then at  $\mathcal{O}(\alpha)$ , the 2-point function (correlator) becomes

$$\frac{i}{(\not{p} - m_{\text{pole}} + i\epsilon) - [\not{p} A^{(1)}(p^2, m_{\text{pole}}) + B^{(1)}(p^2, m_{\text{pole}})] + \delta_{22}(m_{\text{pole}}, \Lambda, \epsilon)} \not{p} - (\delta_{22} m_{\text{pole}} + \delta m)$$

The coefficients of  $\not{p}$  in the denominator:  $[1 - A^{(1)}(p^2, m_{\text{pole}}) + \delta_{22}(m_{\text{pole}}, \Lambda, \epsilon)]$

$$-A^{(1)}(p^2, m_p^2) + A^{(1)}(m_p^2, m_p^2) = \frac{\alpha Q^2}{4\pi} \int_0^1 dx \, x(1-x) \left\{ \ln \left( \frac{(1-x)\Lambda^2 + x m_p^2 - x(1-x)p^2}{+x m_p^2 - x(1-x)p^2} \right) \right. \\ \left. - \ln \left( \frac{(1-x)\Lambda^2 + x m_p^2 - x(1-x)m_p^2}{+x m_p^2 - x(1-x)m_p^2} \right) \right\}$$

→ finite in the  $\Lambda \rightarrow +\infty$  limit

The quantum corrected correlator remains finite

when expressed in terms of observed (physical) parameter  $m_{\text{pole}}$ .

✓ The resulting expression does not depend on the choice of the regularization method.

(strictly speaking, this is a desire/criterion than a theorem.)

insensitive to the high-energy theory.

"divergence from infinitely many UV DOFs  
has been swept under the carpet (rug)"  
(in "pole")