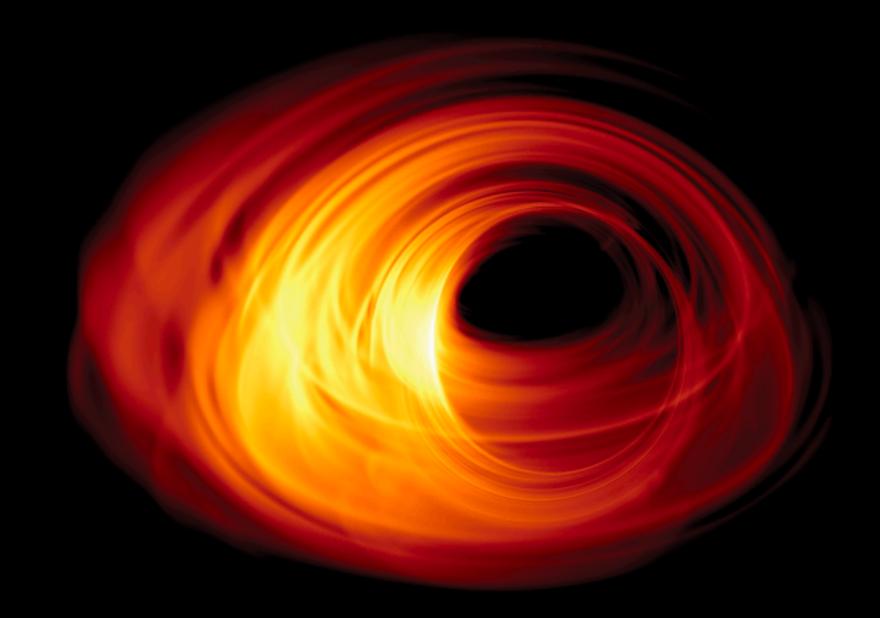


Event Horizon Telescope

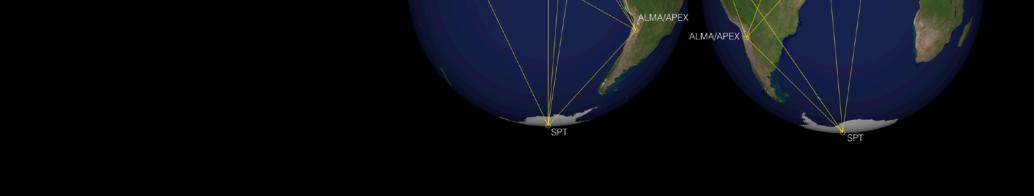


First results and potential African expansion of the Event Horizon Telescope



Roger Deane

University of Pretoria
Visiting Fellow, Rhodes University





Collaboration Teleffills

EXERT FISTIZOR TELESCOPE COMPRESSOR TELESCOPE



Collaborate Teleffills

Coligionation Mariffig

EXELL LIBLIZON TELESCOPE

outline

- EHT Overview: the instrument and science goals
- First imaging results on M87:
- The future: upcoming results and array expansion





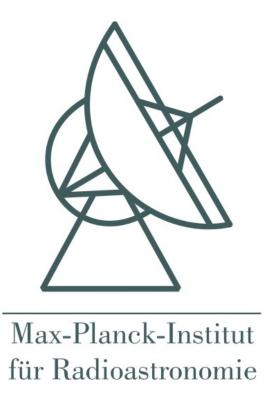
ames Clerk Maxwell Telescope
East Asian Observatory















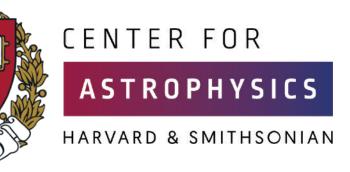
















Brandeis University



Los Alamos Caltech CIFAR















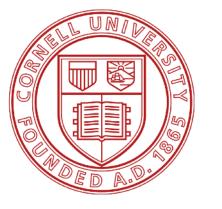






















INADE



































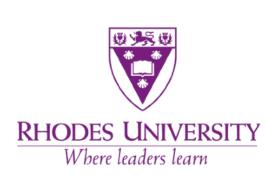


The Institute of Statistical Mathematics



Netherlands Organisation for Scientific Research

















K E N D















S O





WATERLOO













EHT Associated Logos



















del Personal Académico

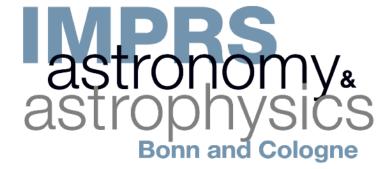














Japan Society for the Promotion of Science 本学術振興会













1986





























Netherlands Organisation for Scientific Research





Foundation of Korea





Innovation, Science and Economic Development Canada

RESEARCH AND INNOVATION

Innovation, Sciences et Développement économique Canada



Swedish Research Council



MINISTERIO DE ECONOMÍA Y COMPETITIVIDAD







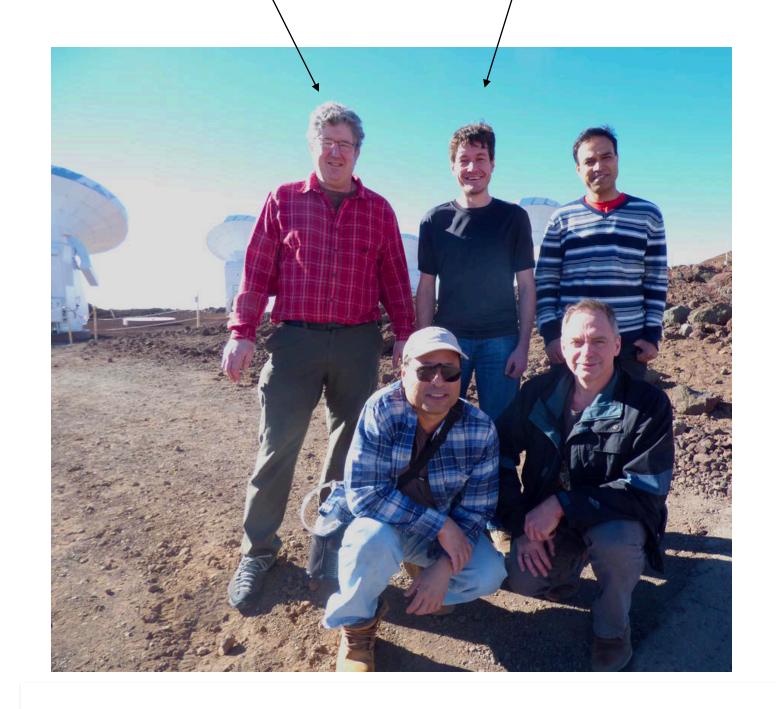




South African EHT involvement

Andre Young

Jonathan Weintroub



Iniyan Natarajan



Tariq Blecher



Micaela Menegaldo







RPD

Heinrich van Deventer

Based at EHT stakeholder institutions

Based in South Africa

Radio interferometry 101

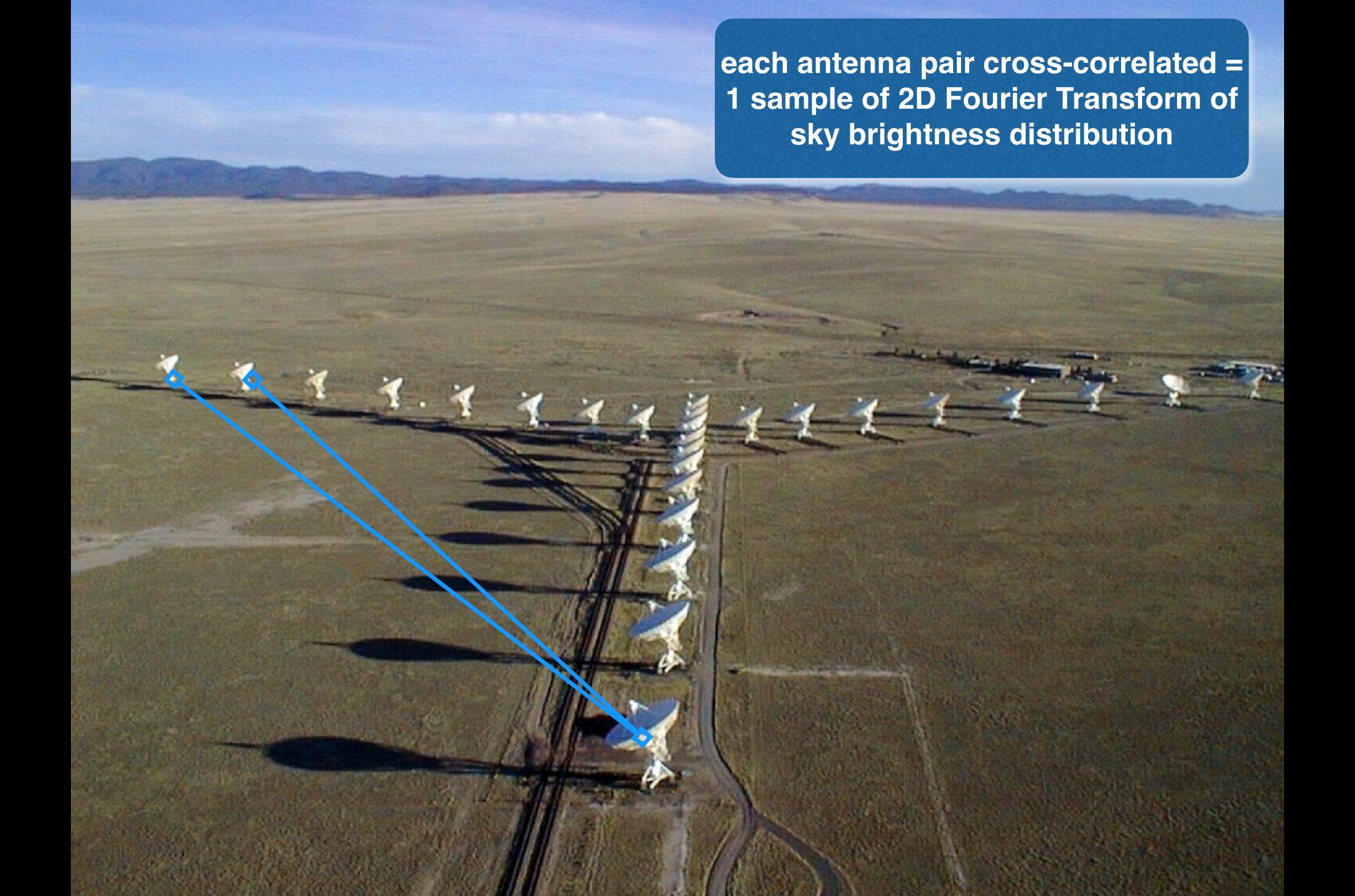
Radio interferometers are not digital cameras

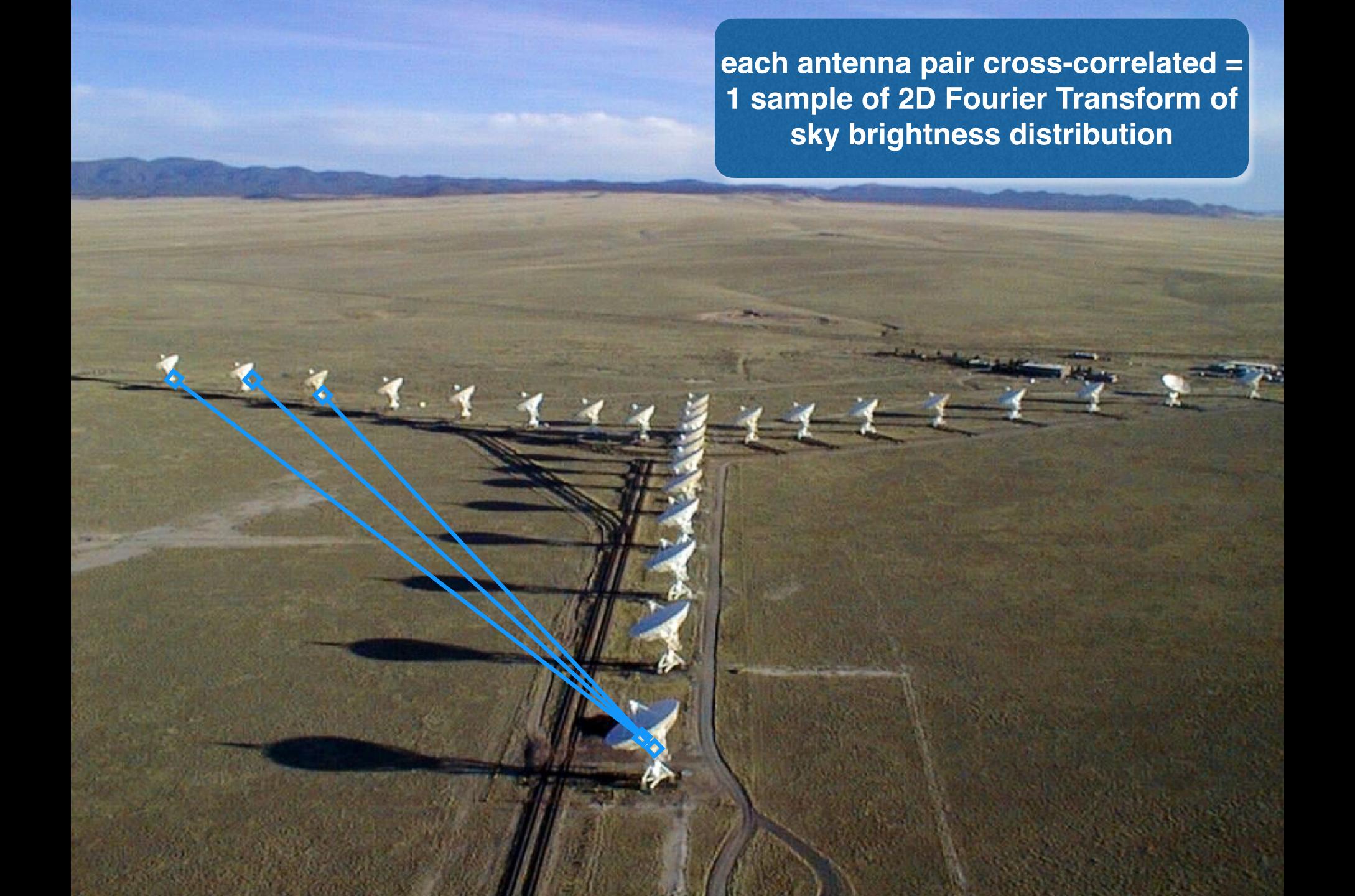


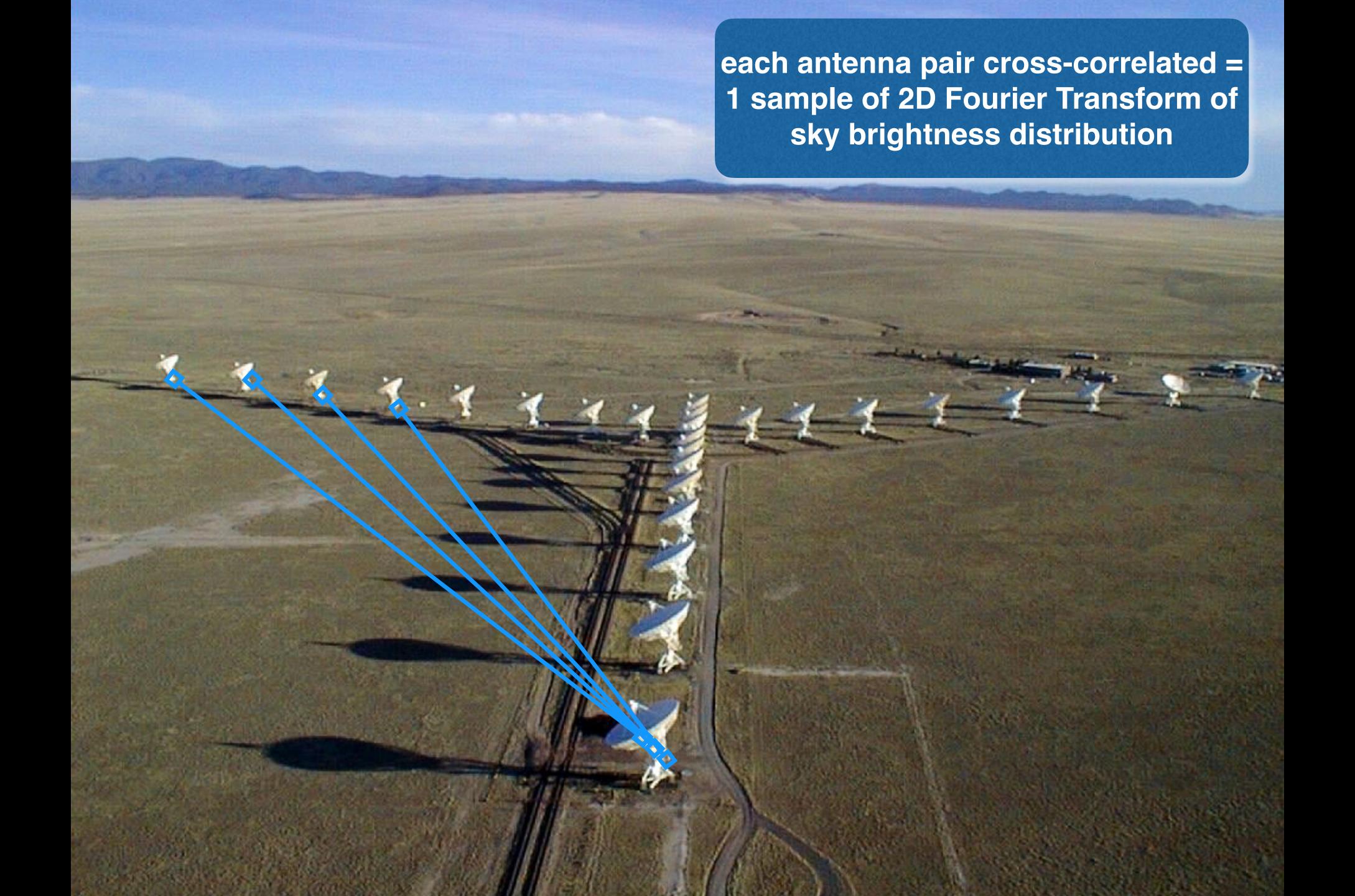


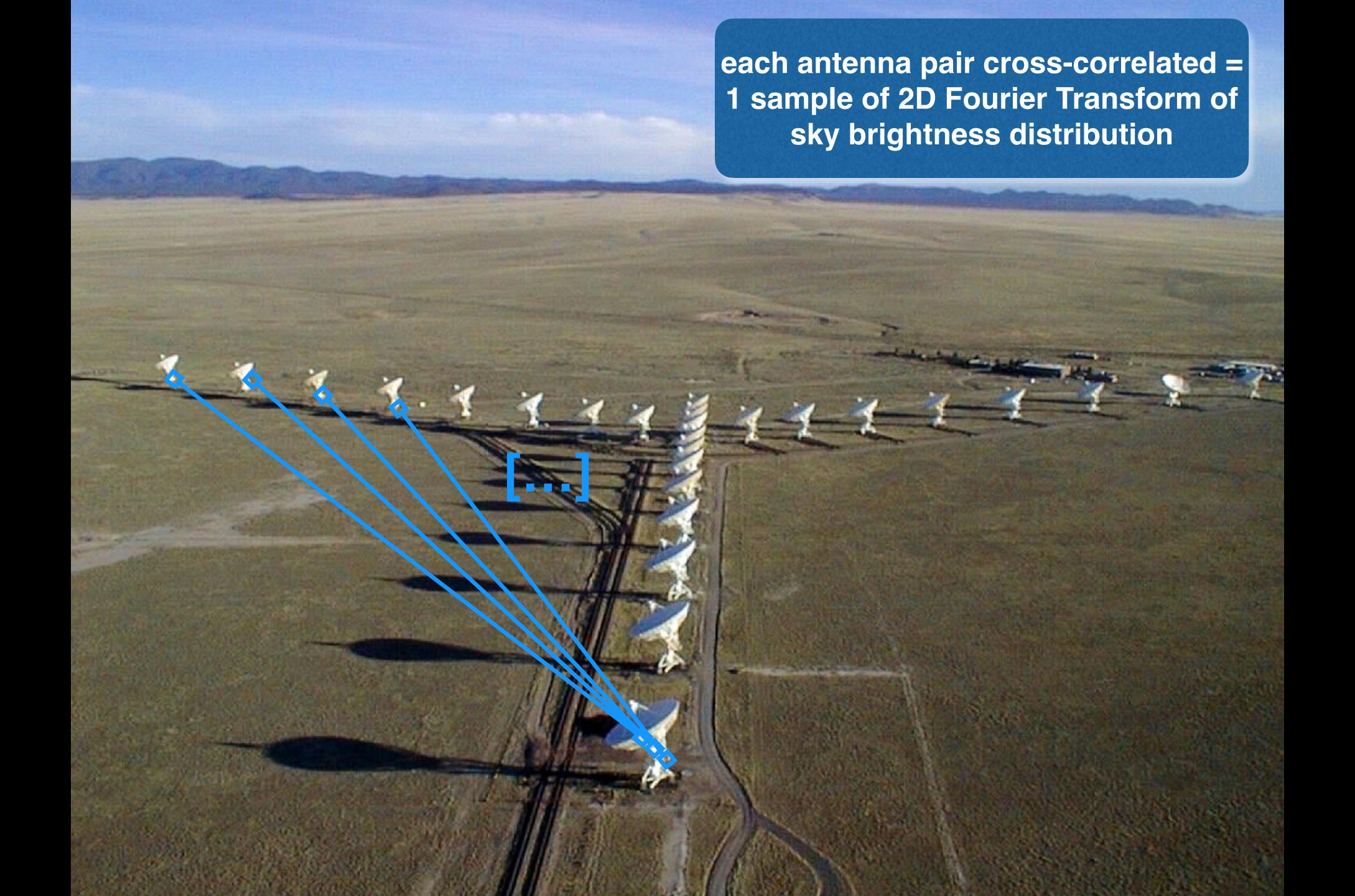


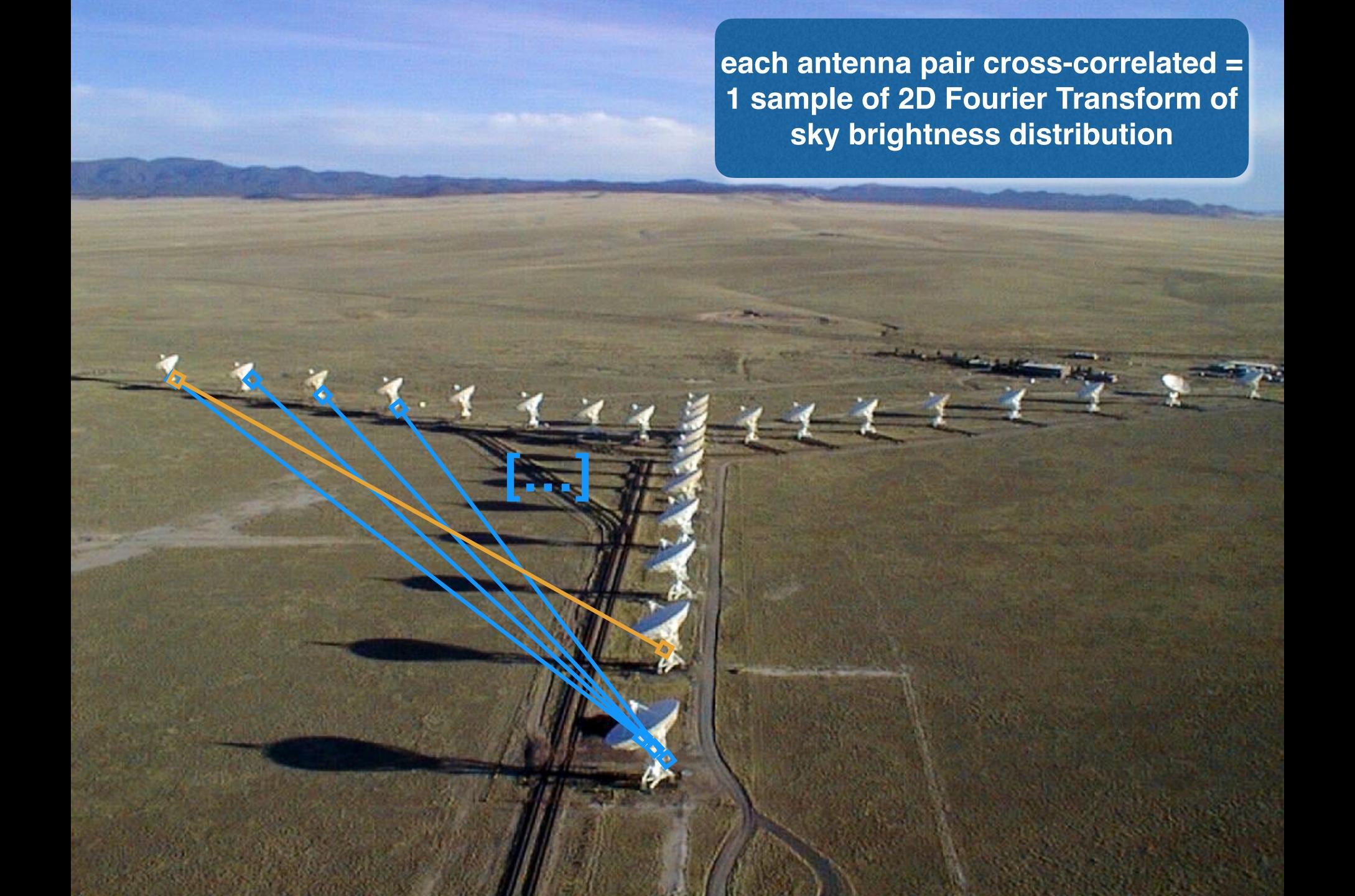


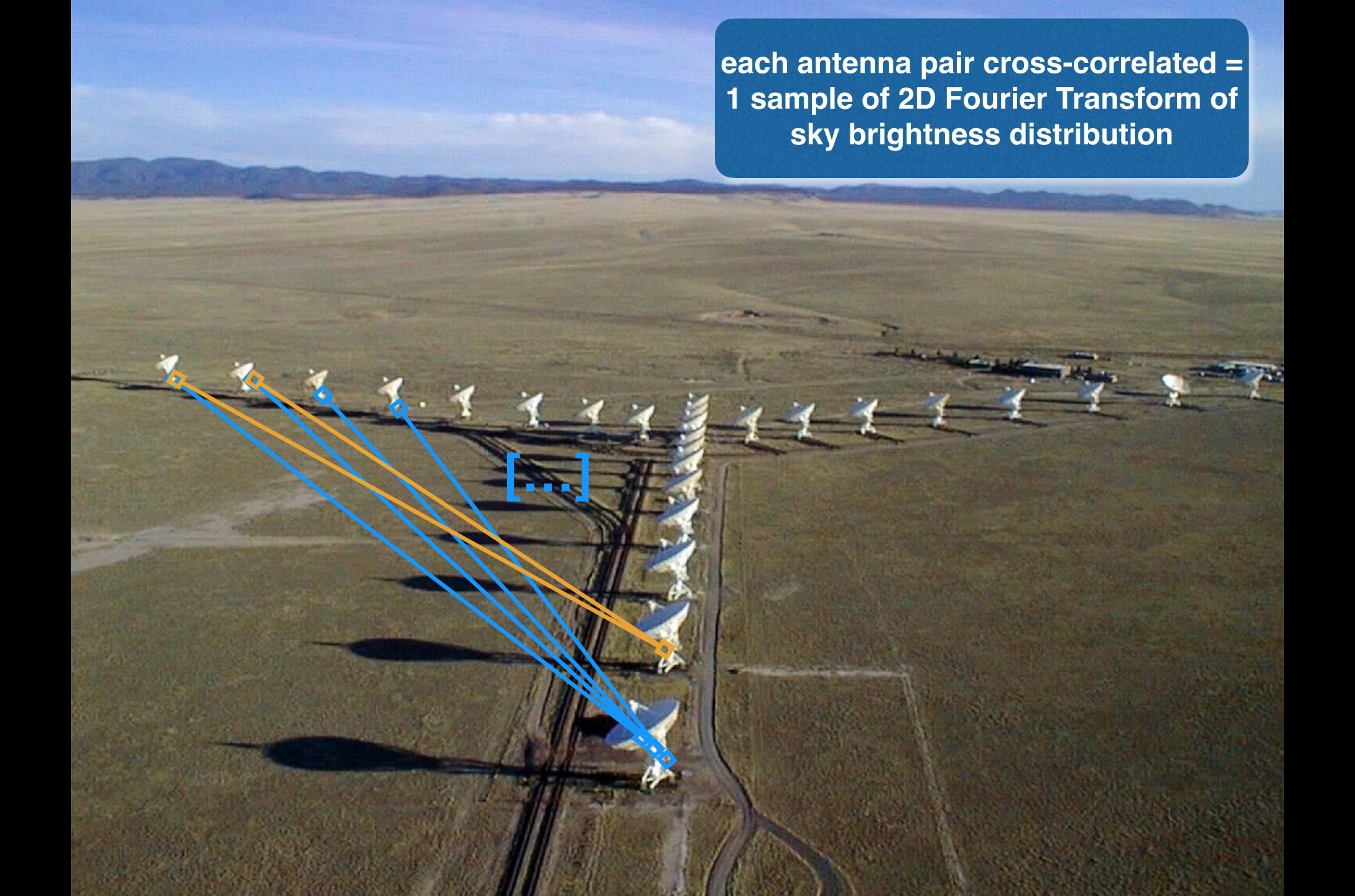


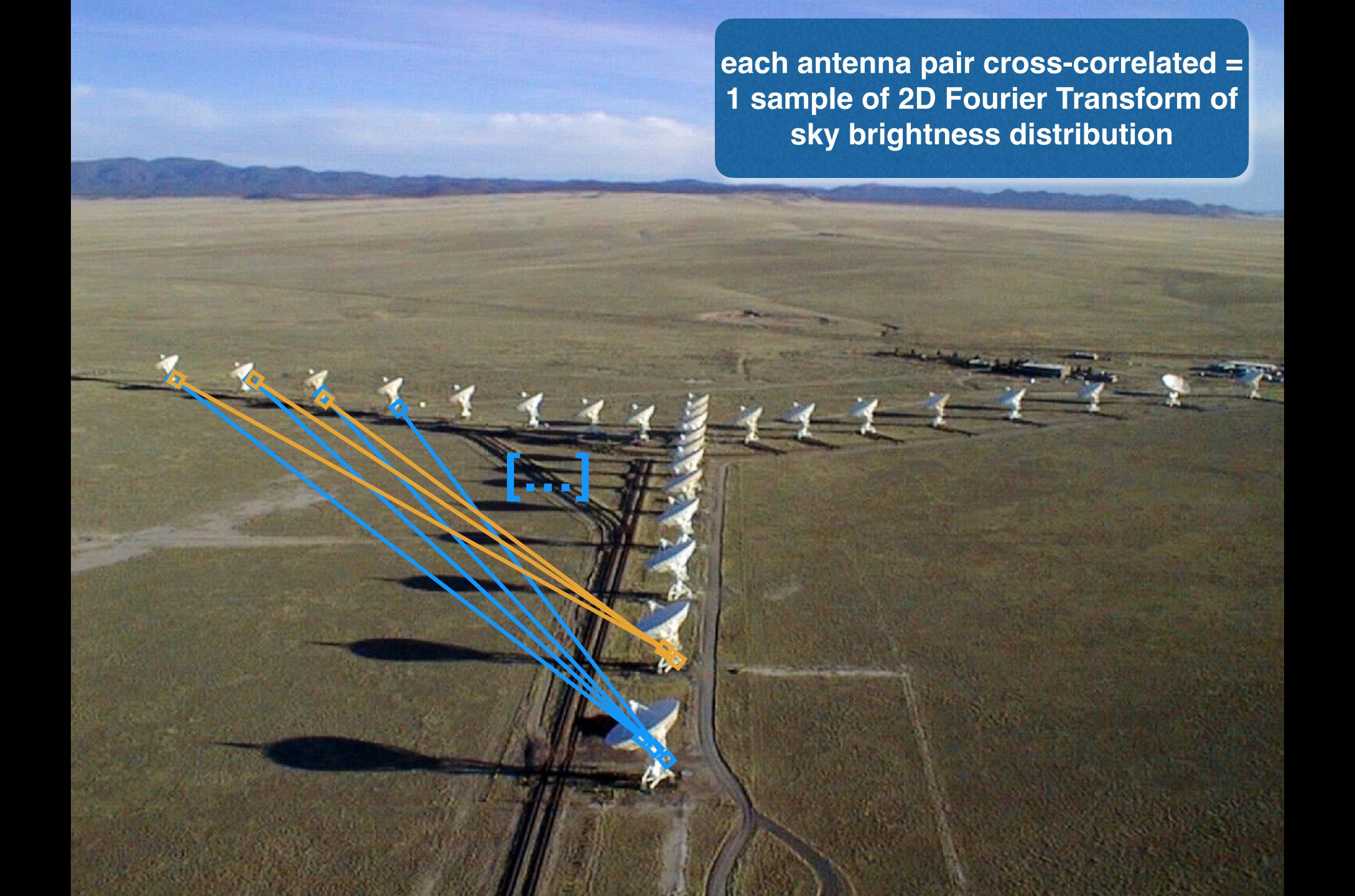


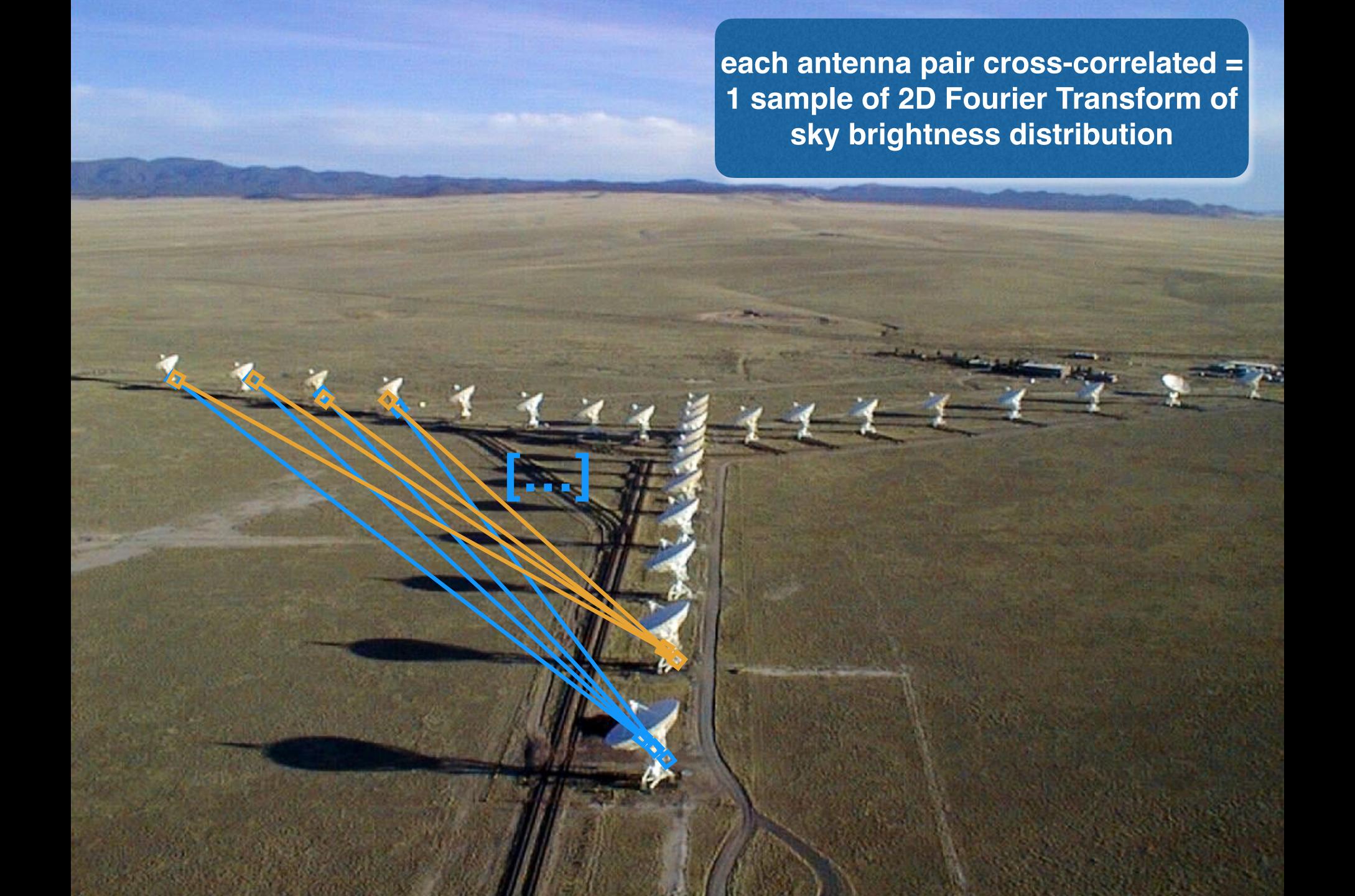


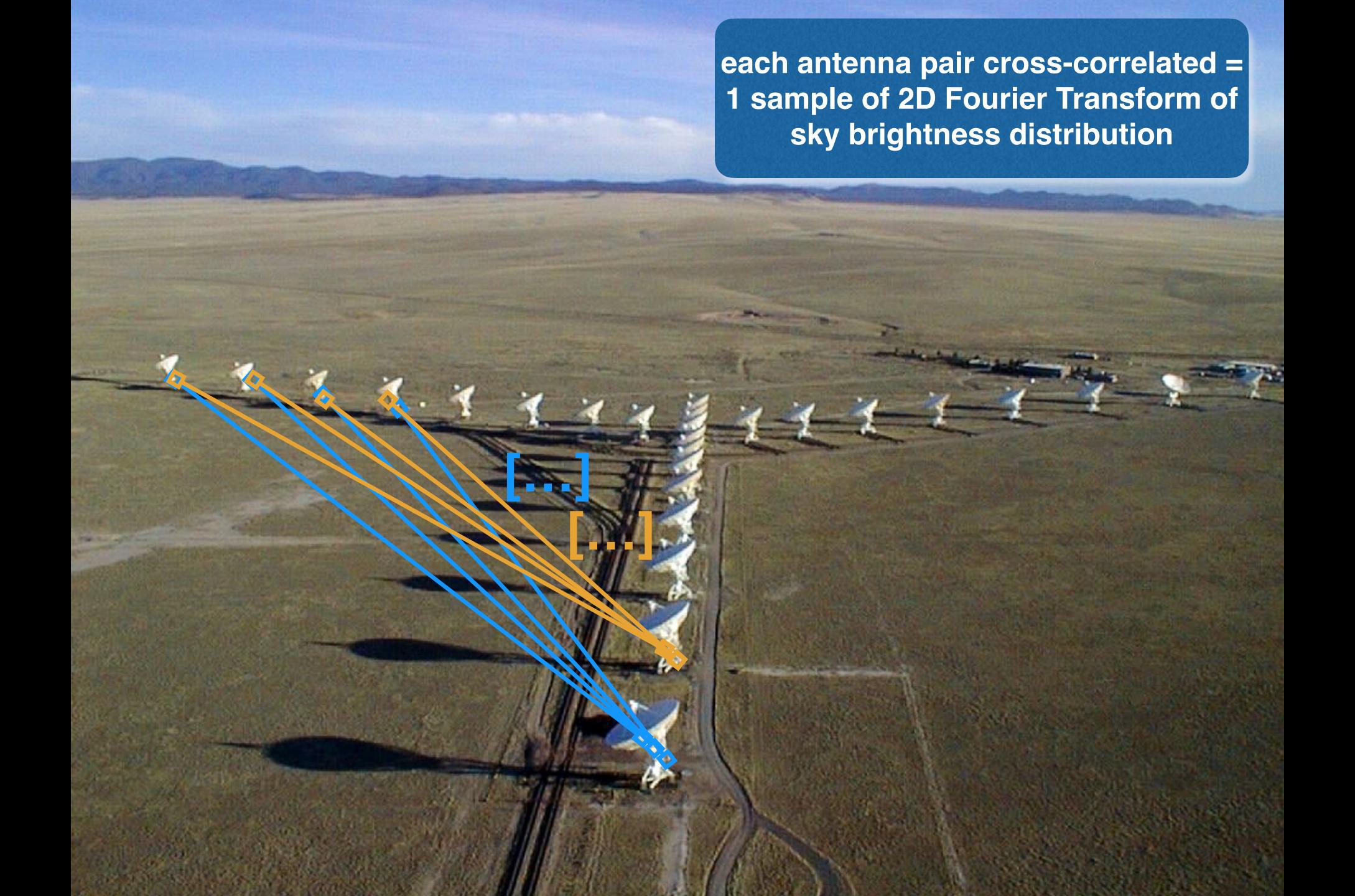


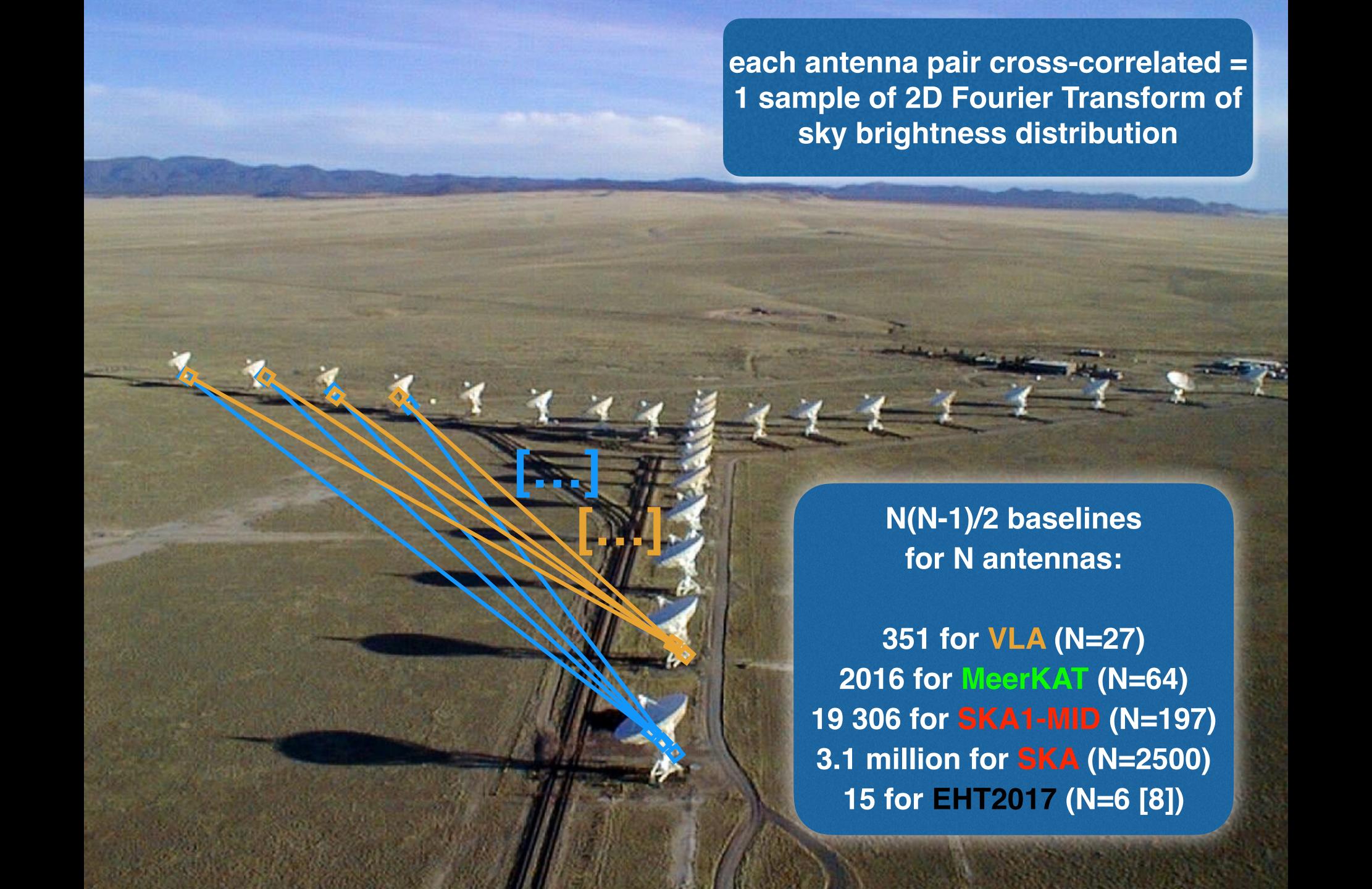


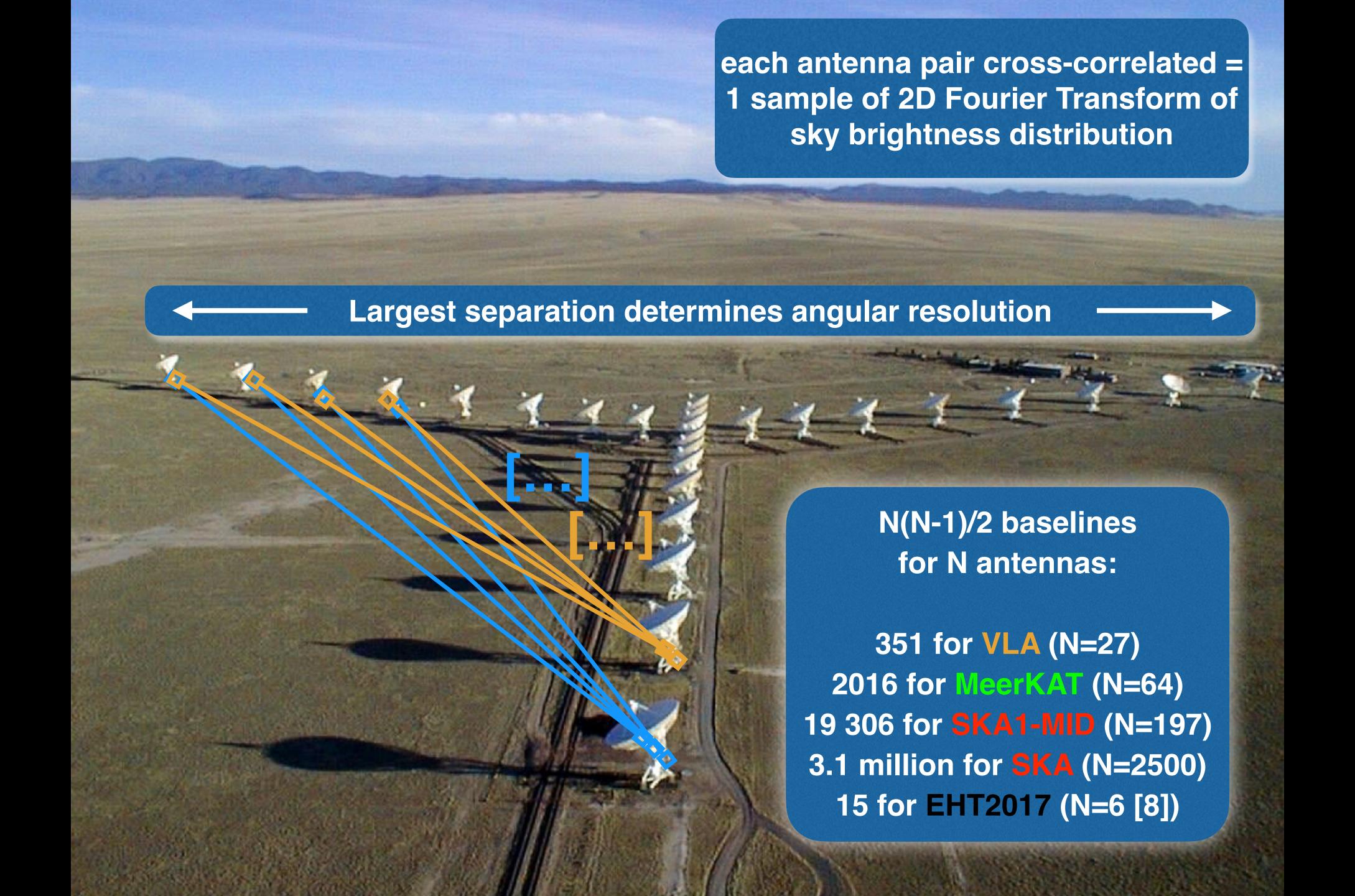




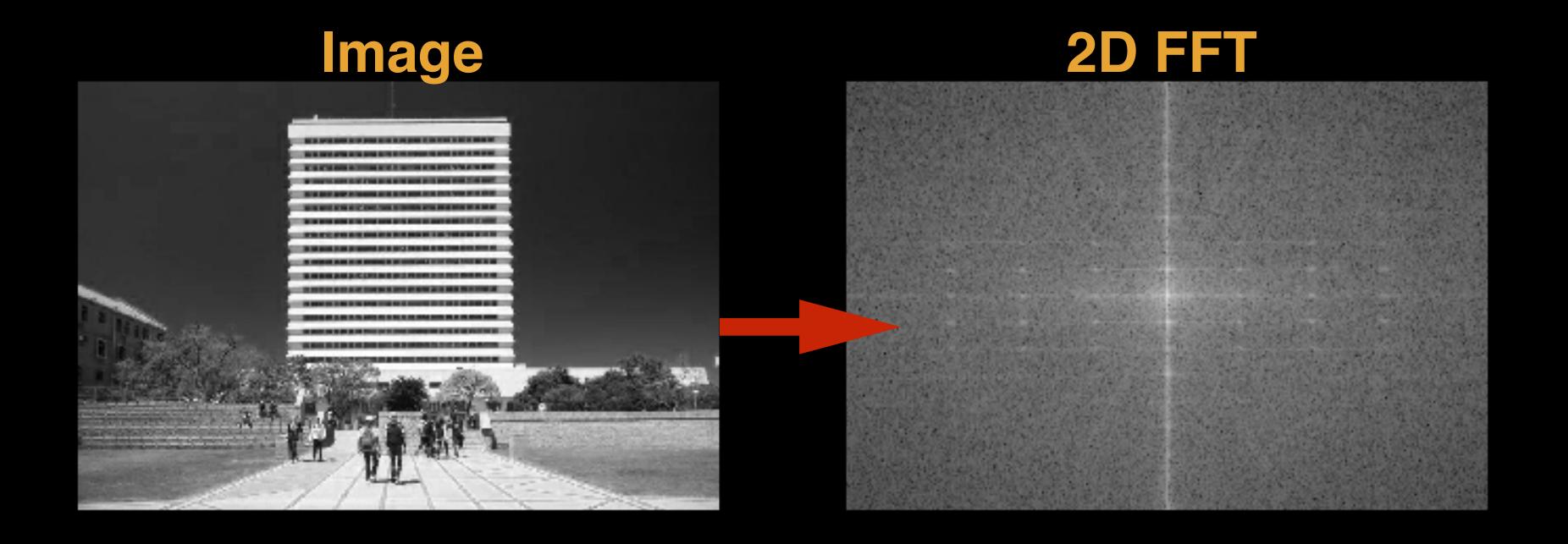




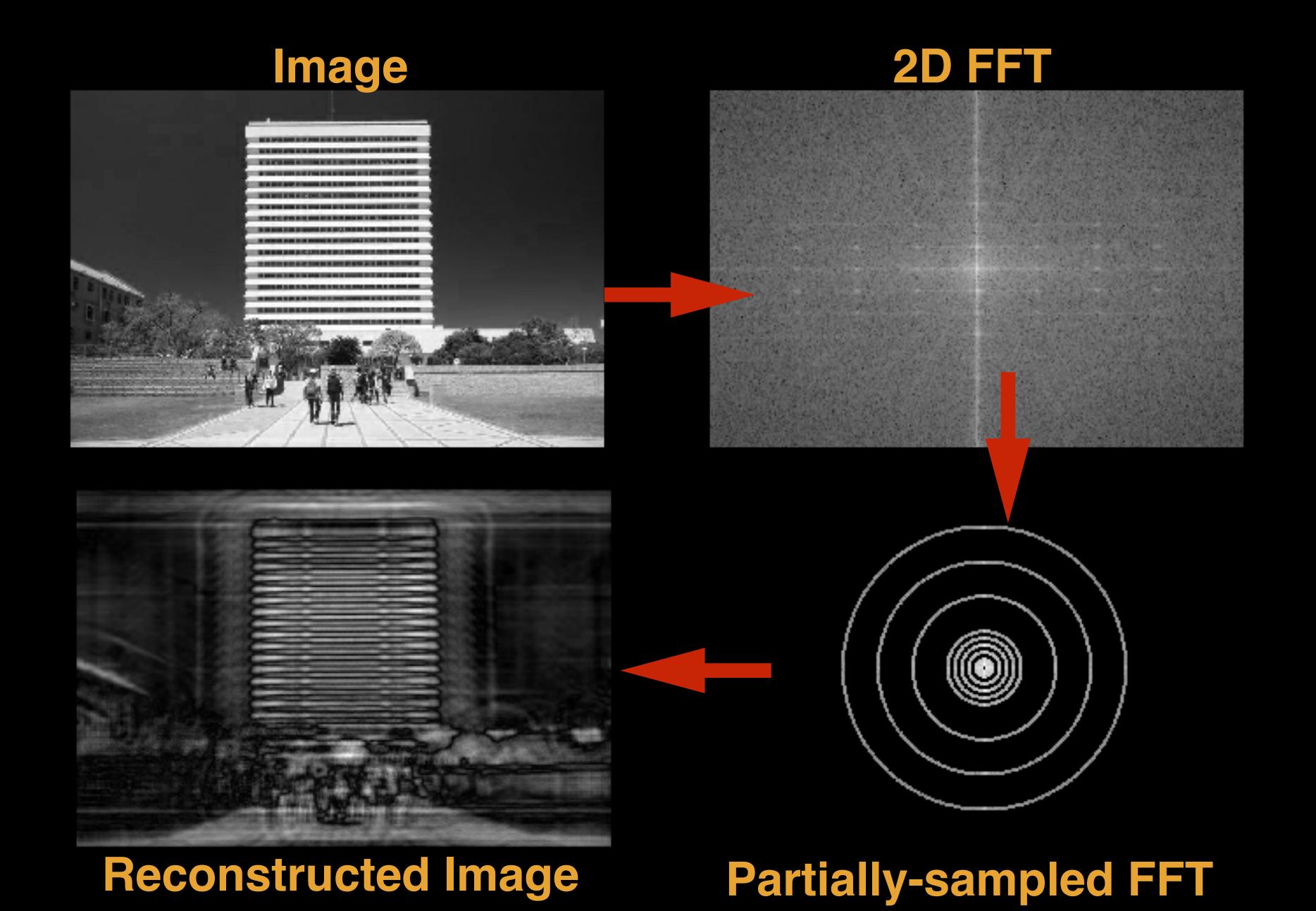




A radio interferometer is a Fourier transform machine



A radio interferometer is a Fourier transform machine



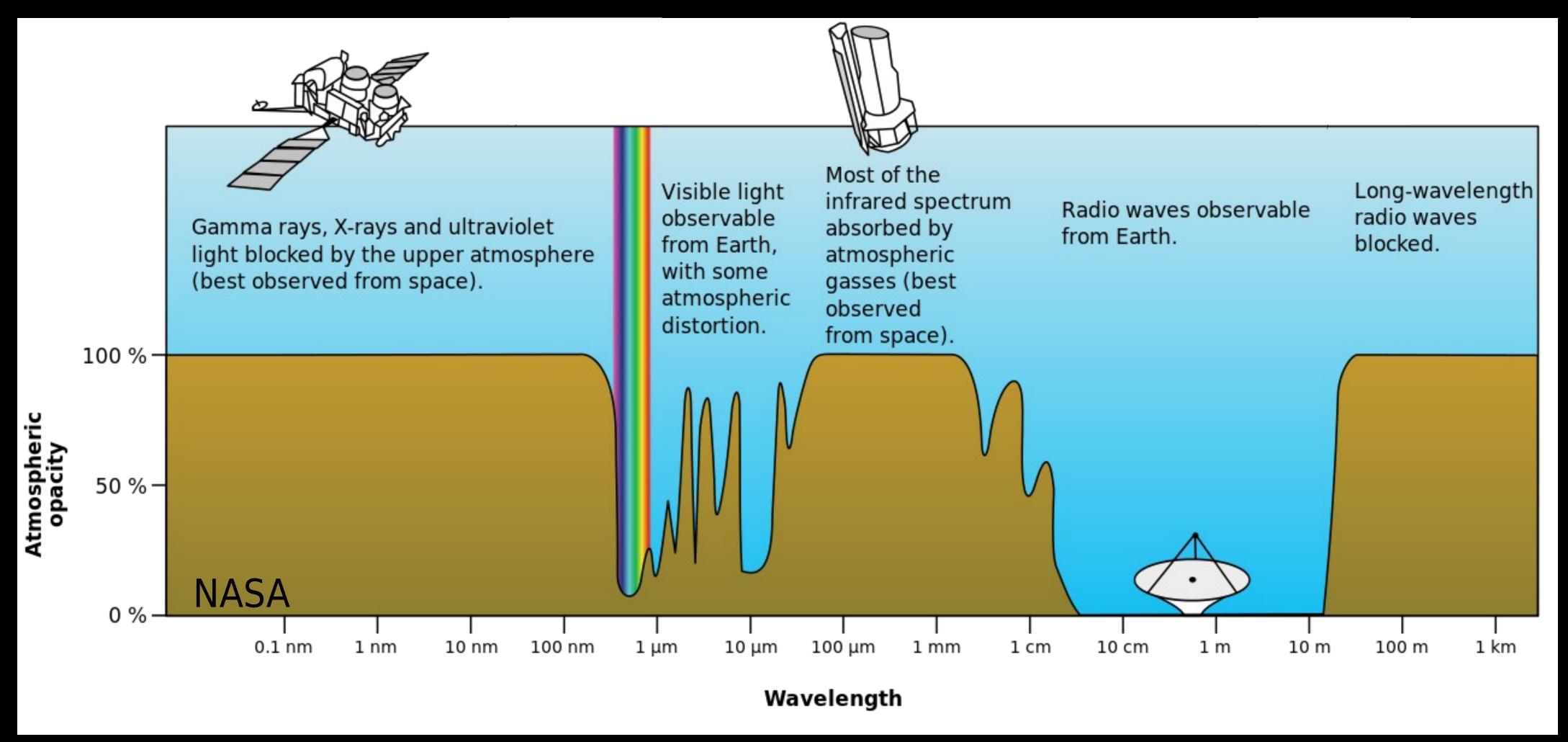
The EHT array



- Earth-sized telescope operating at 230 & 345 GHz
- Achieves an angular resolution of ~20 µas (~1000 times finer than Hubble Space Telescope)
- Primary science goal: spatially resolve event-horizon-scale emission towards Sgr A* and M87
- Test Kerr metric hypothesis (or other theories of gravity) in the strong-field regime and constrain accretion flow / jetlaunch physics
- "extreme interferometry" requires an intense engineering, data processing, calibration, theoretical, and modeling effort

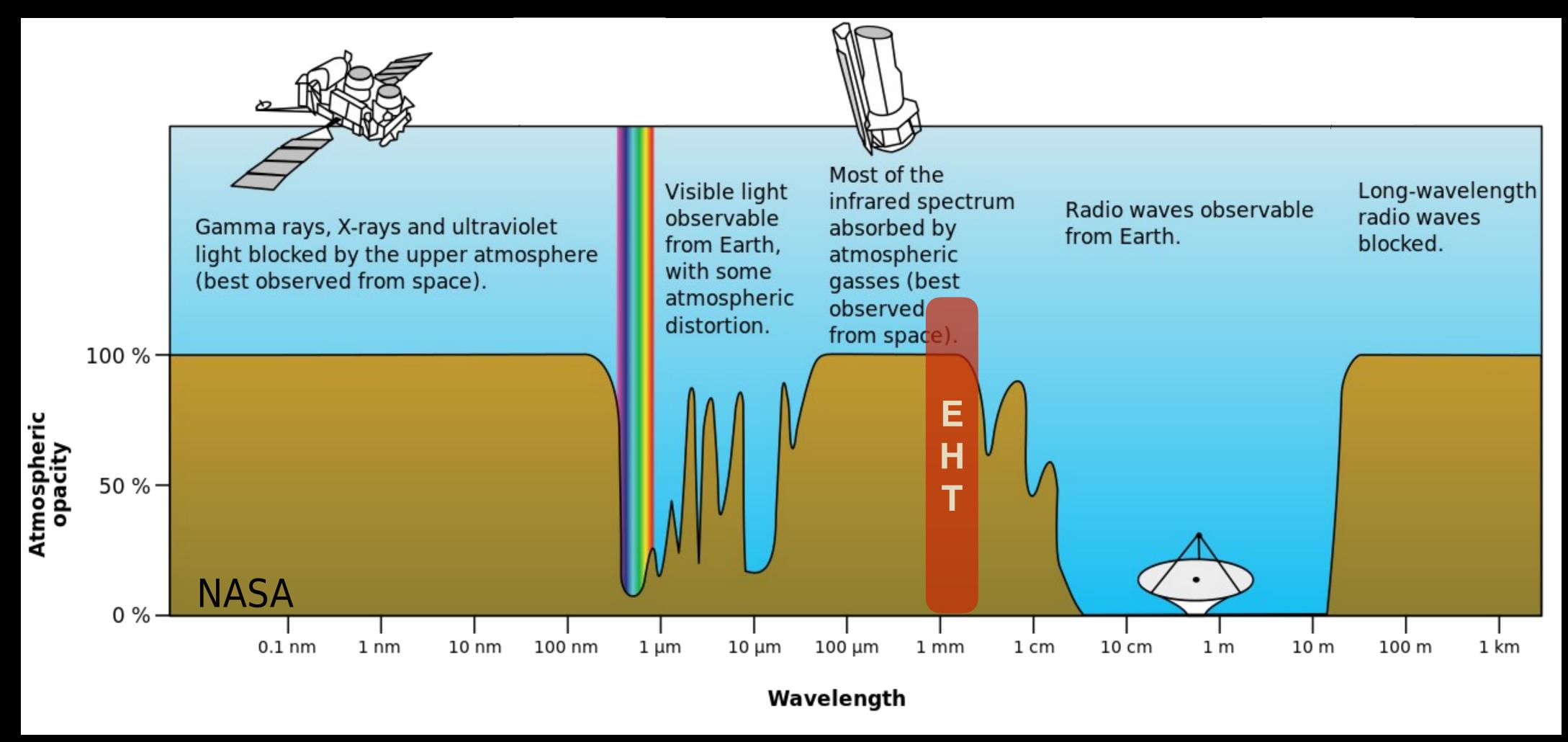


The troublesome (and turbulent) troposphere





The troublesome (and turbulent) troposphere









A Boeing 747's bandwidth

- Data recorded independently at different sites at very high time and frequency resolution and precision
- Disks brought together at supercomputers at MIT and MPIfR to correlate the signals and form visibilities (fundamental measurement of an interferometer)
- ~4 PB required for entire EHT 2017 observing run
- ~0.5 PB on M87 alone
- This will increase by a factors of a few in the coming years



A Boeing 747's bandwidth

 Data recorded independently at different sites at very high time and frequency resolution and precision

 Disks brought together at supercomputers at MIT and MPIfR to correlate the signals and form visibilities (fundamental measurement)

Final image size ~few kiloBytes

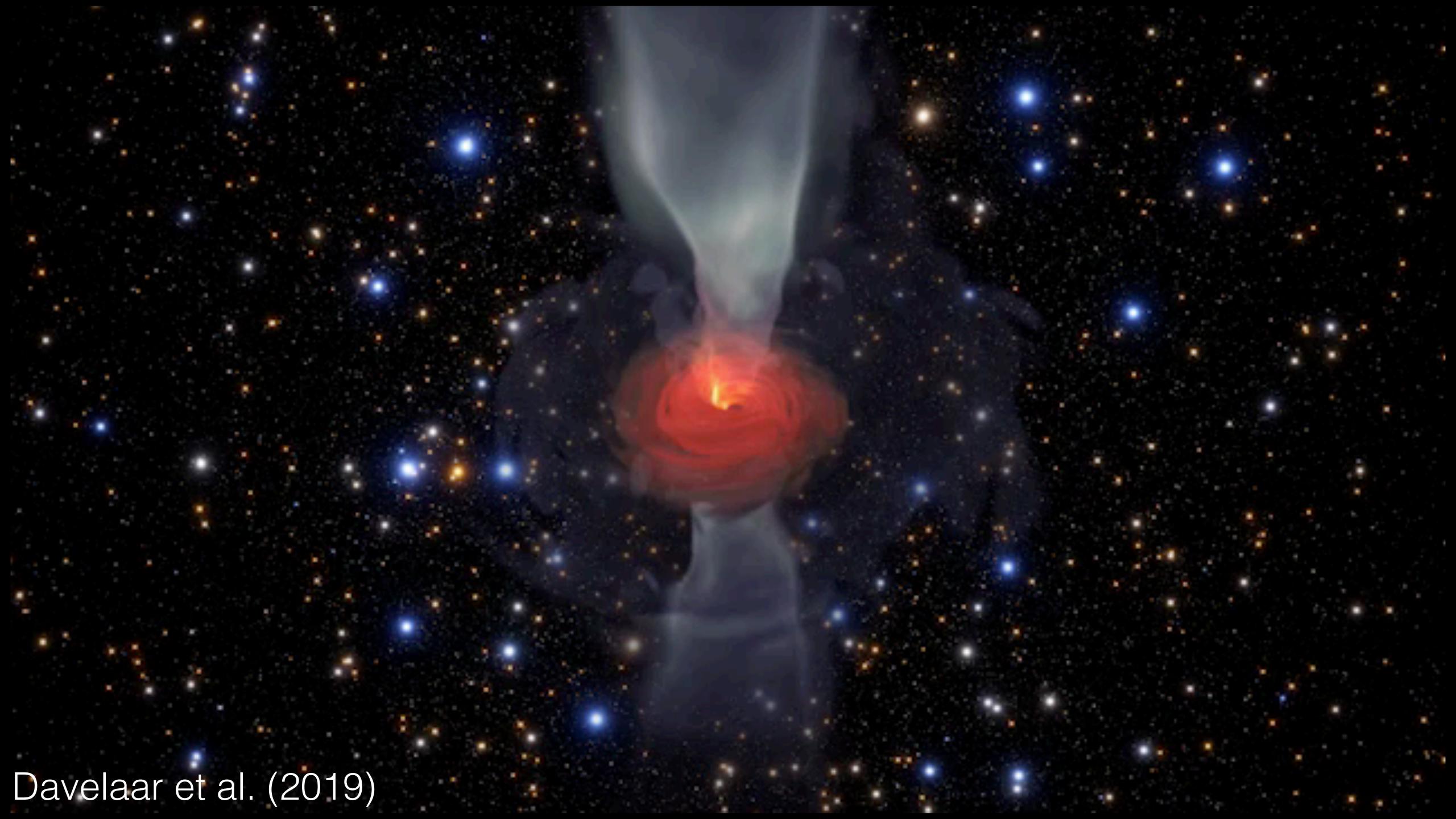
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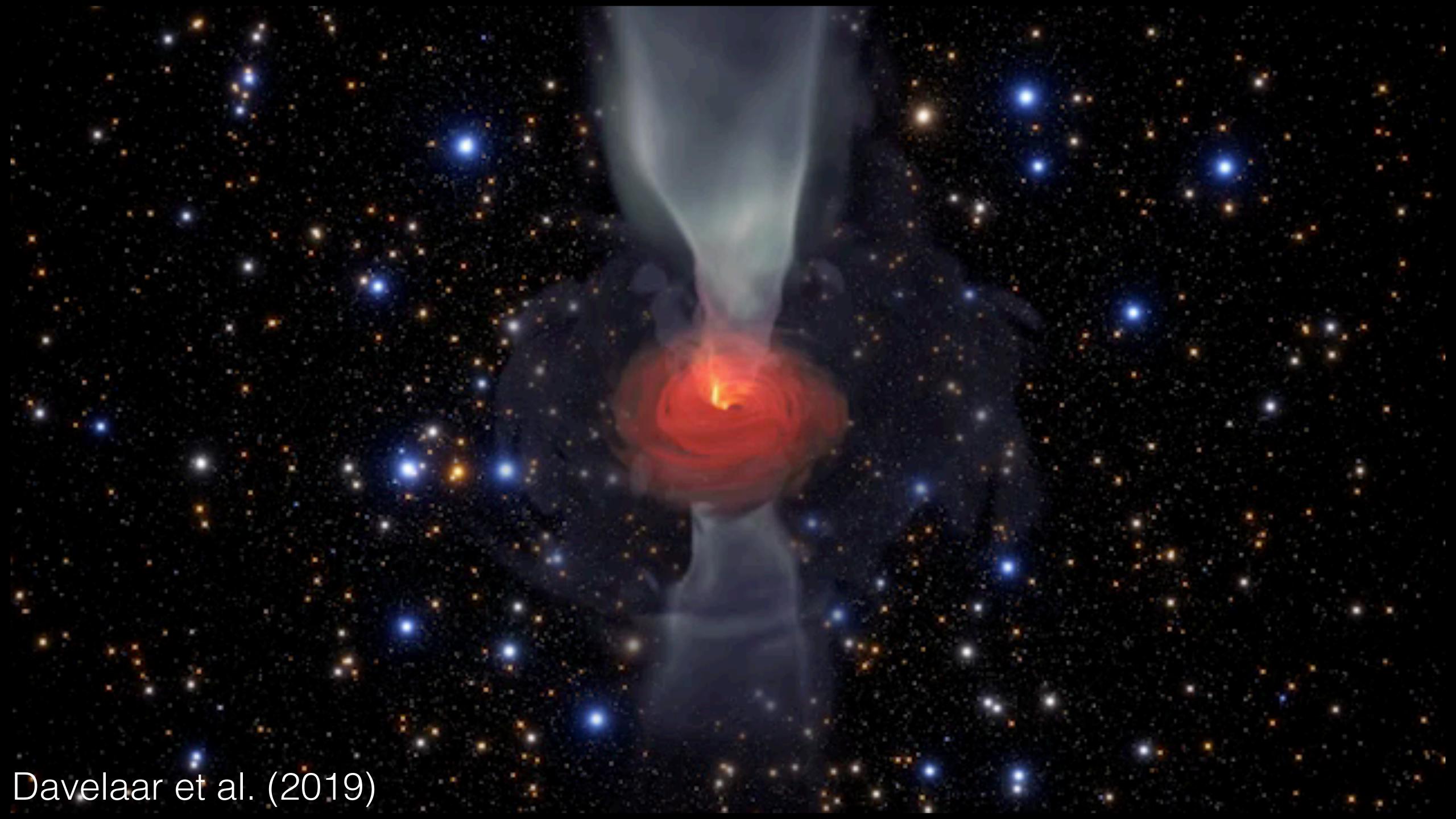




Primary science goal:

Spatially-resolve event horizon scales of nearby supermassive black holes

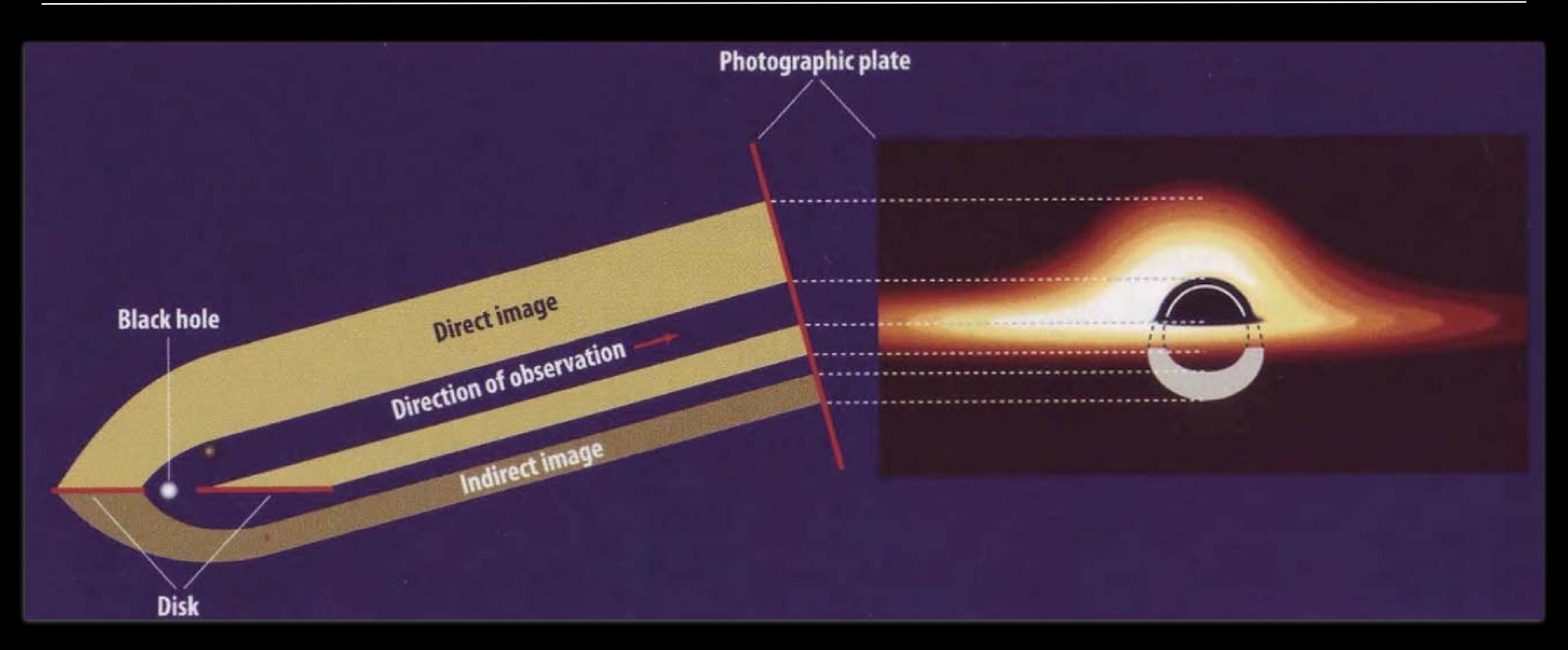




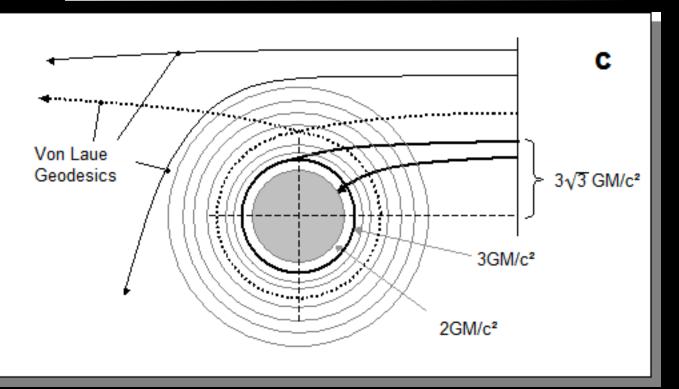
A black hole shadow?



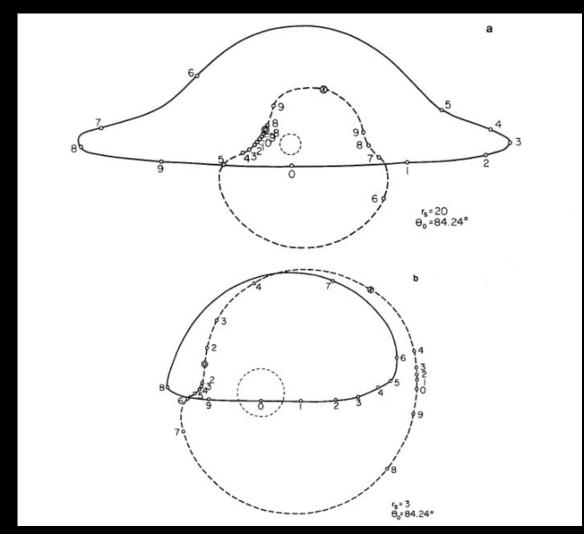
A black hole shadow?



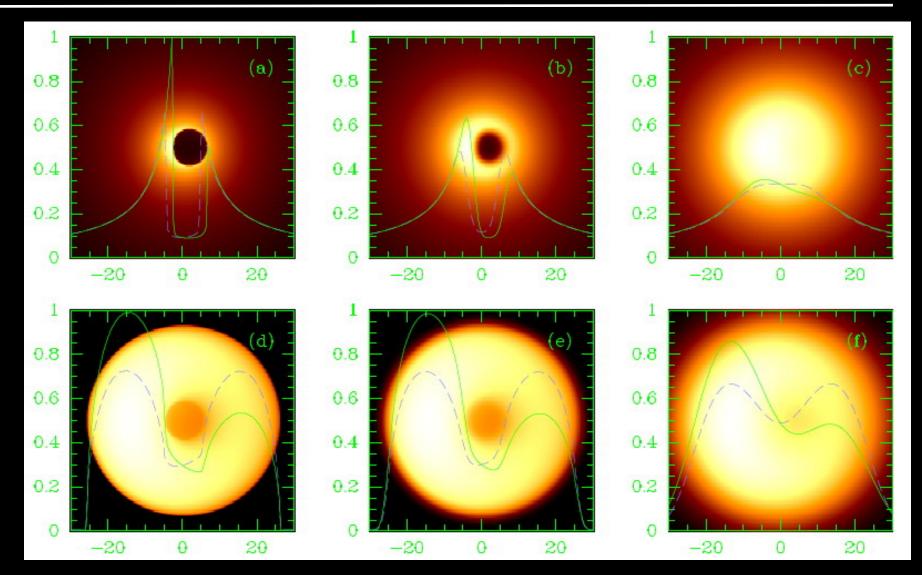
Black hole "shadow" history



Von Laue (1921) based on Hilbert (1916) Lights paths that form perfect circles



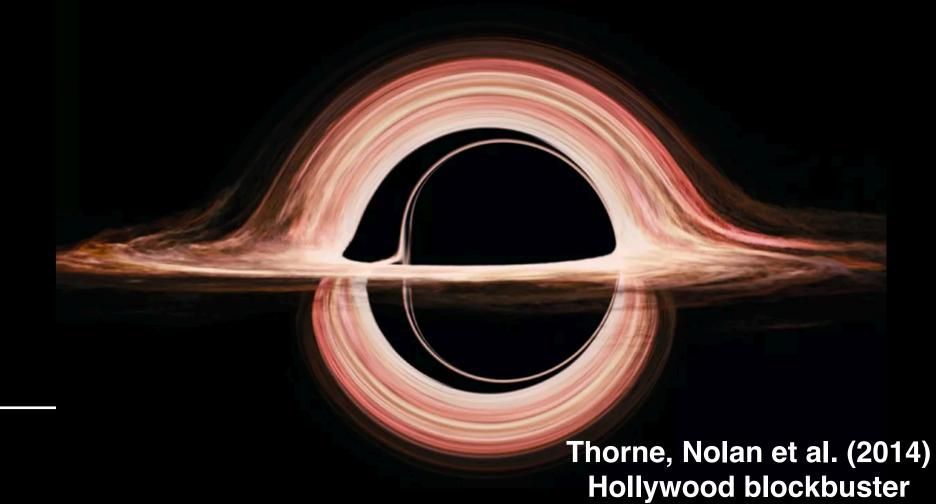
Cunningham & Baarden (1983) Stellar orbit around a black hole also, Baarden (1973) lensed star



Falcke et al. (2000)
Optically thin emission, including inside ISCO
Predicted mm-VLBI



Luminet (1979)
Optically thin accretion disk beyond the ISCO

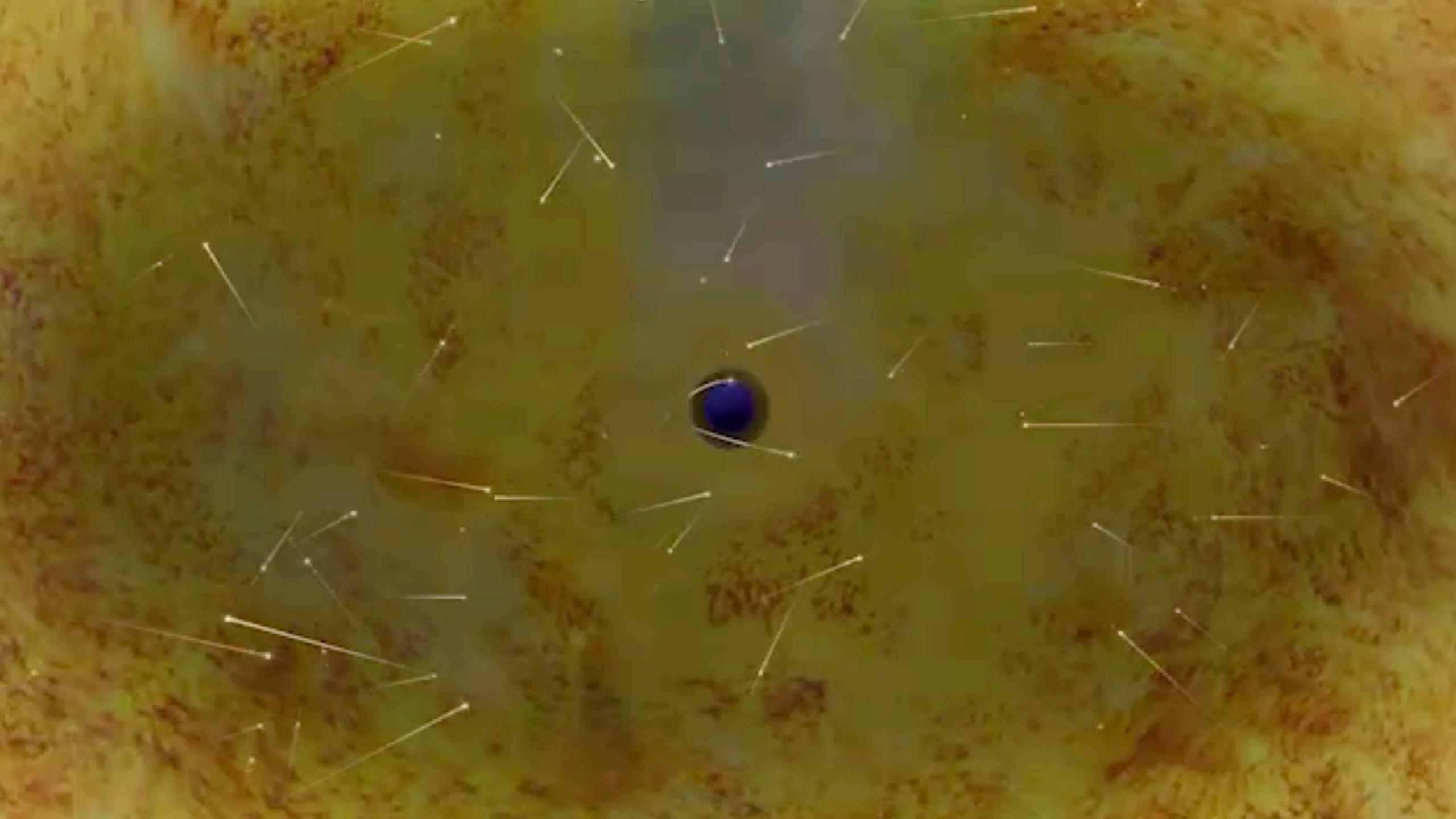


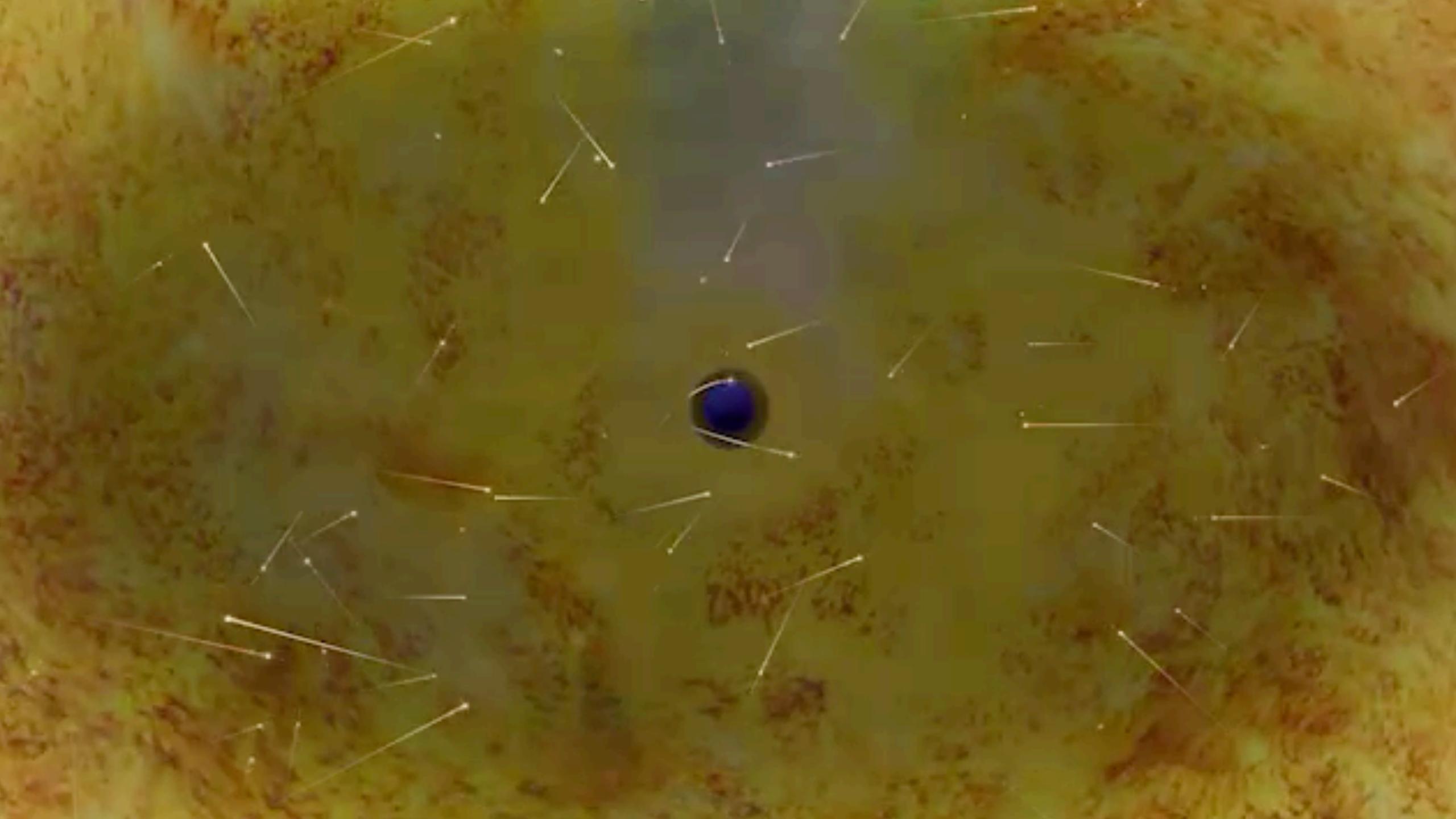
A black hole shadow?

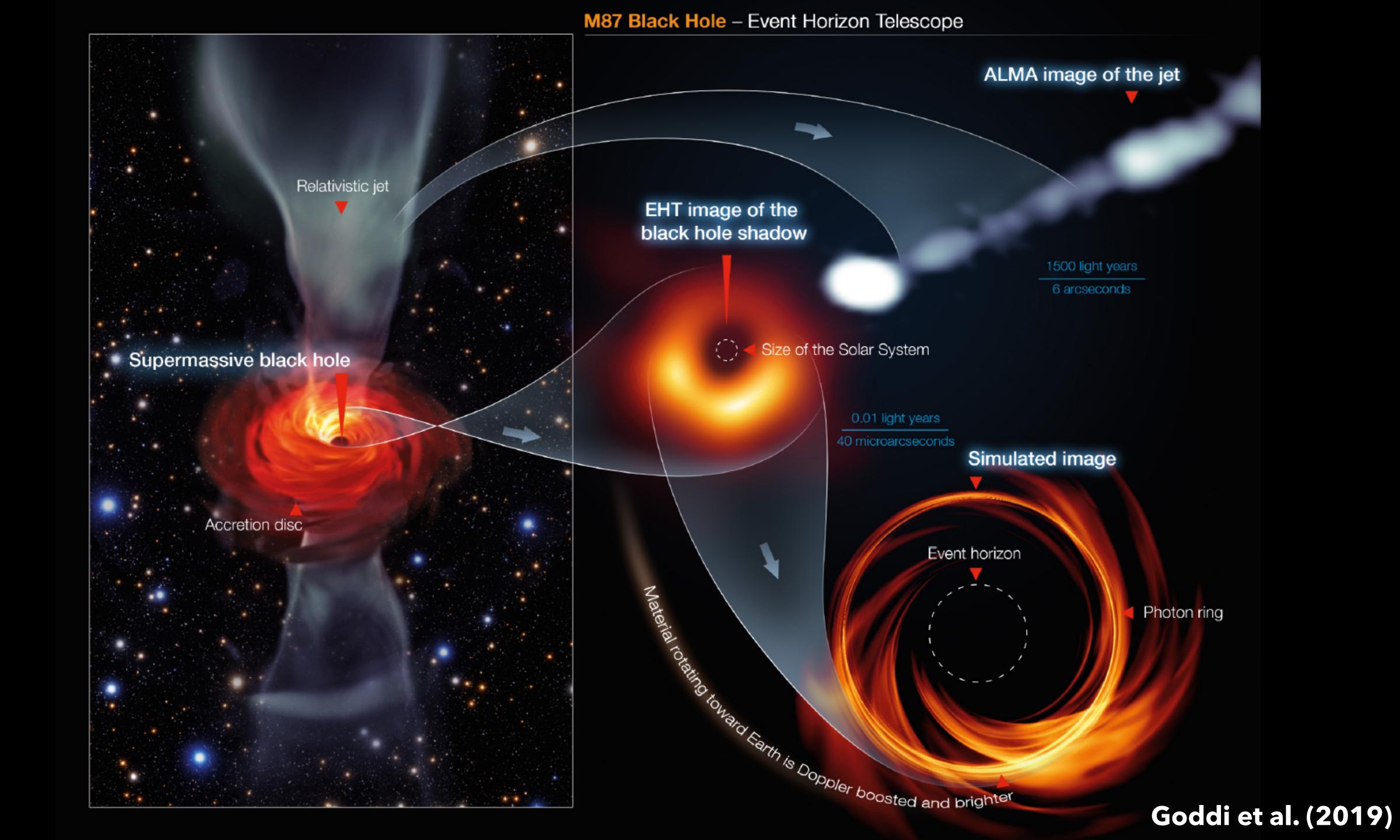
Two key concepts:

- 1. Light bending around a black hole
- 2. Light captured by a black hole









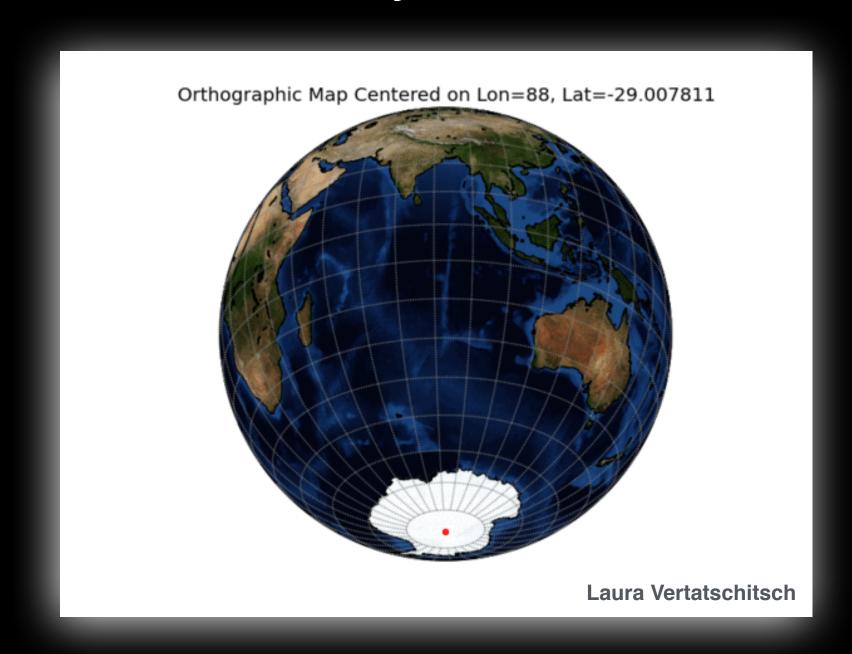
$$\theta_p \equiv \frac{\sqrt{27} \, GM}{c^2 D}$$

for a Schwarzschild metric, ~10% change for spin

Zooming into the Event Horizon

Using this:

Global array of antennas



To image this:

Apparent size of a doughnut on the moon



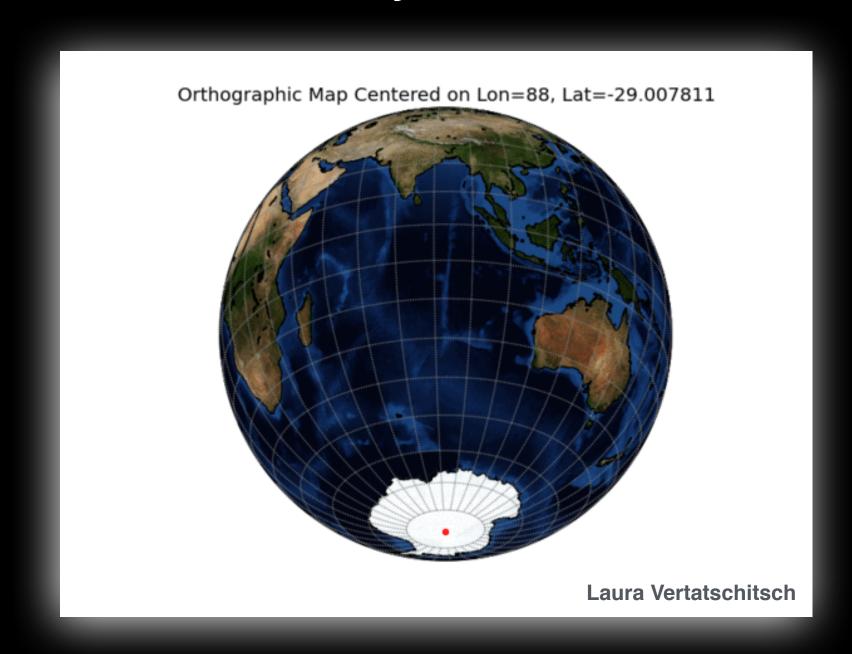




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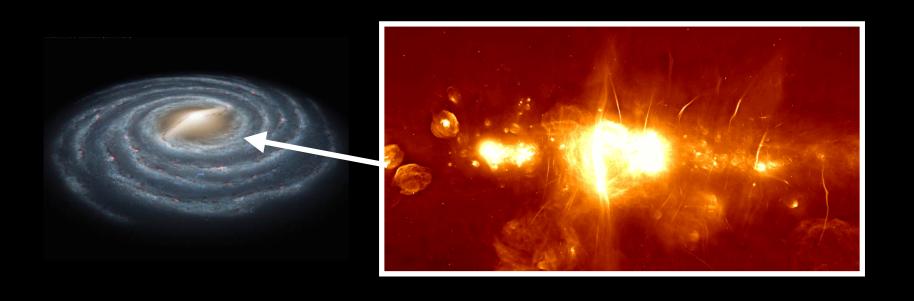






EHT's two primary targets

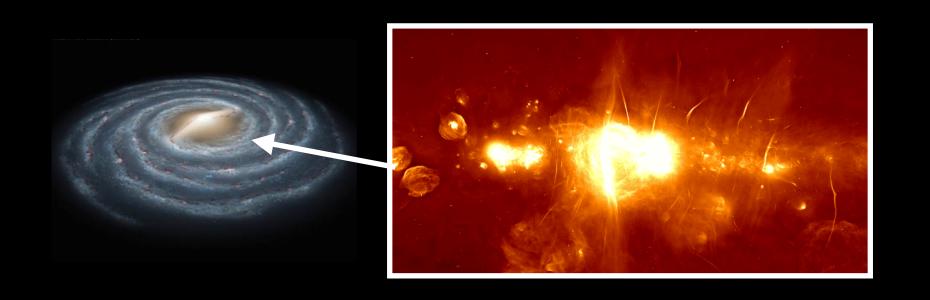
Sgr A*



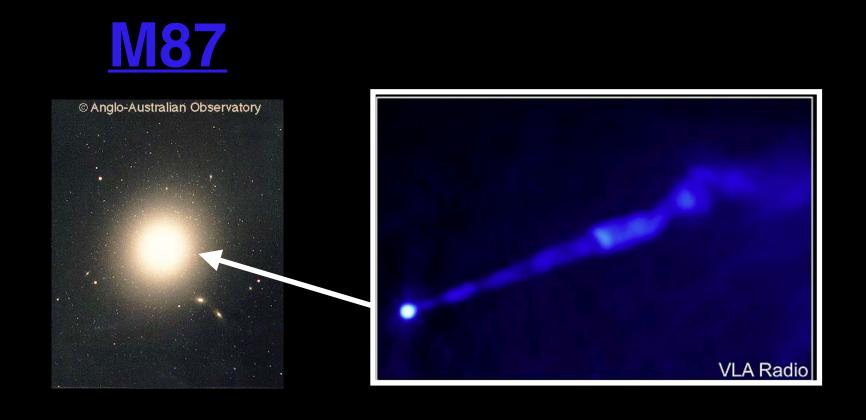
- $M_{BH} \sim 4 \times 10^6 M_{\odot}$
- Predicted photon ring ~50 μas
- Shortest orbital timescale: ~minutes

EHT's two primary targets

Sgr A*



- $M_{BH} \sim 4 \times 10^6 M_{\odot}$
- Predicted photon ring ~50 μas
- Shortest orbital timescale: ~minutes



- $M_{BH} \sim 6 \times 10^9 M_{\odot}$
- Predicted photon ring ~20-40 µas
- Shortest orbital timescale: ~weeks

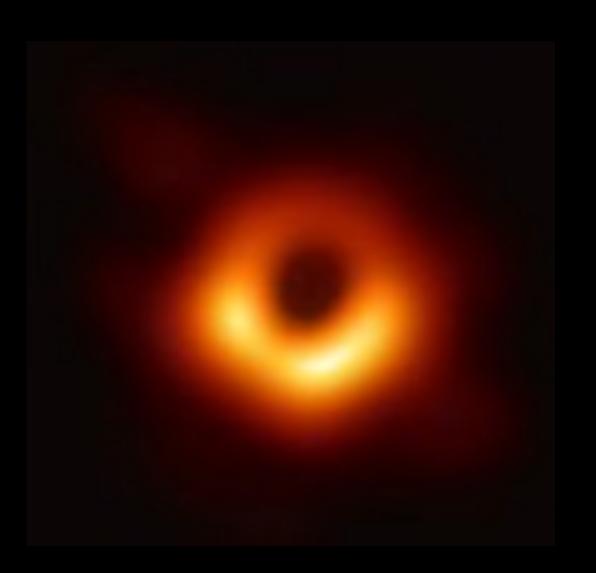
outline

- EHT Overview: the instrument and science goals
- First imaging results on M87:
- The future: upcoming results and array expansion



outline

First imaging results on M87:





But first some contrast of the imaging challenges with MeerKAT

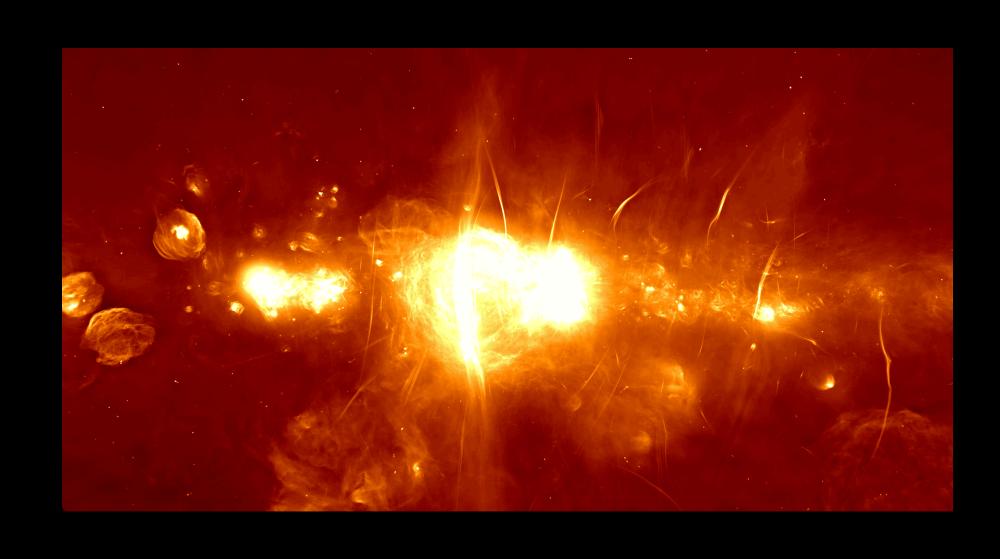
MeerKAT image of the centre of our Galaxy **Image credits: SARAO**

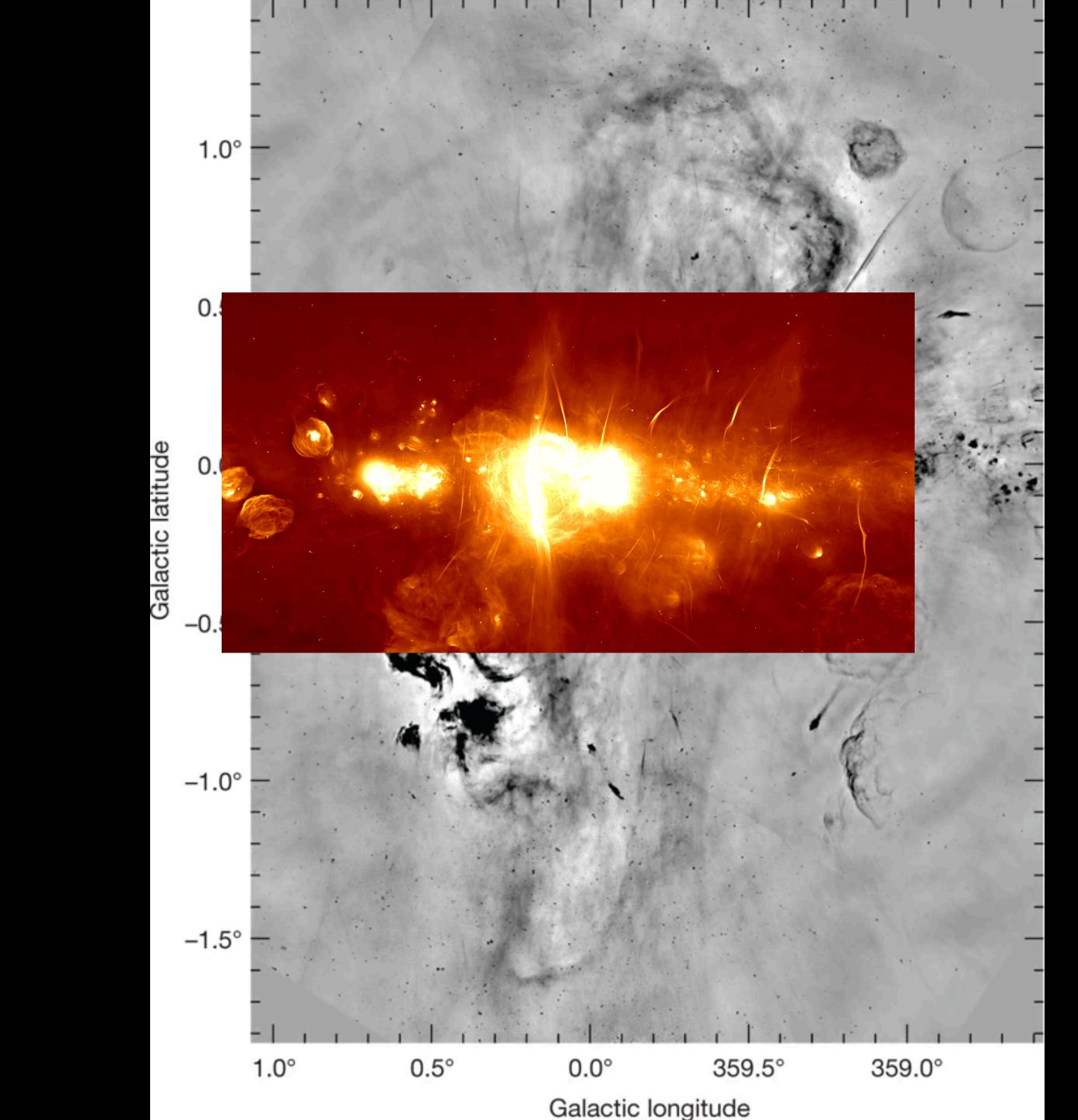
MeerKAT image of the centre of our Galaxy

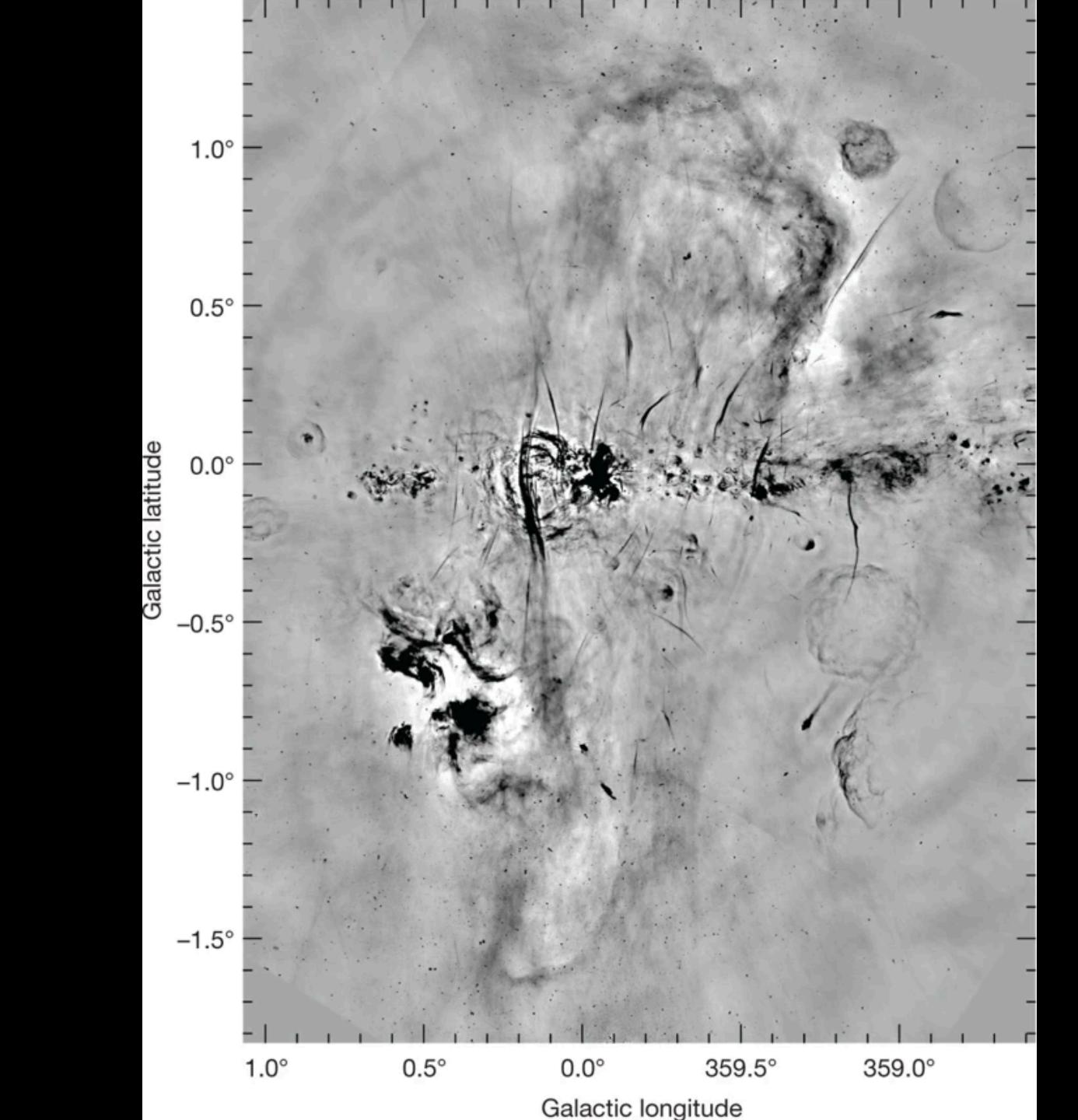




Image credits: SARAO







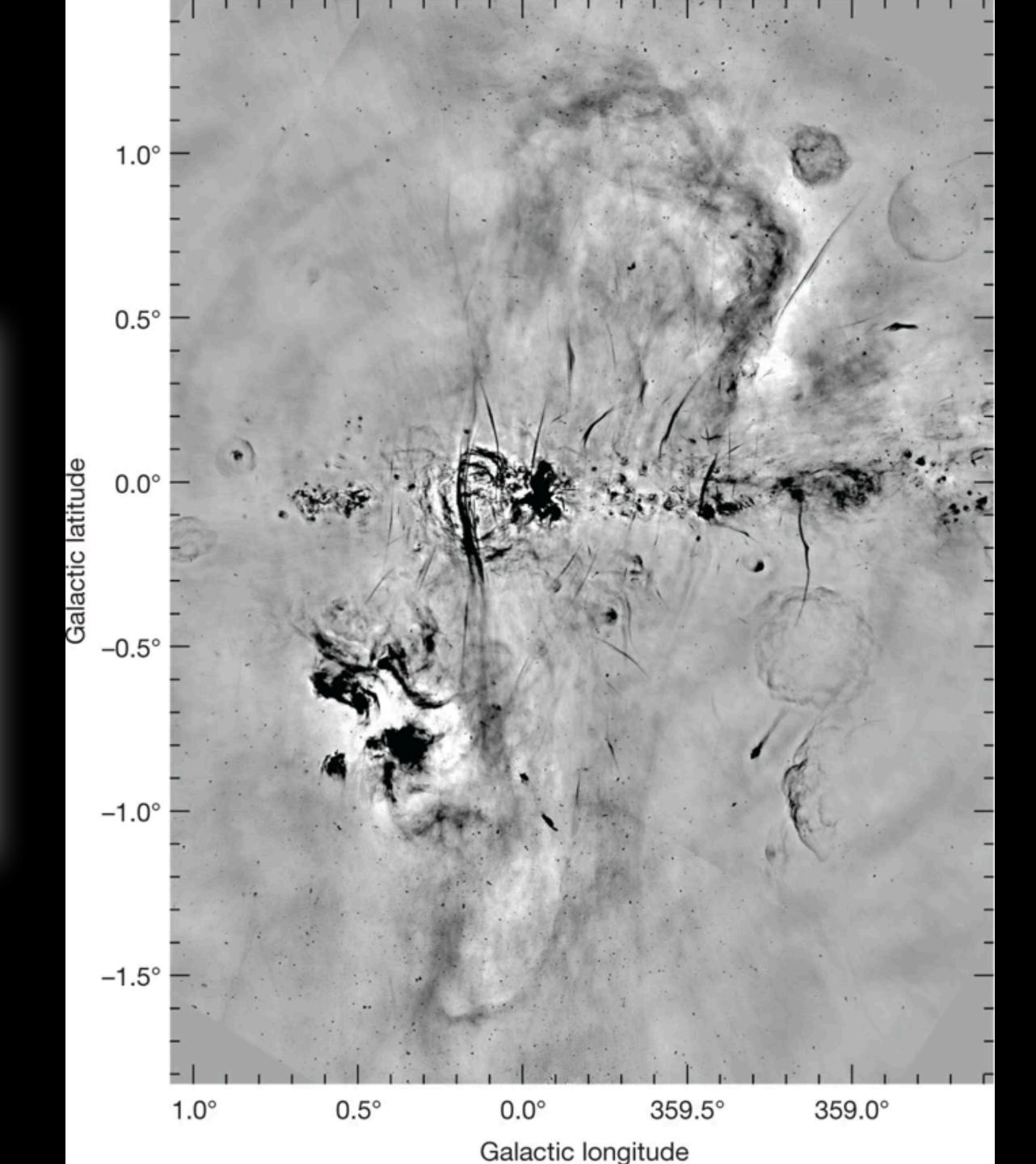


Letter Published: 11 September 2019

Inflation of 430-parsec bipolar radio bubbles in the Galactic Centre by an energetic event

I. Heywood [™], F. Camilo [™], […] L. P. Williams

Nature **573**, 235–237 (2019) | Download Citation **±**



A more typical MeerKAT map:

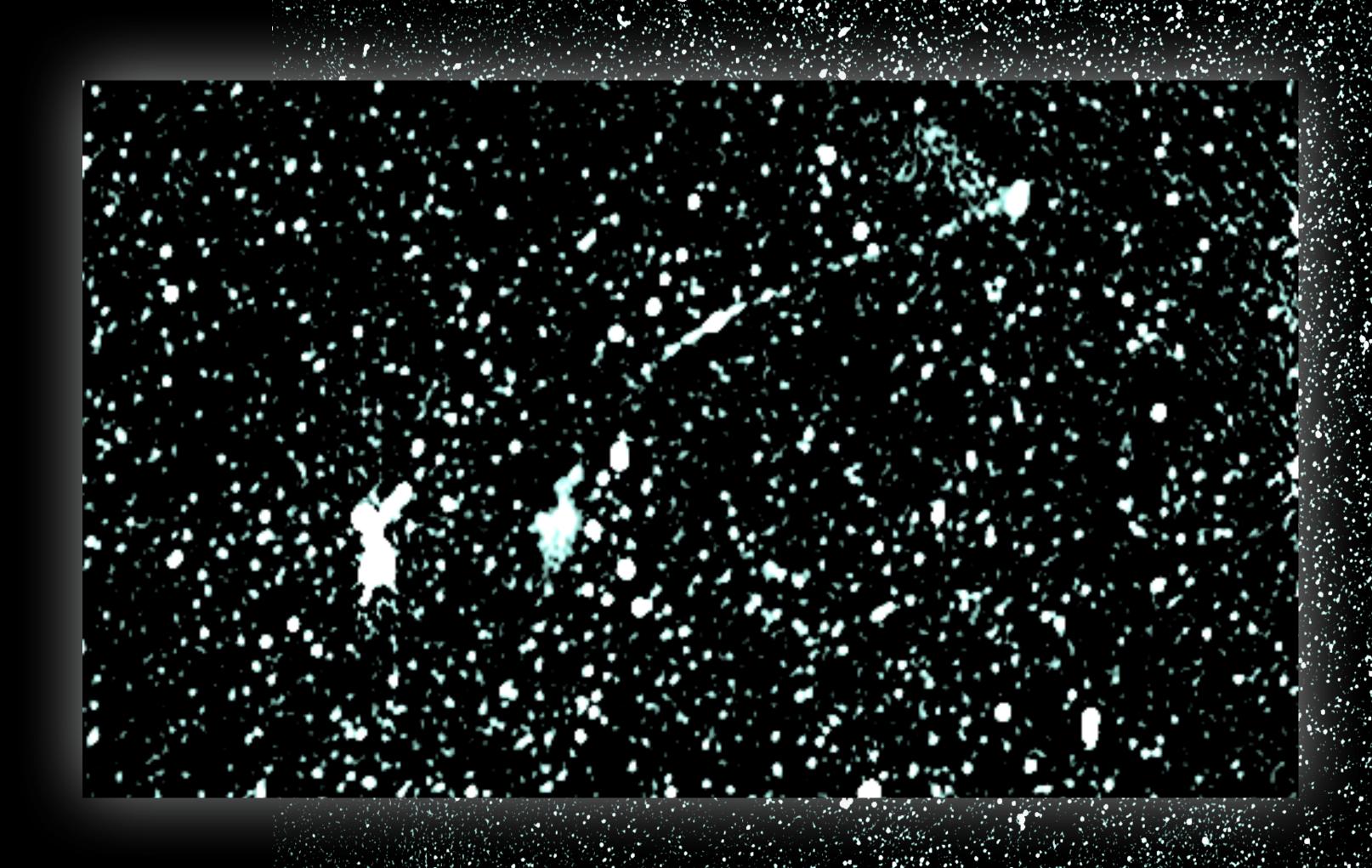
MeerKAT COSMOS

Single pointing

~3 uJy/beam; > 1 deg²

~5000 sources; ~16 hours

A more typical MeerKAT map:

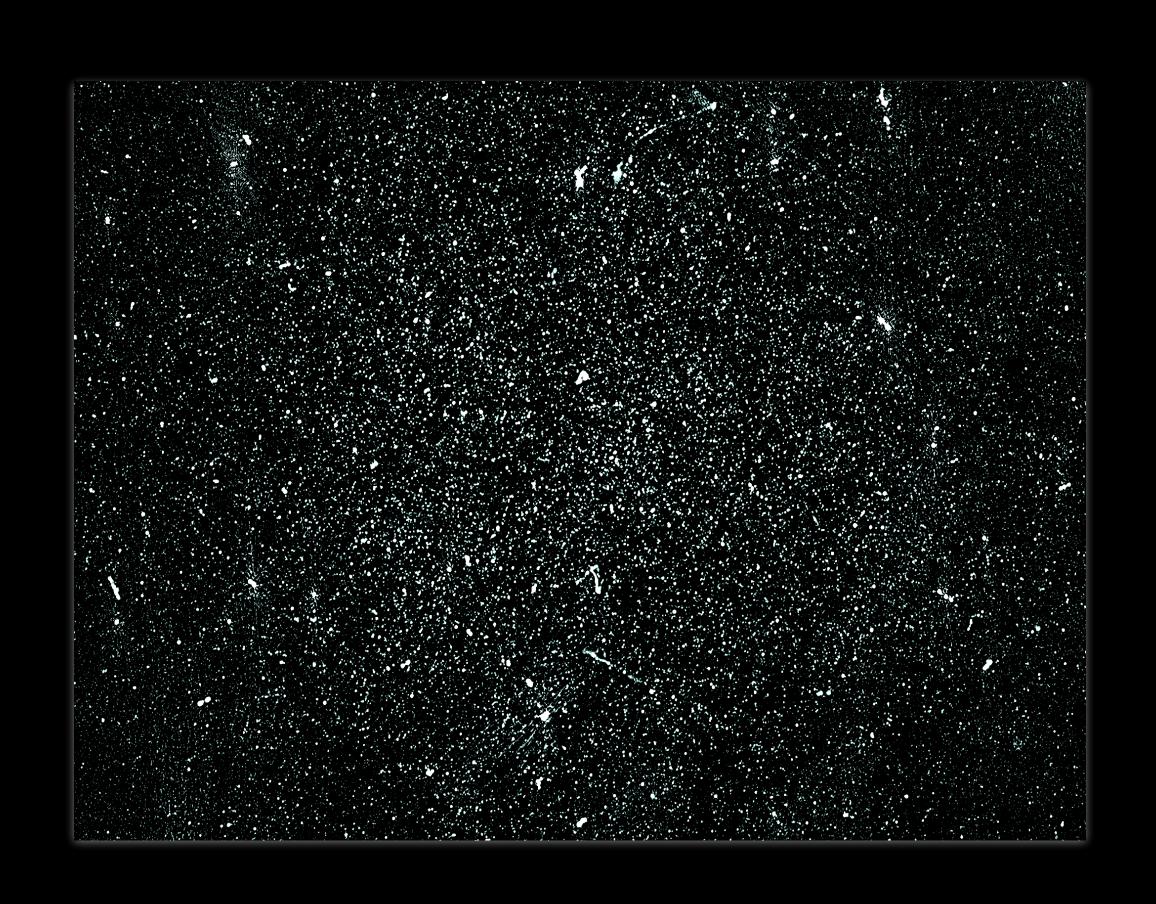


MeerKAT COSMOS

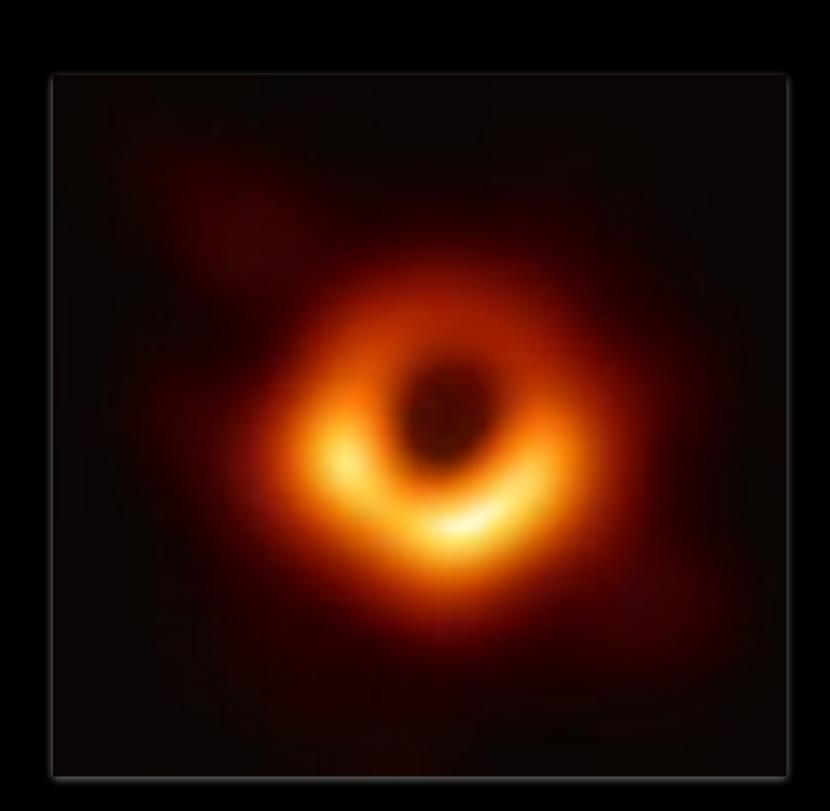
Single pointing
~3 uJy/beam; > 1 deg²
~5000 sources; ~16 hours

EHT has it easy and hard

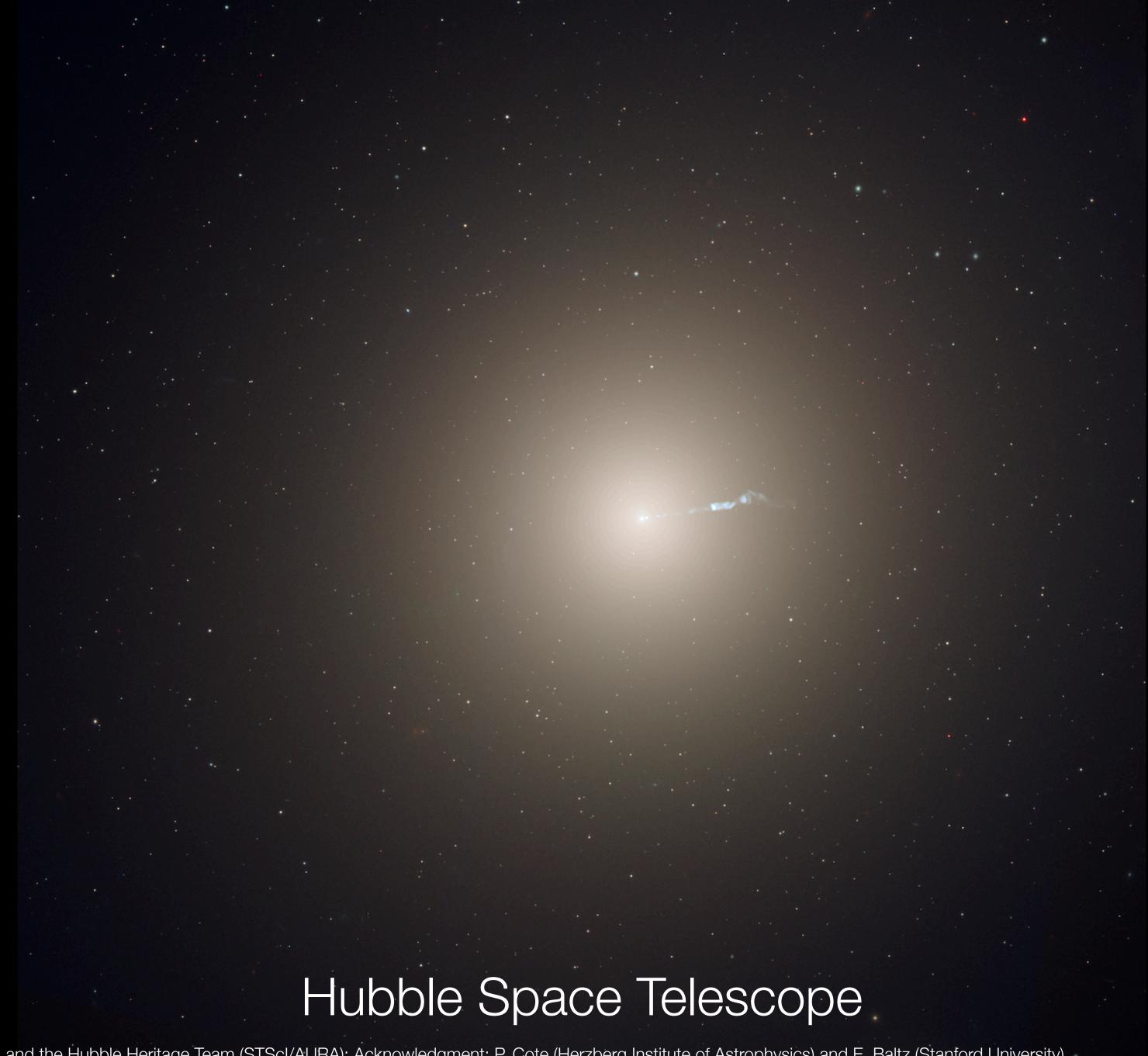
(time-variable source, gains, propagation effects, but just ~1 source)

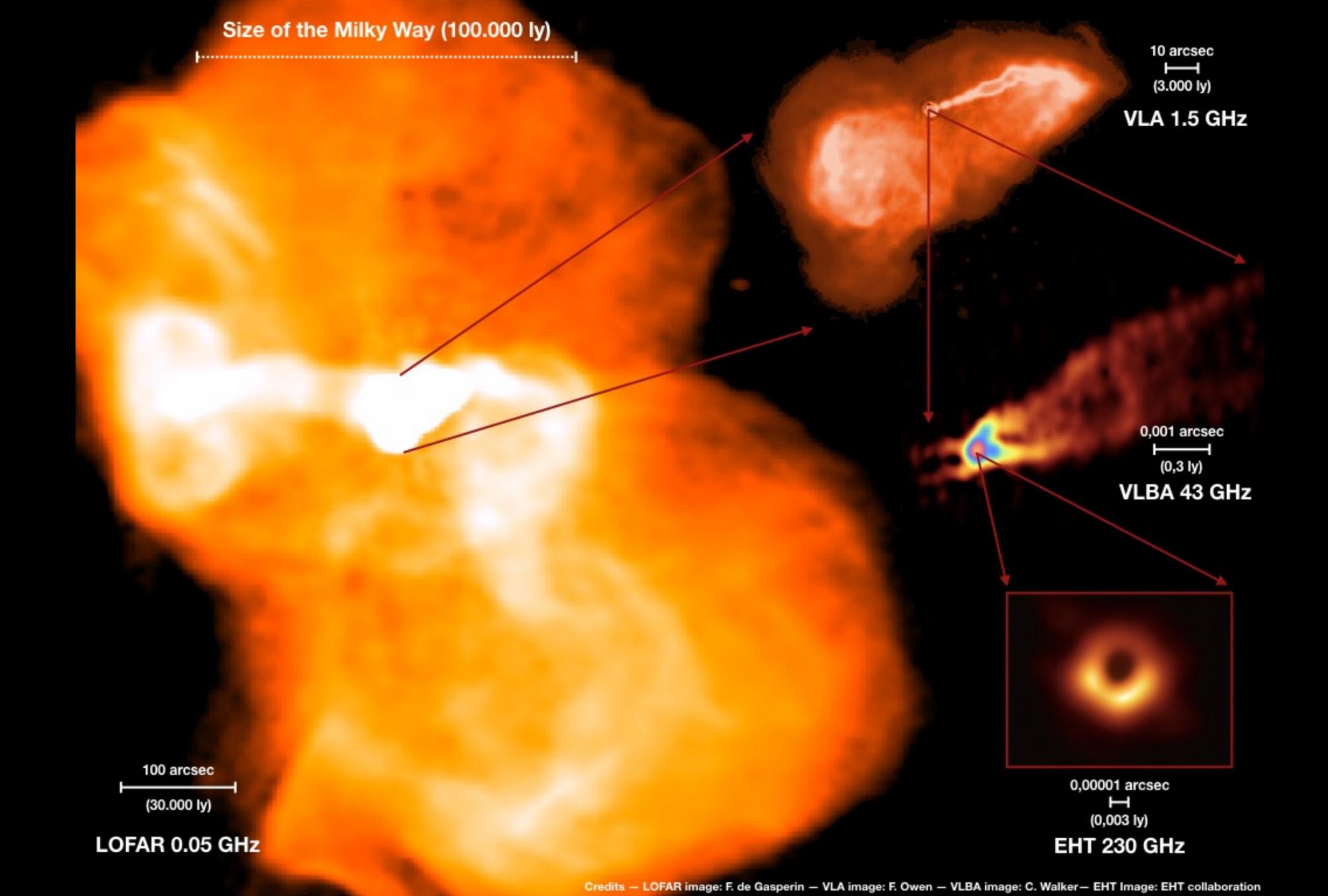


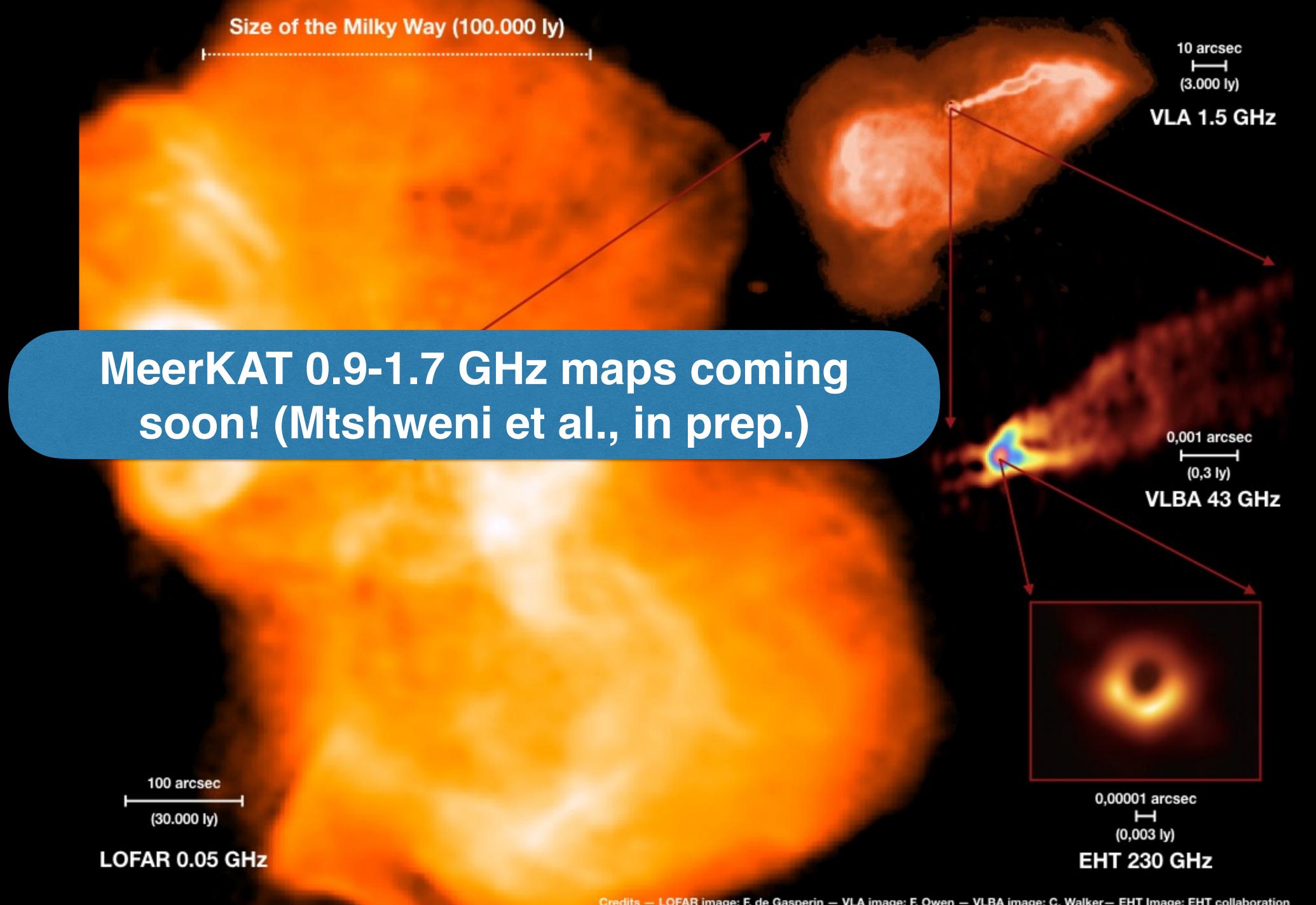
versus

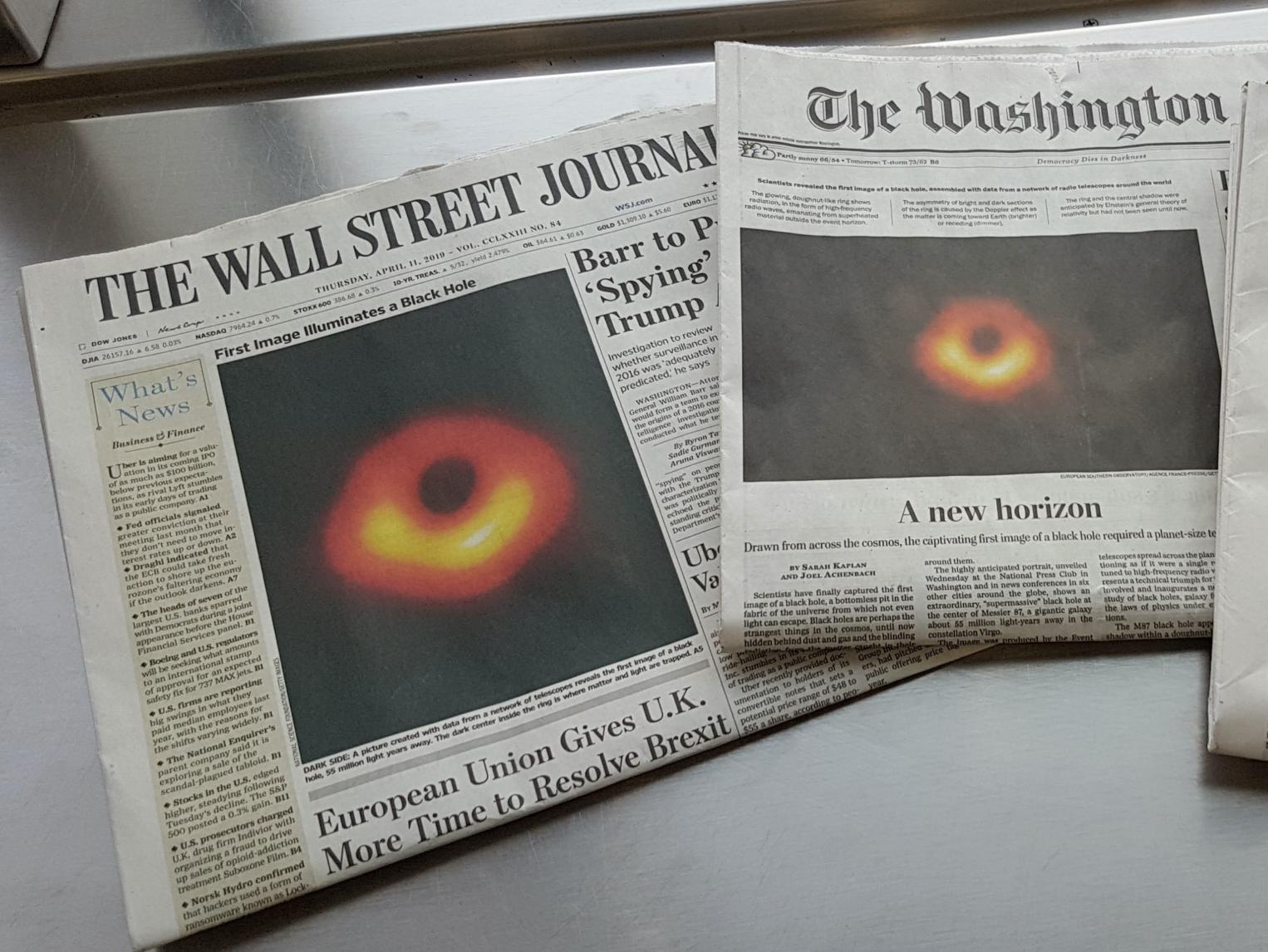


First imaging results on M87 Very Large Telescope credit: Chris Mihos (Case Western Reserve University)/ESO









A new horizon

Drawn from across the cosmos, the captivating first image of a black hole required a planet-size te

BY SARAH KAPLAN

Scientists have finally captured the first image of a black hole, a bottomless pit in the

The highly anticipated portrait, unveiled tioning as if it were a single refund to high-frequency radio v Washington and in news conferences in six resents a technical triumph for

That's Fit to Print.

Could be New Machine and some clouds.

THURSDAY, APRIL 11, 2019

Washington Edition
Today, Surables and some clouds.

THURSDAY, APRIL 11, 2019

NEWS ANALYSIS

Israelis Lean On Stability

With Leader Netanyahu Is Symbol of Cherished Security

Peering Into Light's Graveyard: The First Image of a Black Hole

By DENNIS OVERBYE

Astronomers announced on captured an image of the unobadys so deep and dense that not even light can escape it.

For years, and for allulus.

By DENNIS OVERBYE

Astronomers announced on captured an image of the unobady that at last they had servable: a black hole, a cosmic even light can escape it.

For years, and for allulus.

Astronomers announced on captured an image of the unobady that at last they had servable: a black hole, a cosmic even light can escape it.

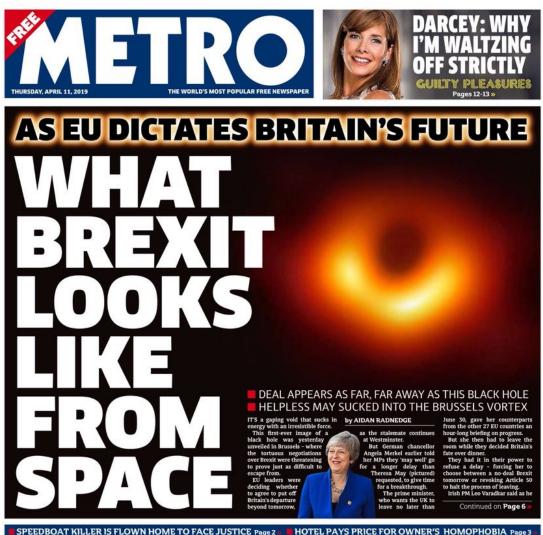
For years, and for allulus.

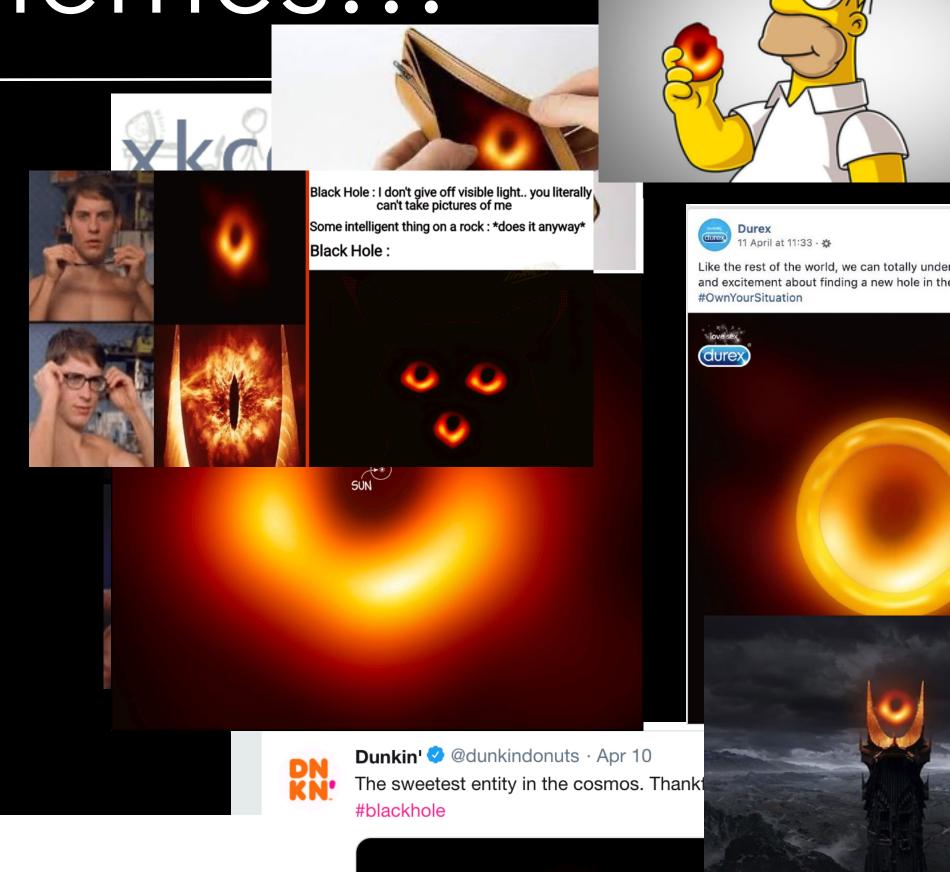
For years, and for a trophysics, and director of the effort to capture the image, during a lion light-years away from Earth, resembled the Eye of Sauron, a re-



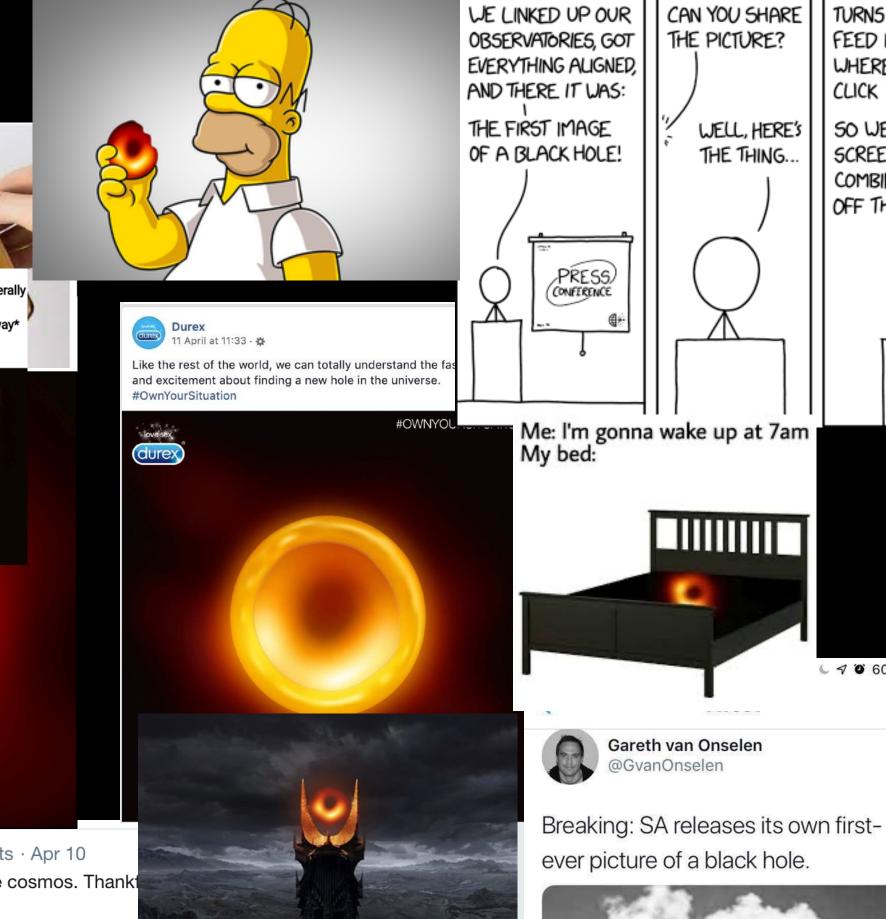


Commercial VS Reality





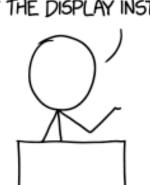
Second picture of black hole.







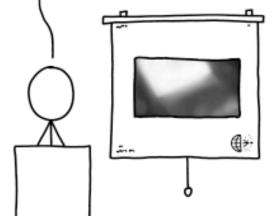
SCREENSHOT, BUT THE KEY COMBINATION KEPT TURNING OFF THE DISPLAY INSTEAD.



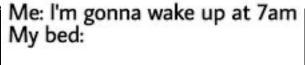
TO TAKE A PICTURE OF THE SCREEN, BUT I WAS TOO SLOW. THE OBSERVATION HAD ENDED. WE'RE PLANNING TO TRY AGAIN

I GRABBED MY PHONE AND TRIED

NEXT YEAR, AND WE'LL DEFINITELY RECORD THE SCREEN THIS TIME.



© कि.⊪ 69% **व**





Gareth van Onselen

Mogawatt Park

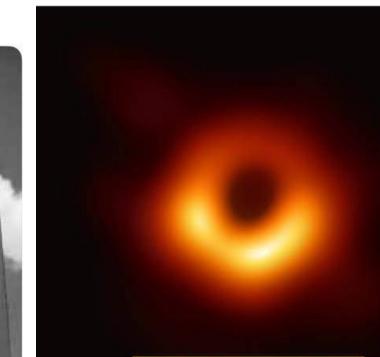
@GvanOnselen





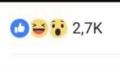
22:07 🎜 🗯 🛂

Congratulations to SA's own Professor Roger Deane and Dr Iniyan Natarajan who are among the group of international researchers who captured the first image of a black hole. 💥



Your stomach at 11am when you skip breakfast.

88 Comments • 163 Shares











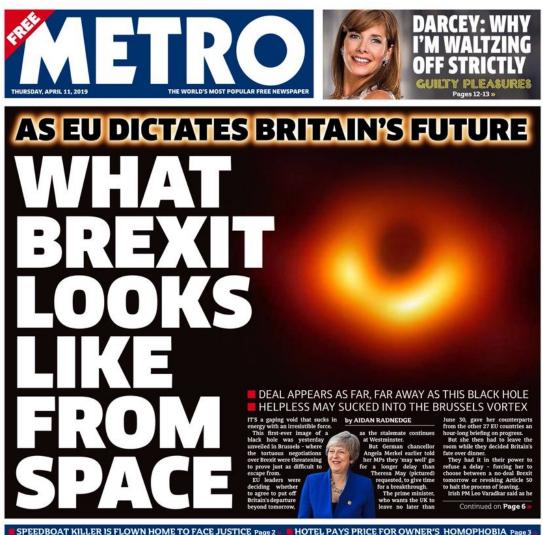
111

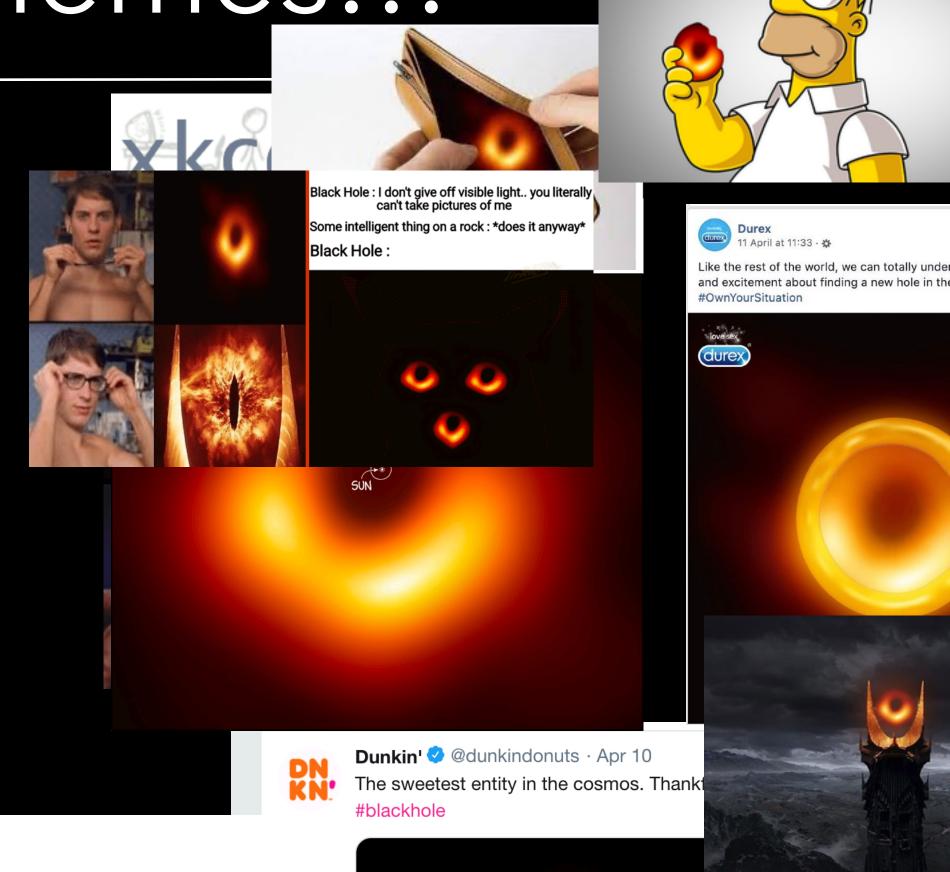
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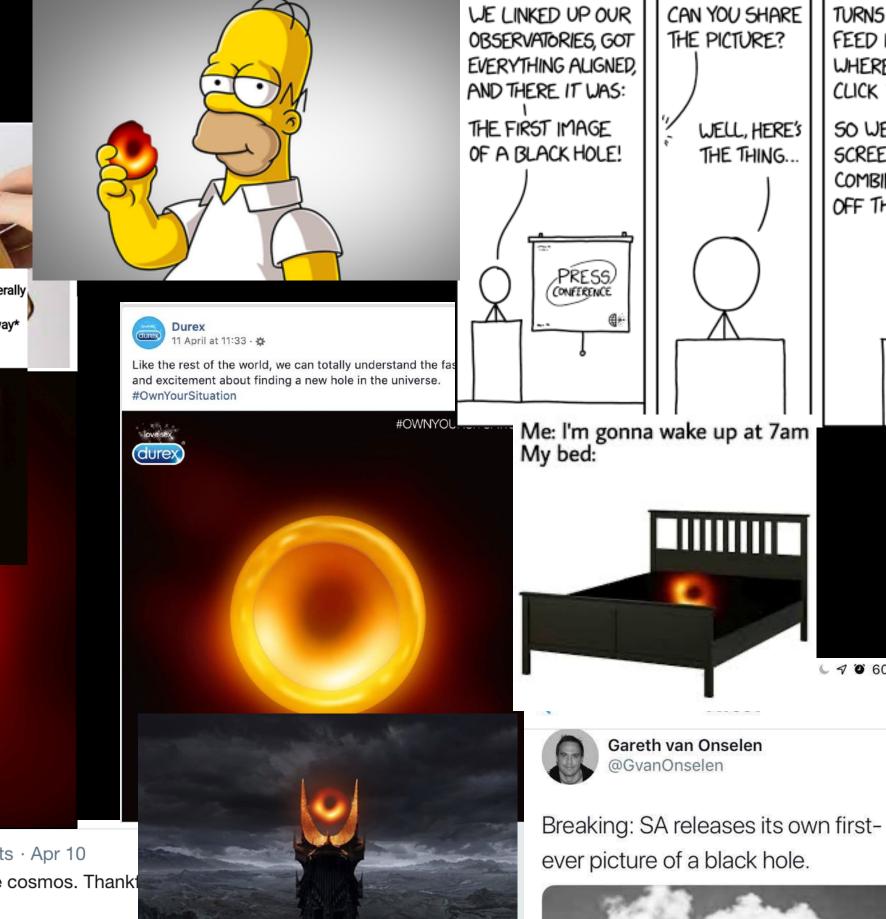


Commercial VS Reality





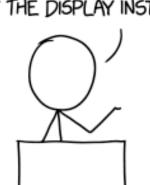
Second picture of black hole.







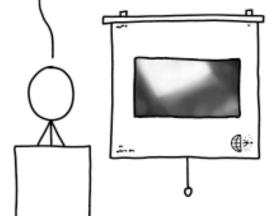
SCREENSHOT, BUT THE KEY COMBINATION KEPT TURNING OFF THE DISPLAY INSTEAD.



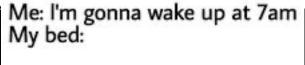
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Gareth van Onselen

Mogawatt Park

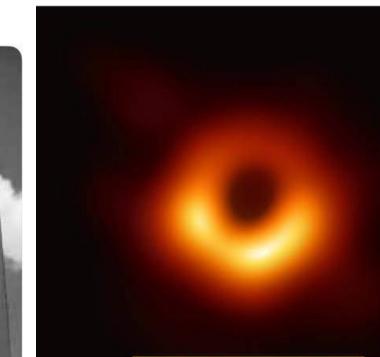
@GvanOnselen





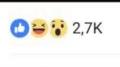
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...and six peer-reviewed journal articles

First M87 Event Horizon Telescope Results. I. The Shadow of the Supermassive Black Hole

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L1

First M87 Event Horizon Telescope Results. II. Array and Instrumentation

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L2

First M87 Event Horizon Telescope Results. III. Data Processing and Calibration

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L3

First M87 Event Horizon Telescope Results. IV. Imaging the Central Supermassive Black Hole

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L4

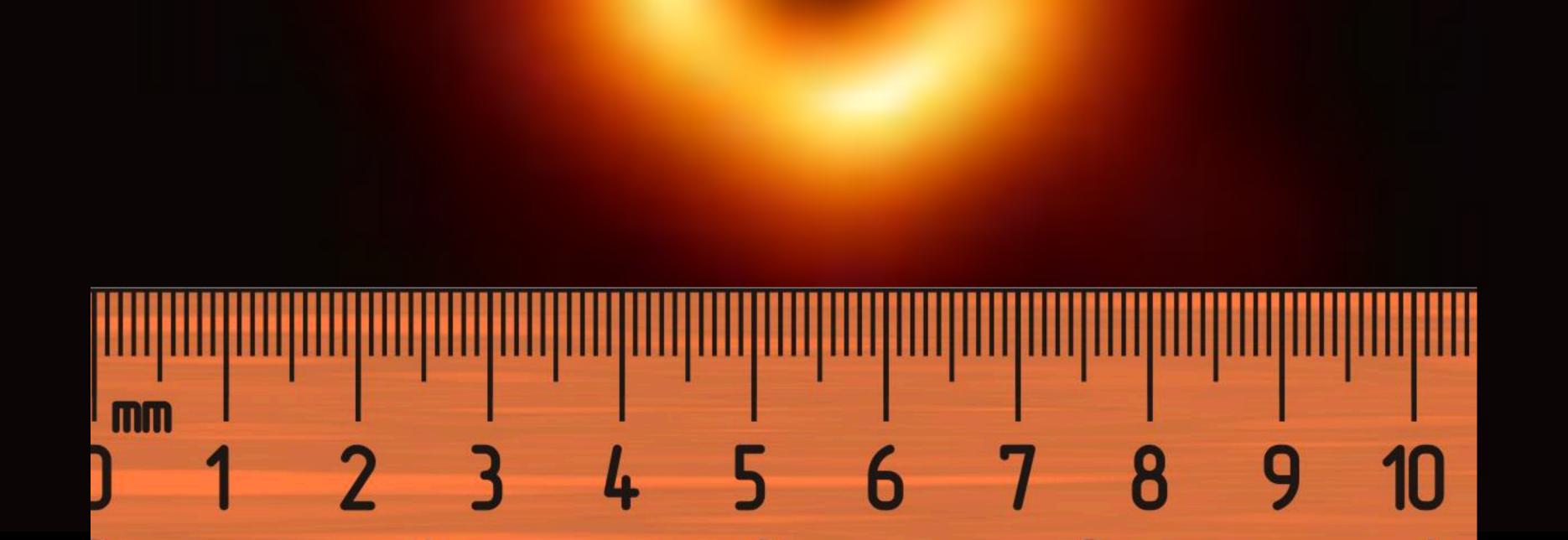
First M87 Event Horizon Telescope Results. V. Physical Origin of the Asymmetric Ring

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L5

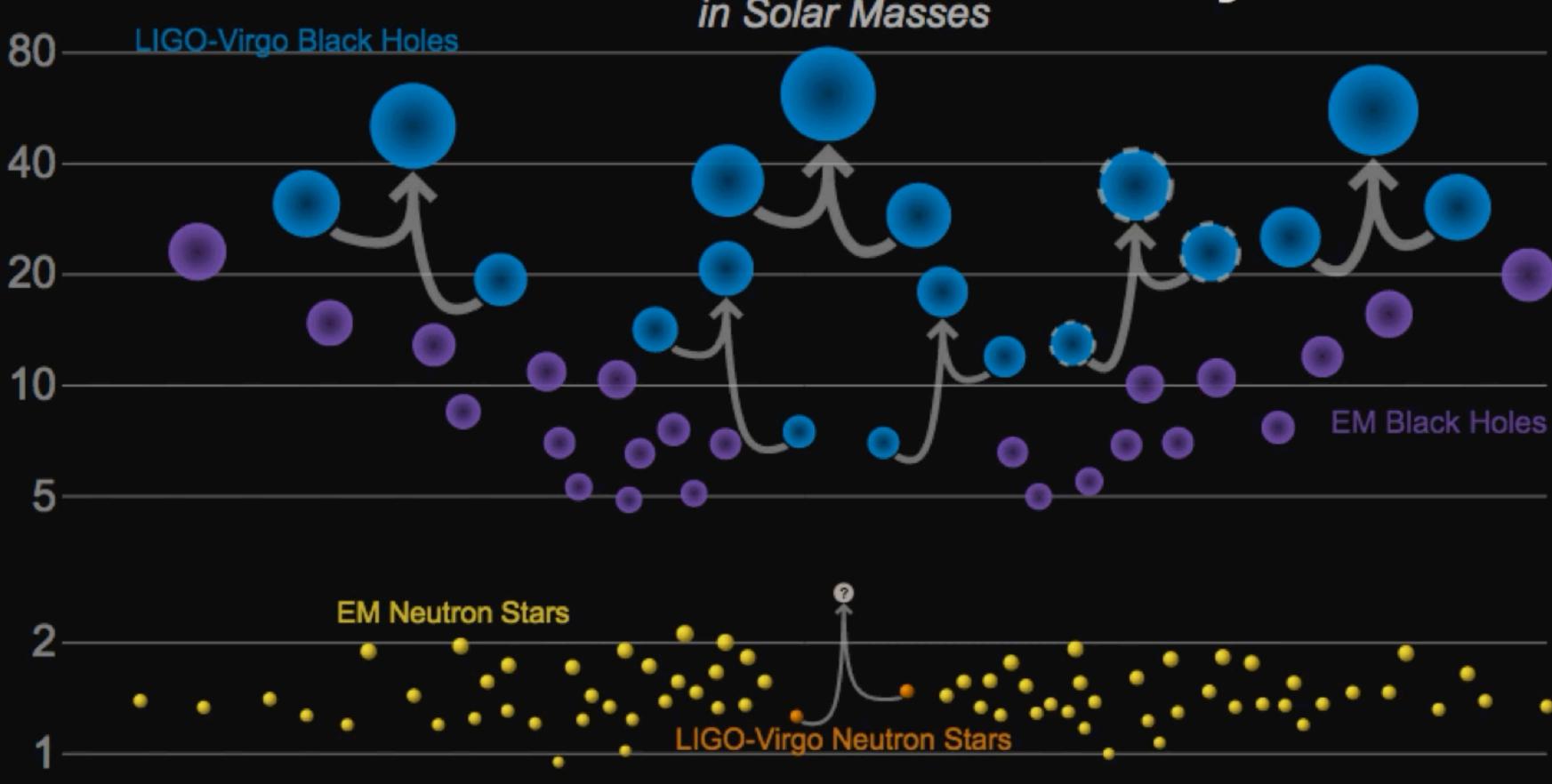
First M87 Event Horizon Telescope Results. VI. The Shadow and Mass of the Central Black Hole

The Event Horizon Telescope Collaboration et al. 2019 ApJL 875 L6

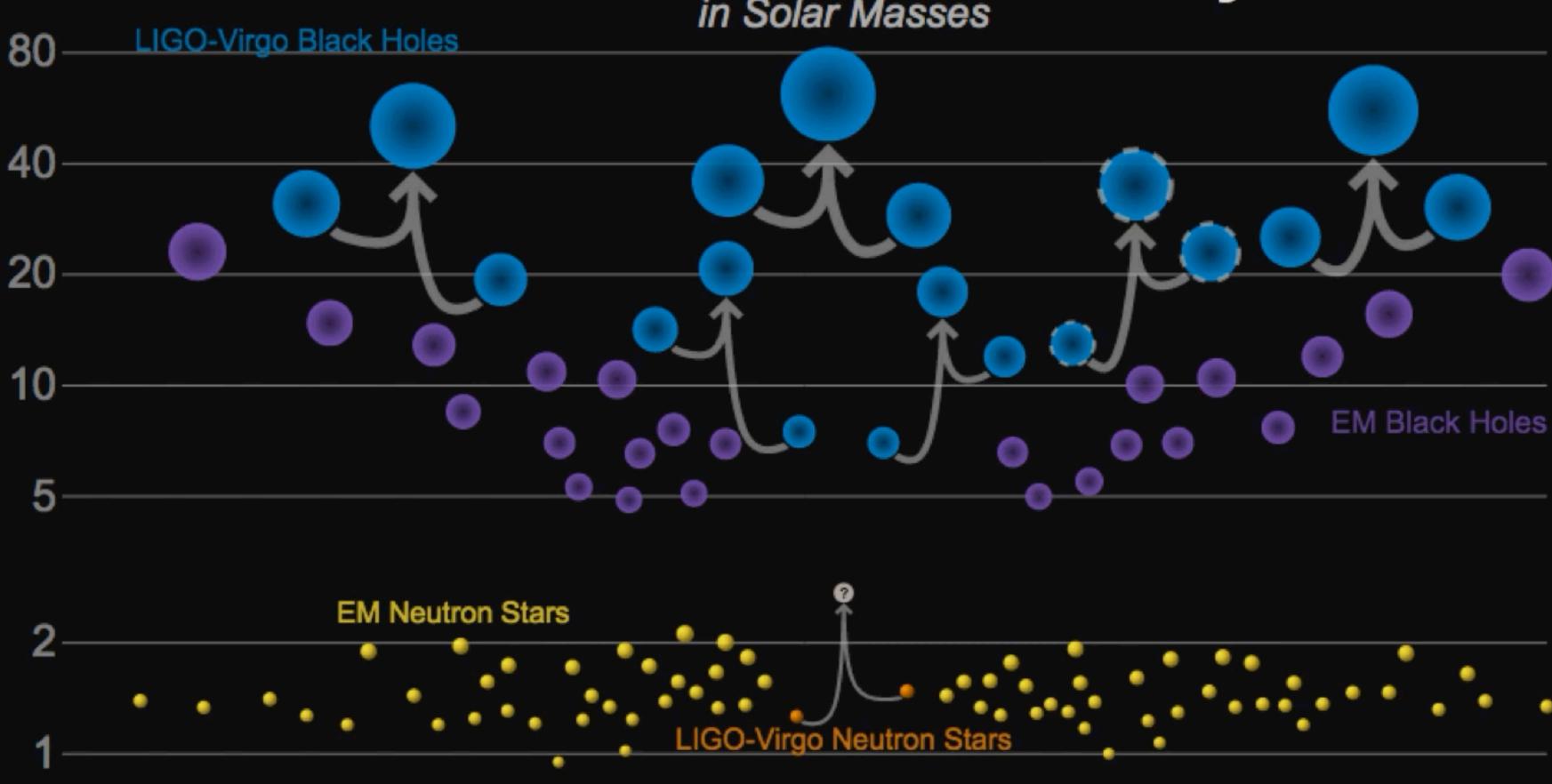




Masses in the Stellar Graveyard



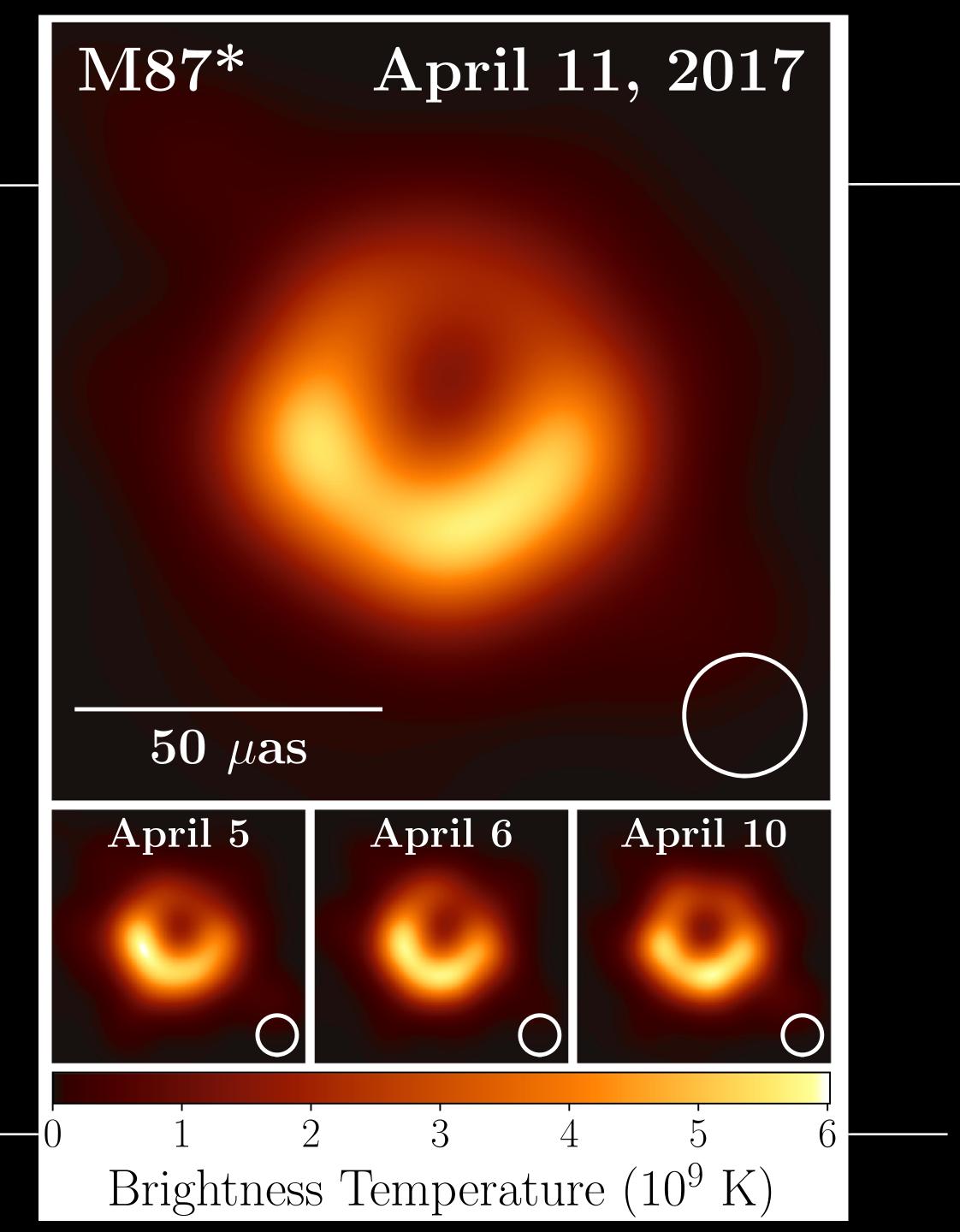
Masses in the Stellar Graveyard



Not just 1 image!

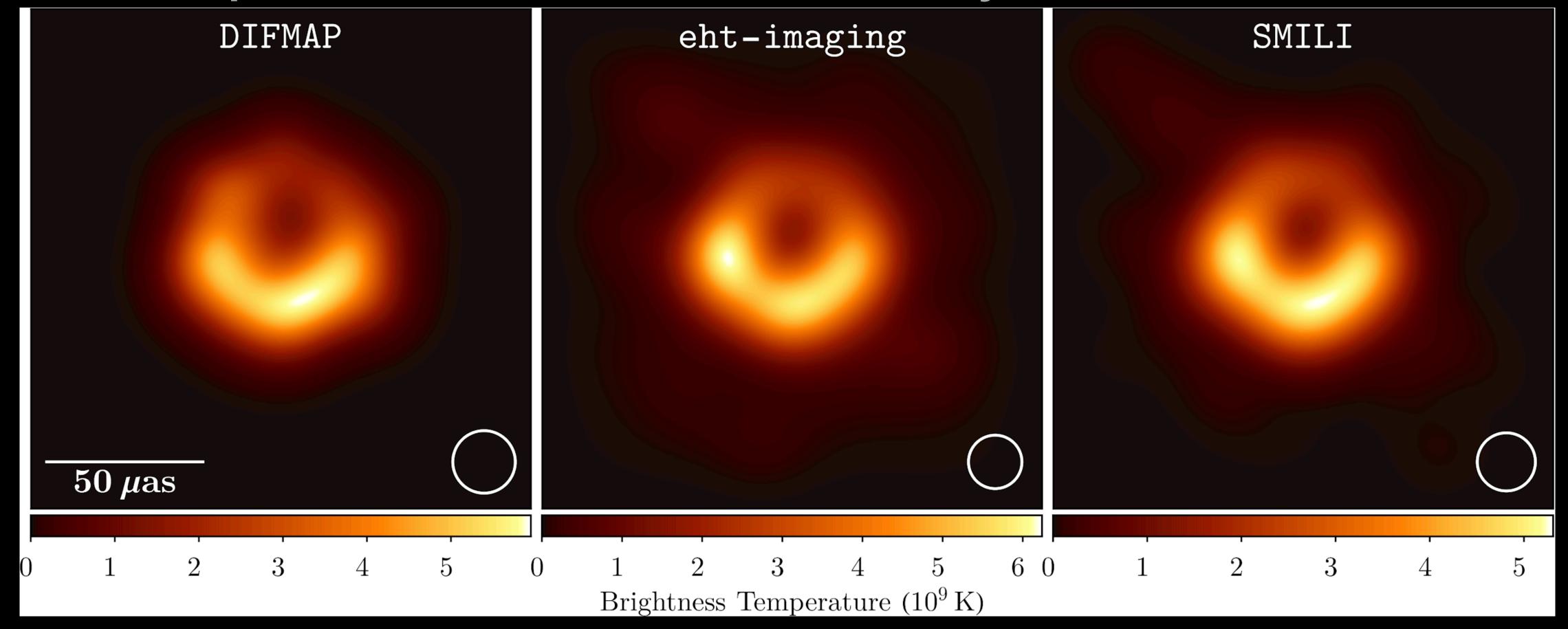
- Average of three different imaging algorithms = consensus image
- Convolved with a 20 uas beam
- Consistent structure over four days, although some super-resolved structure be real
- A next step is to compare static versus dynamic source structure models





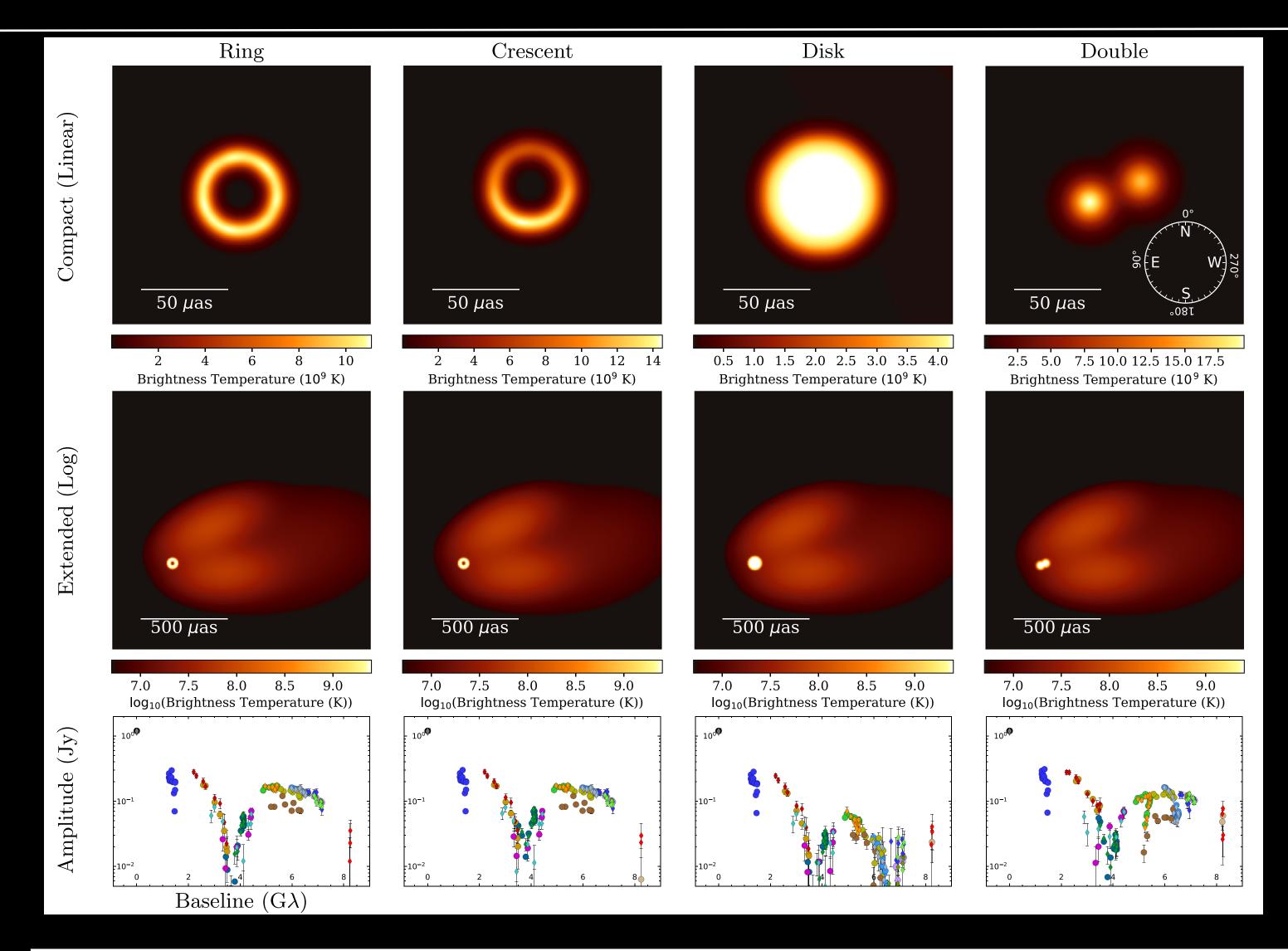
Agreement with a trio of imaging algorithms

Fiducial images of M87 for April 11 restored to an equivalent resolution show remarkably similar structure

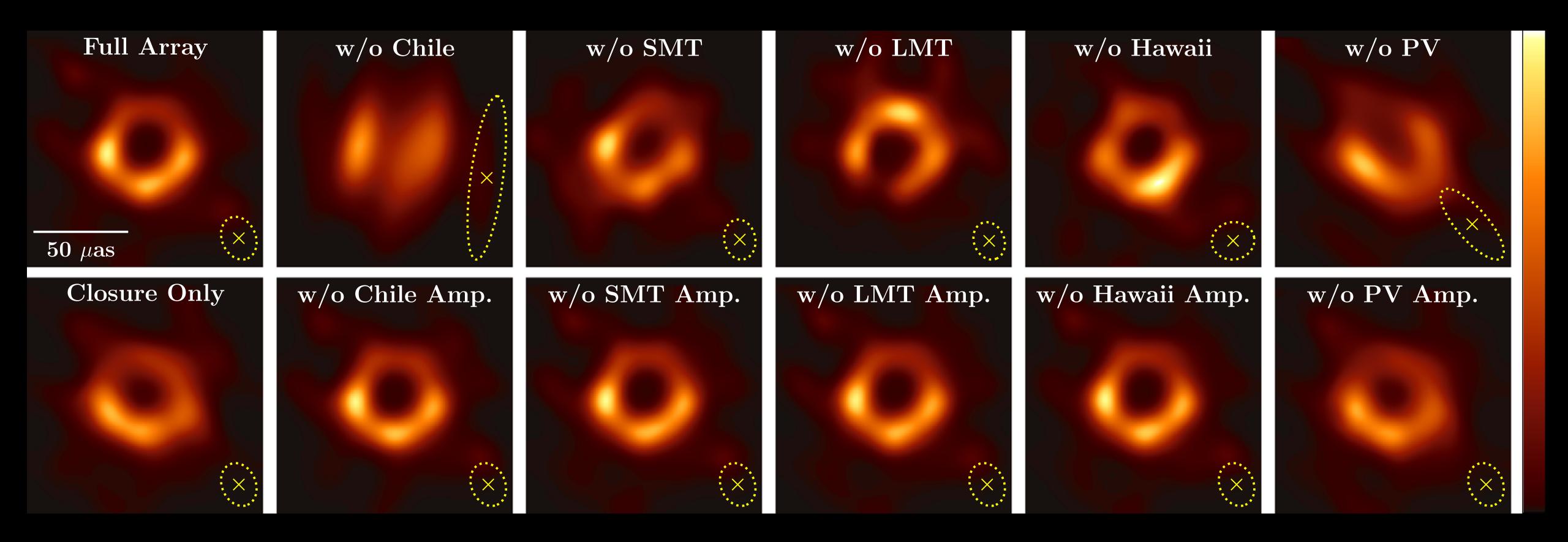




Imaging tests



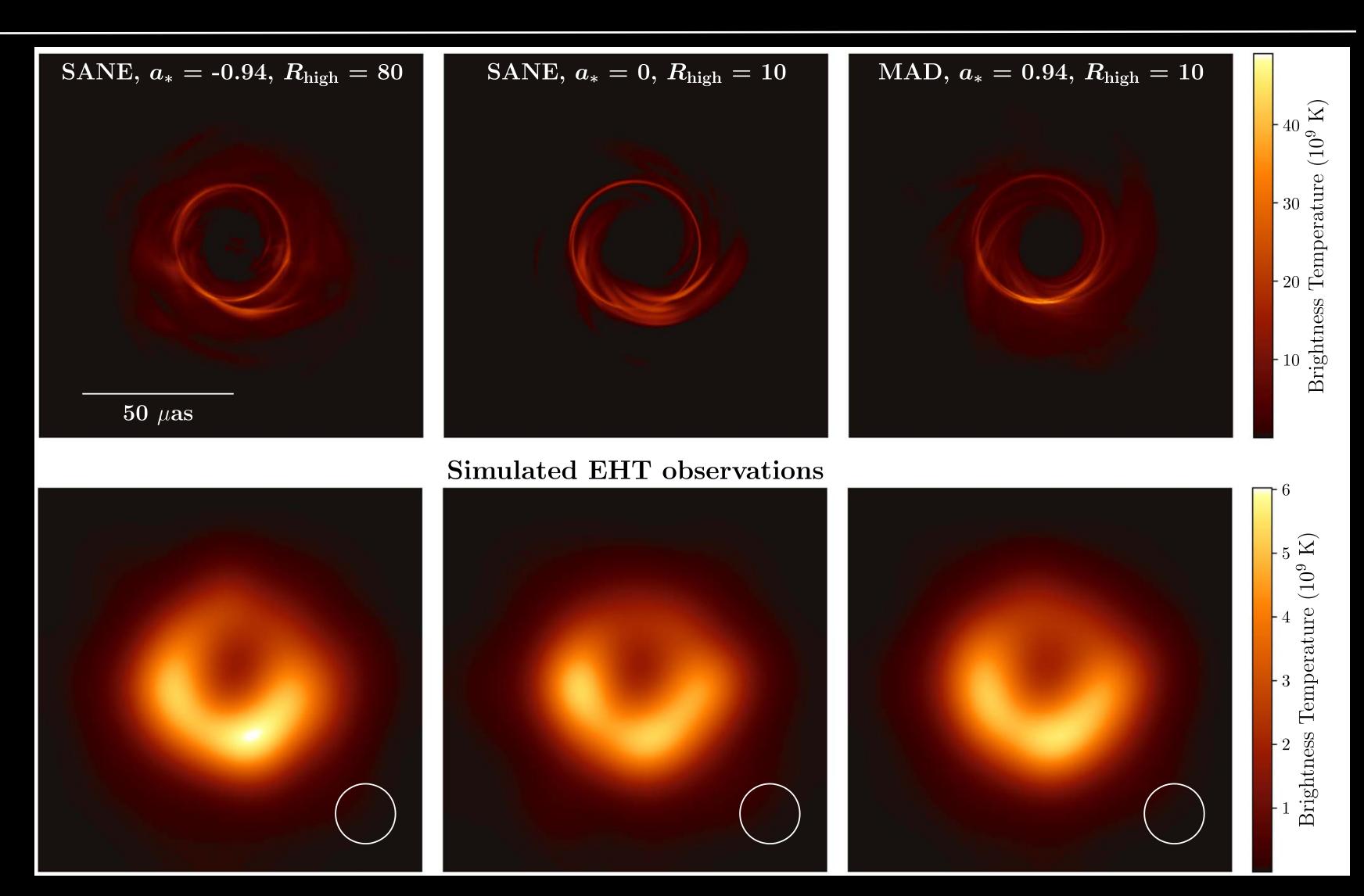




Comparison with simulation

GRMHD simulations generally consistent with image (despite wide range of BH spin, T_e/T_p ratio, magn field model)

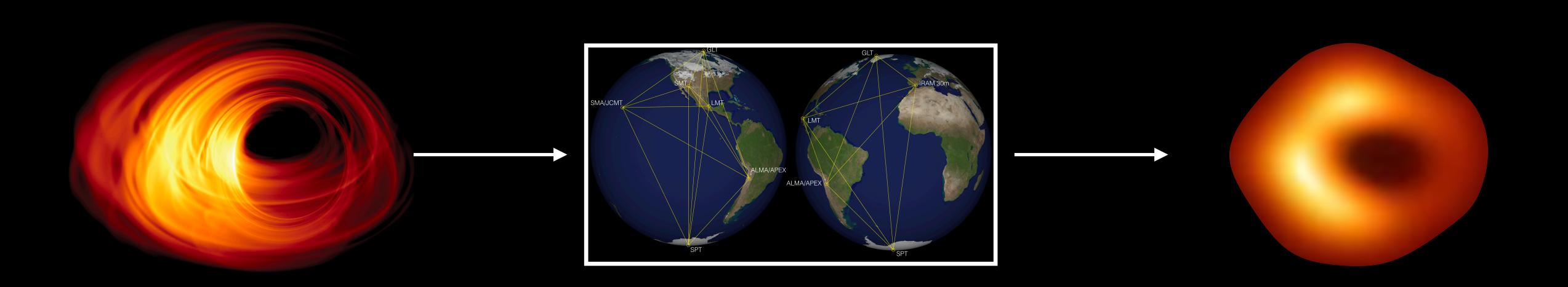
Passed through instrument simulator + official EHT calibration and imaging pipeline





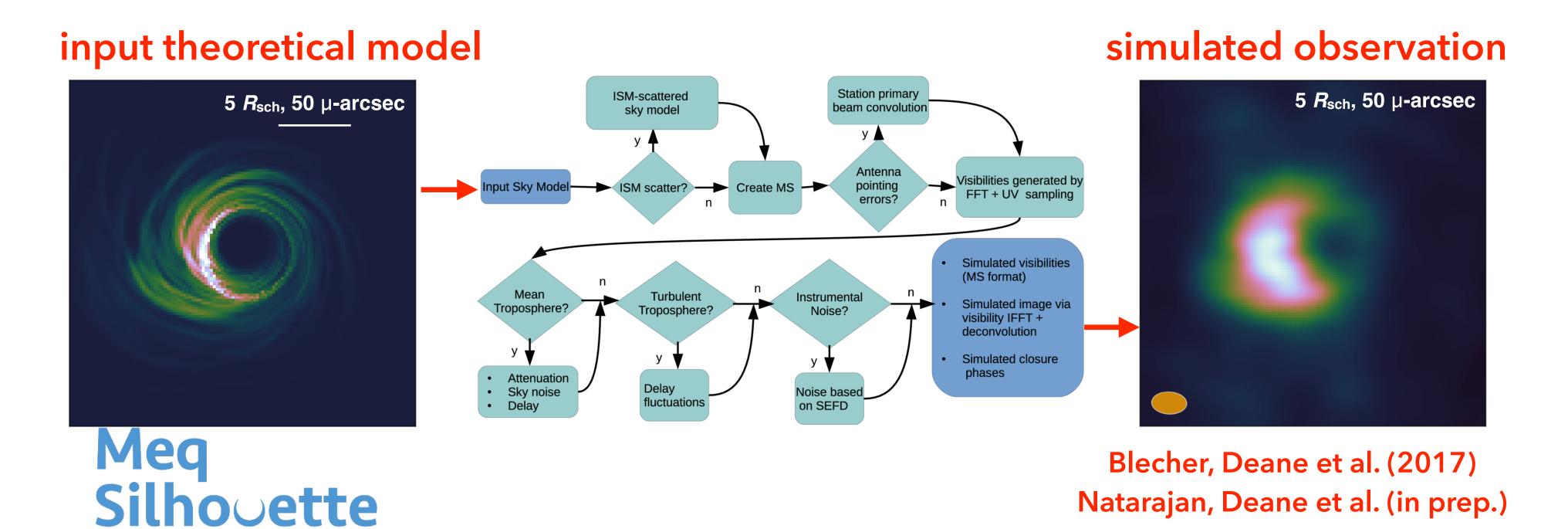
Suite of new tools for VLBI (not just for EHT)

Understanding instrumental and propagation transfer functions

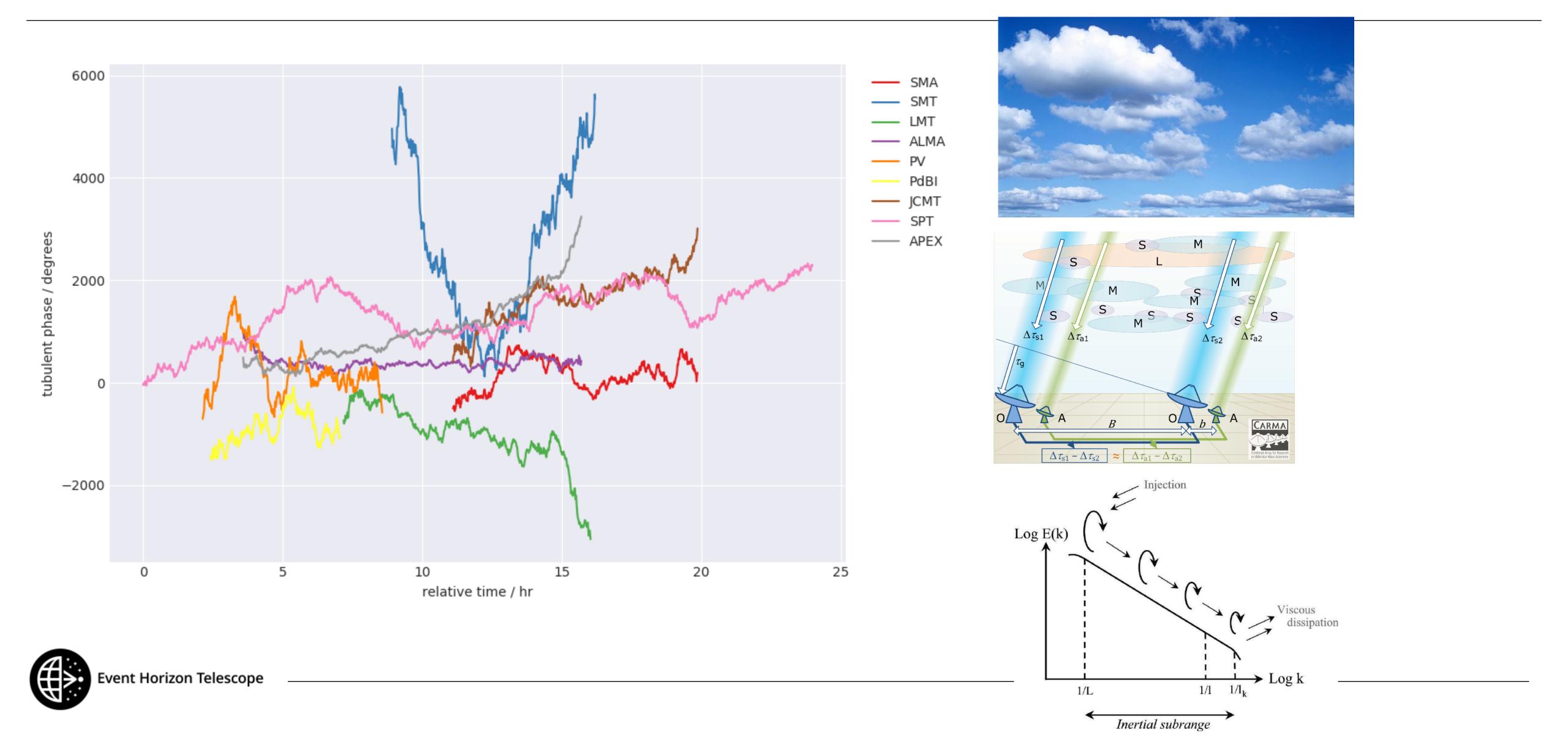


EHT synthetic data

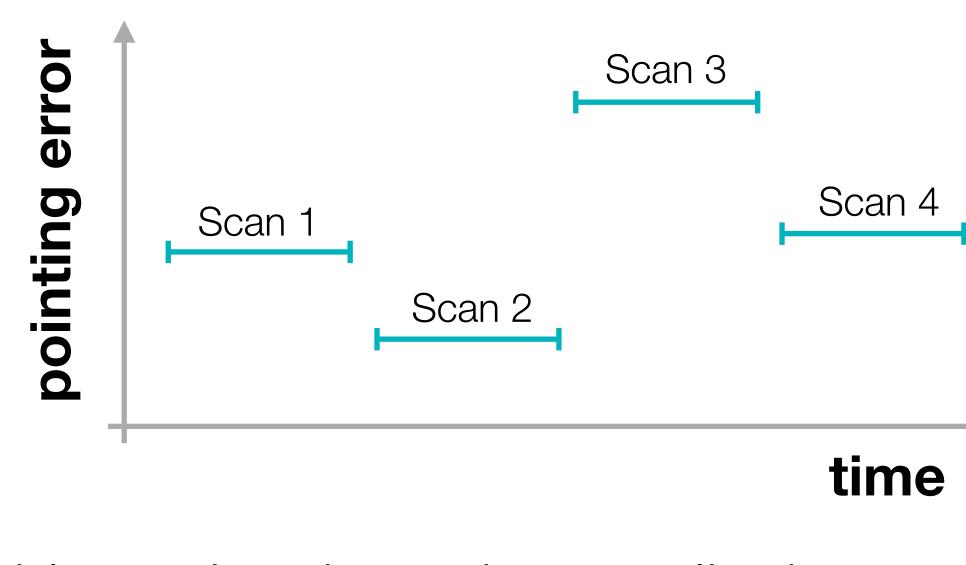
- Given the wide range of stations, propagation effects, calibration uncertainties, and complex source structure, sophisticated synthetic data is a critical resource for testing purposes
- These same synthetic data engines are used in probabilistic modeling of physical parameters from observables
- My group leads MeqSilhouette development for use in the EHT Consortium



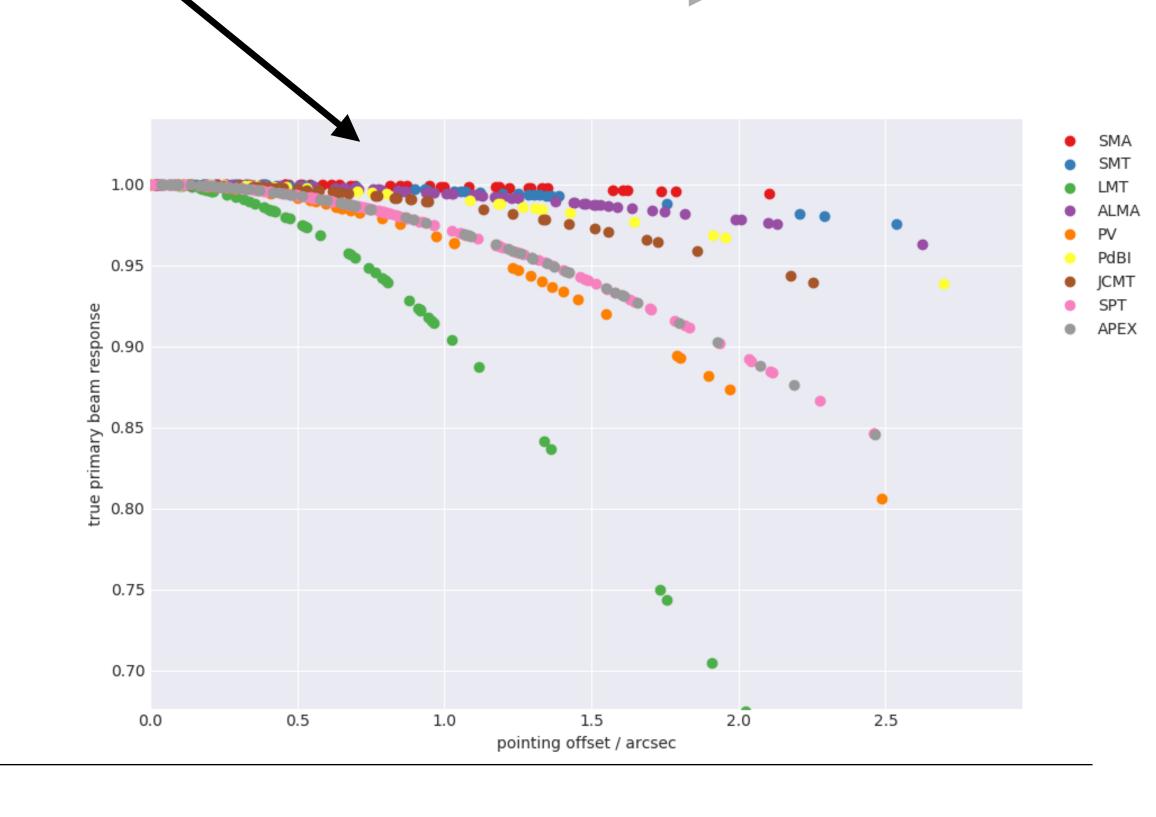
Turbulent troposphere



Antenna pointing error



- Time-variable, station-dependent amplitude errors
- Worst off for LMT (biggest dish)
- Introduction of systematics that must be understood
- Using realistic parameter enables feedback into potential array improvements



Scan 7

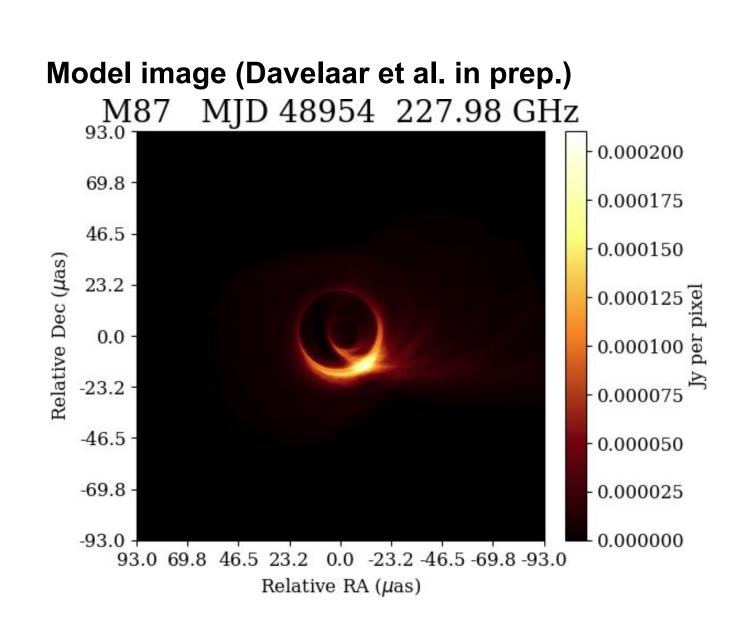
Scan 6

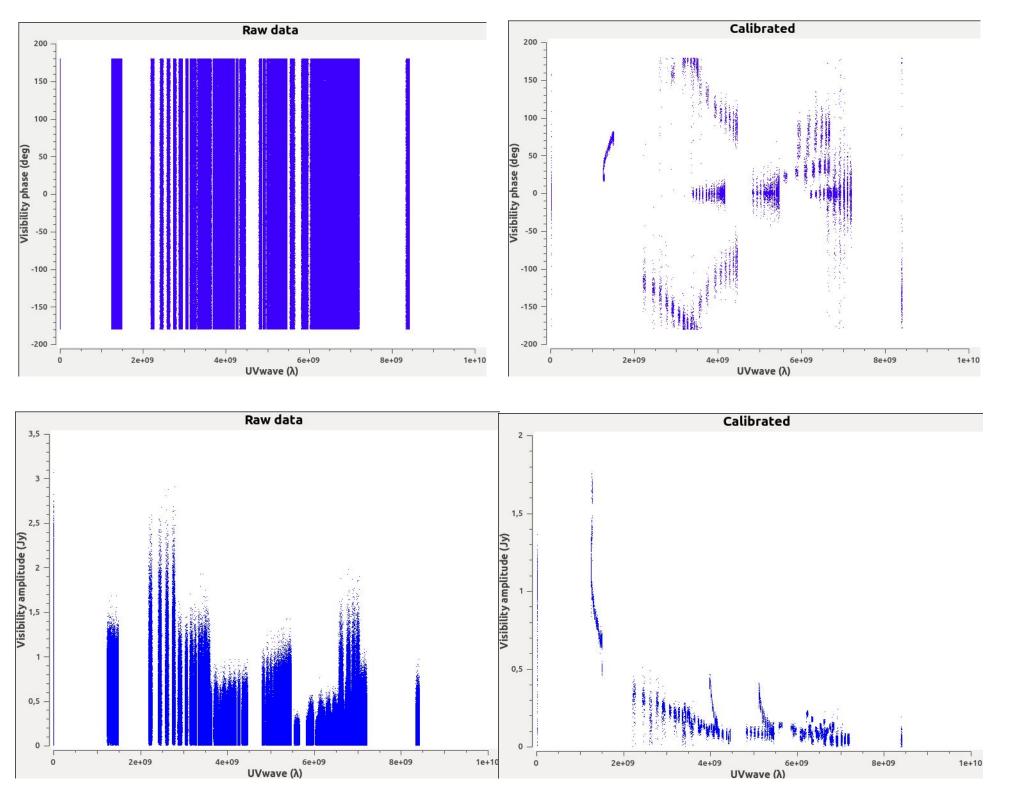
Scan 5

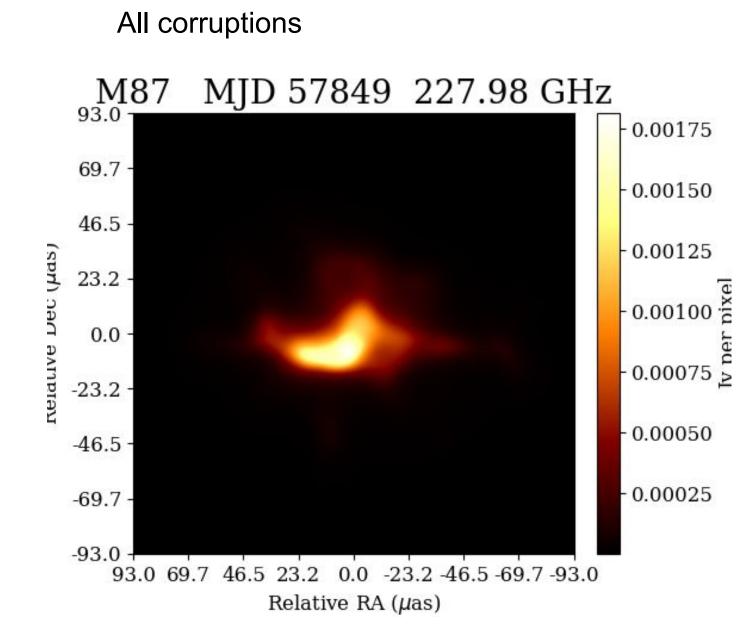


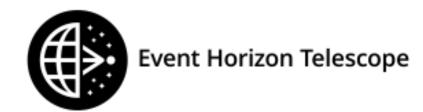
MeqSilhouette + rPICARD = SYMBA

Our synthetic data fed through the real EHT post-processing pipeline









Model Comparison: which is most consistent with the data?

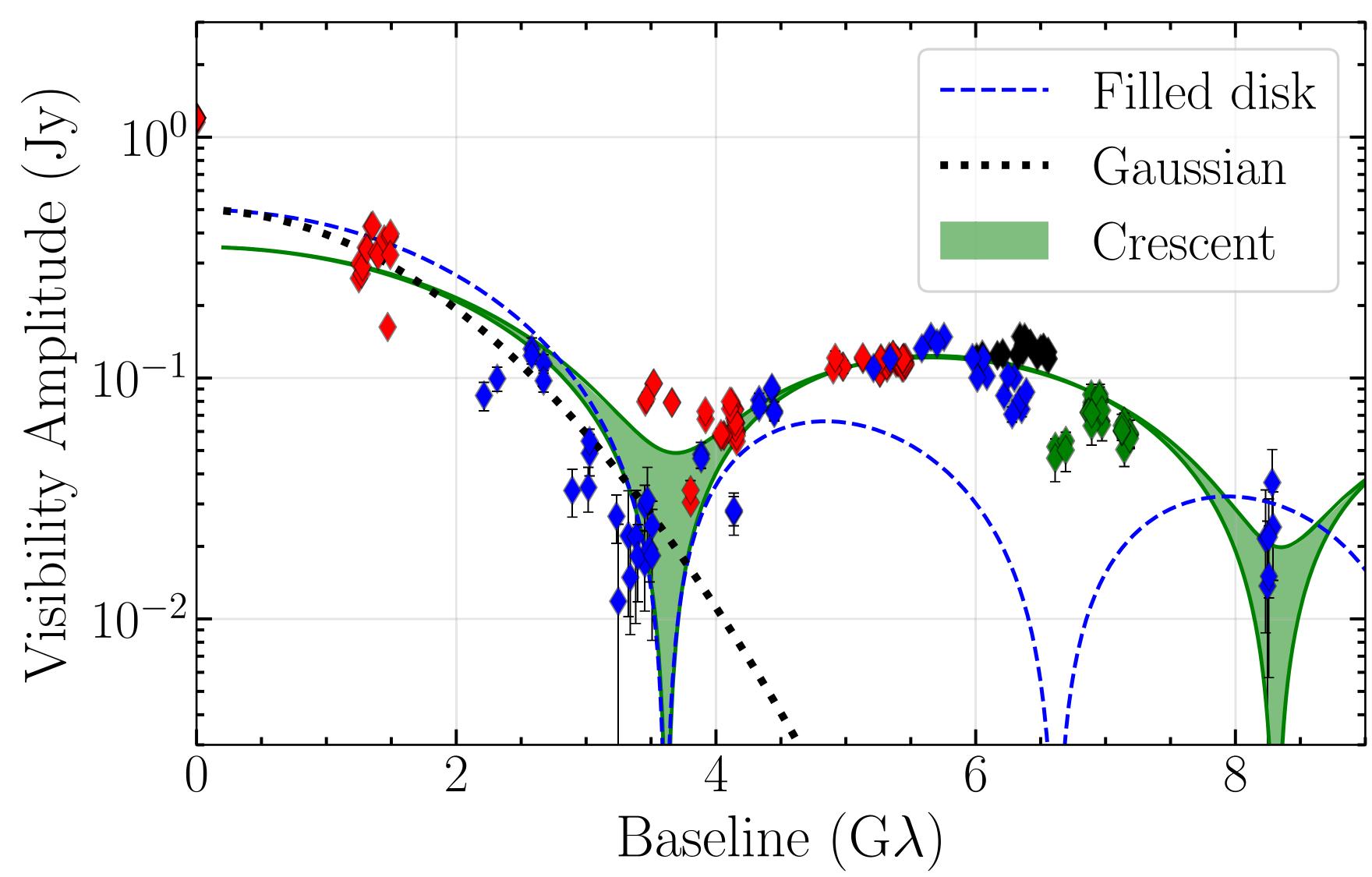
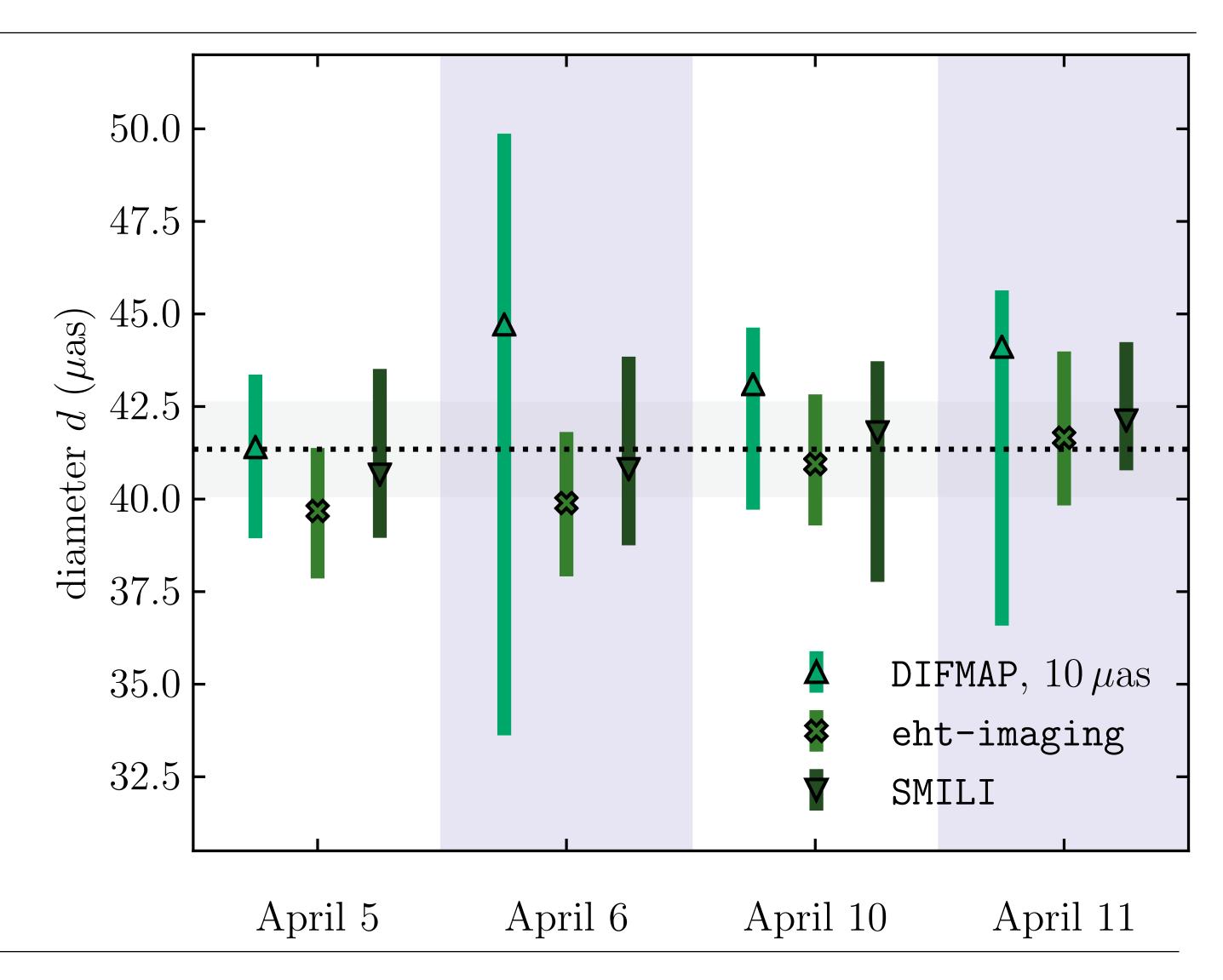


image-based fitting results

correlated noise

Consistency across parameter estimation algorithms, frequency bands, observing days

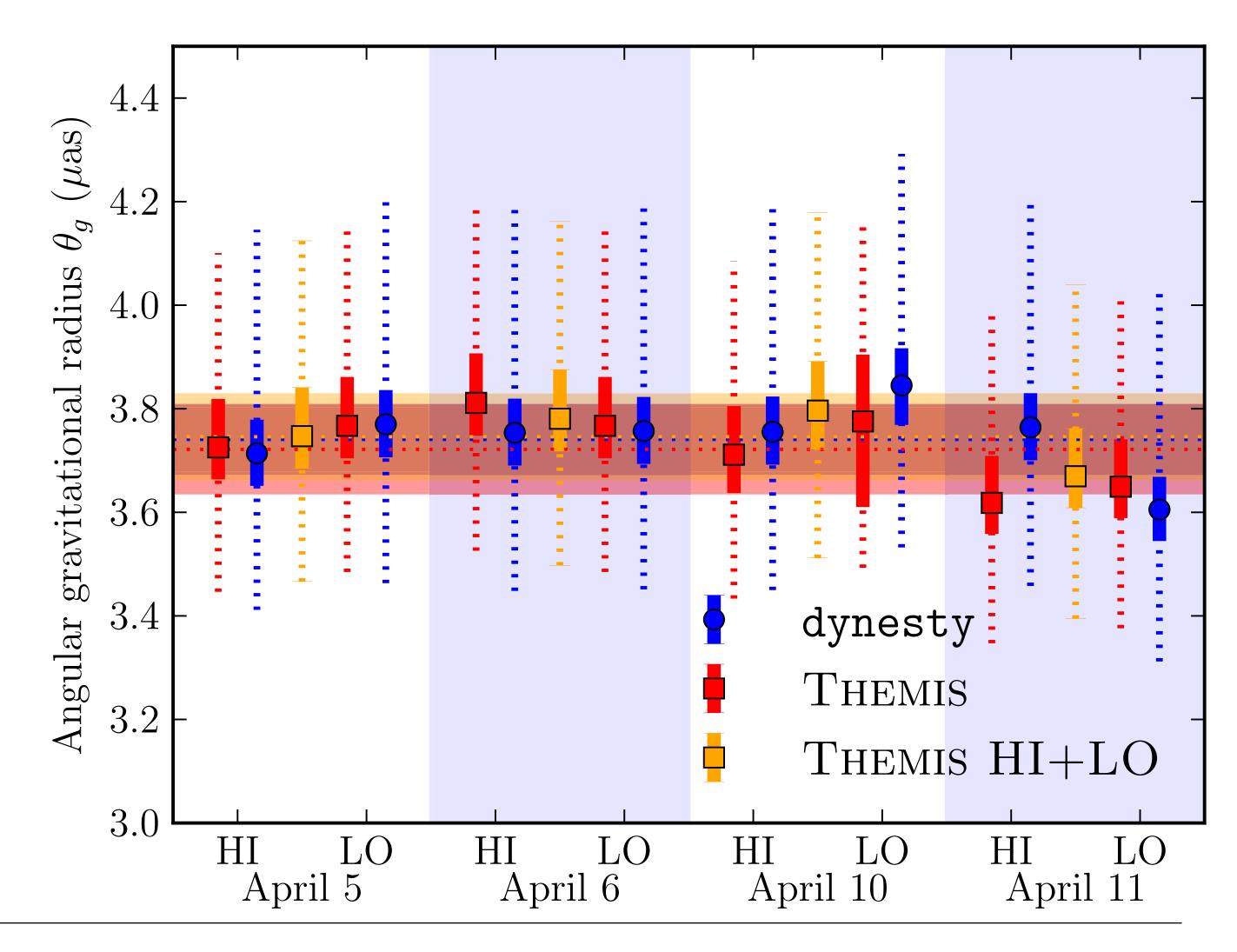


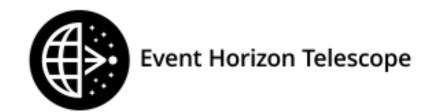


Visibility-based fitting results

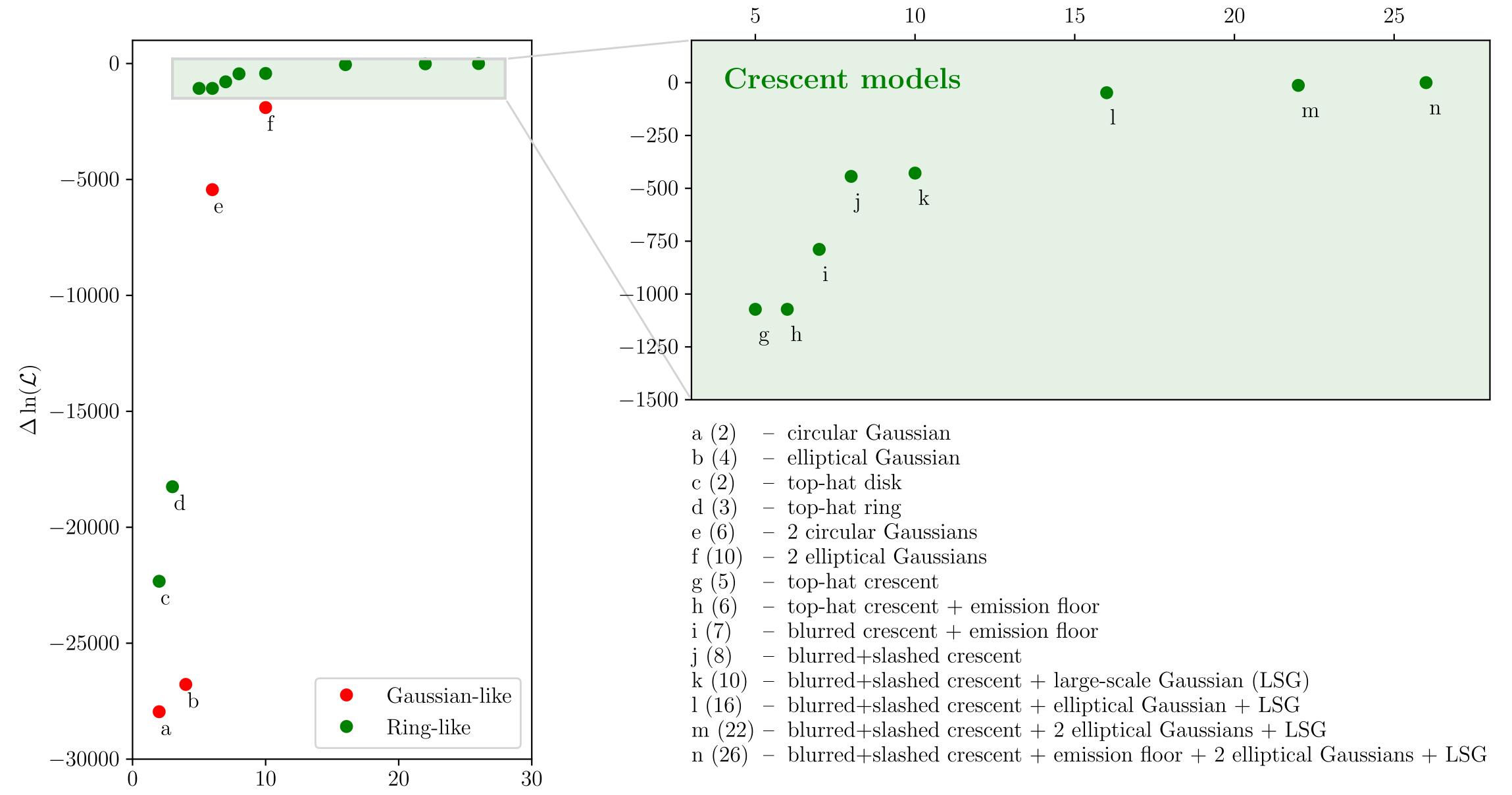
uncorrelated noise

Consistency across parameter estimation algorithms, frequency bands, observing days



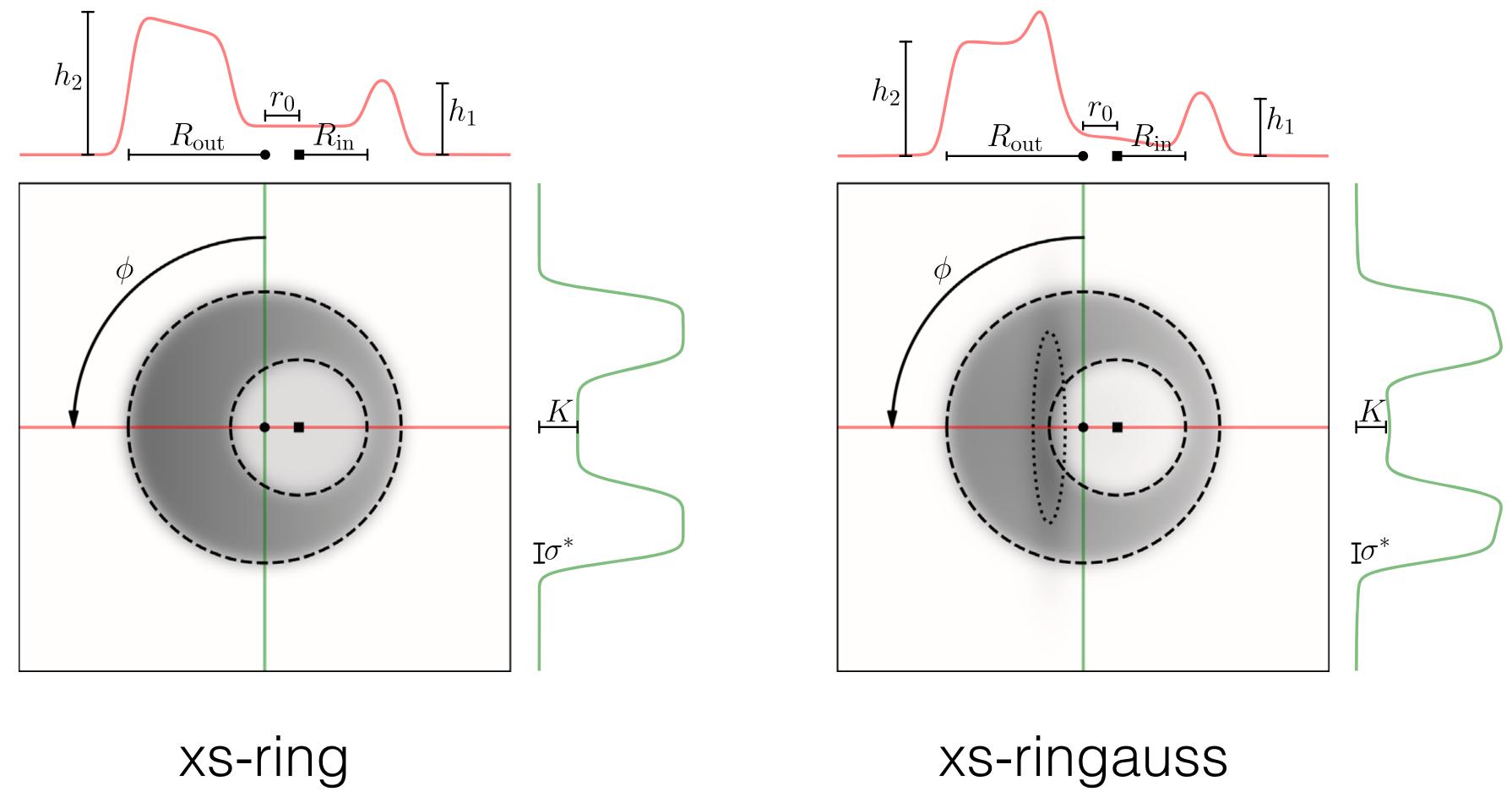


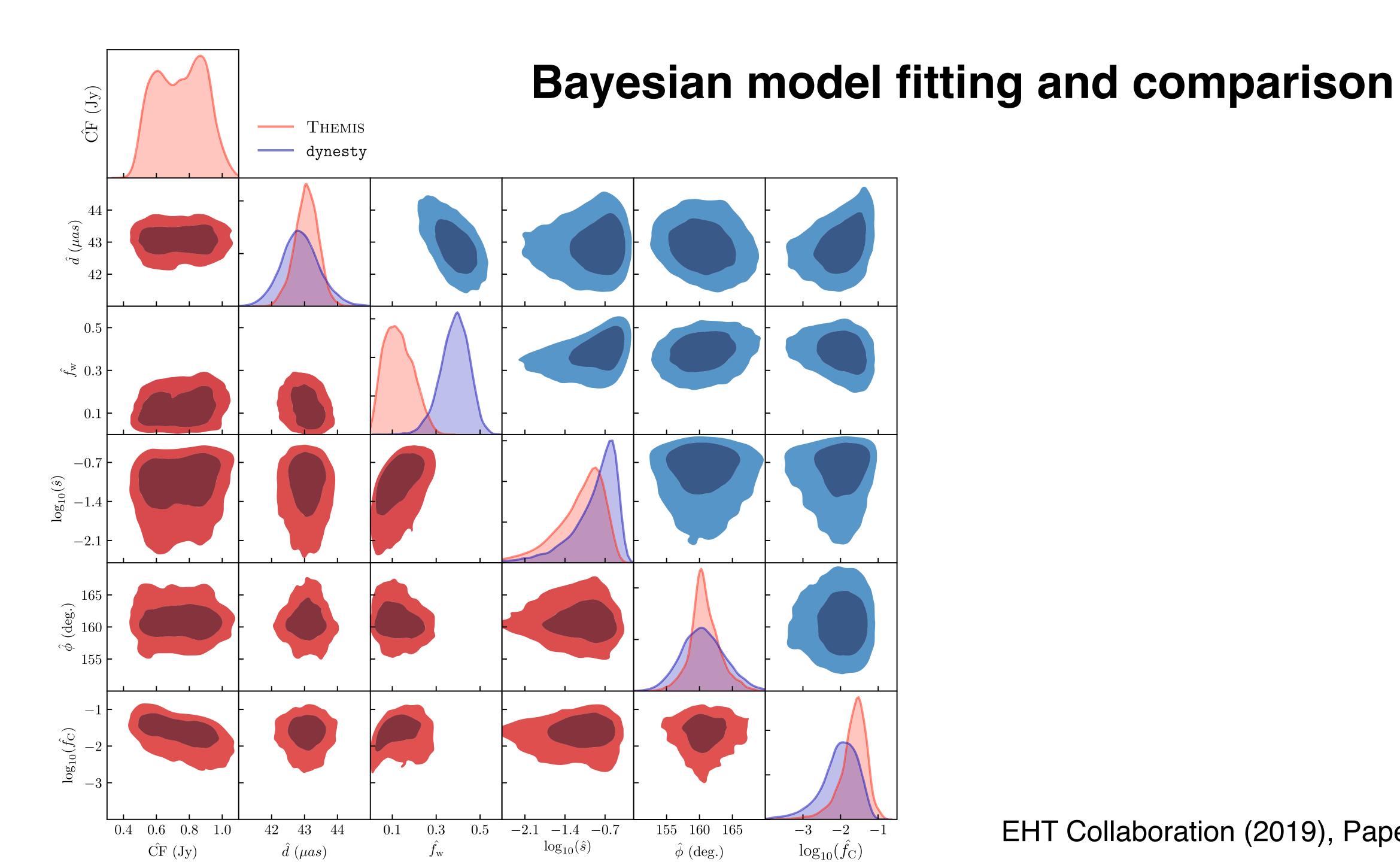
Model Comparison: which is most consistent with the data?



Number of parameters

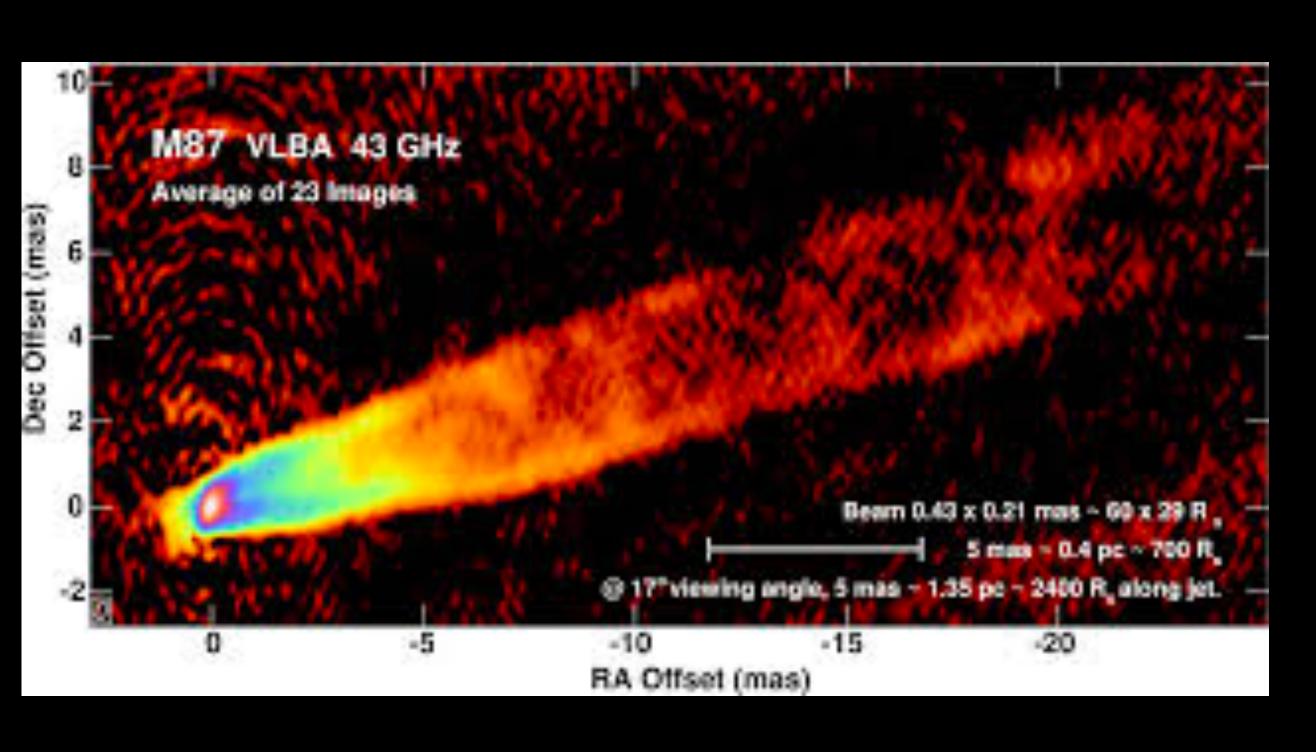
Two primary geometric models used





M87 consistency with other observations

Inferred spin vector consistent with jet axis and ionised gas rotation



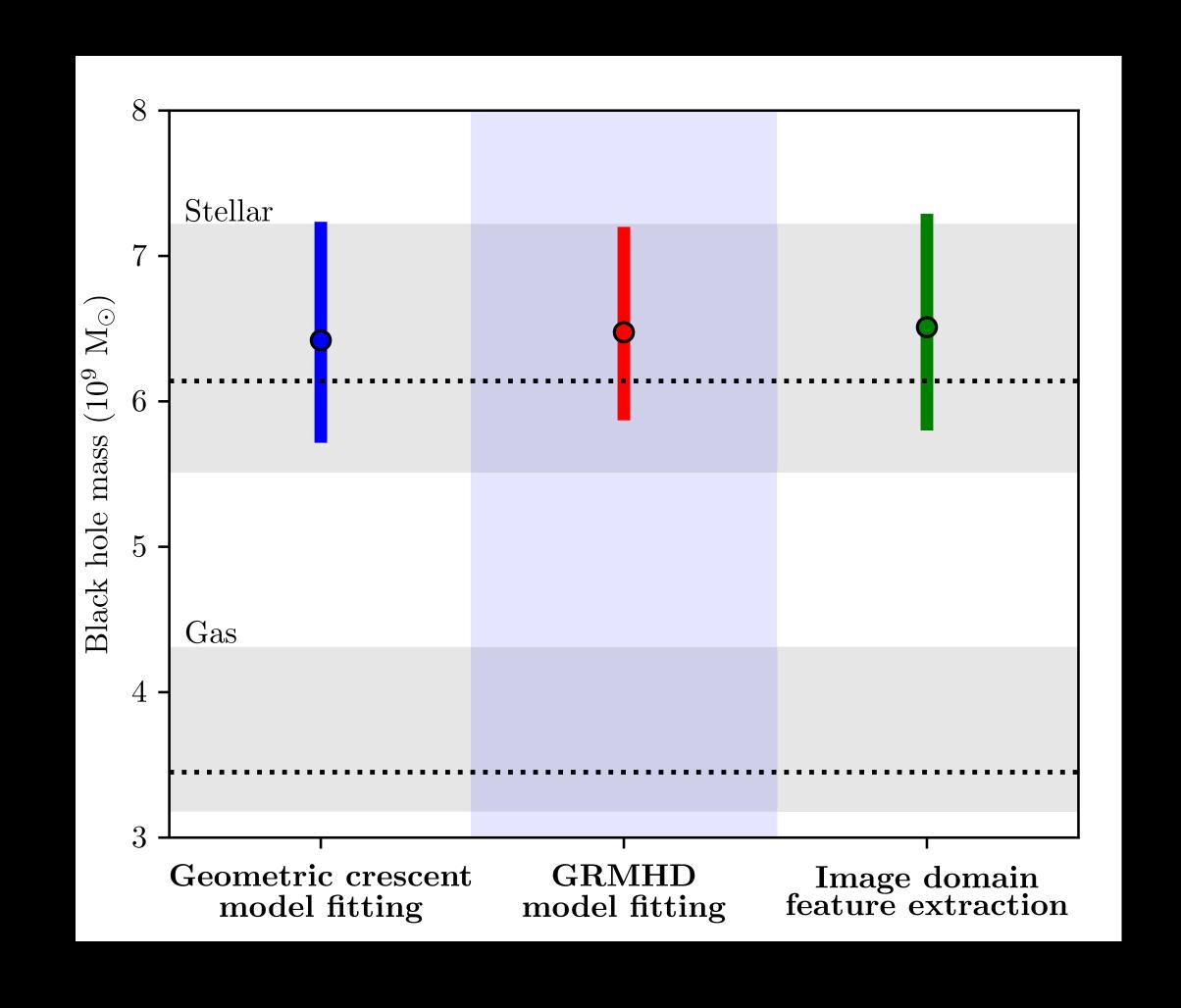
 $a_* > 0, i > 90^{\circ}$ $a_* < 0, i > 90^{\circ}$ approaching jet accretion flow $a_* < 0, i < 90^{\circ}$ $a_* > 0, i < 90^{\circ}$

Walker et al. (2008)



Consistency with black hole mass estimates

- Factor of ~2 discrepancy between M87 black hole mass estimates based on stellar velocity dispersion and gas dynamics
- Black hole shadow consistent with the larger mass suggested by stellar kinematics
- Assumes 230 GHz emission is not well beyond the shadow radius
- All three methods of black hole mass estimation with EHT data (images and visibility data) are consistent



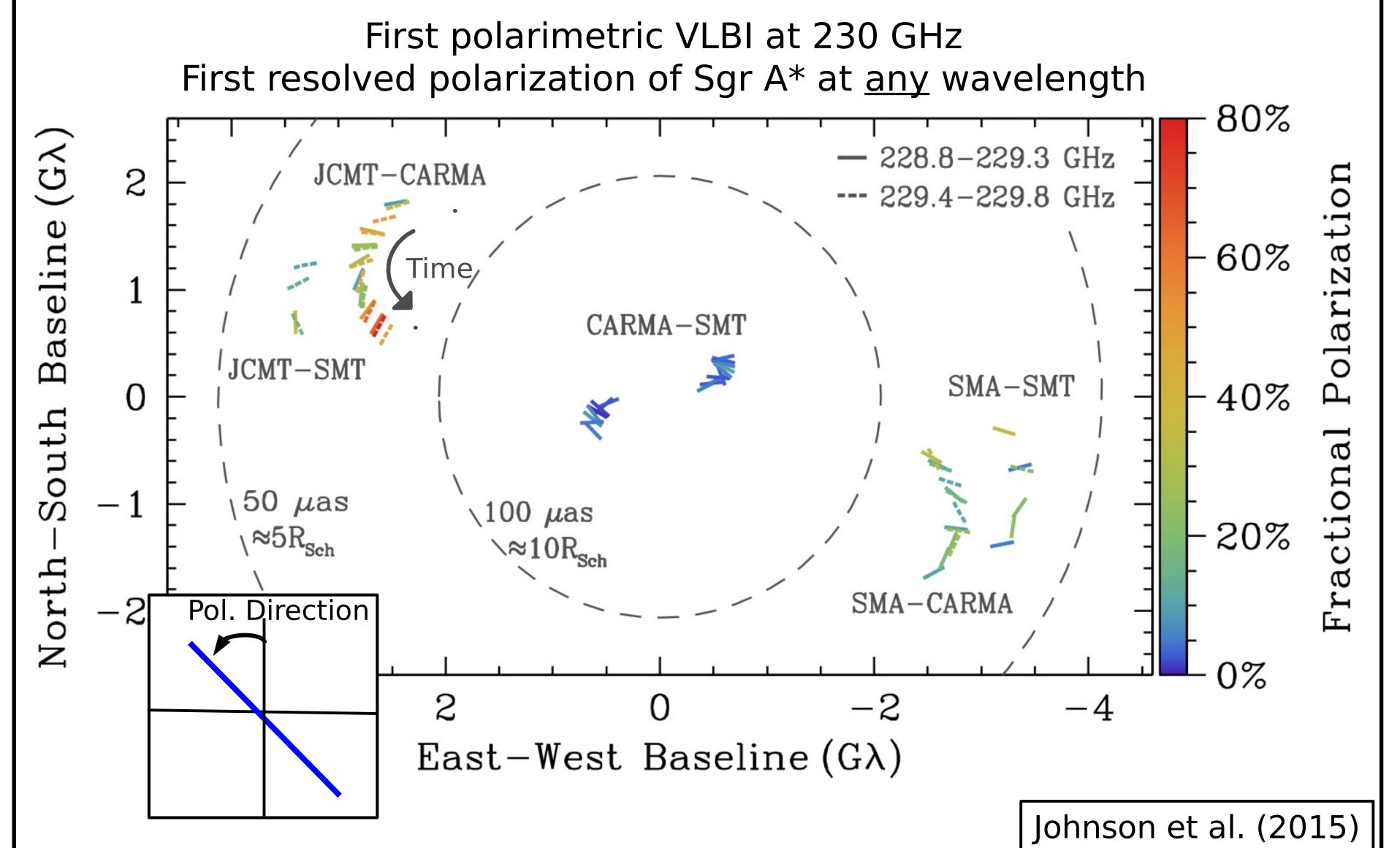


What's next for the EHT?

- Repeat experiment to test if ring stable on ~1 year timescales
- Make an image movie of Sgr A*
- Polarimetry to probe magnetic field structure
- "non-Horizon" science targets (e.g. 3C279 recently published)
- Telescope array expansion

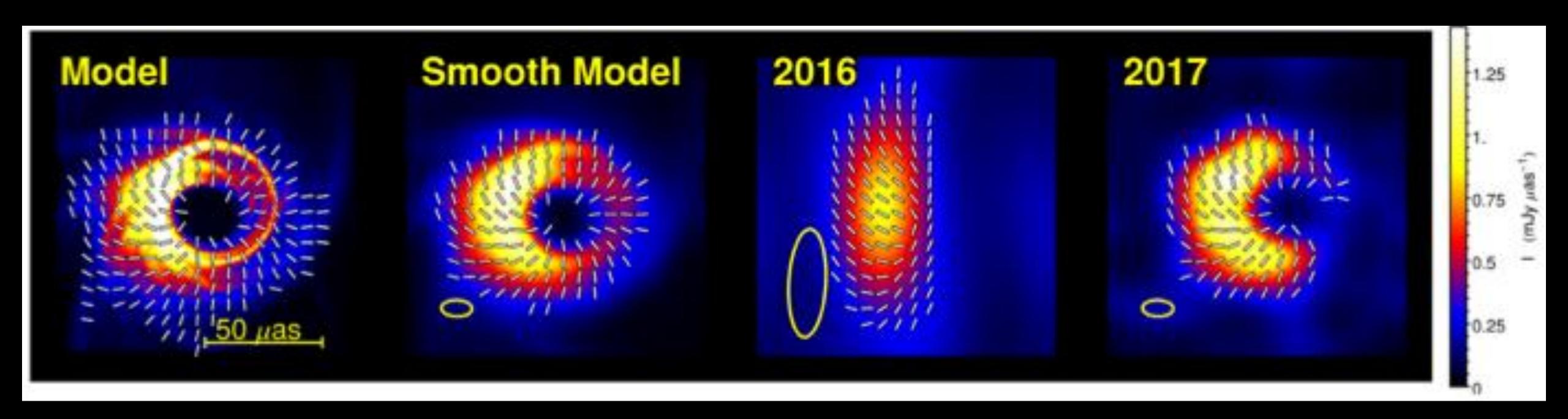


Resolving Sgr A* with the EHT





Imaging polarimetry of M87 (in progress)

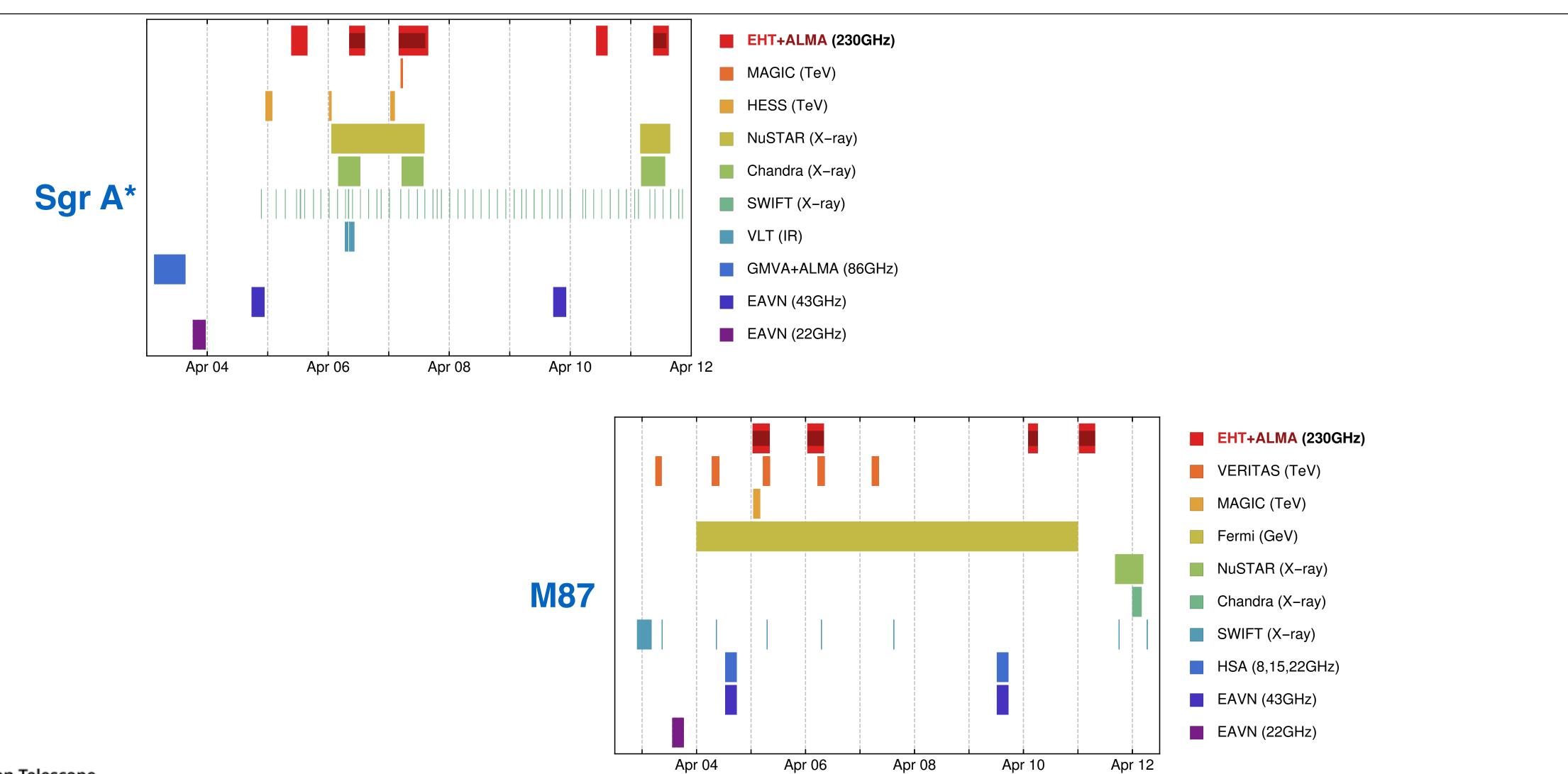


GRMHD simulations: Jason Dexter

Imaging: Chael+2016



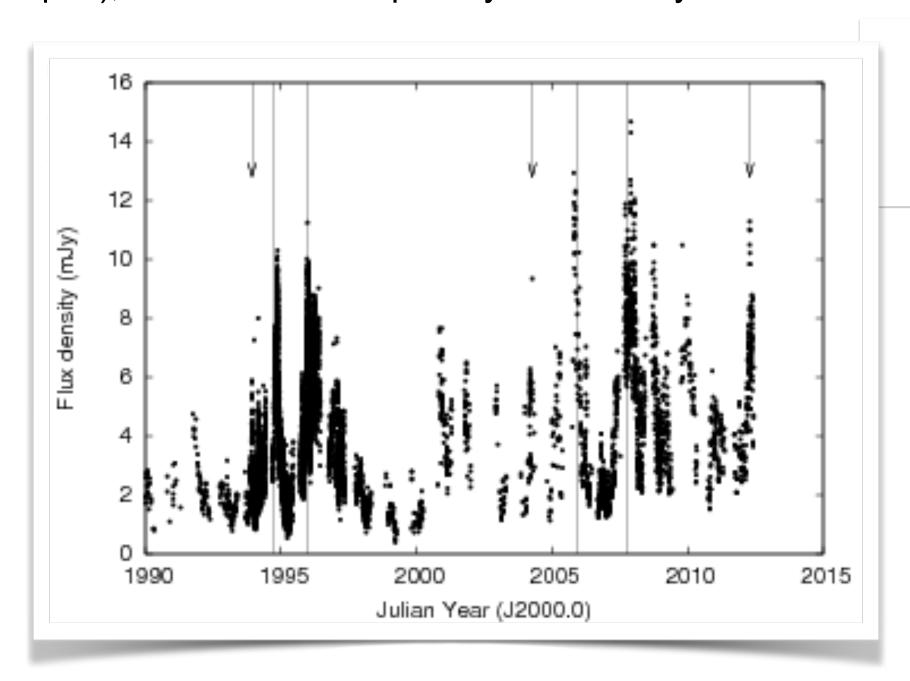
Multi-wavelength coverage (April, 2017)



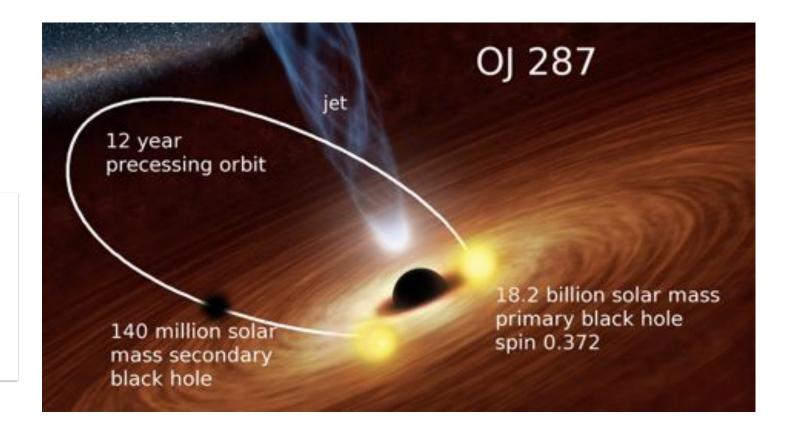
OJ 287 - binary SMBH candidate

OJ287

OJ287 is one of the best candidates for hosting a supermassive binary black hole system (i.e., Valtonen+2016) in an eccentric (ϵ =0.7) orbit with a major axis of 0.1 pc (~26 µas), which could be spatially resolved by the EHT.



PI: J.L. Gomez



Science goals:

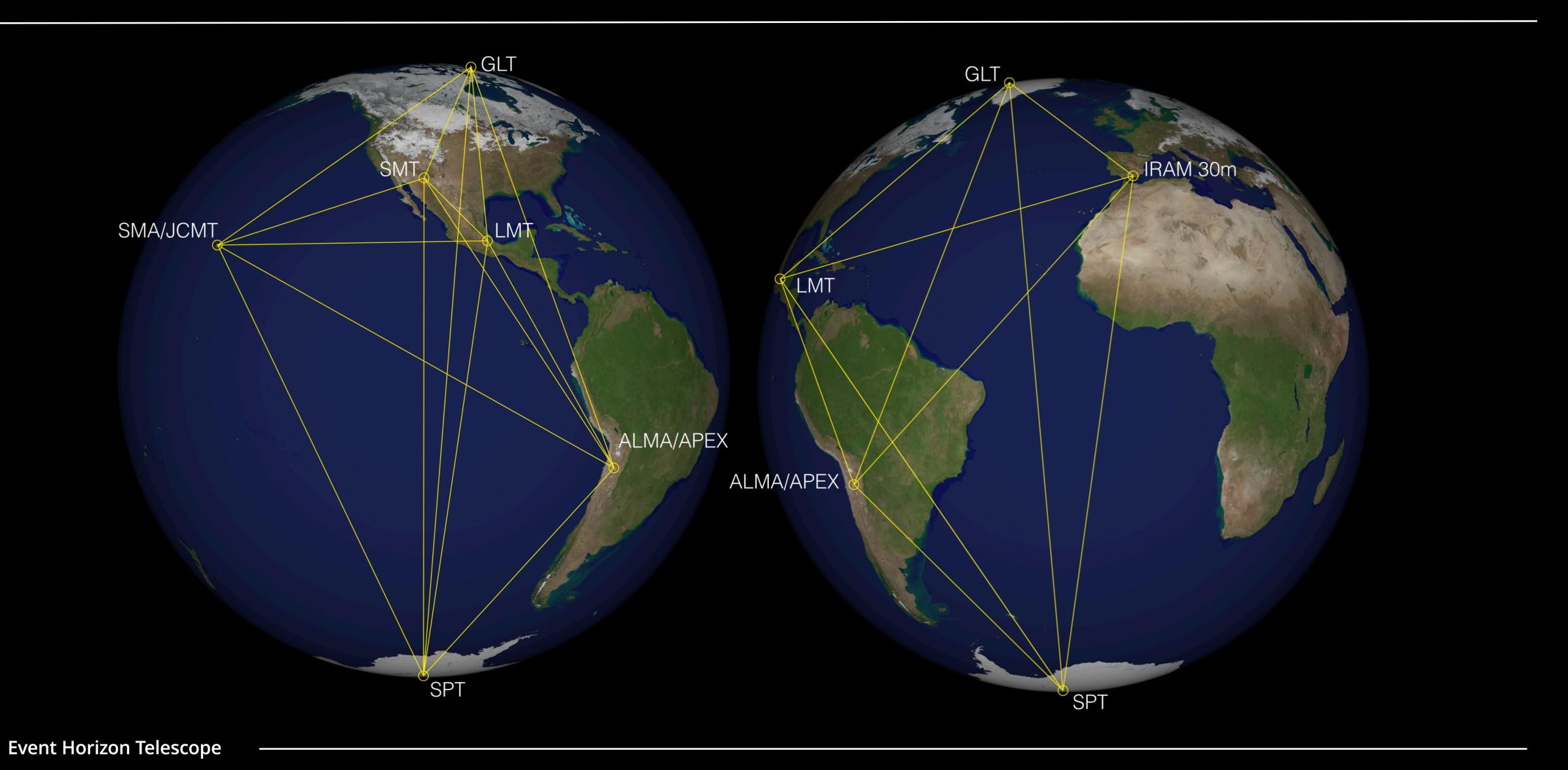
- Confirmation of binary BH system
- Determine magnetic field 3D structure to test jet formation models
- Study collimation profile
- Probe accretion flow

14

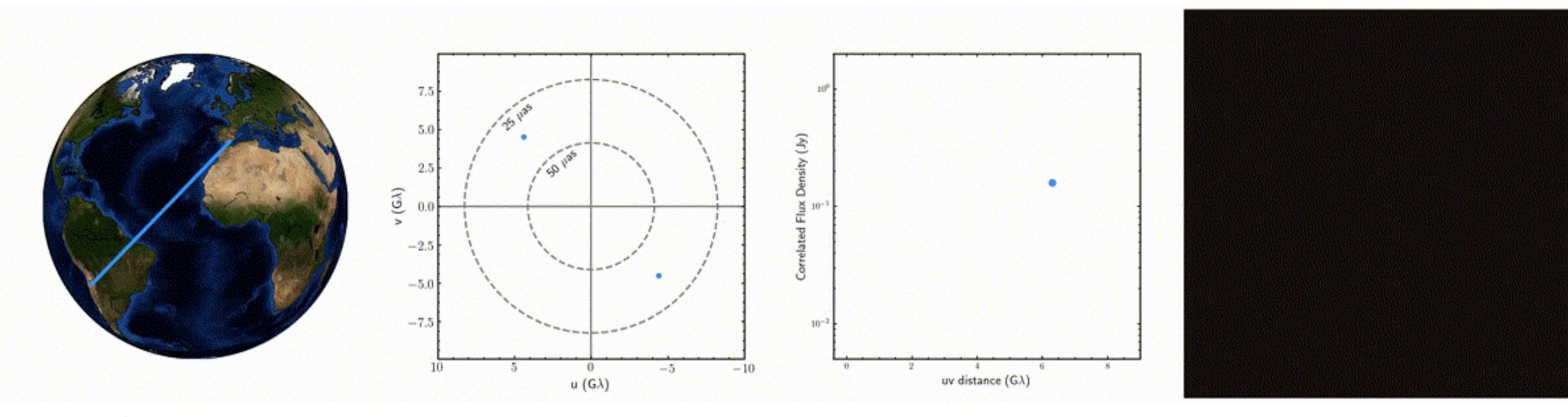


EHT array expansion

The case for expanding the EHT into Africa

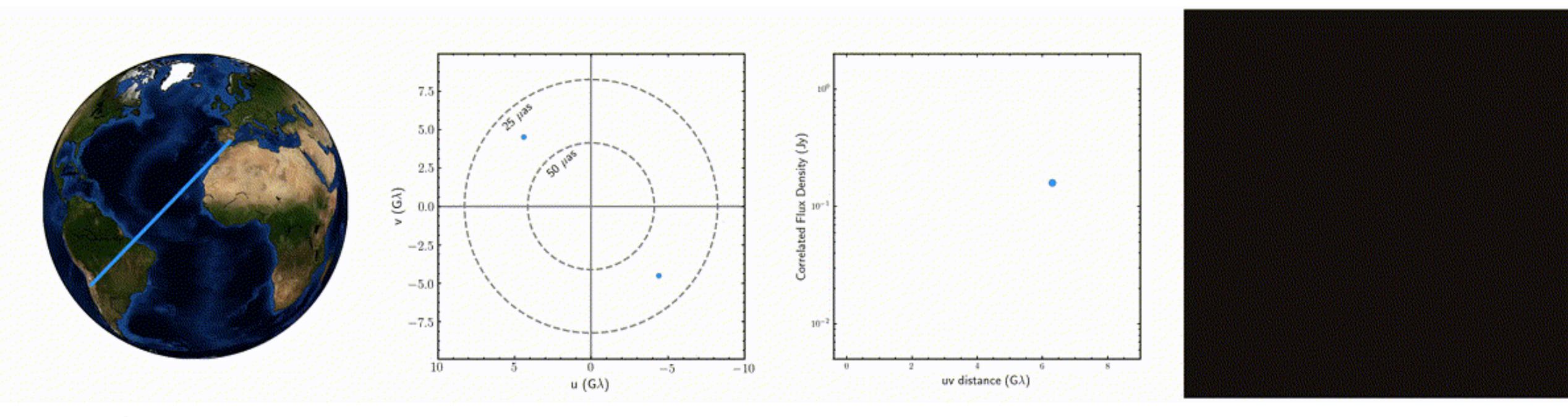


Lo-band eht-imaging on April 11: slowly building up data



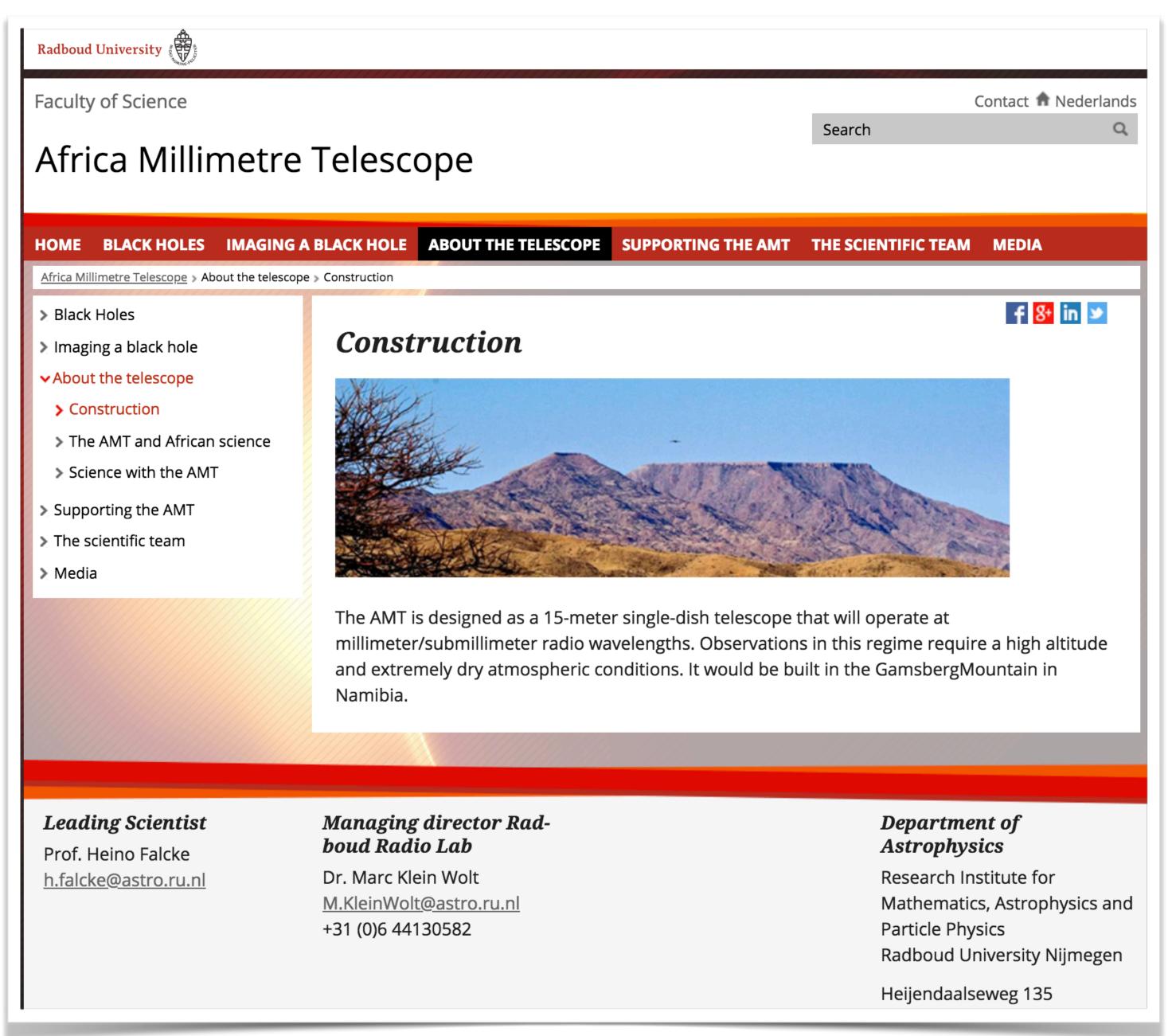


Lo-band eht-imaging on April 11: slowly building up data





Co-led by Radboud and University of Namibia

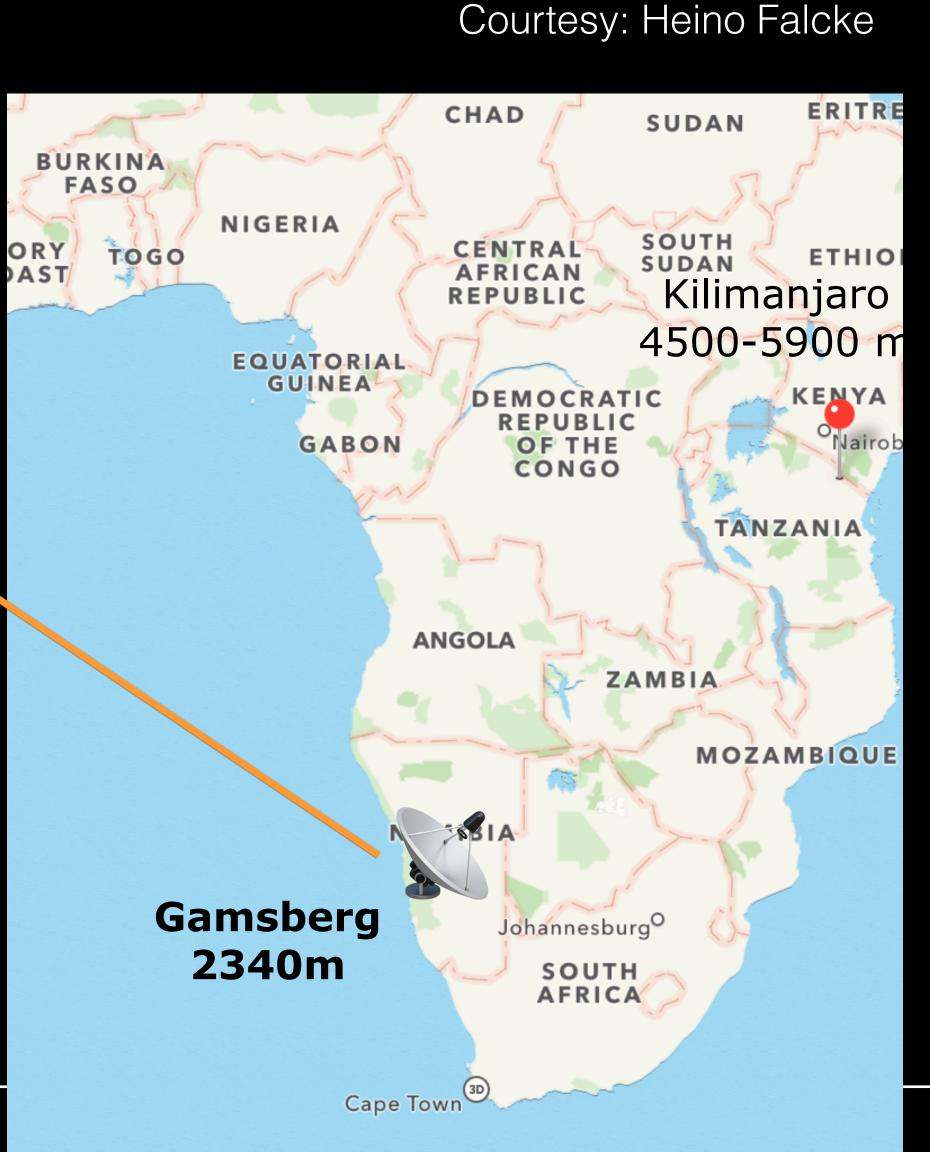


Prospective site for the African Millimetre Telescope



Proposed site for AMT project

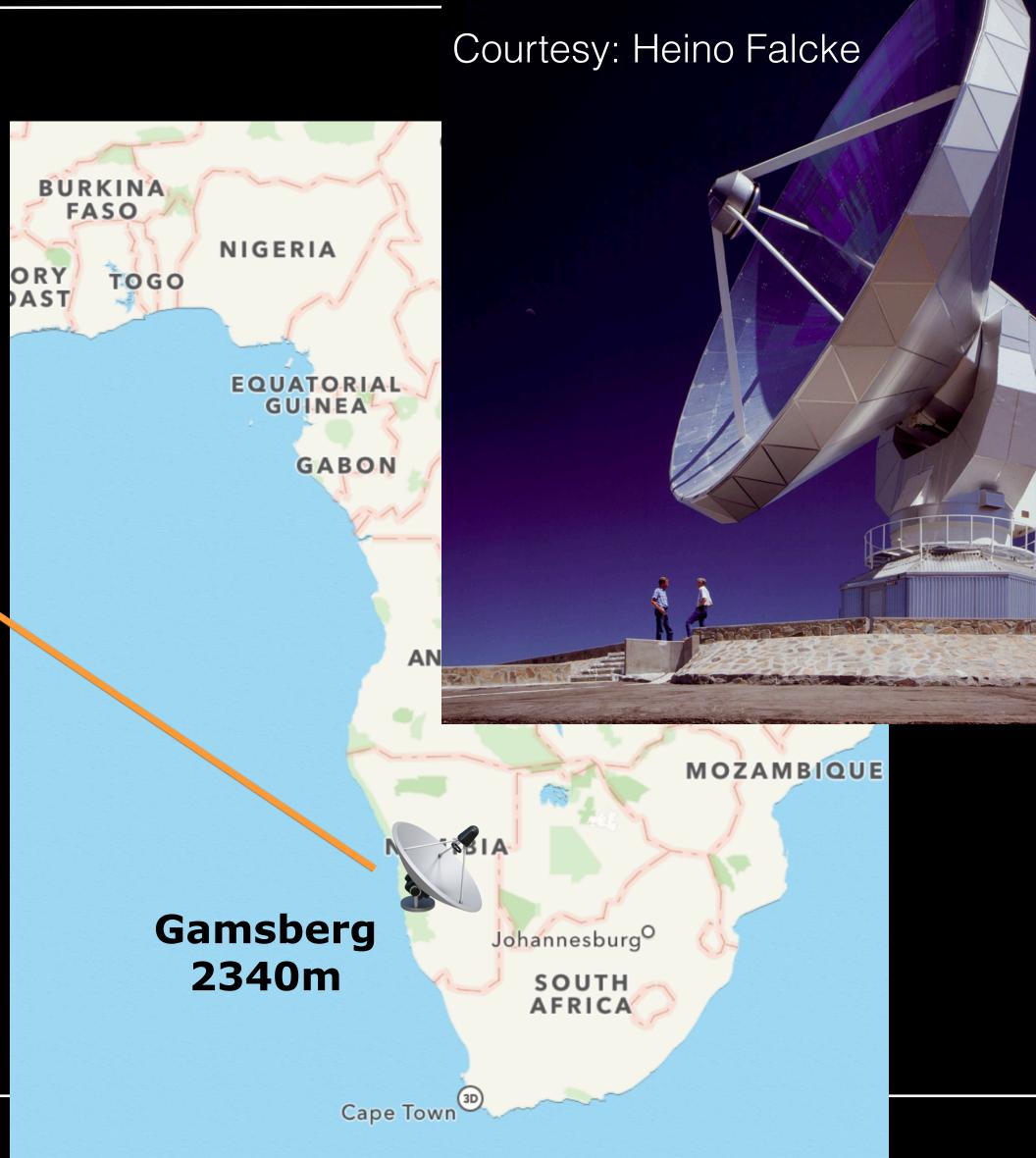




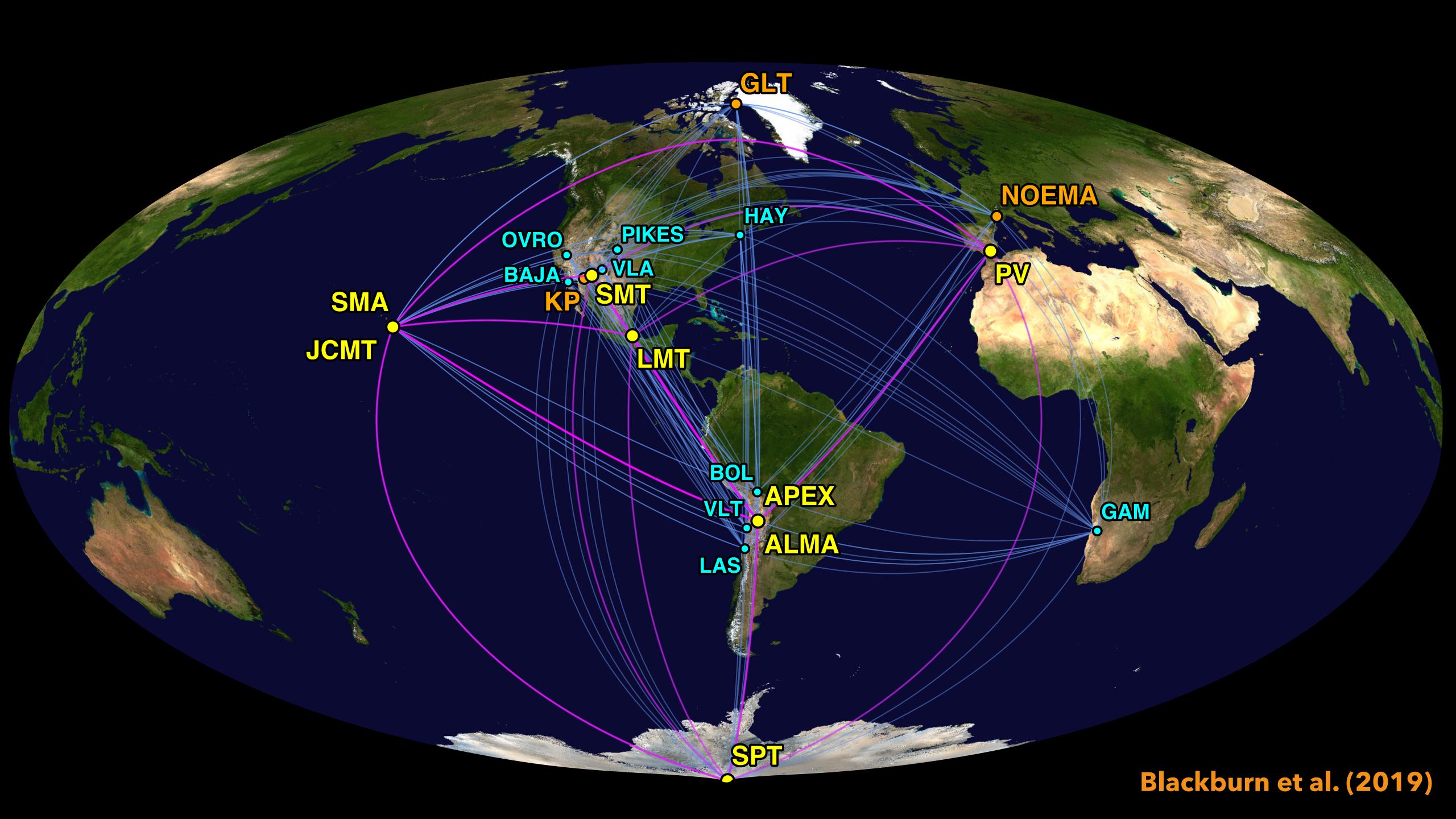


Proposed site for AMT pro



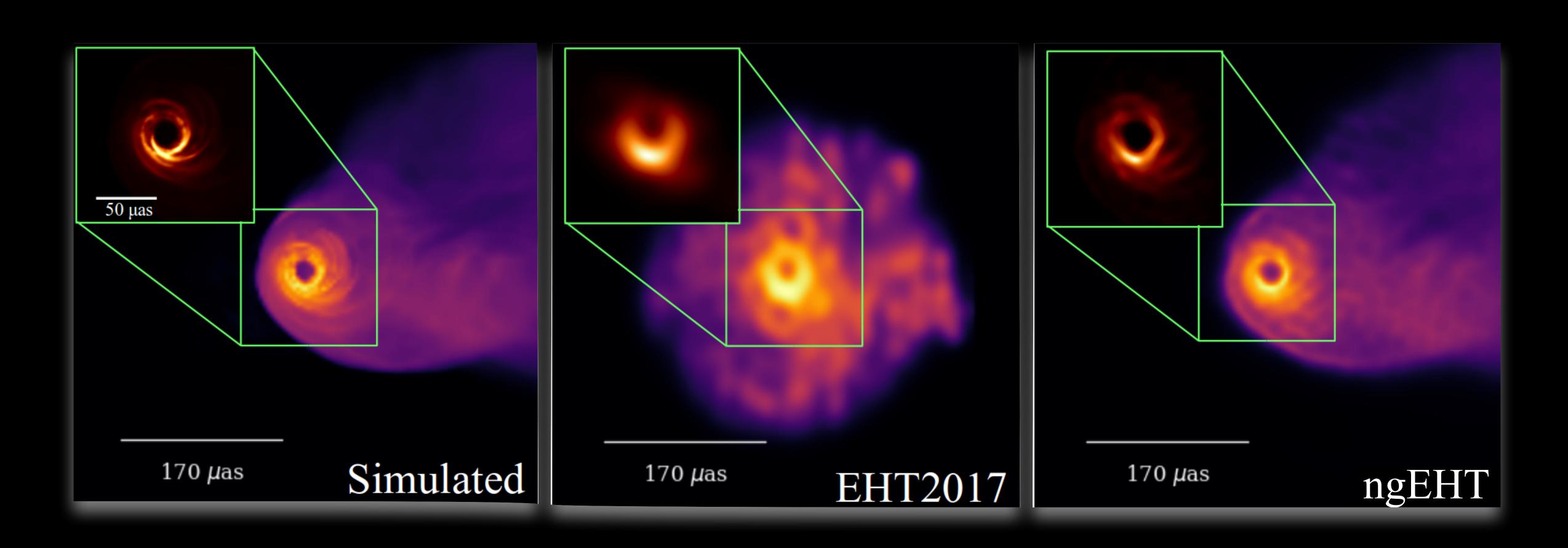






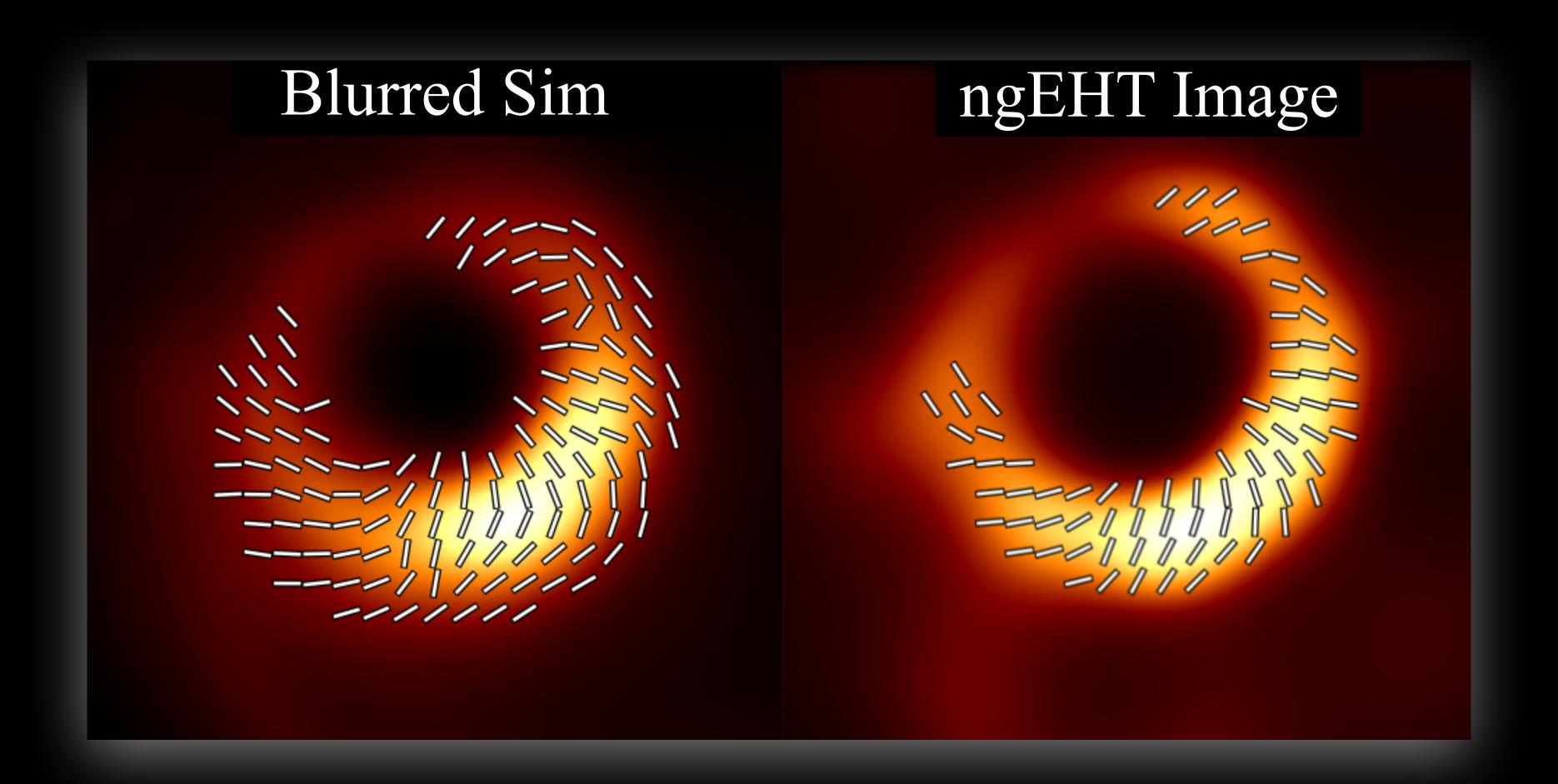
Next-generation EHT

(as SKA-VLBI is to SKA, ngEHT is to ngVLA)



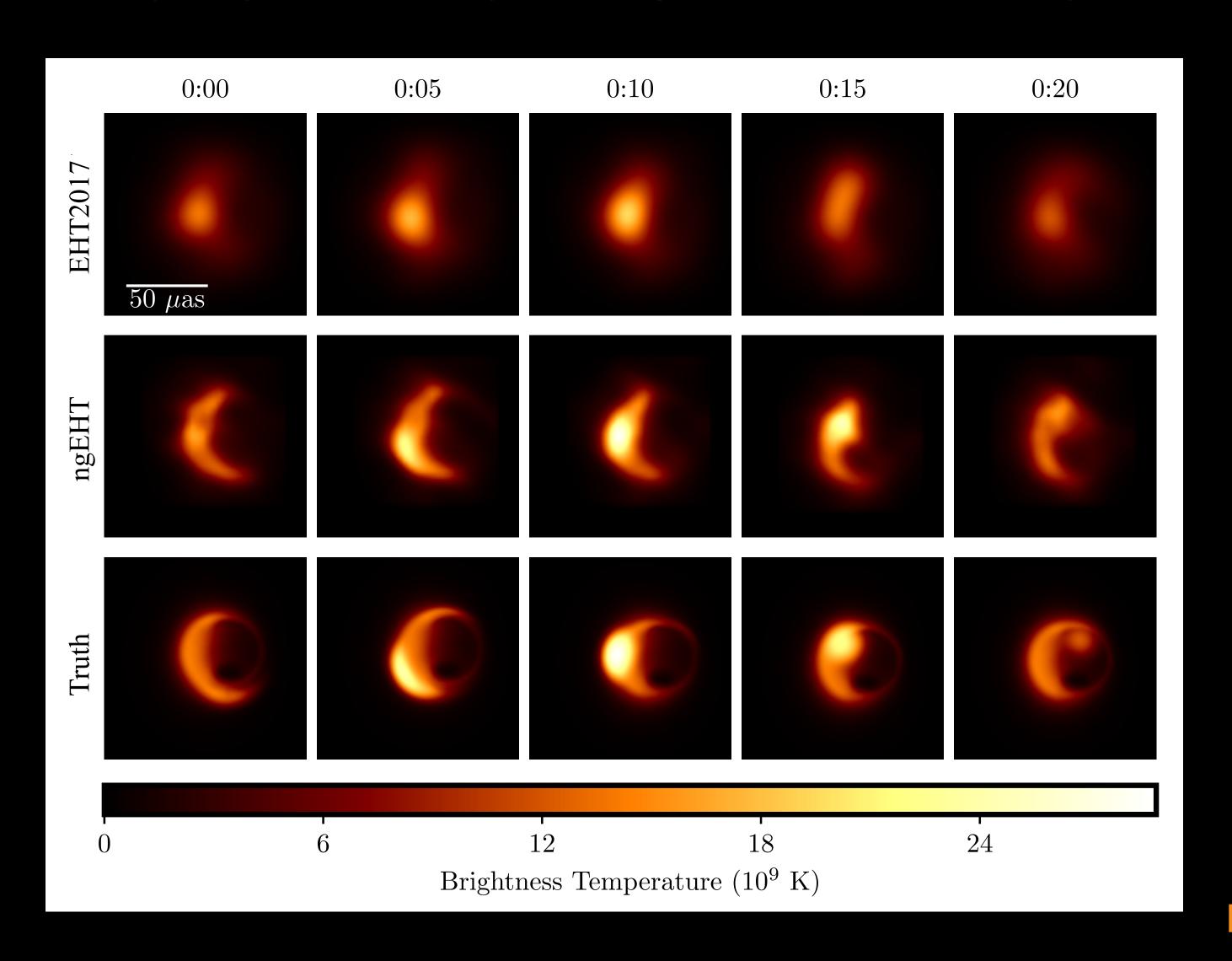
Next-generation EHT

Polarimetric imaging capability



Next-generation EHT

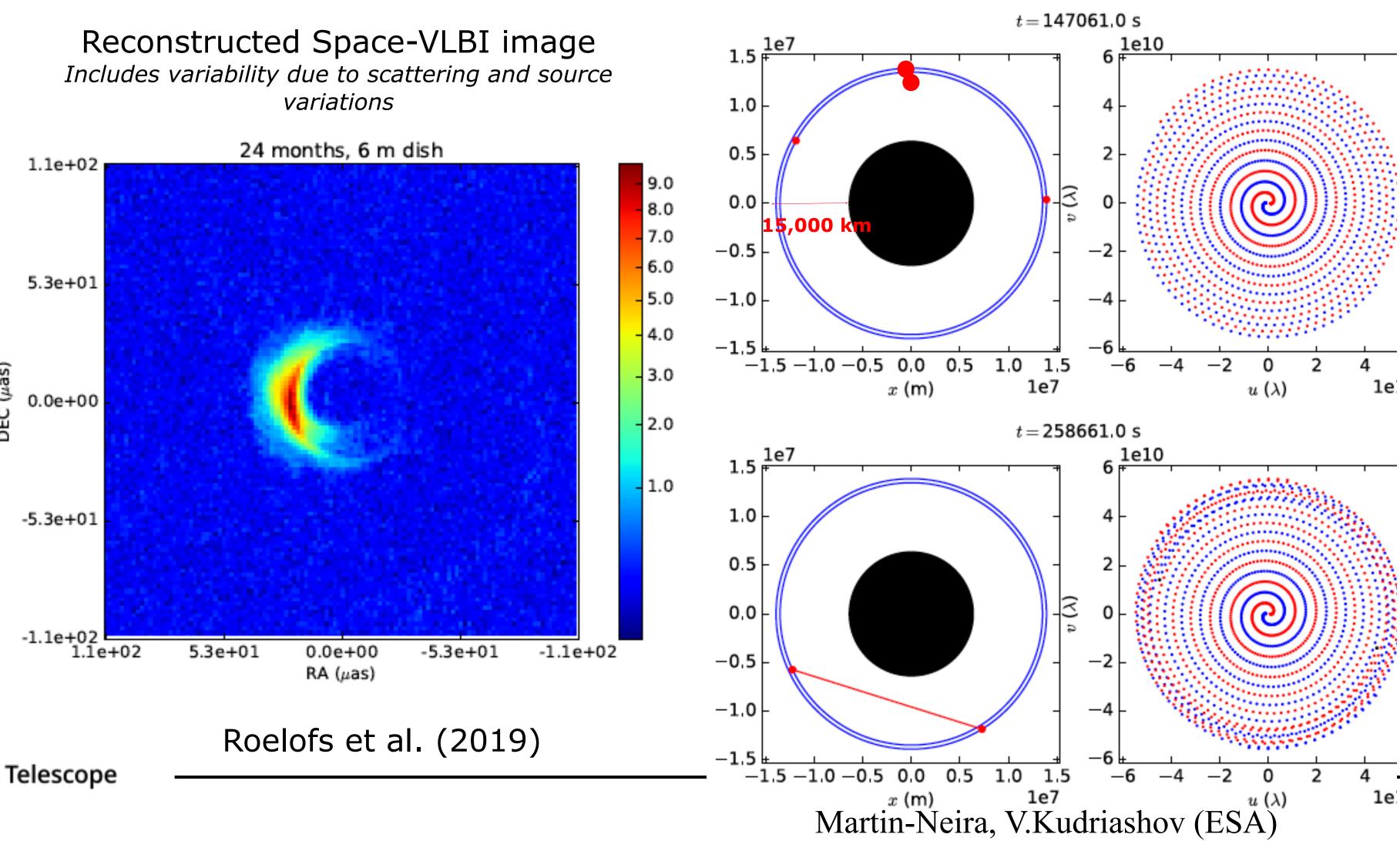
Imaging orbiting hotspots around Sgr A*



To space!

ESA-Radboud study: Event Horizon Imager (EHI)



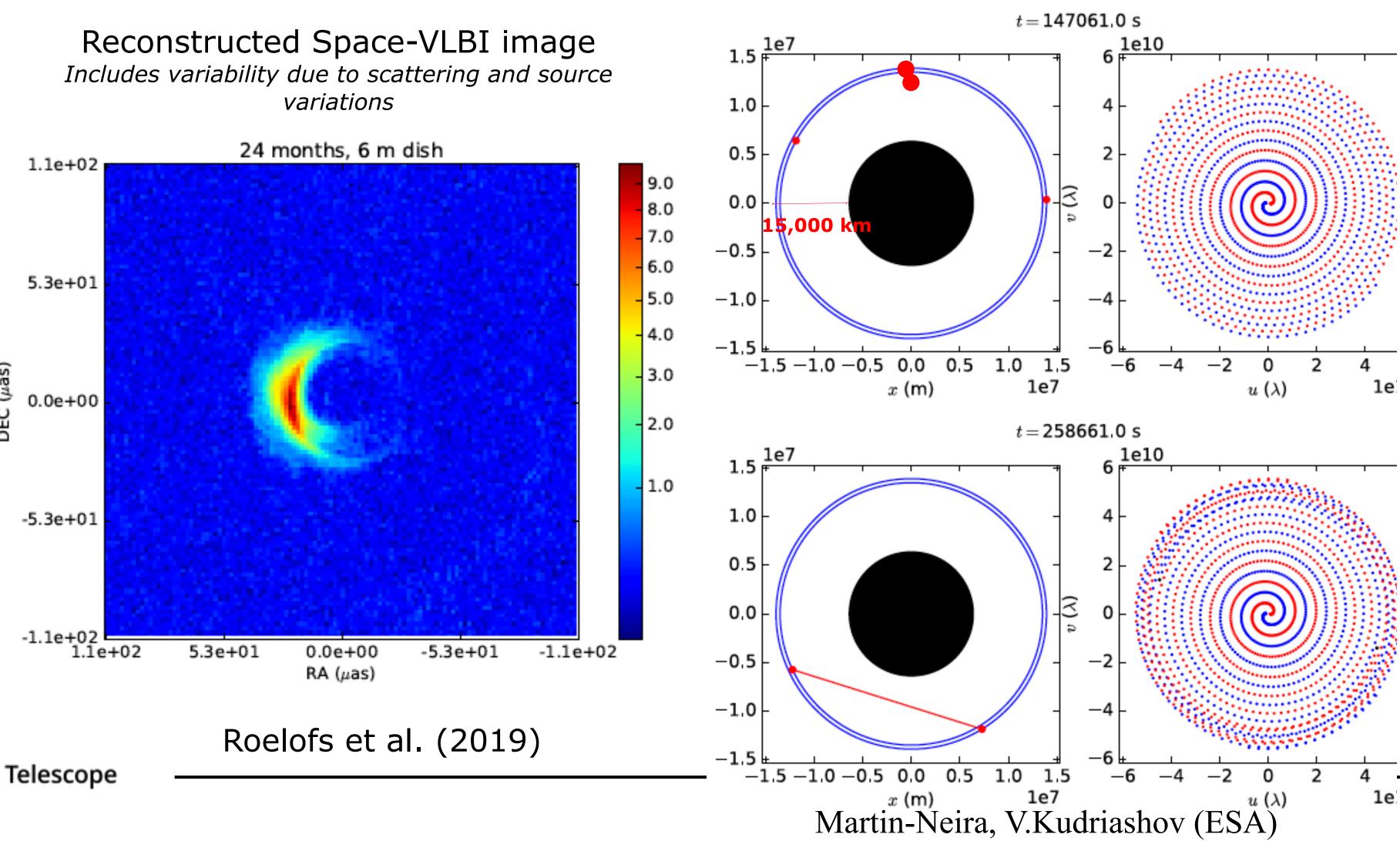




To space!

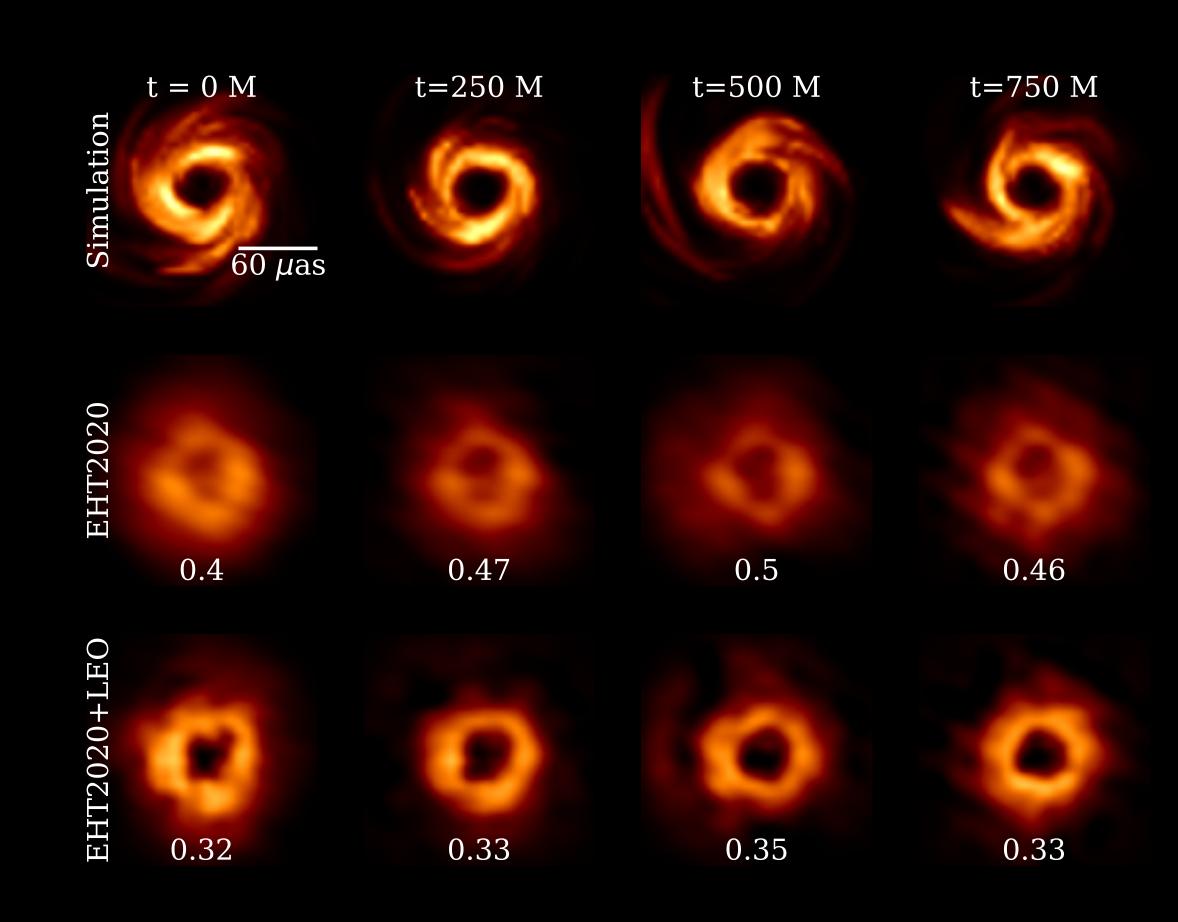
ESA-Radboud study: Event Horizon Imager (EHI)





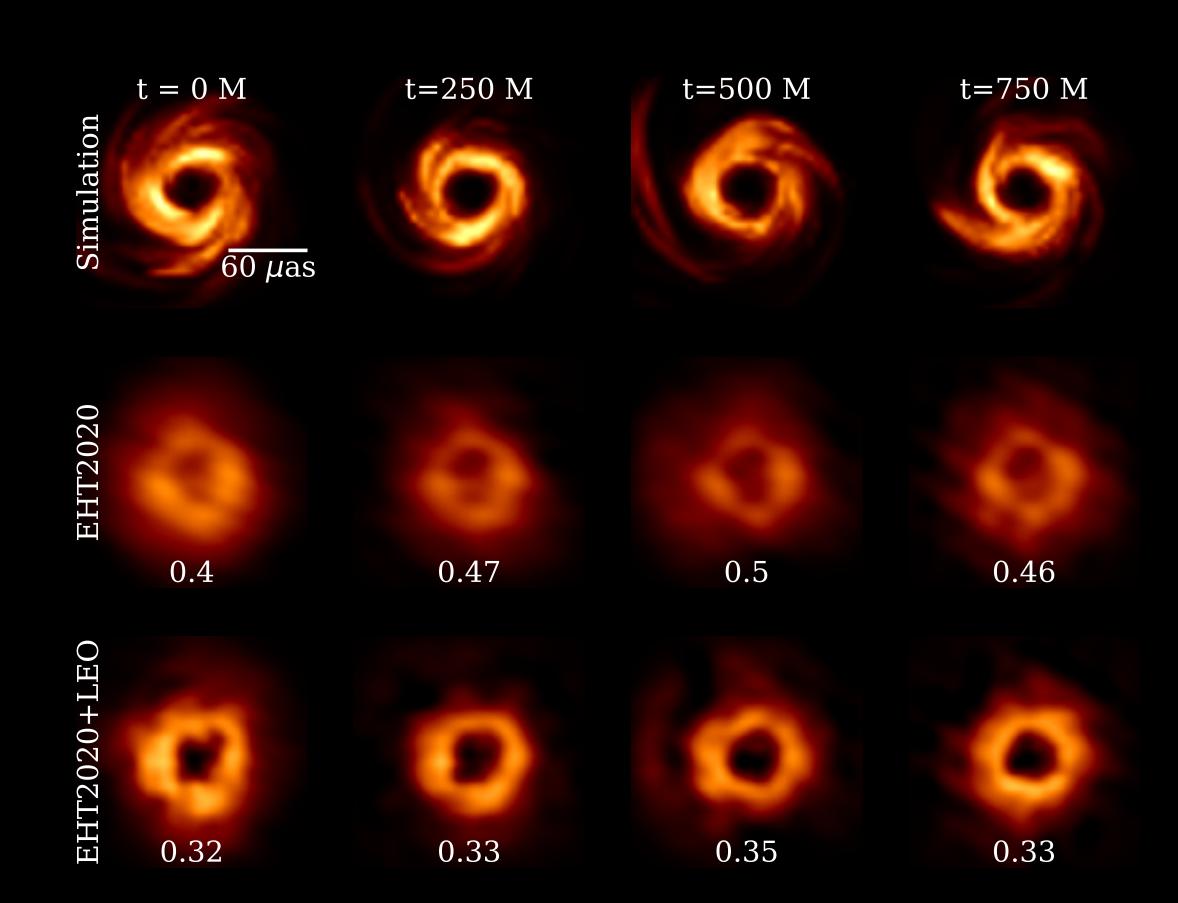


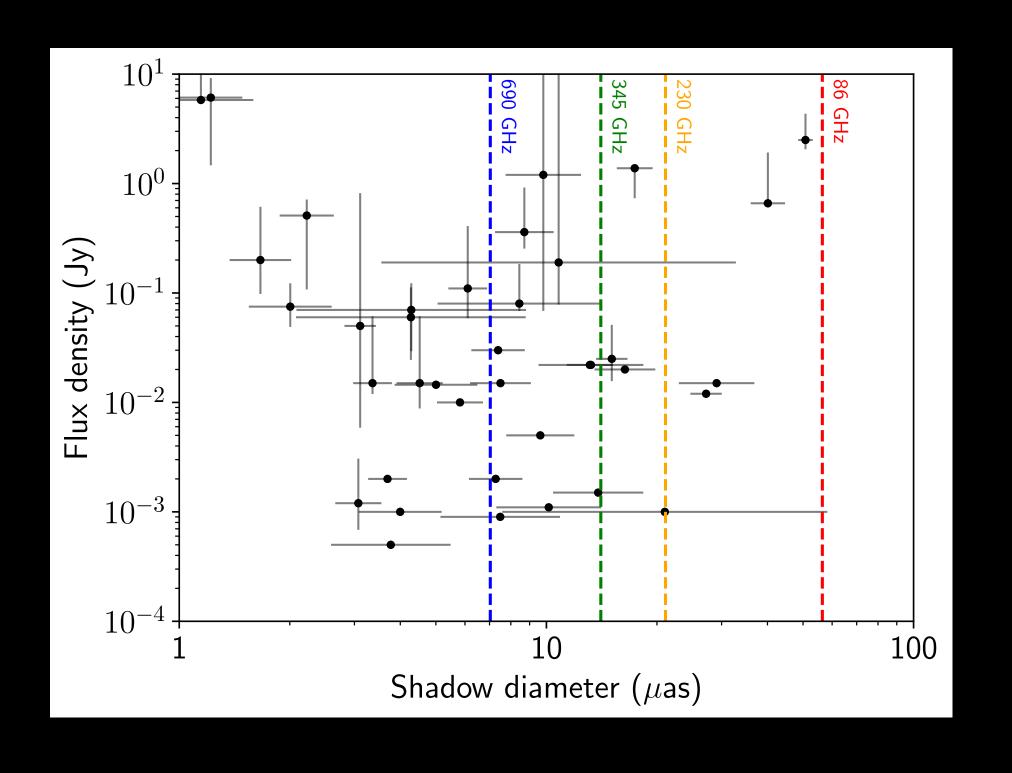
The enormous potential of mm-space VLBI





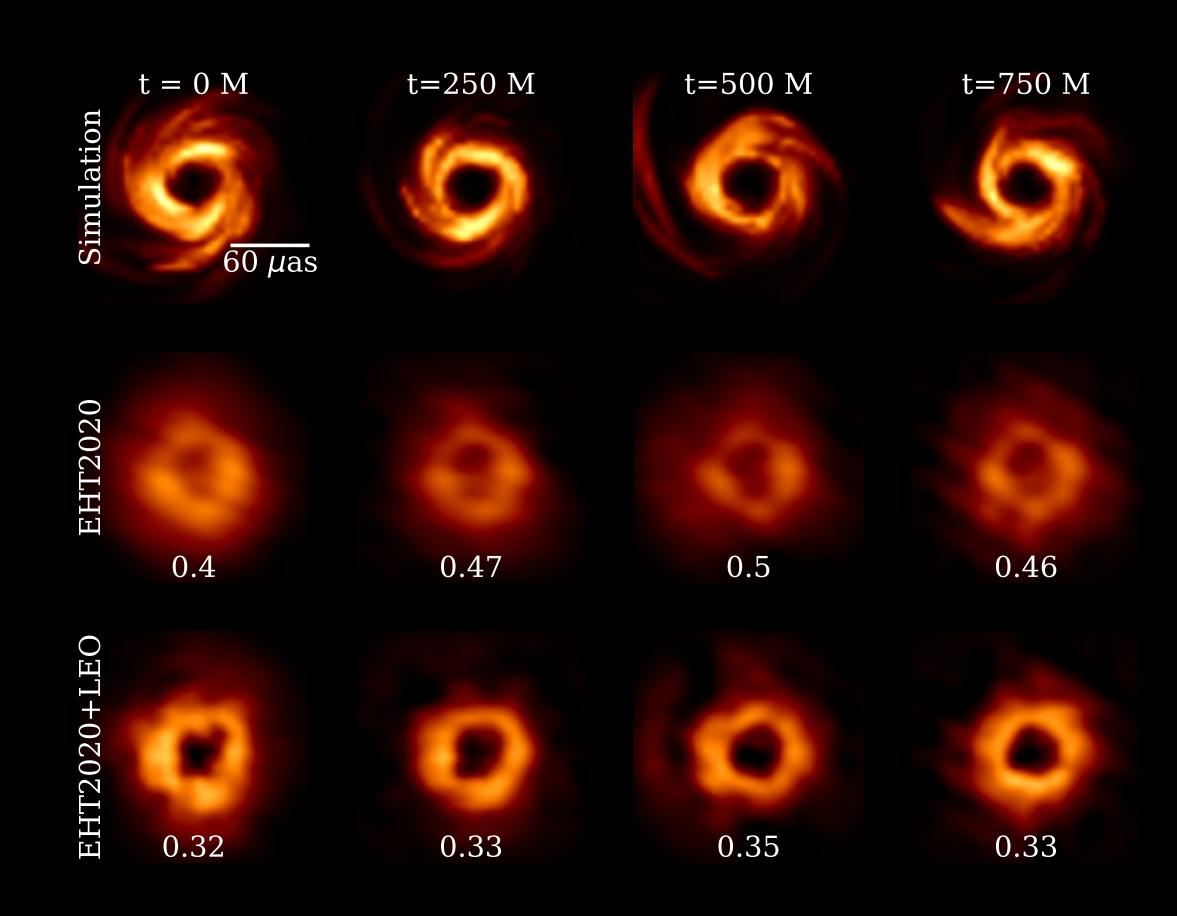
The enormous potential of mm-space VLBI

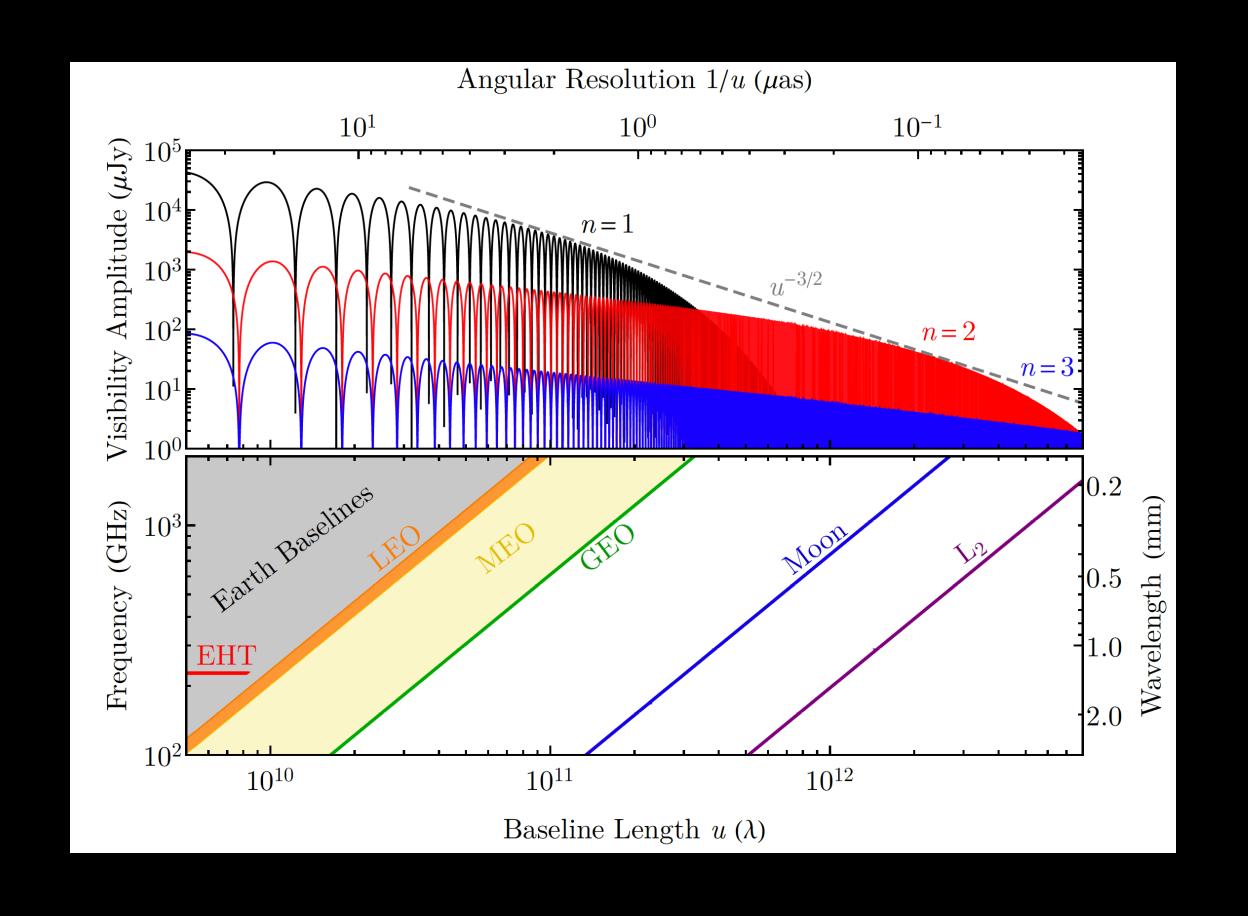




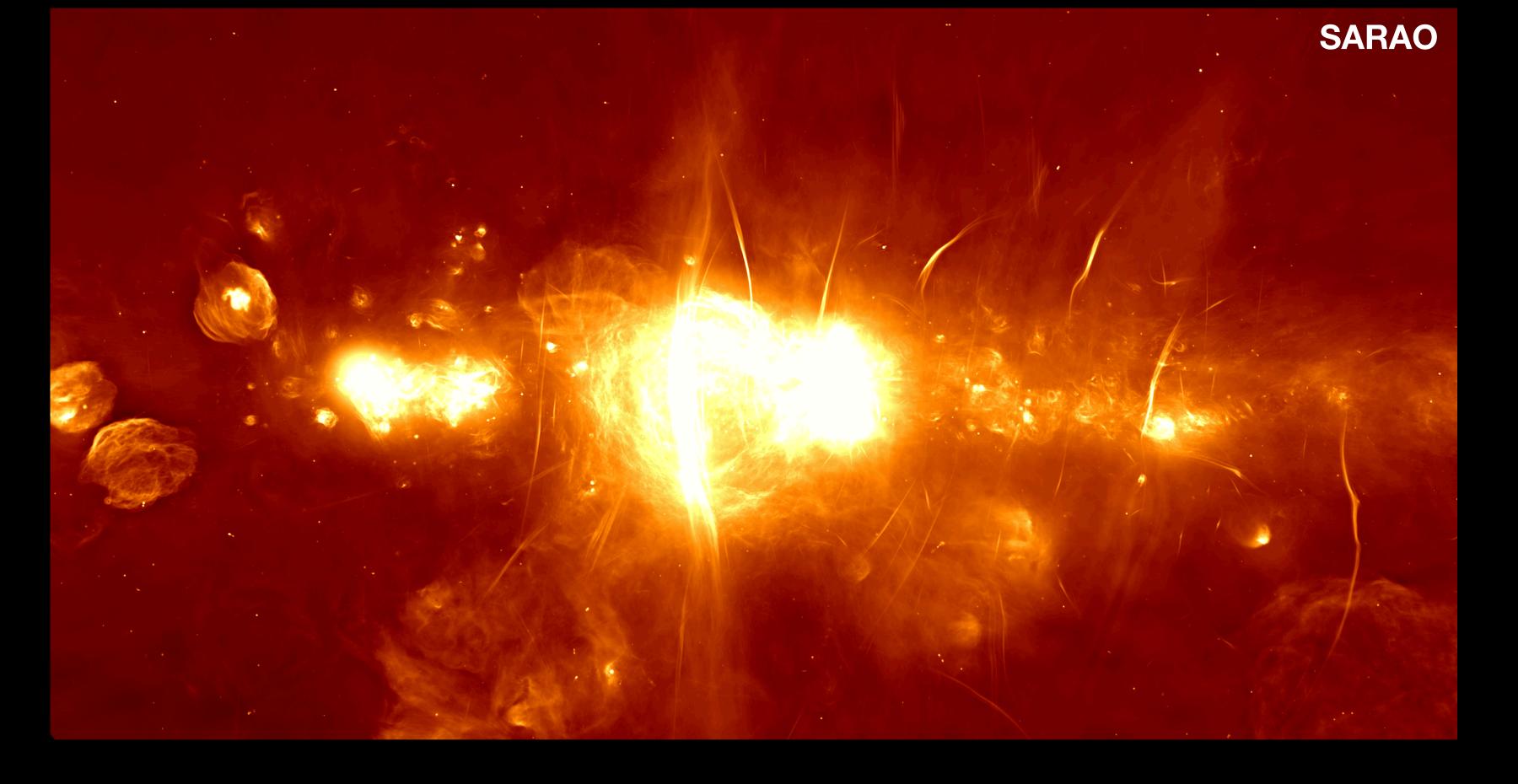


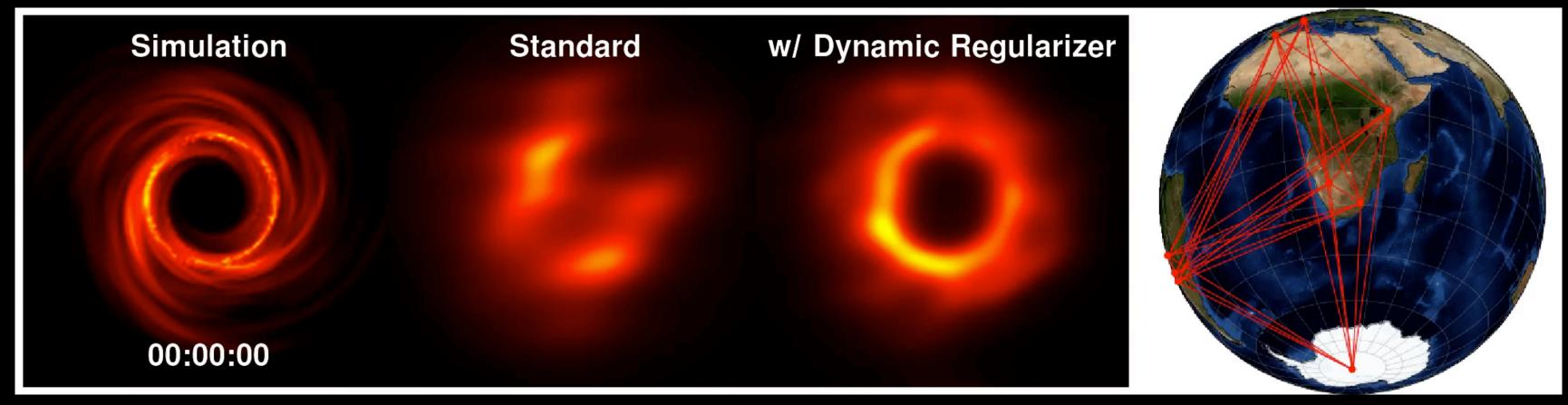
The enormous potential of mm-space VLBI



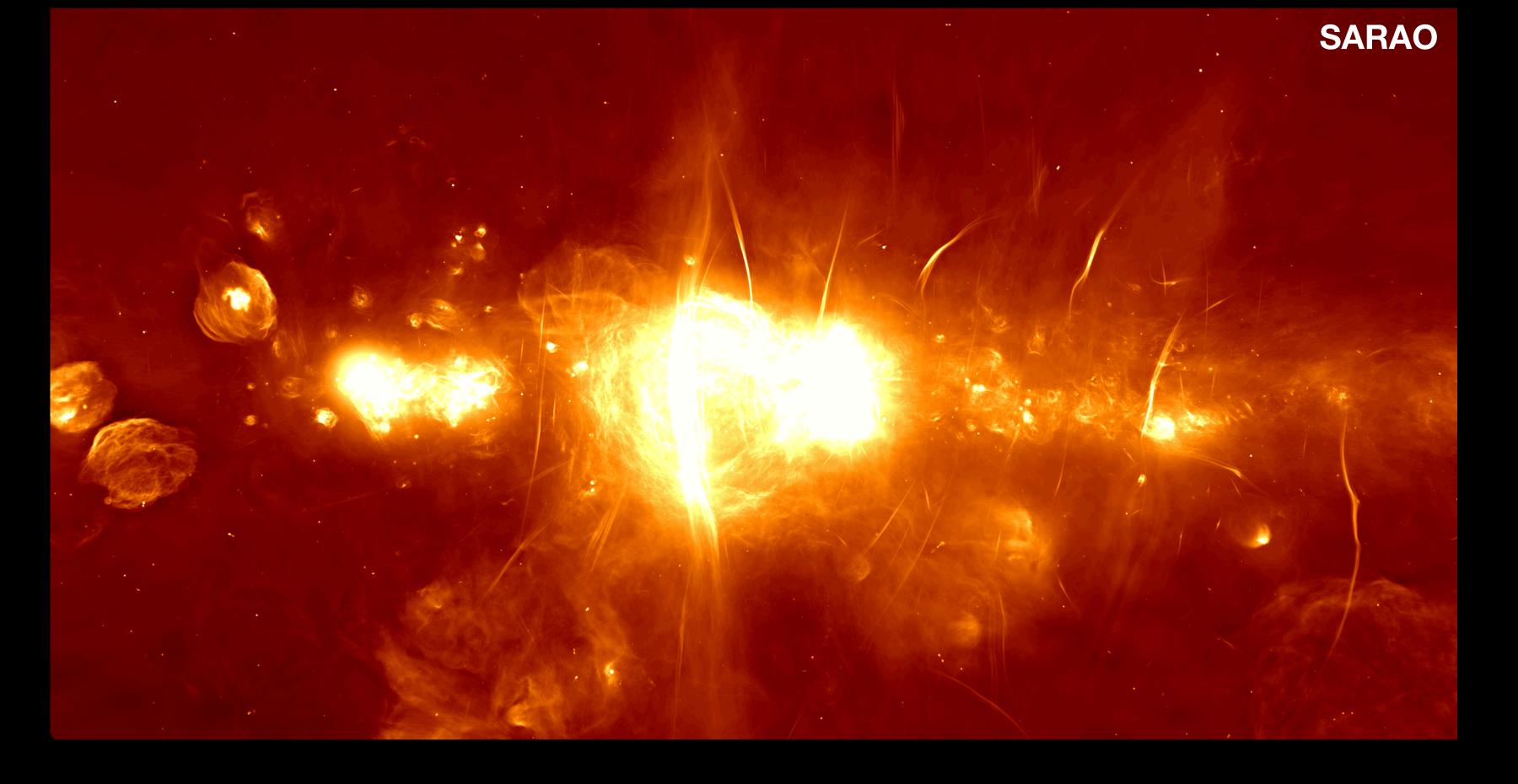


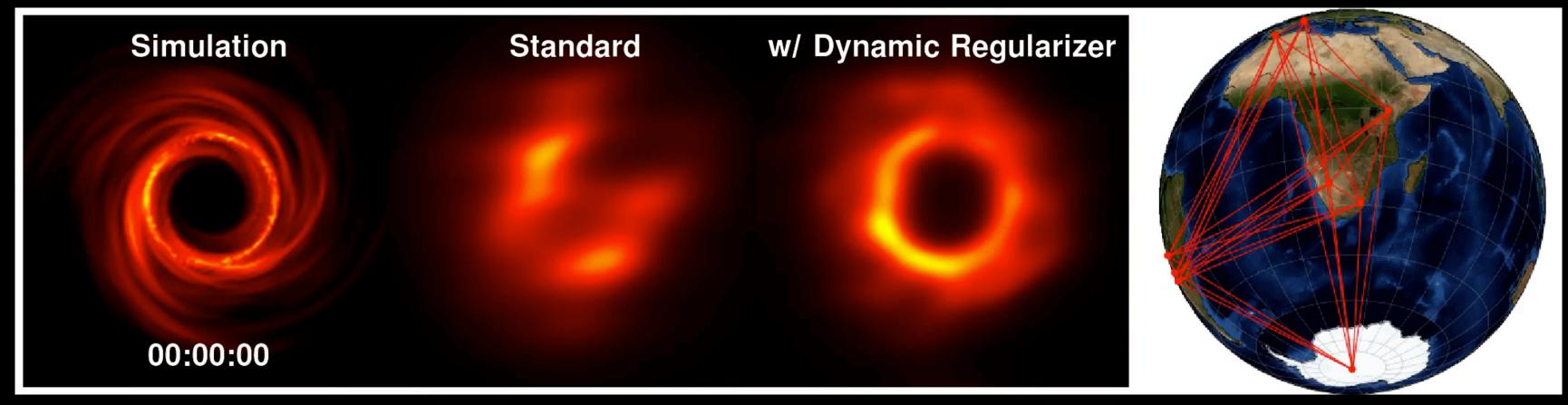




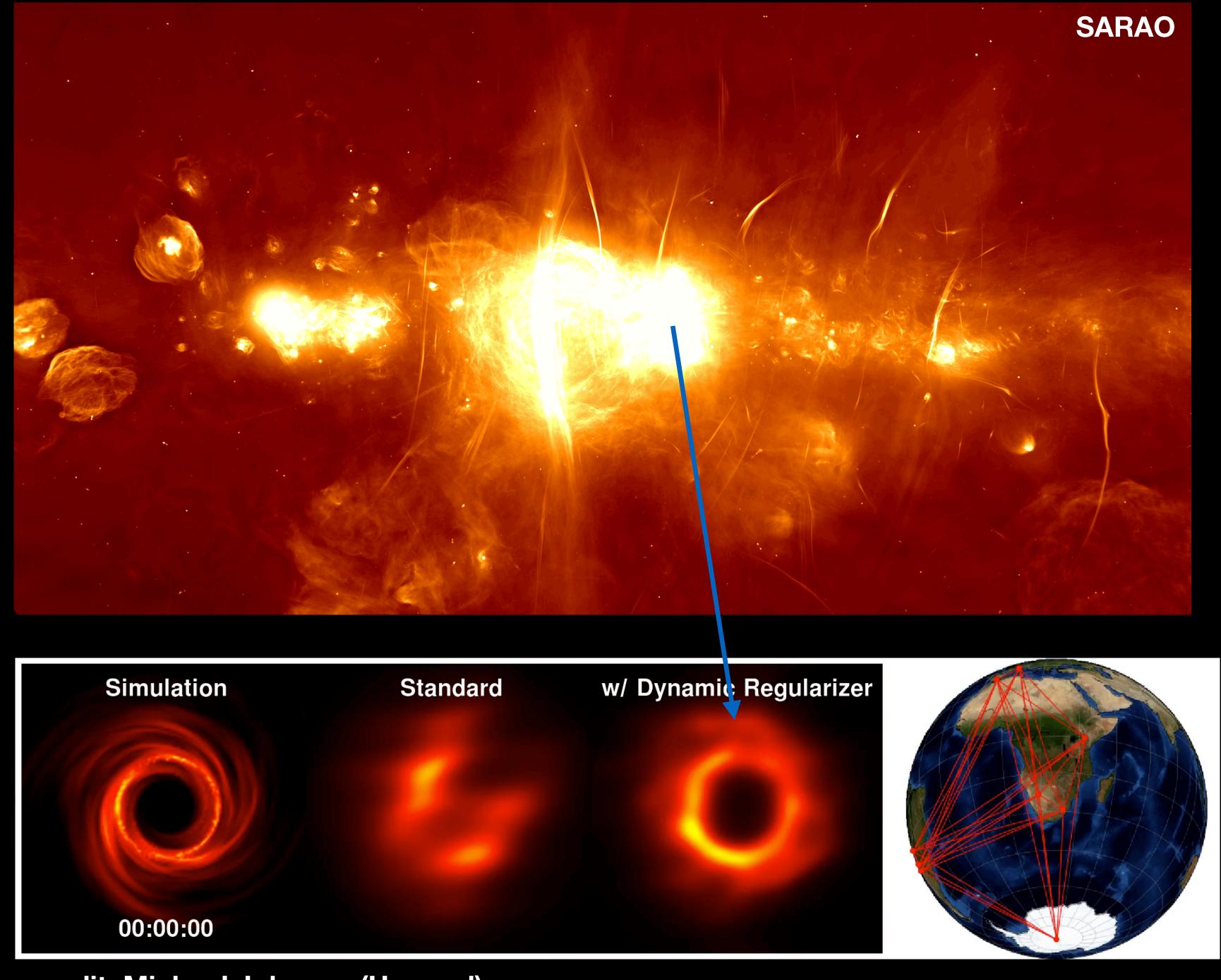


credit: Michael Johnson (Harvard)





credit: Michael Johnson (Harvard)



credit: Michael Johnson (Harvard)

summary

- The EHT has a strong record of high impact science built on engineering excellence
- Achieved its primary goal: captured the first image of a black hole!
 But much more to come...
- With current imaging quality, black hole mass consistent with stellar kinematics; and the shape is consistent with GR
- Experiment mode for now (shadow imaging), but large range of unique science on many sources possible, especially as array expands
- The tools and techniques developed with the EHT project will have have had a much broader impact on VLBI
- EHT expansion (including in Africa and space-VLBI) will significantly sharpen tests of gravity

