

CHIRAL GRAVITY IN THREE DIMENSIONS

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^{plus}
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So far, string theory is the only known consistent quantum theory of gravity complex enough to contain black holes. Simpler examples - e.g. in 2 or 3 dimensions - would be useful. Oddly, despite 30 years of research, it remains unclear if such exist!

In this talk we explore a new possibility:
"CHIRAL GRAVITY"

At first, 3D Einstein gravity³

$$S_E = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left(R + \frac{2}{l^2} \right)$$

appears trivial because

$$R_{abcd} = R_{ac}g_{bd} + \text{symm}$$

But then BTZ found ●.

In the late 80's an exact solution was proposed, but it did not explain ● entropy and was reconsidered last year.

OPEN QUESTION

Is there a quantum theory of pure 3D Einstein gravity?

No viable current proposal.

Another theory in 3D ⁴

Topologically Massive Gravity

Deser Jackiw & Templeton

$$S_{\text{TMG}} = \frac{1}{16\pi G} \left(R + \frac{3}{\lambda^2} + \frac{1}{\mu} I_{\text{CS}} \right)$$

\uparrow graviton mass

$$I_{\text{CS}} = -\frac{1}{2} \epsilon^{\lambda\mu\nu} \Gamma_{\lambda\sigma}^{\rho} \left[\partial_{\mu} \Gamma_{\rho\gamma}^{\sigma} + \frac{2}{3} \Gamma_{\mu\zeta}^{\sigma} \Gamma_{\nu\rho}^{\zeta} \right]$$

E.O.M.

$$G_{\mu\nu} + \frac{1}{\lambda^2} g_{\mu\nu} = \frac{1}{\mu} \epsilon_{\mu}{}^{\sigma\gamma} \nabla_{\alpha} (R_{\sigma\gamma} - \frac{R}{4} g_{\sigma\gamma})$$

Every solution of Einstein's equation

$$G_{\mu\nu} = -\frac{1}{\lambda^2} g_{\mu\nu}$$

is a solution of TMG,
but there are more....

MASSIVE GRAVITONS

Consider an expansion about flat space with $\frac{1}{\lambda^2} = 0$. The higher derivative Ics adds propagating degrees of freedom: a massive graviton with positive mass μ but negative energy $E < 0$. Its like a wrong-sign scalar

$$S = \int dt \int d^3x (\nabla \phi)^2 + m^2 \phi^2$$

~~Take $G < 0$?~~ Instead \rightarrow

AdS₃ TMG

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We wish to study the case of negative c.c. expanded about AdS₃

$$ds^2 = \ell^2 (-\cosh^2 \rho d\tau^2 + d\rho^2 + \sinh^2 \rho d\phi^2)$$

with Brown-Henneaux b.c.s.

This has 6 $SL(2, \mathbb{R})_L \times SL(2, \mathbb{R})_R$

Killing vectors

$$L_0 = \partial_\tau - \partial_\phi$$

$$\bar{L}_0 = \partial_\tau + \partial_\phi$$

$$L_1 = \dots$$

$$\bar{L}_1 = \dots$$

$$L_{-1} = \dots$$

$$\bar{L}_{-1} = \dots$$

Gravitons

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linearized soln's
fall into $SL(2, \mathbb{R})_R \times SL(2, \mathbb{R})_L$
irreps (highest weight)
for generic m, l .

HWS

$$L_0 |4\rangle = h_L |4\rangle$$

$$\bar{L}_0 |4\rangle = h_R |4\rangle$$

$$L_{\pm 1} |4\rangle = \bar{L}_{\pm 1} |4\rangle = 0$$

After much work, one
finds

$$(h_L, h_R)$$

$(0, 2)$ right boundary graviton

$(2, 0)$ left bdy graviton

$(\frac{1}{2} + \frac{3m^2}{2}, \frac{m^2}{2} - \frac{1}{2})$ massive graviton

Note $m^2 \rightarrow 1$ degeneracy!!
Also $E \rightarrow 0$.

Asymptotic Symmetry Group

Given any set of b.c.s,
in our case Brown-Henneaux

$$g_H = -e^2 P + \mathcal{O}(1)$$

⋮
⋮
⋮

there is a set of **allowed** diffeos which preserve ^{them}. Given some dynamics, a subset can be labelled as **trivial** if they vanish, including surface terms, when the constraints are applied. Then

$$ASG \equiv \frac{\text{allowed diffeos}}{\text{trivial diffeos}}$$

Physical states are in reps of the **ASG**, and are annihilated by trivial diffeos.

TMG ASG

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Generically, the ASG for TMG is generated by

Left Virasoro

$$c_L = \frac{3g}{2G} \left(1 - \frac{1}{\mu l}\right)$$

Right Virasoro

$$c_R = \frac{3g}{2G} \left(1 + \frac{1}{\mu l}\right)$$

Brown Henneaux
Kraus Larsen

with explicit generators

$$G(\epsilon^+, \epsilon^-) = \oint_{\infty} dz^+ h_{++} \epsilon^+ \left(1 + \frac{1}{\mu l}\right) + \oint_{\infty} dz^- h_{--} \epsilon^- \left(1 - \frac{1}{\mu l}\right)$$

metric fluctuation \nearrow \nwarrow diffeo parameter on boundary

$$\epsilon^{\pm} = \epsilon^{\pm}(z \pm \phi)$$

FOR $\mu l = 1$, ϵ^- TRANSFORMATIONS ARE TRIVIAL!

So

$M_2 \rightarrow 1$
 $ASG \rightarrow$ Right Virasoro

Left + Massive ^{Also} \rightarrow Null states
 Gravitons
 BTZ black holes $\rightarrow E = J$
 $C_L \rightarrow 0$

basis of LSS conjecture

TMG becomes chiral.

There is no left Virasoro for states to transform under.

// CHIRAL GRAVITY //

Open Question:
 $E+J > 0$???

Conjecture

Chiral gravity
exists as a quantum
theory and is dual to
a $c_R = \frac{3g}{G}$ holomorphic
boundary CFT.

N.B. If the CFT is unitary,
its microstates explain the
black hole entropy via the
Cardy formulae as in A.S.'98.

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A number of
purported counterexamples
appeared in the literature
after the chirality conjecture
was made ^{Lij Song, A.S.} but before it
was proven, ^{A.S.}. These all
fail to be true counter-
examples and do not
take into account the
fact that left Virasoro
transformations are
pure gauge at the chiral
point.

ASIDE

Warped AdS₃ ^{salvation¹²} for $\mu \neq 1$?

For $\mu \neq 1$, TMG on

AdS₃ is unstable. Could there be another ground state? \exists interesting warped WAdS₃ solutions

$$ds^2 = L^2 \left(-\frac{dt^2 + dy^2}{y^2} + \lambda^2 (d\phi + \frac{dt}{y})^2 \right)$$

← Warp factor

These have interesting black hole solutions which are quotients of WAdS₃ just as BTZ = AdS₃/Z. A rich and interesting story is unfolding. . . .

- Gurles Nuthu
- Clement Leygnac
- Padi Aninos Li Song A.S. Ait Mousse
- Guennoune Bouchateb. . . .

Relation to Witten's 2007¹⁴
"3D Gravity Reconsidered"

Proposed end run for pure gravity on AdS_3

Consistency requirements

Assumptions

1. Modular invariance for $Z(\tau, \bar{\tau})$

2. No primaries below lightest mass

$$1. Z(\tau) = Z(\tau)Z(\bar{\tau})$$

← CHIRAL GRAVITY

⇒ exact formula for Z in terms of Hecke transforms of j -functions.

Sum over Euclidean

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3-geometries can be done exactly and yields

Yin Maloney Witten

$Z \neq Z(\tau)Z(\bar{\tau})$ for pure gravity.

For chiral gravity, though some details remain to be worked out, the Euclidean action is complex and appears to yield precisely the desired result. Moreover, the integrality of the coefficients in the q -expansion of Z is nontrivial evidence for the existence of quantum chiral gravity.

But what does
this have to do
with the real
world?

The Kerr-CFT
Correspondence

Monica Guica, Tom Hartman,
Wei Song & A.S.

in progress

An extreme 4D

Kerr  black hole
has angular momentum

$$J = GM^2$$

Hawking temperature

$$T_H = 0,$$

and Bekenstein-Hawking
entropy

$$S_{BH} = \frac{2\pi J}{\hbar}.$$

GRS 1915+105 has

$$M \sim 14 M_{\odot}$$

$$\frac{J}{GM^2} > .98$$

McClintock, Shafee, Narayan,
Remillard, Davis & Li (2006)

The near-horizon region¹⁹
of extreme Kerr



is like chiral gravity:
all excitations must move
counterclockwise at the
speed of light!

Bardeen & Horowitz (99)
found the Near-Horizon
Extreme-Kerr Nhek geometry

$$ds^2 = 2GJ\Omega^2 \left[-\frac{dt^2 + d\psi^2}{\gamma^2} + d\theta^2 + \Lambda^2 \left(d\phi + \frac{d\psi}{\gamma} \right)^2 \right]$$

$$\Lambda^2 = \frac{2\sin\theta}{1+\cos^2\theta}$$

$$\Omega^2 = \frac{1+\cos^2\theta}{2}$$

and found an $SL(2, \mathbb{R}) \times U(1)$
isometry. Cross sections
at fixed polar angle θ
are precisely the $WAdS_3$
geometries discussed
earlier! Expect CFT...

We have found that for suitable b.c.s the **ASG** contains a Virasoro with

$$c_L = \frac{12J}{\hbar}$$

Further, the Frolov-Thorne vacuum is a thermal state with energy precisely E_0 of the Virasoro at temperature

$$T_L = \frac{1}{2\pi}$$

the Cardy formula

$$S_{\text{CFT}} = \frac{\pi^2 c_L T_L}{3}$$

gives

$$S_{\text{CFT}} = \frac{2\pi J}{\hbar} = S_{\text{BH}}!!!$$

Conjecture

The black hole
GRS 1915+105 is
dual to a $c_L = 2.0 \times 10^{76}$
two-dimensional conformal
field theory.