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The Holographic Principle for General Backgrounds

hep-th/9905177
hep-th/9906022

What is N_{dof} in nature?

- QFT + Planck cutoff $\rightarrow N_{\text{dof}} \sim V$
- thermodynamic entropy cannot exceed number of degrees of freedom: $S \leq N_{\text{dof}}$
- therefore $S_{\text{max}} \sim V$? No!
- with gravity, most states collapse to black holes
- Bekenstein (1981): $S \leq A/4$
- 't Hooft (1993): $N_{\text{dof}} \leq A/4$

Holographic Principle: $N_{\text{dof}} \leq A/4$

“the world is two-dimensional”

- required by quantum gravity
- should be manifest in the fundamental theory
- implies that space and time are derived concepts?
- strongly supported by AdS/CFT correspondence

Maldacena; Gubser, Klebanov+Polyakov; Witten;
Susskind+Witten; ...

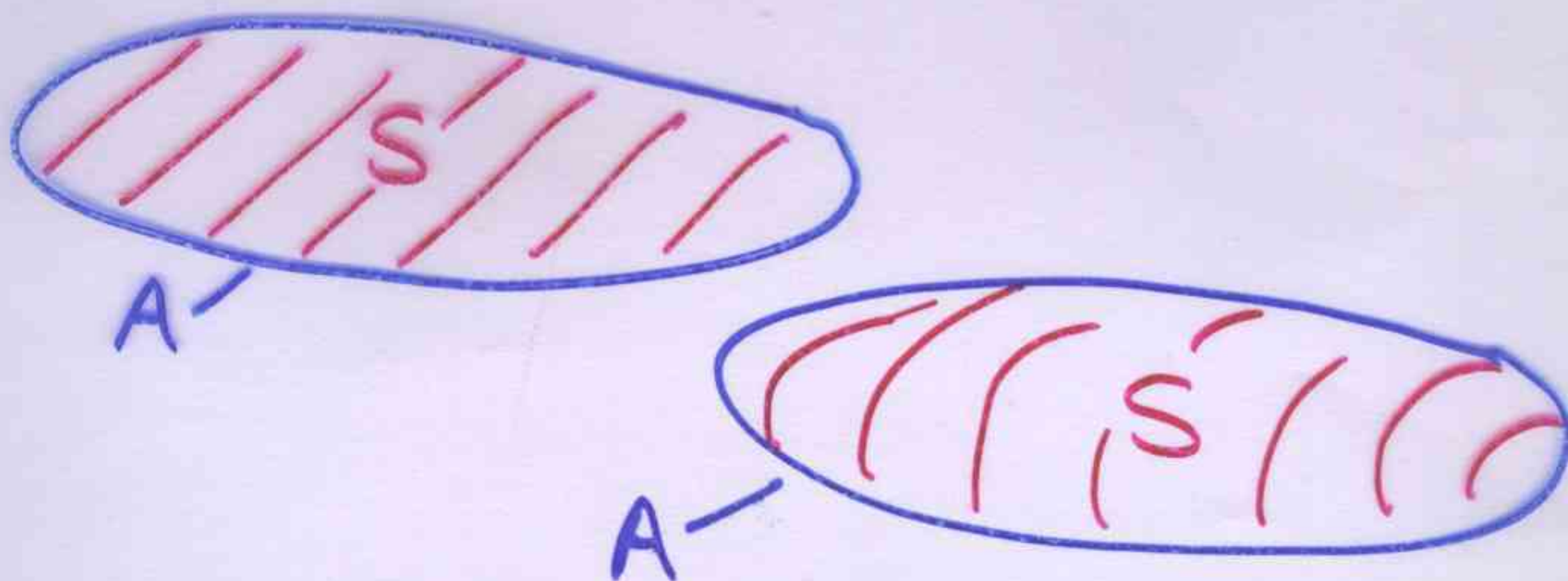
The Problem

- if truly fundamental, H.P. should apply to all solutions of Einstein's equations
- but Bekenstein's bound is valid only for "regions of limited self-gravity."
- holographic principle violated in
 - gravitationally collapsing systems
 - large cosmological regions

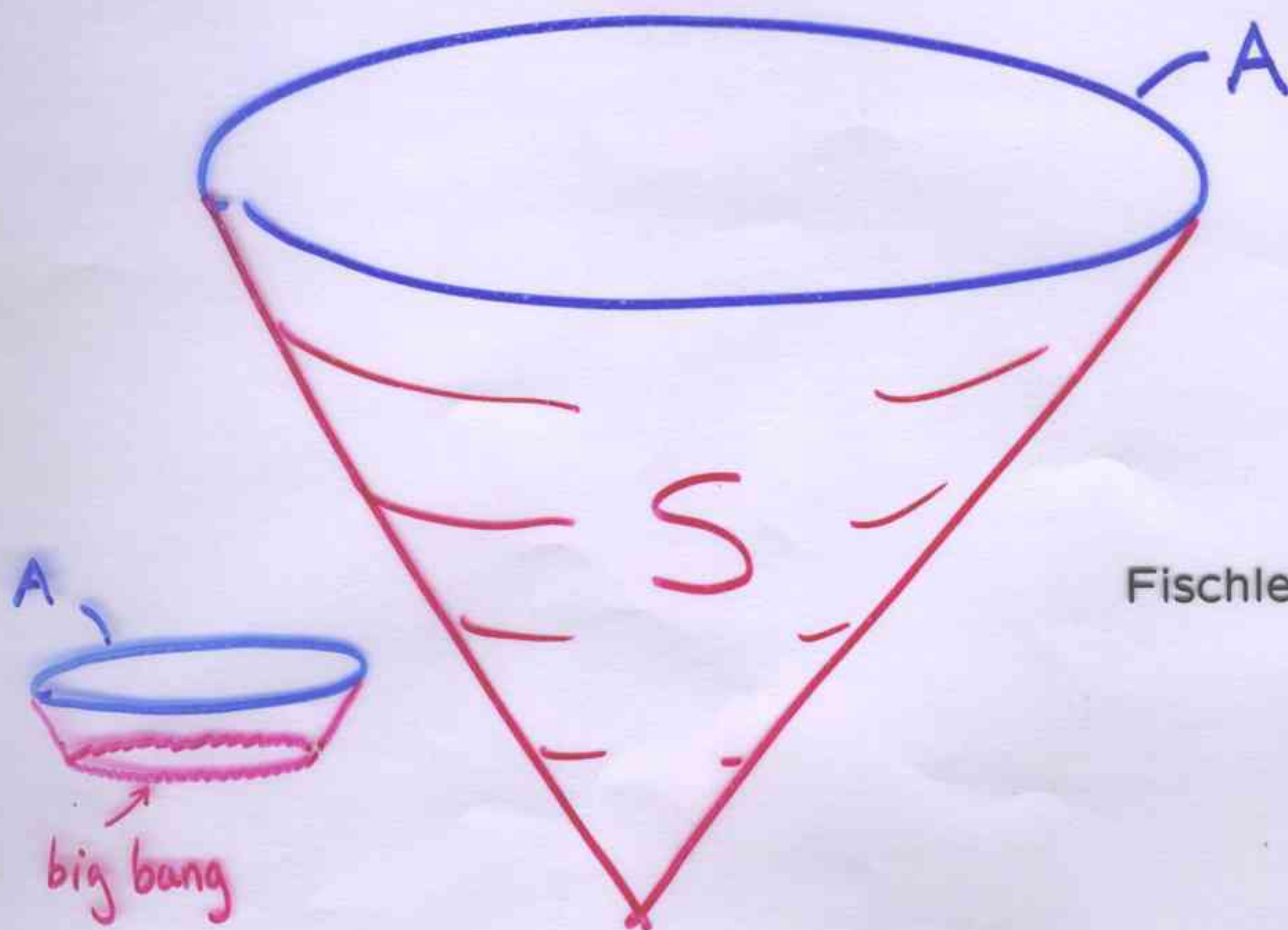
Fischler+Susskind, hep-th/9806039

- \Rightarrow a more general formulation is needed

Demand Covariance!

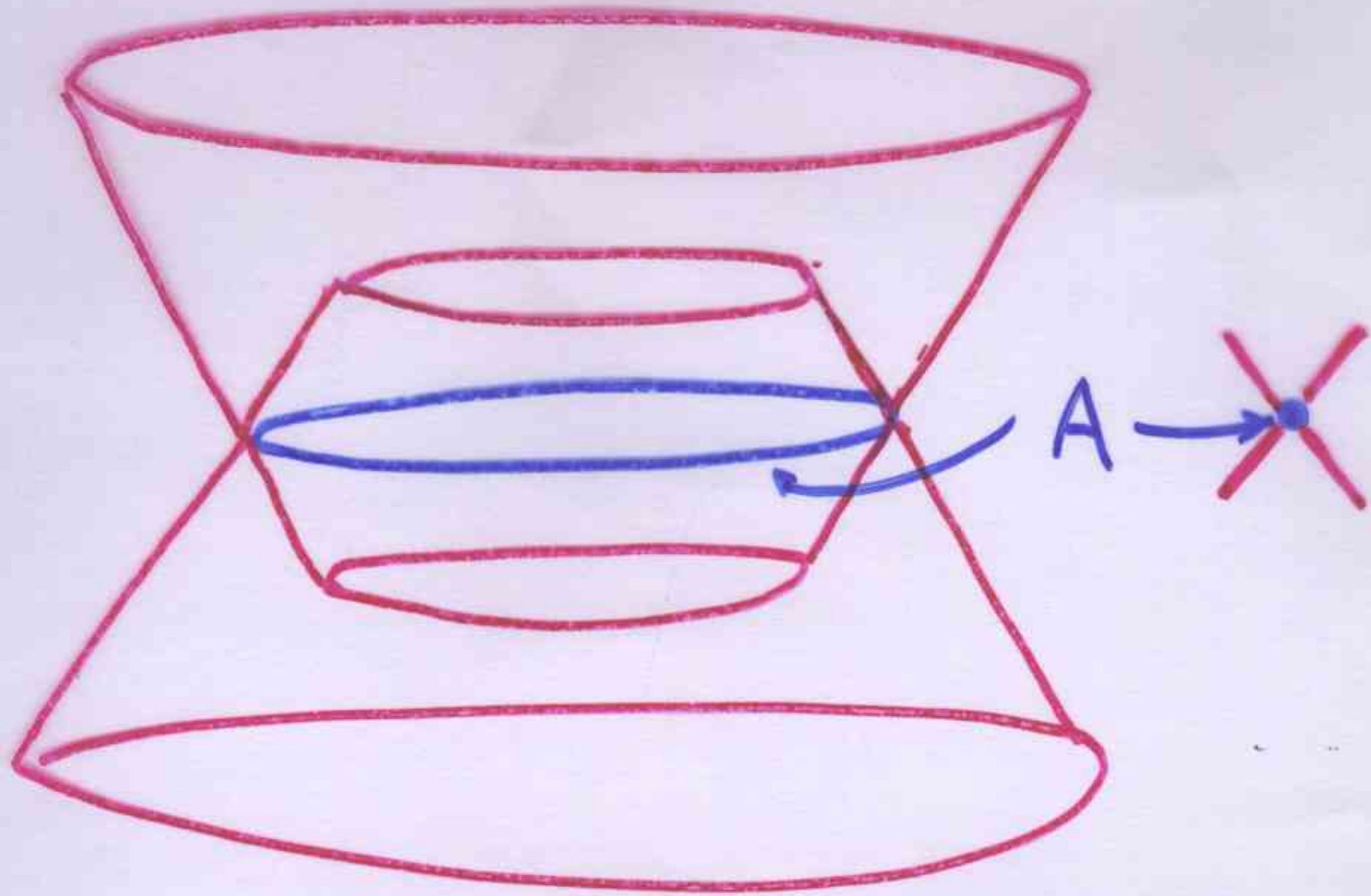


- spacelike hypersurfaces excluded
- try light-like hypersurface



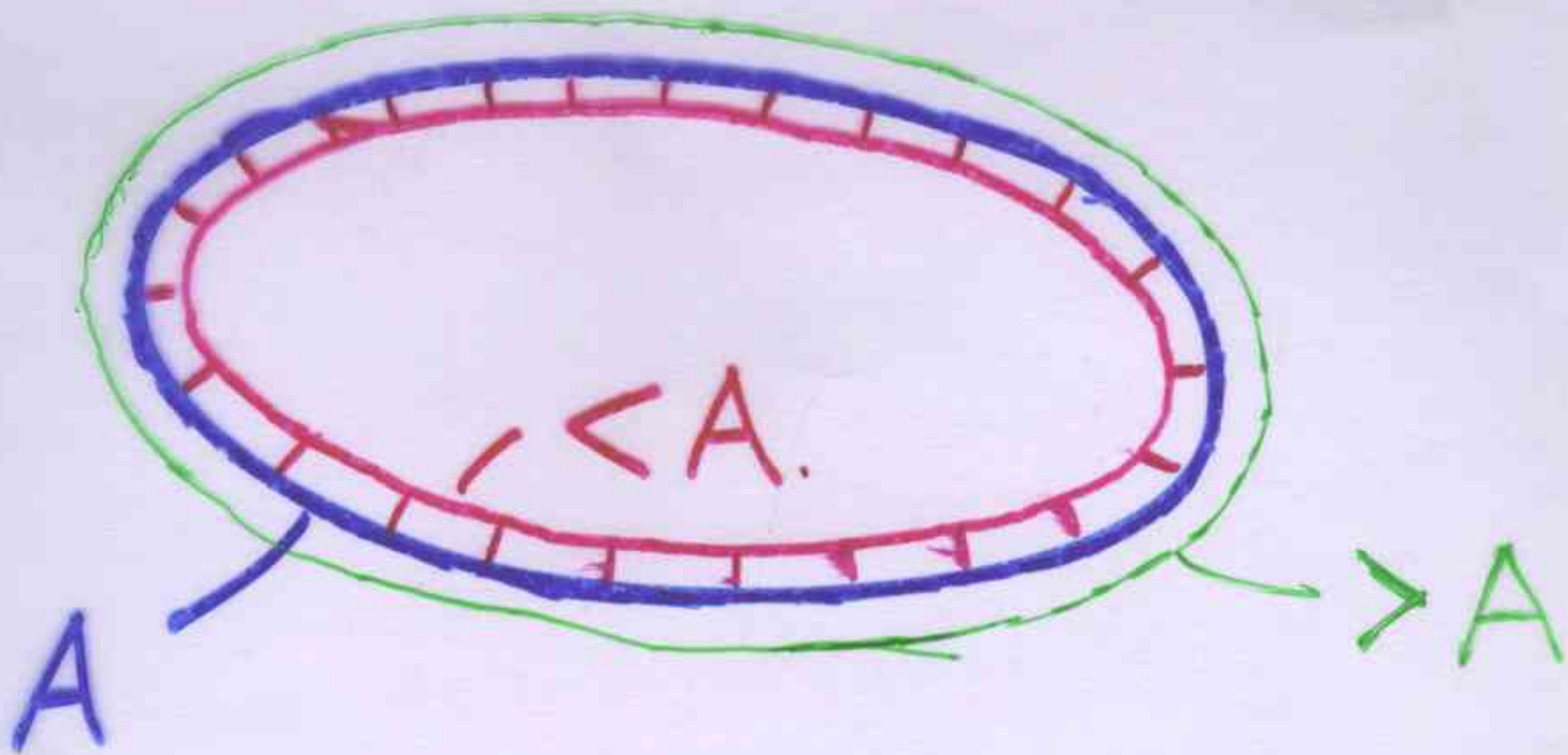
Fischler+Susskind

4 choices







- orthogonal light-rays generate 4 null hypersurfaces
- select a region *inside* the area
- need covariant definition of "inside"

Where is "inside"?



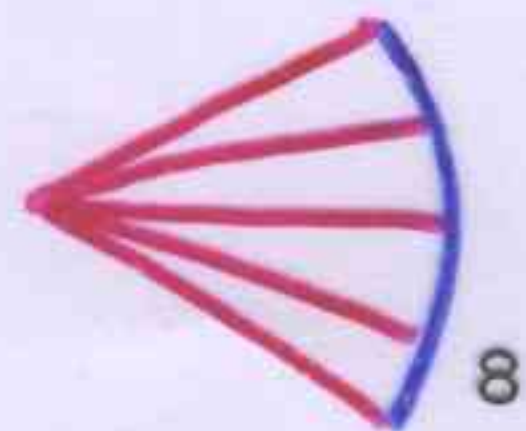
- cross-sectional area must decrease: $\theta \leq 0$

-  trapped:  anti-trapped:  normal: 

- $\theta \leq 0 \Rightarrow$ must stop at caustics



- surface need not be closed



Covariant Entropy Bound

- pick a surface, area A
- find the “allowed” families of light-rays
(at least 2 out of 4)
- pick an allowed family
- construct **light-sheet** by following
each light-ray until $\theta > 0$

$$S \text{ (light-sheet)} \leq A/4$$

Properties of the Bound

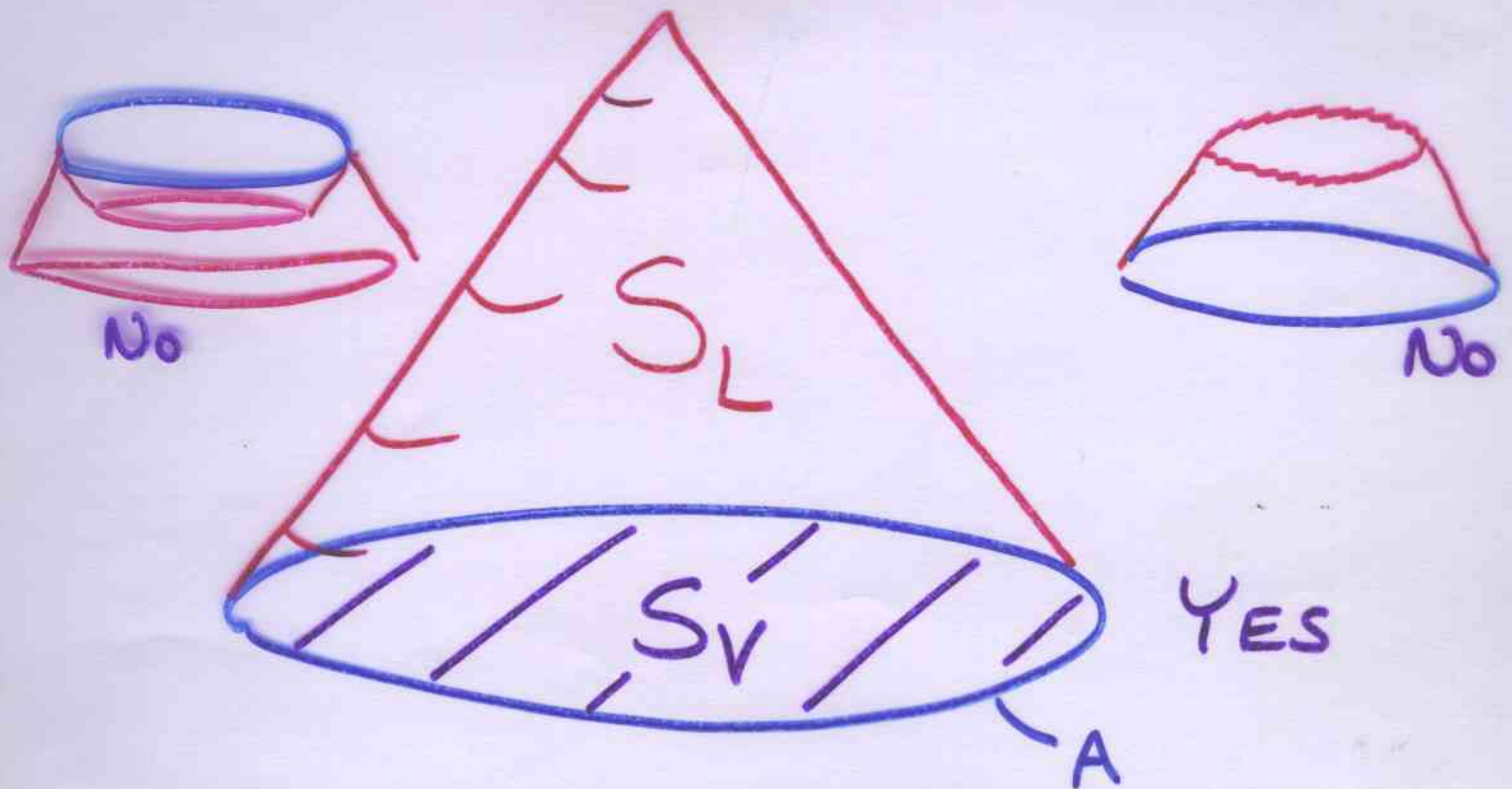
- democratic treatment of all 4 null directions
- no reference to inside/outside
- no reference to past/future:
time reversal invariant

The Holographic Principle for General Spacetimes

$$N_{\text{dof}} (\text{light-sheet}) \leq A/4$$

The Entropy in Spatial Regions

- When are spacelike hypersurfaces permitted?



- causality + second law $\Rightarrow S_V \leq S_L \leq A/4$

- \Rightarrow recover Bekenstein's bound

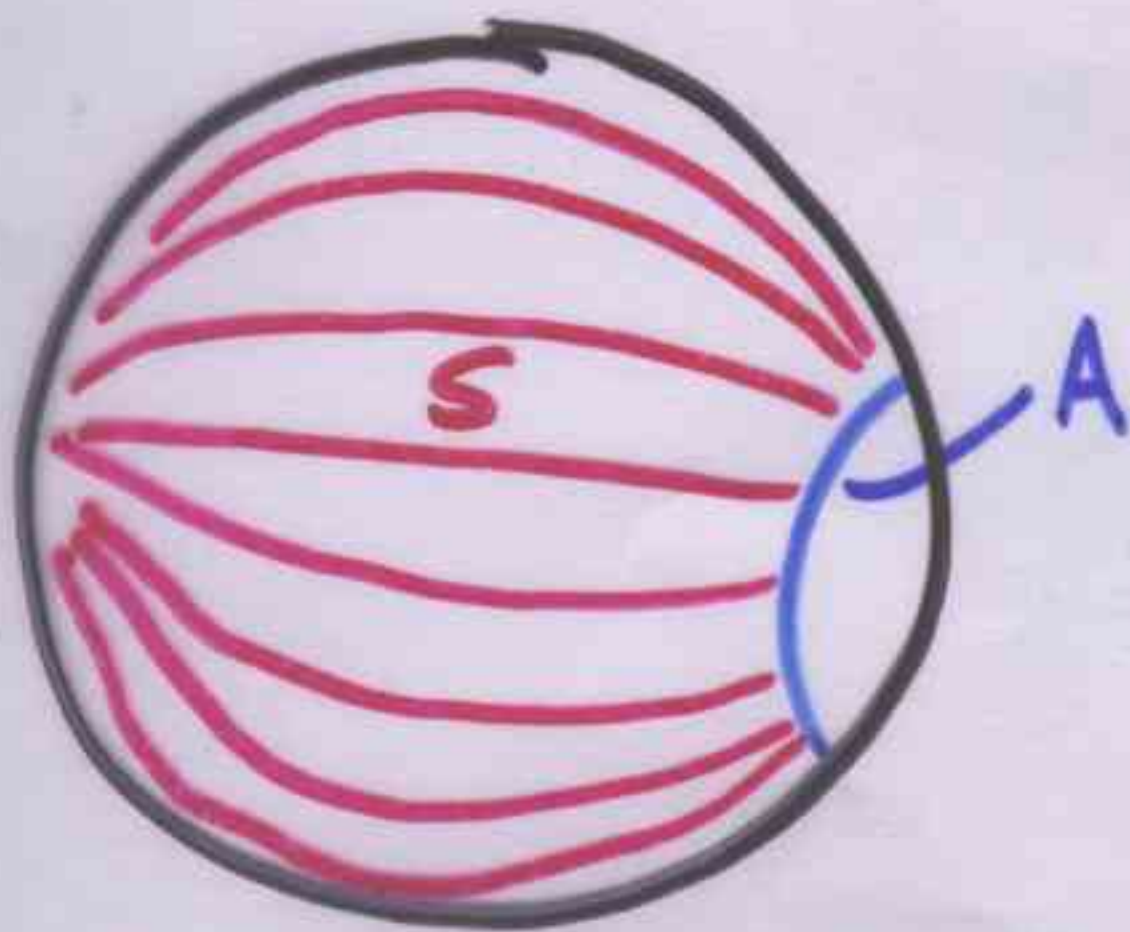
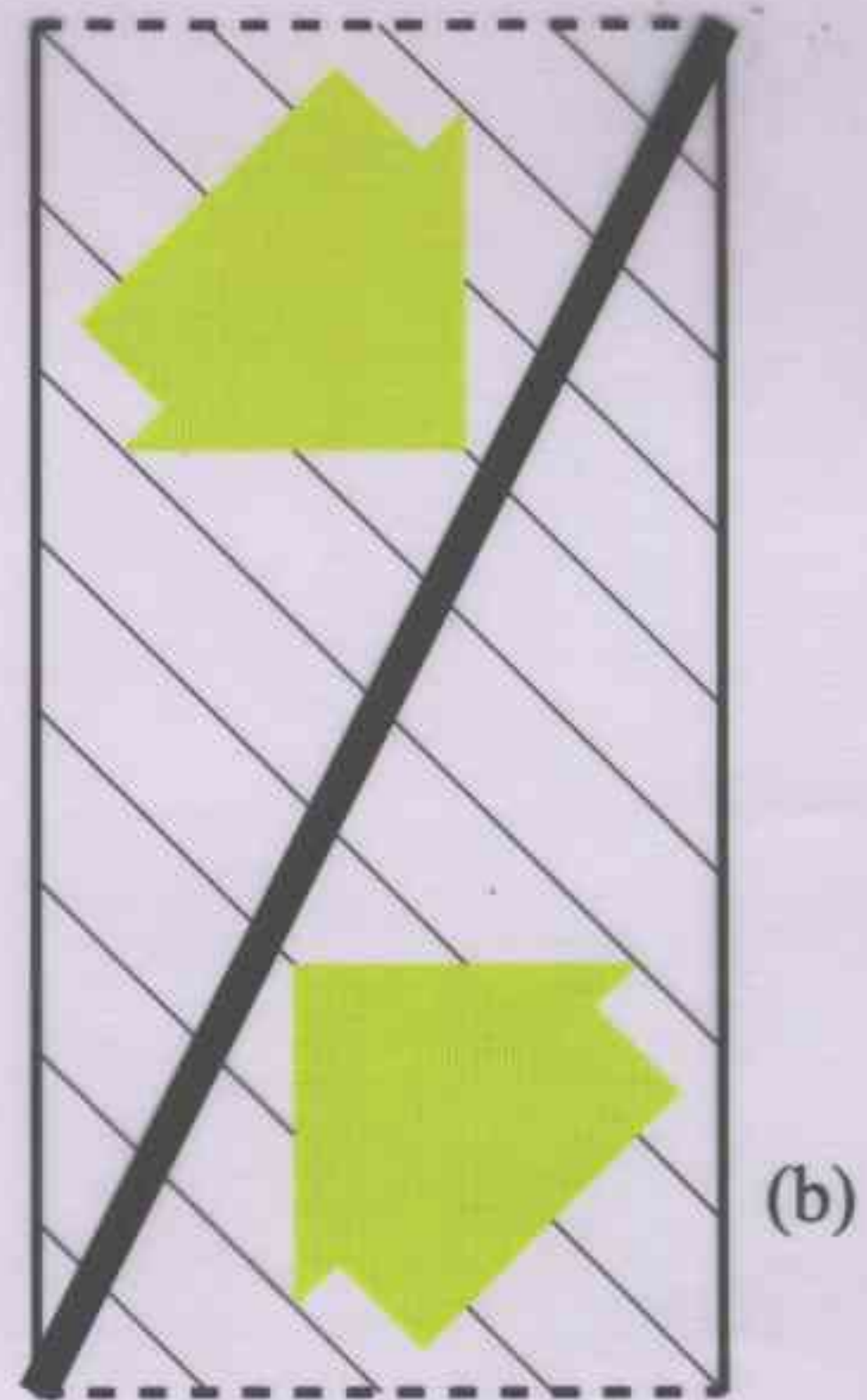
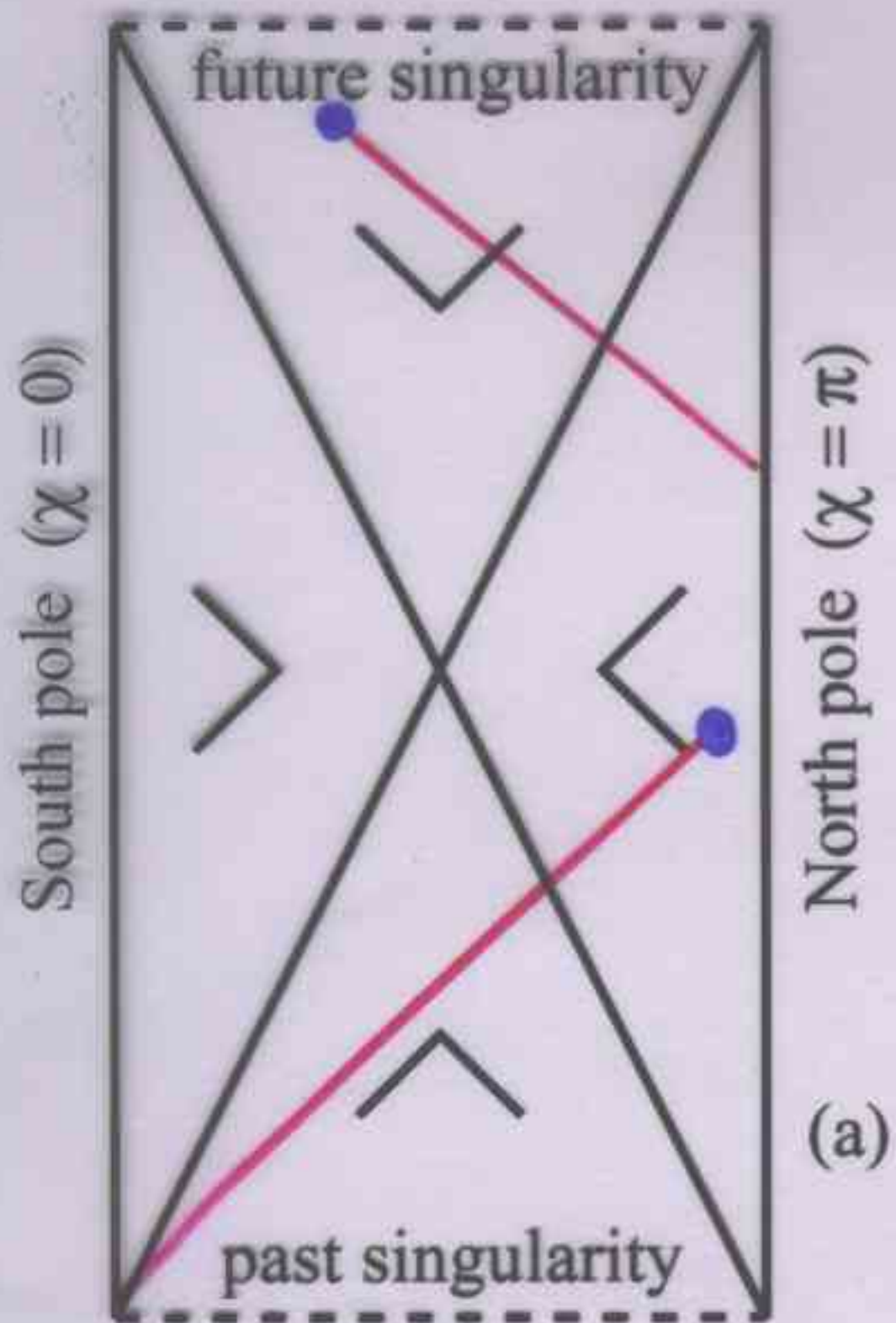
Does it work?

- no straightforward proof
- need comprehensive, non-trivial tests, especially for gravitational collapse

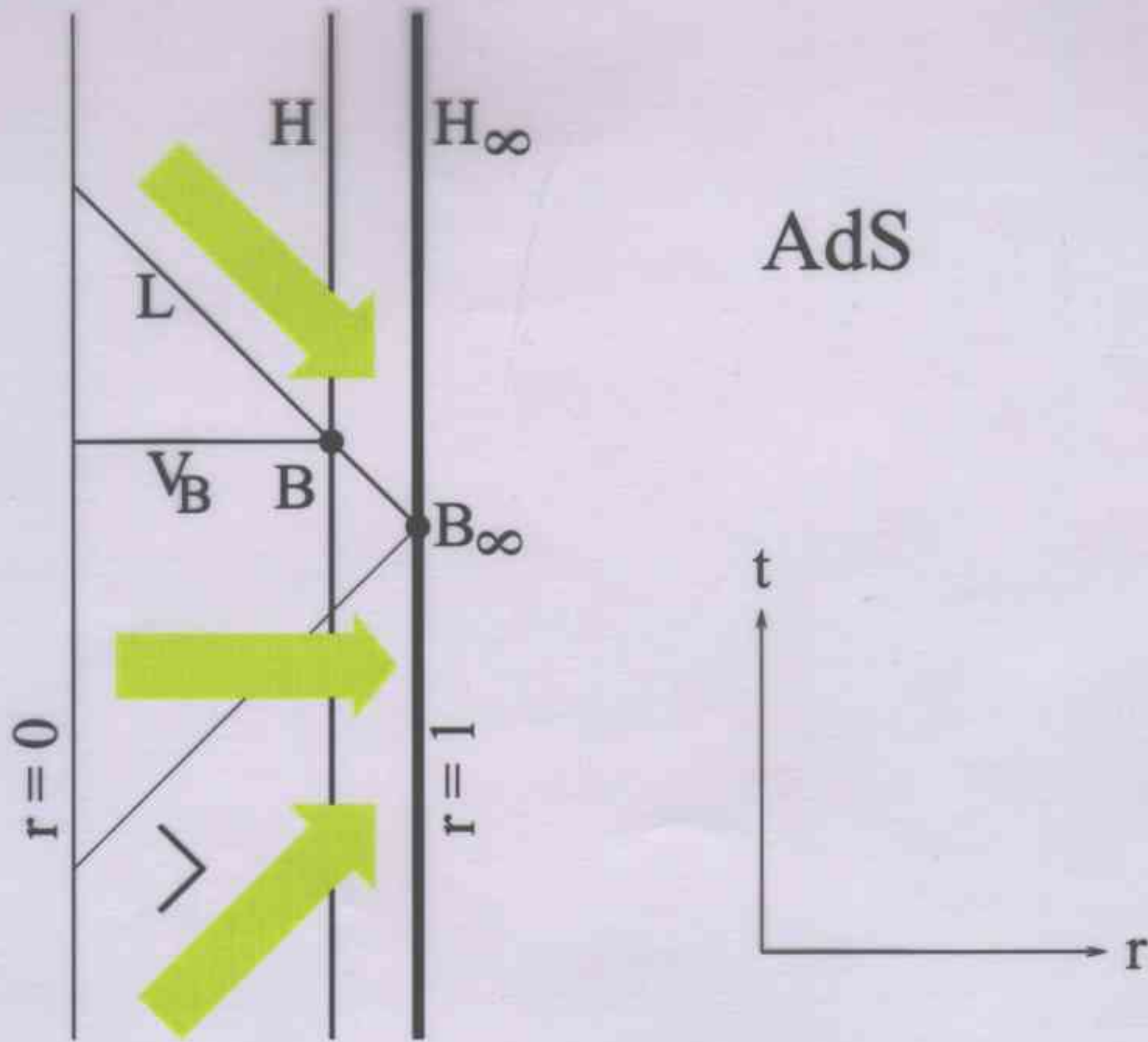
R.B., hep-th/9905177

- worst case assumptions \rightarrow saturation

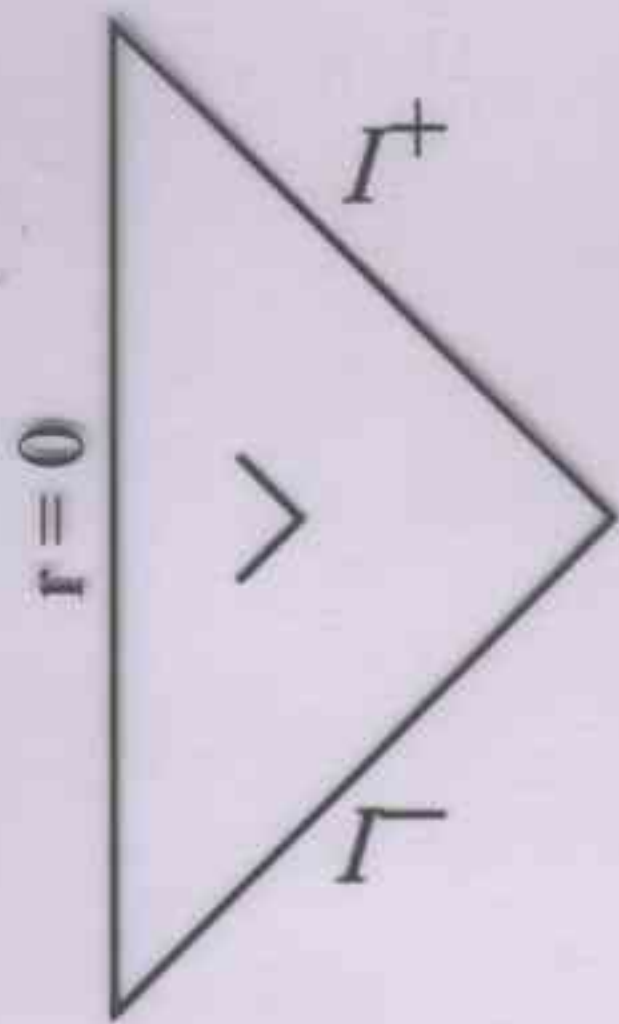
Closed FRW Universe



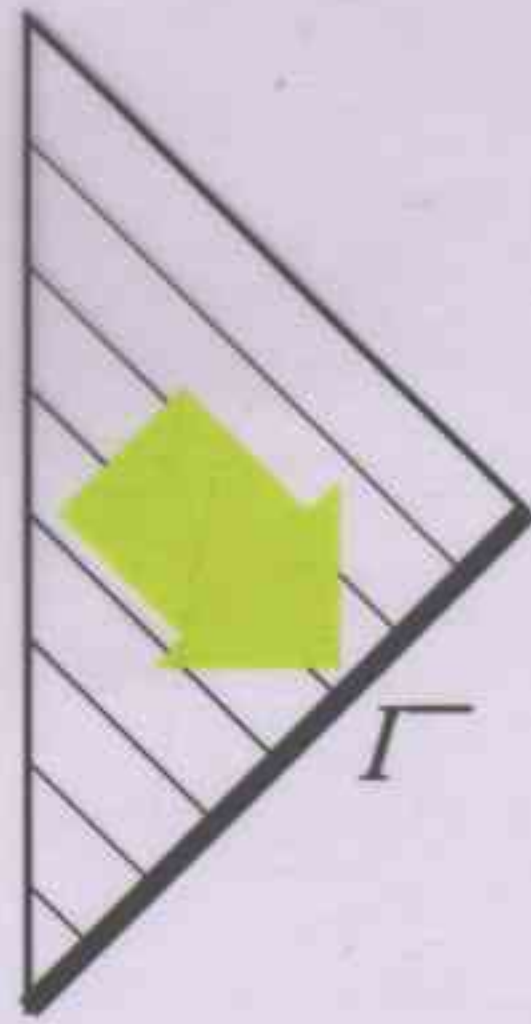
Anti-de Sitter Space



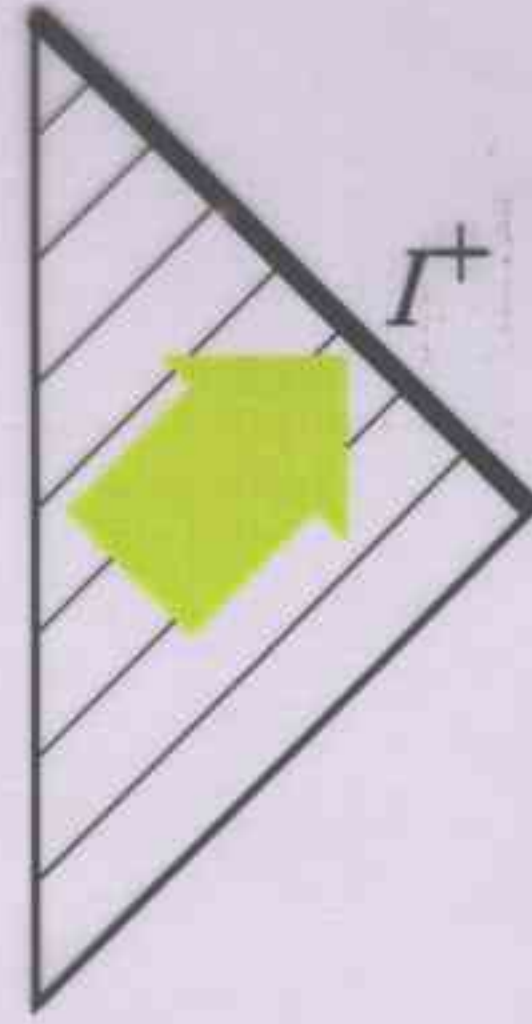
Minkowski Space



(a)

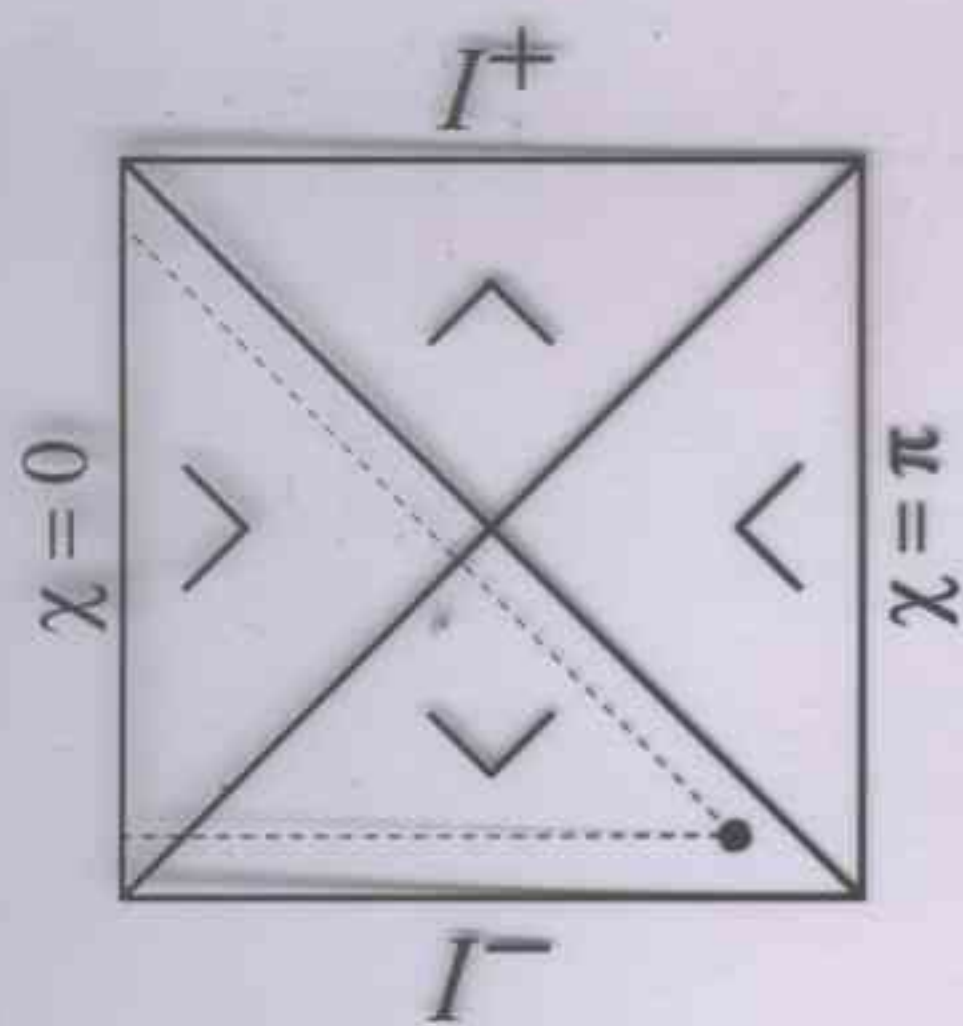


(b)



(c)

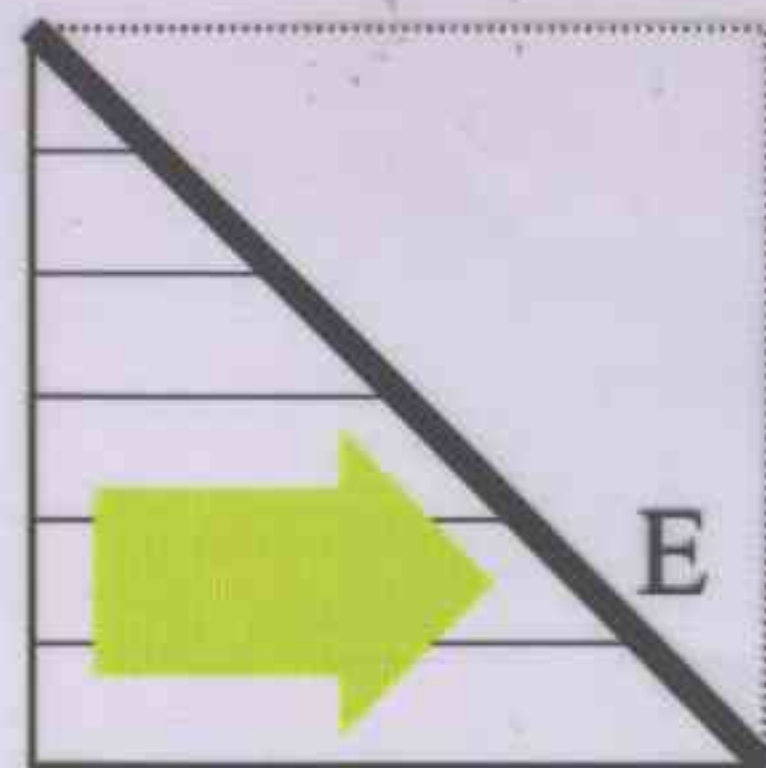
de Sitter Space



(a)



(b)



(c)

Properties of Screens

- can be timelike, null, spacelike, or mixed
- area can be time-dependent
- there may be several disconnected screens
- spacelike projection not normally possible
- AdS very special

What is the Holographic Theory?

- a quantum field theory on the screen, as in AdS/CFT?
- generically, this possibility excluded by the screen properties
- unconventional “theory” on screen?
 - = has varying N_{dof}
 - = lives variously in Euclidean/Lorentzian time

Speculations

- screen-theory approach cannot be fundamental (presupposes geometric structure)
- turn holographic principle around?
 - = hol. principle: "geometry \rightarrow entropy"
 - = hol. theory: "entropy \rightarrow geometry"
- two tasks:
 - = construct theory without gravity with varying N_{dof}
 - = reconstruct space-time from N_{dof}