

Connecting the CGM and galaxy properties with large surveys



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with Brice Ménard and Guangtun Zhu

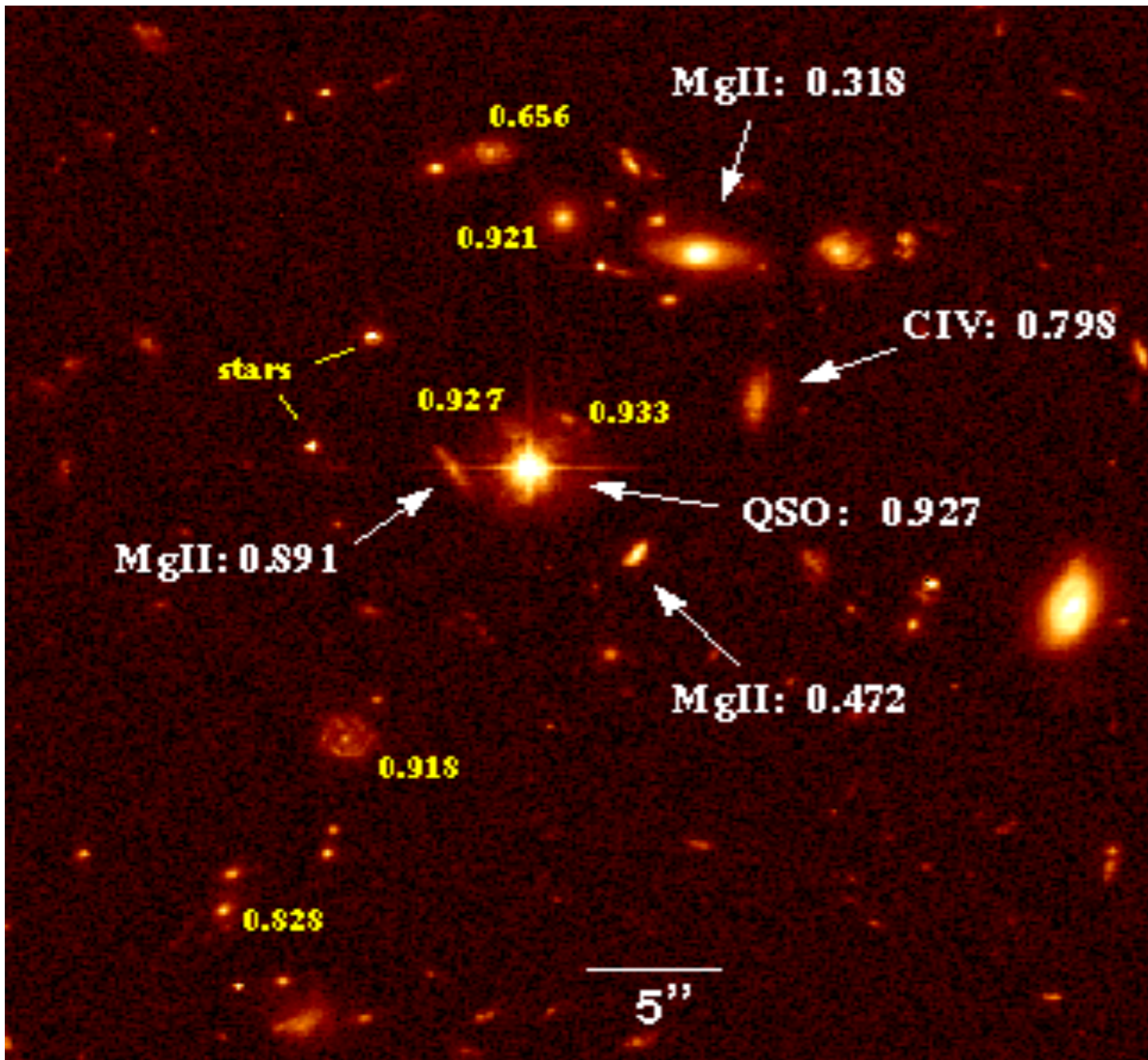


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Searching for galaxy - absorber pairs



1969	0	Bahcall & Spizter
1986	1	Bergeron
1995-2010	50	Steidel et al. Churchill et al. Chen et al.
2013	200	Nielsen et al.

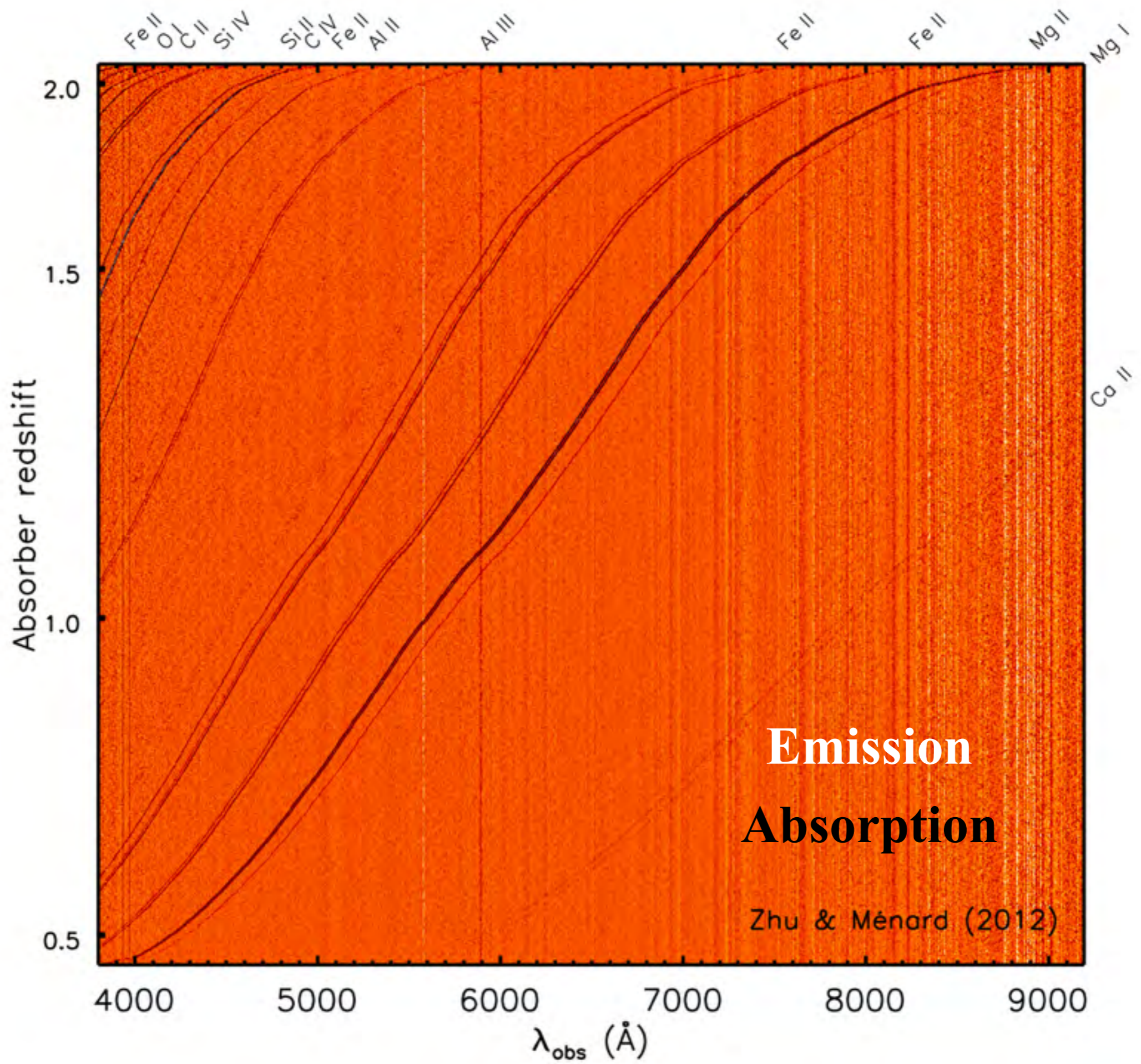
Steidel et al. (1997)

Is there cool gas around both types of galaxies?

**Do the properties of the cool CGM depend
on the galaxy properties?**



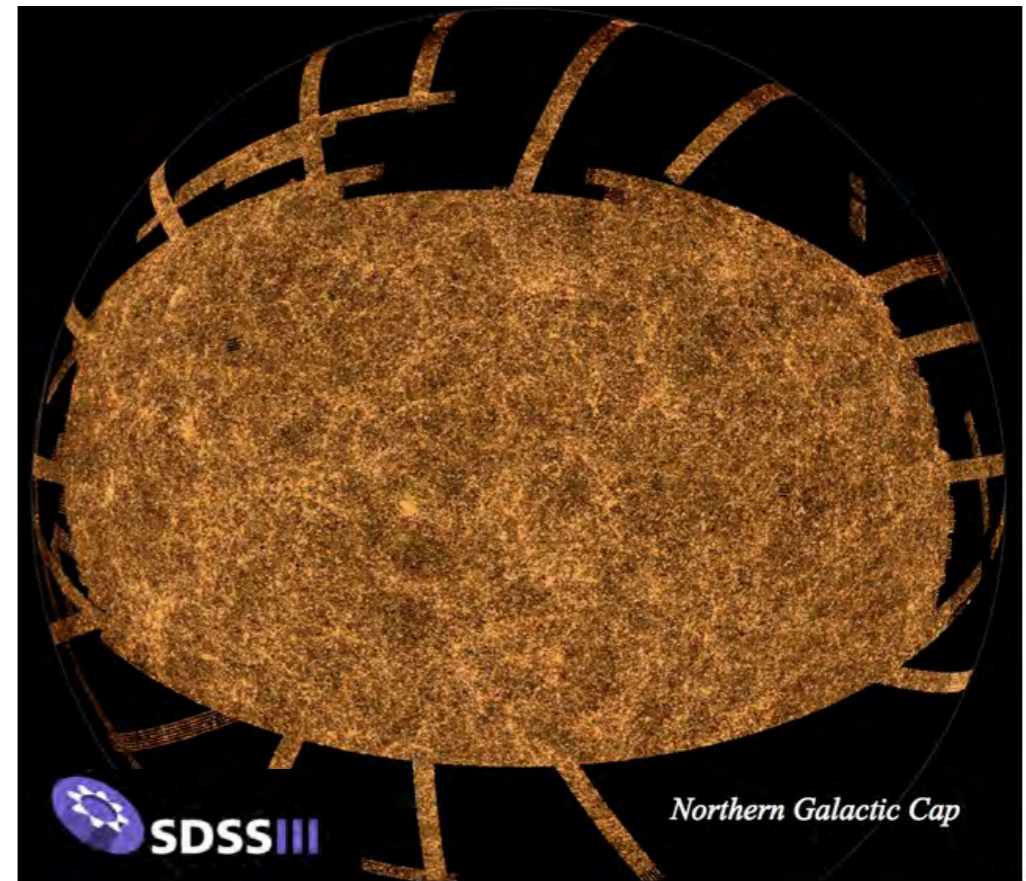
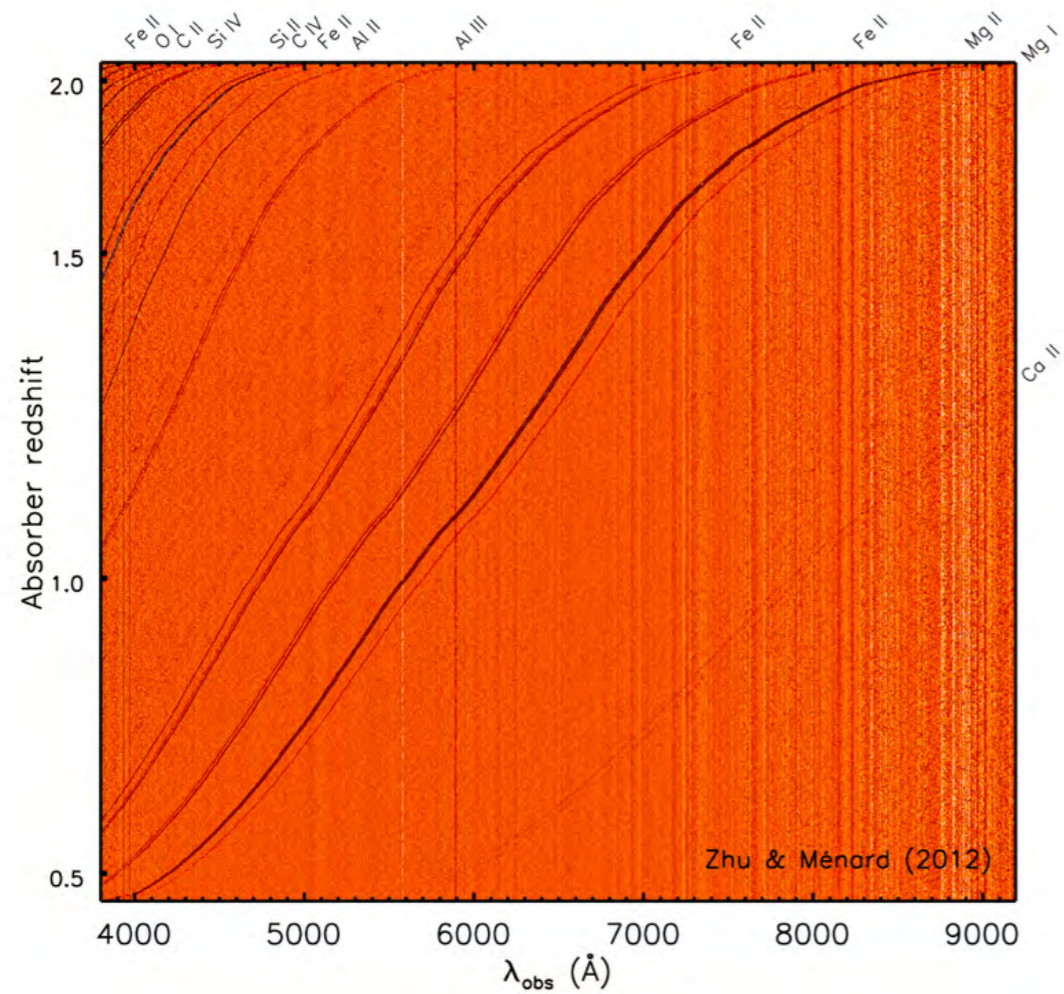
100,000 MgII absorbers **Publicly available**



Millions of photometric galaxies



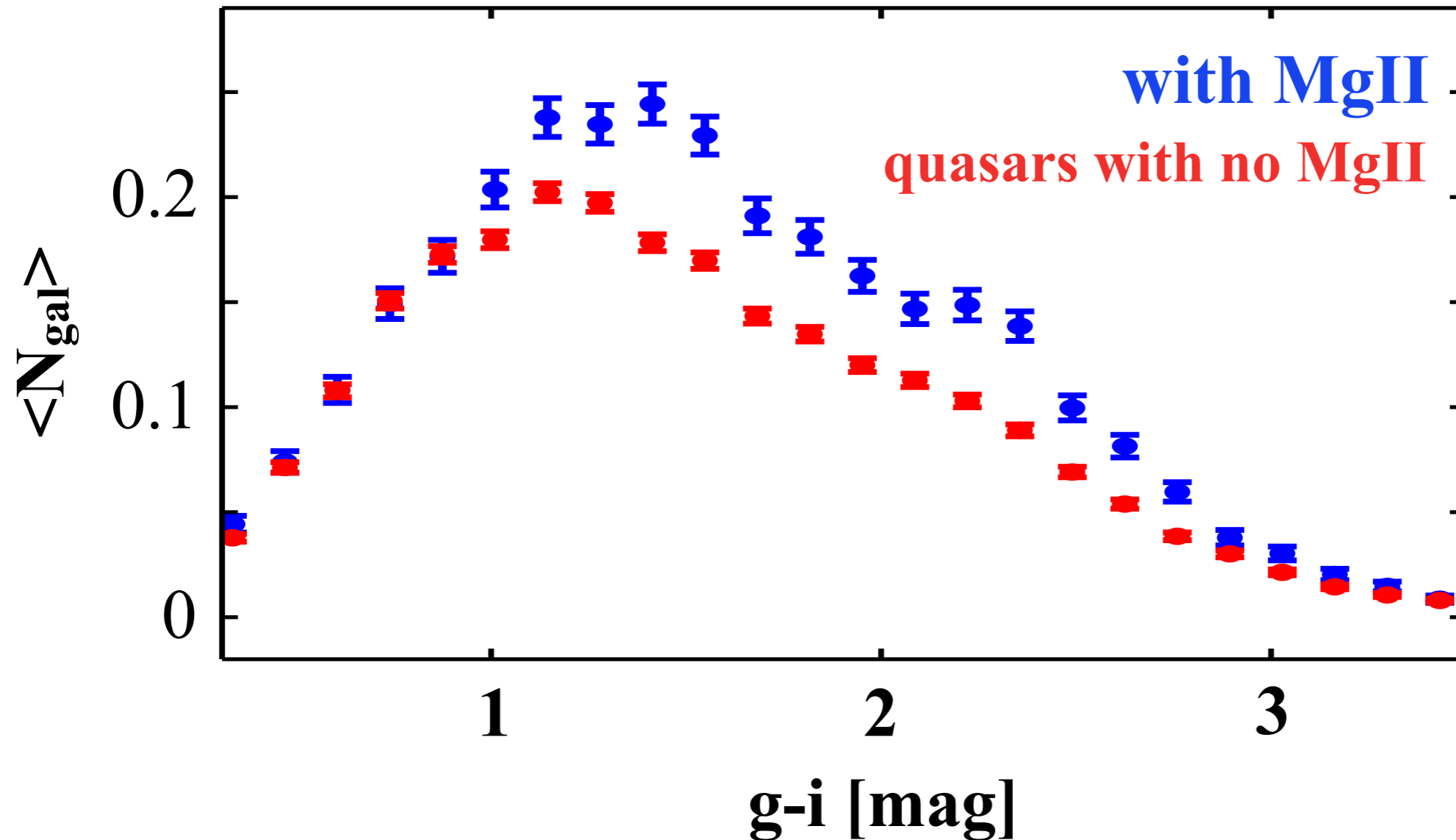
$$\langle \delta_{\text{MgII}} \cdot \delta N_{\text{gal}} \rangle$$



Focus on $\sim 3,000$ absorbers at $z \sim 0.5$

Galaxy number count around MgII absorbers

SDSS galaxies (<200 kpc) around 3,000 absorbers at $z \sim 0.5$



randomly-distributed
background galaxies

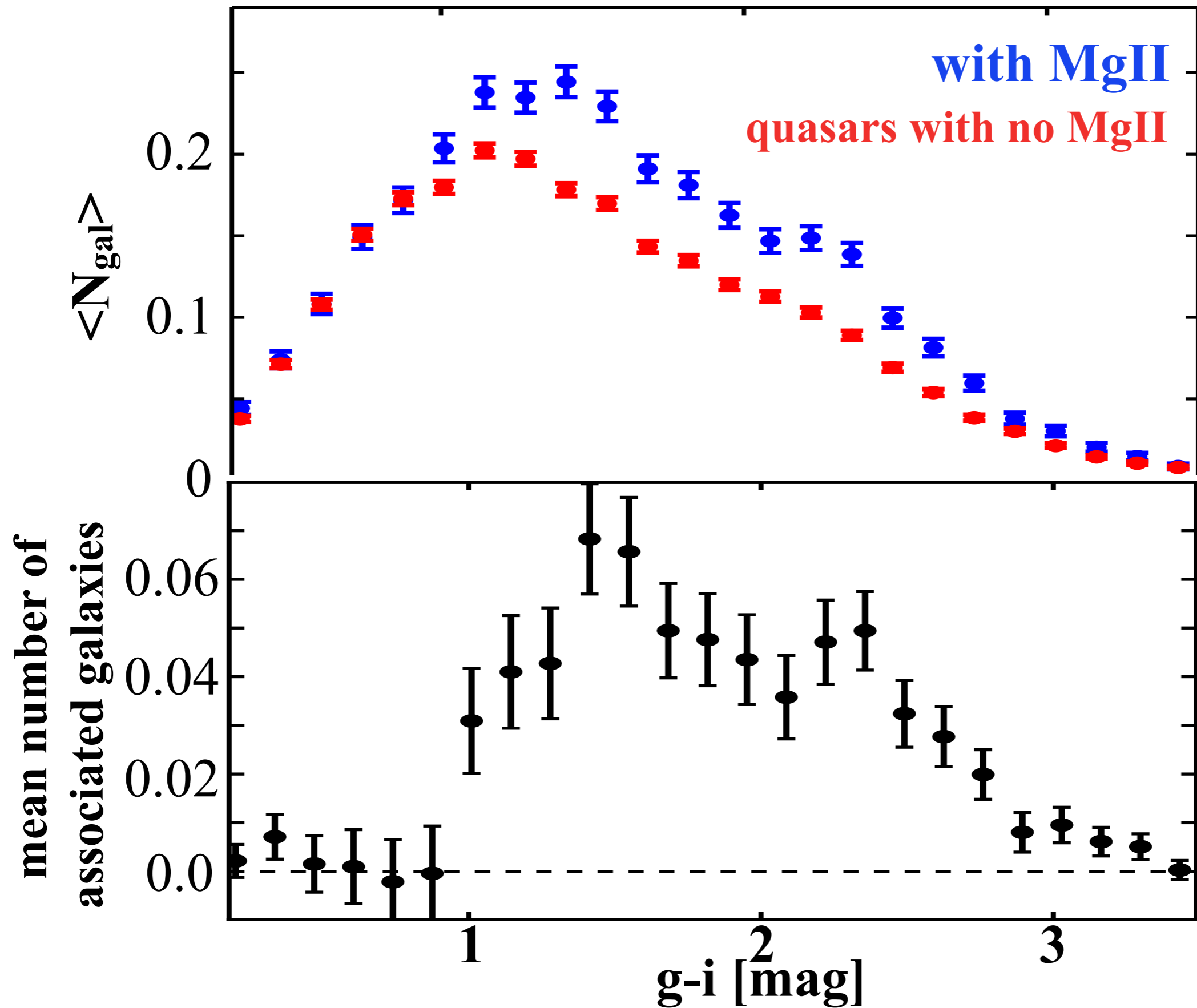
+

galaxies associated with MgII

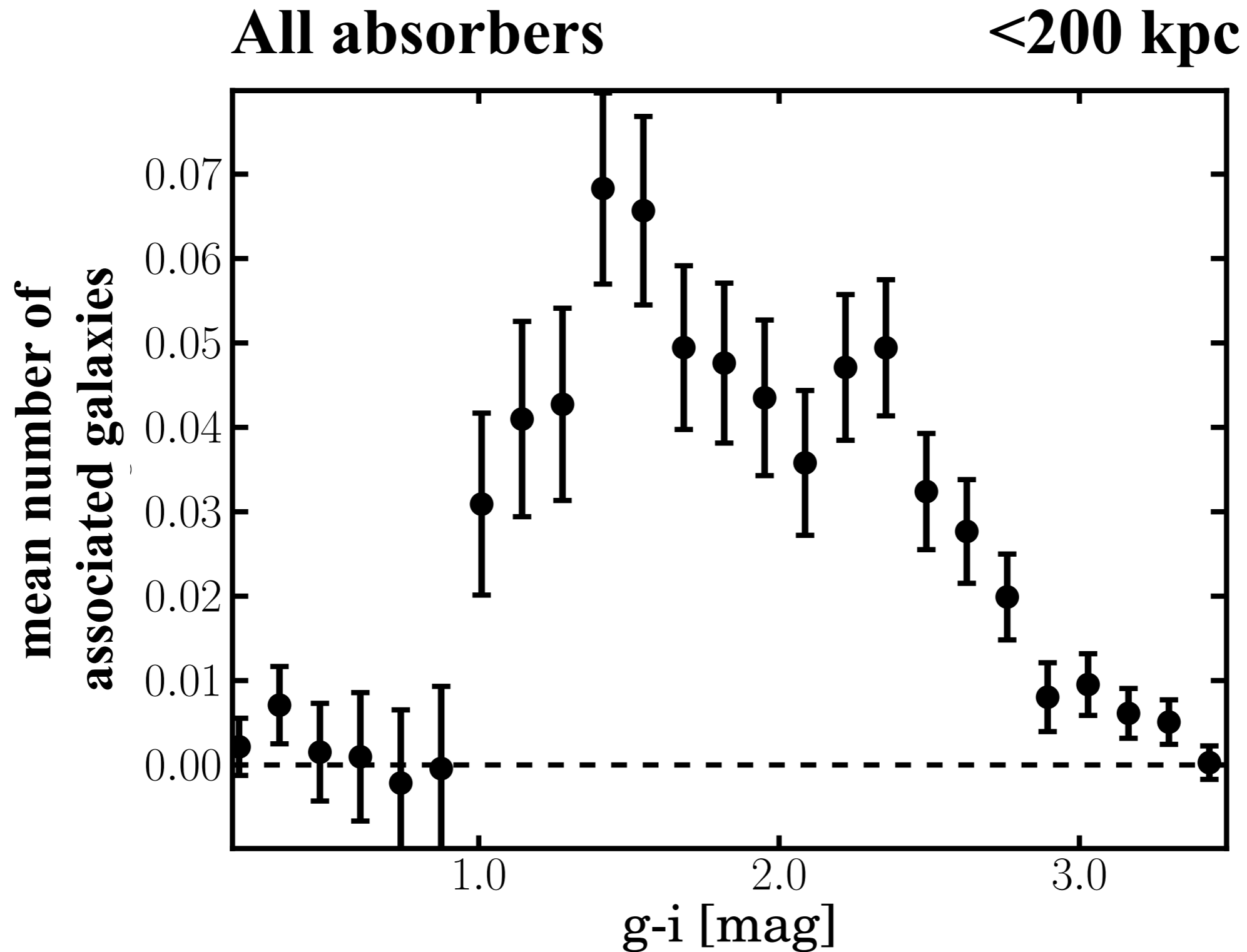
randomly-distributed
background galaxies

Galaxy number count around MgII absorbers

SDSS galaxies (<200 kpc) around 3,000 absorbers at $z \sim 0.5$

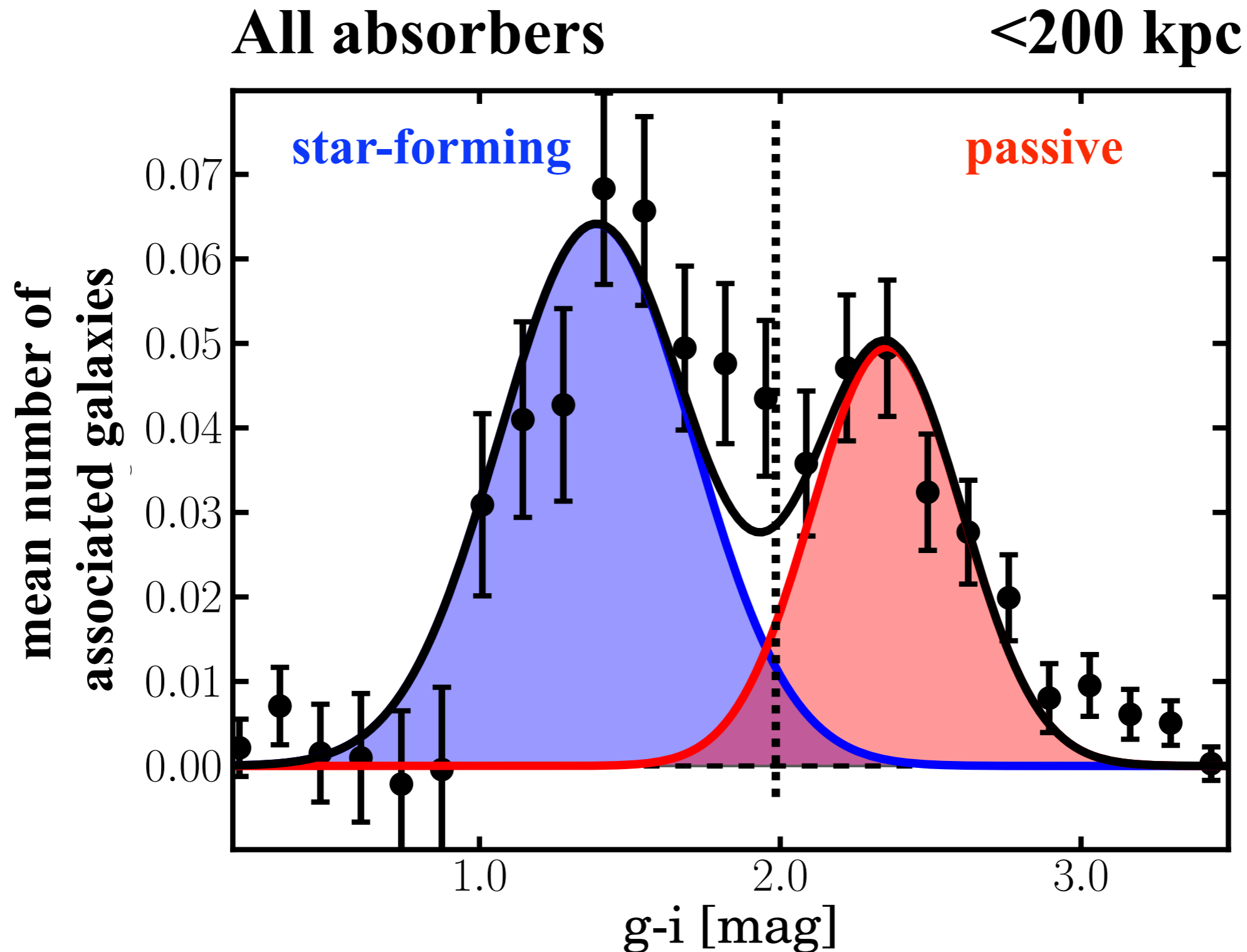


The color distribution of galaxies associated with absorbers



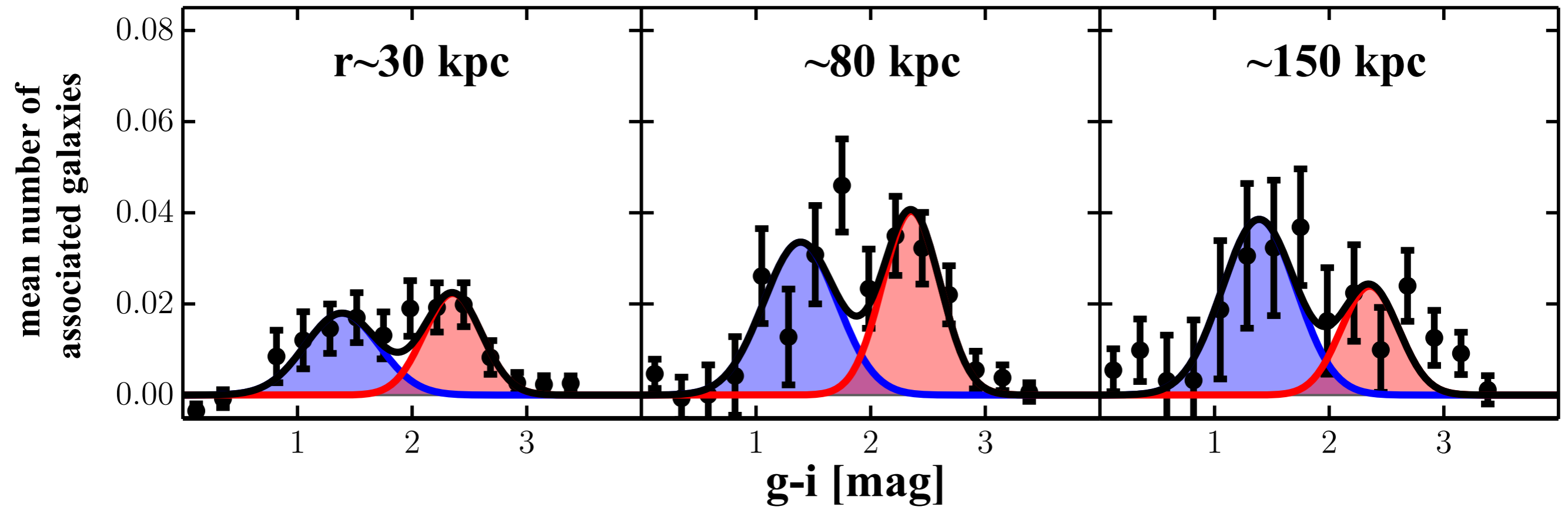
~2,000 galaxy-absorber pairs statistically.

The color distribution of galaxies associated with absorbers

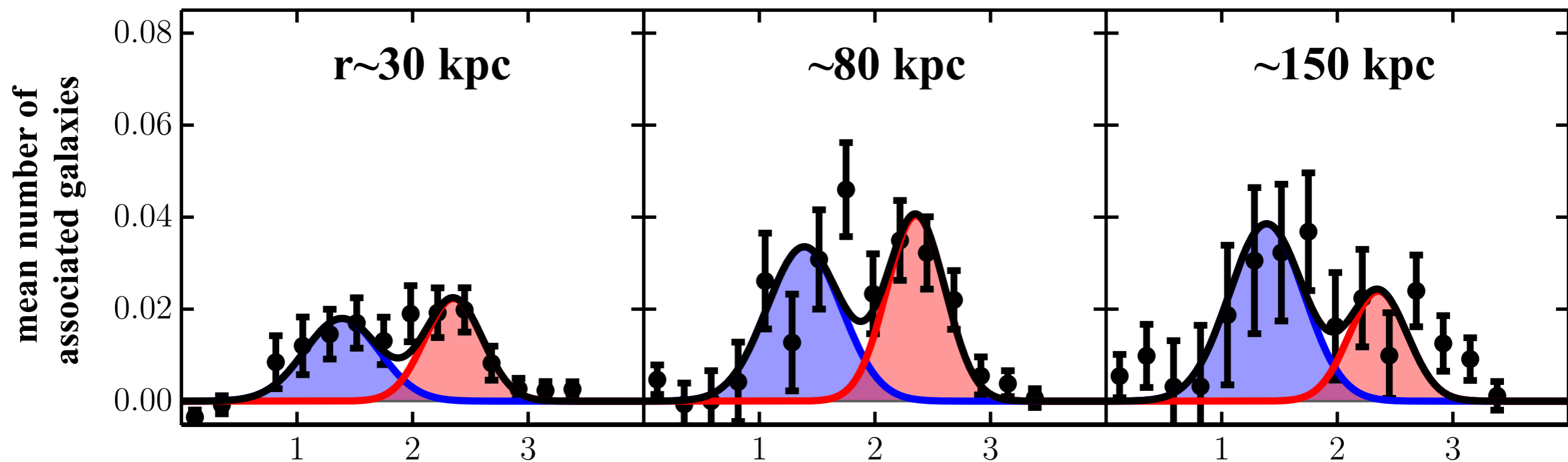


Cool gas is around both
star-forming and **passive** galaxies.

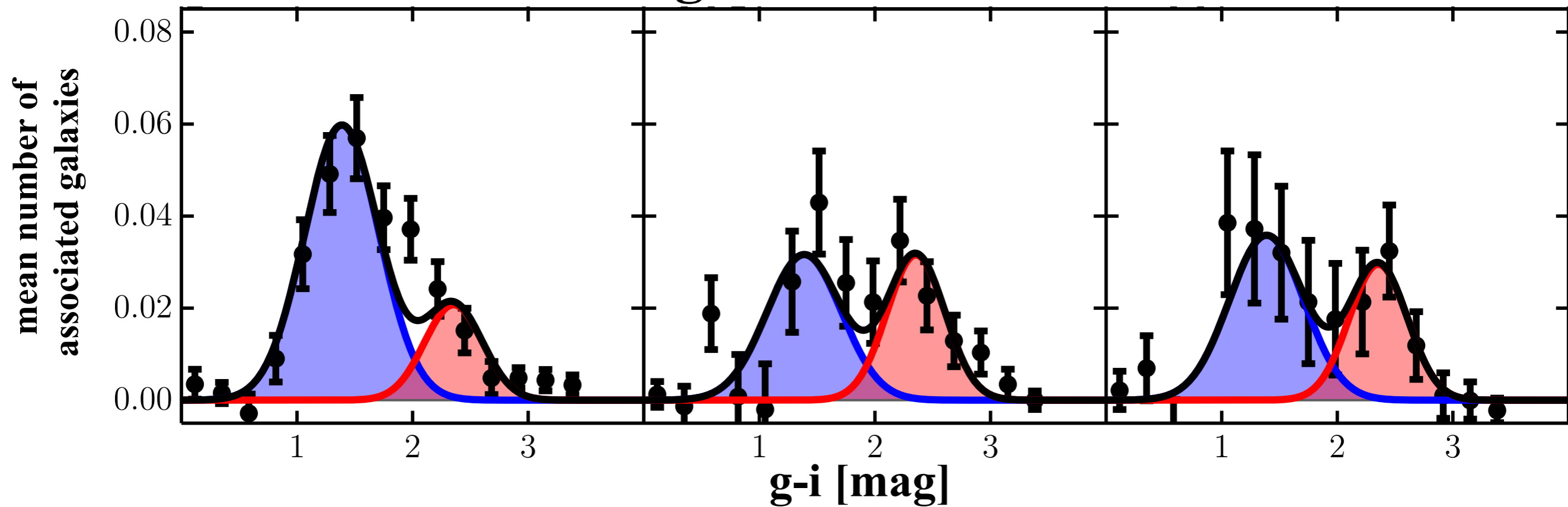
weak absorbers $W < 1 \text{ \AA}$



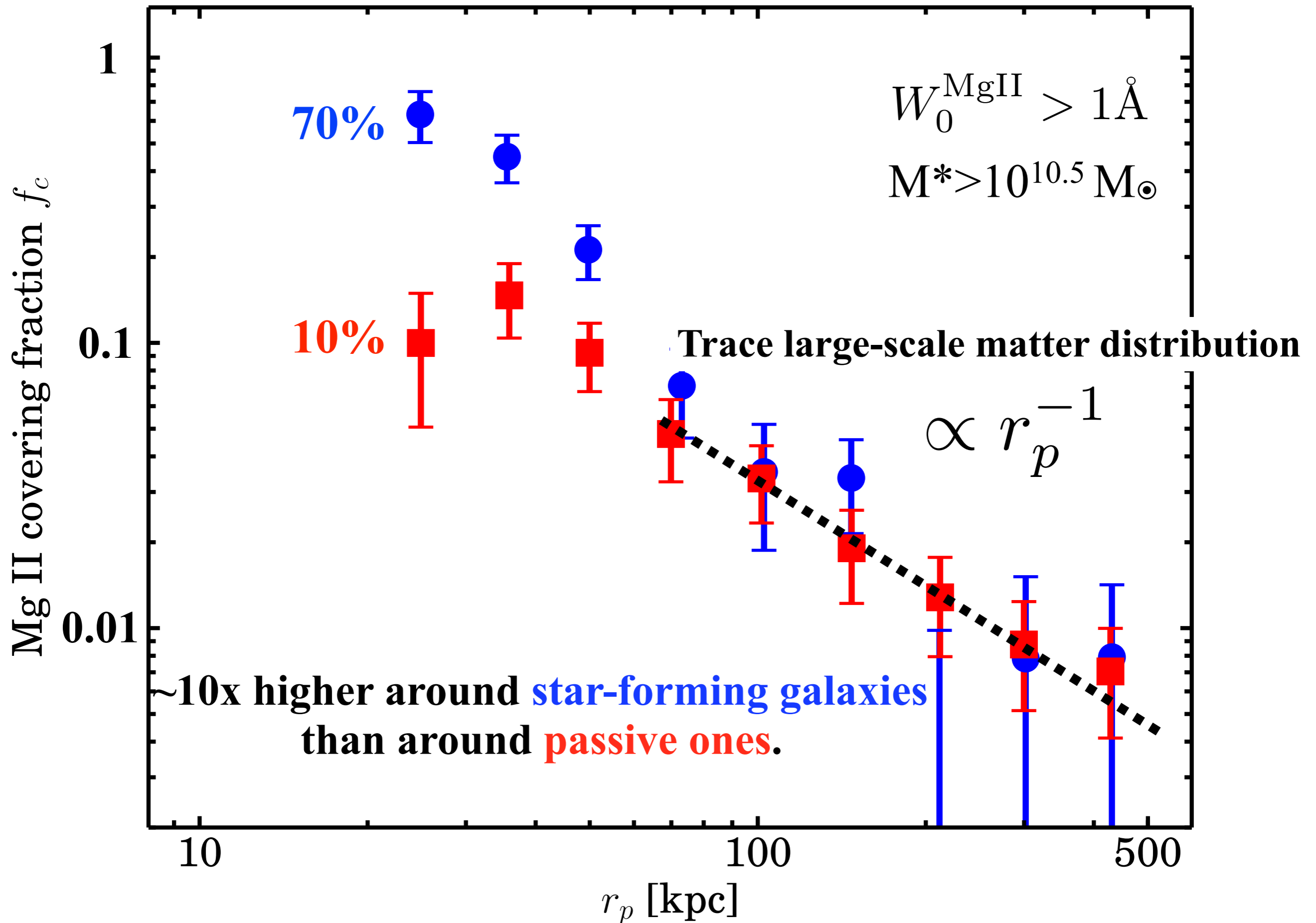
weak absorbers $W < 1 \text{ \AA}$



strong absorbers $W > 1 \text{ \AA}$



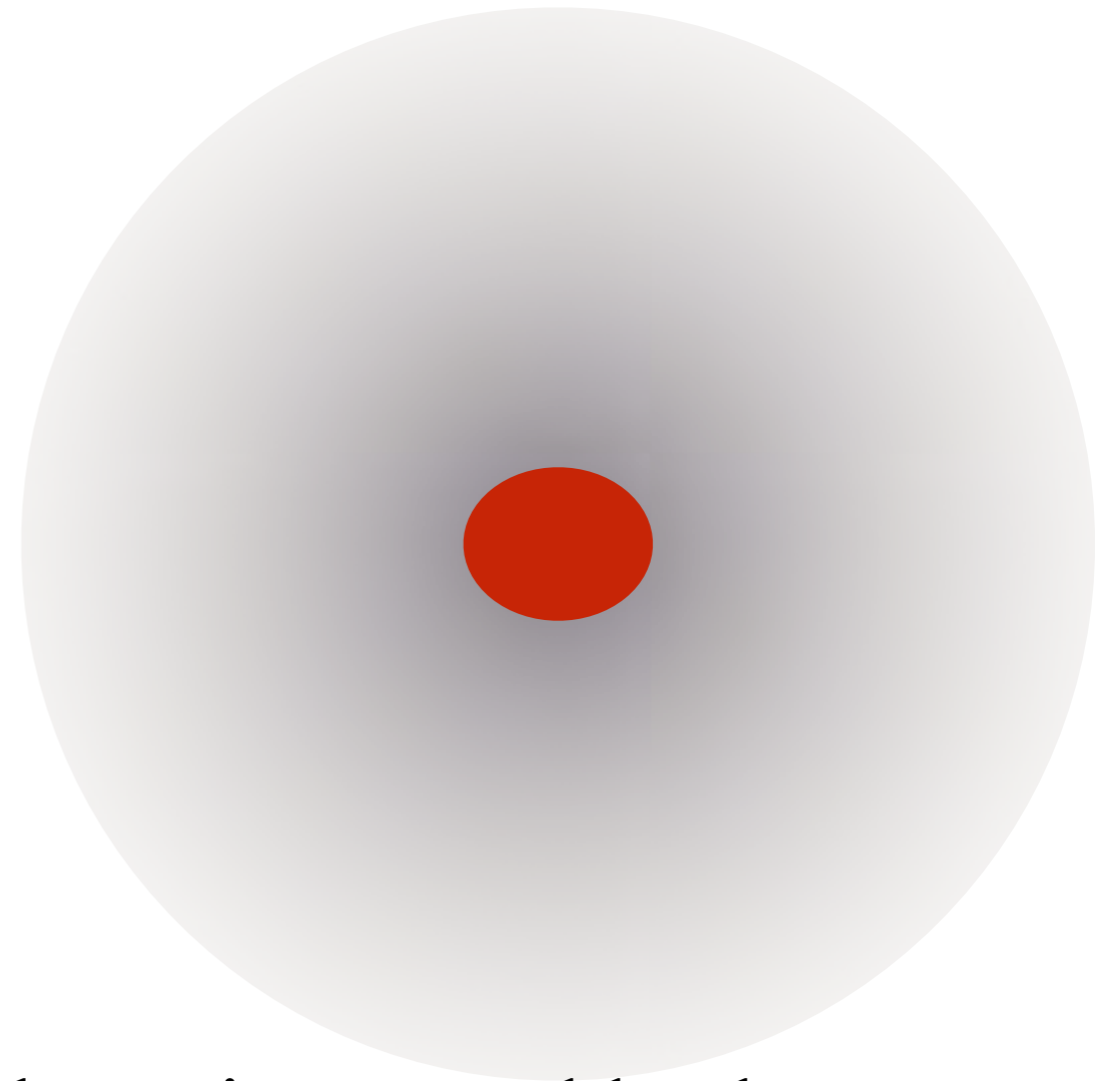
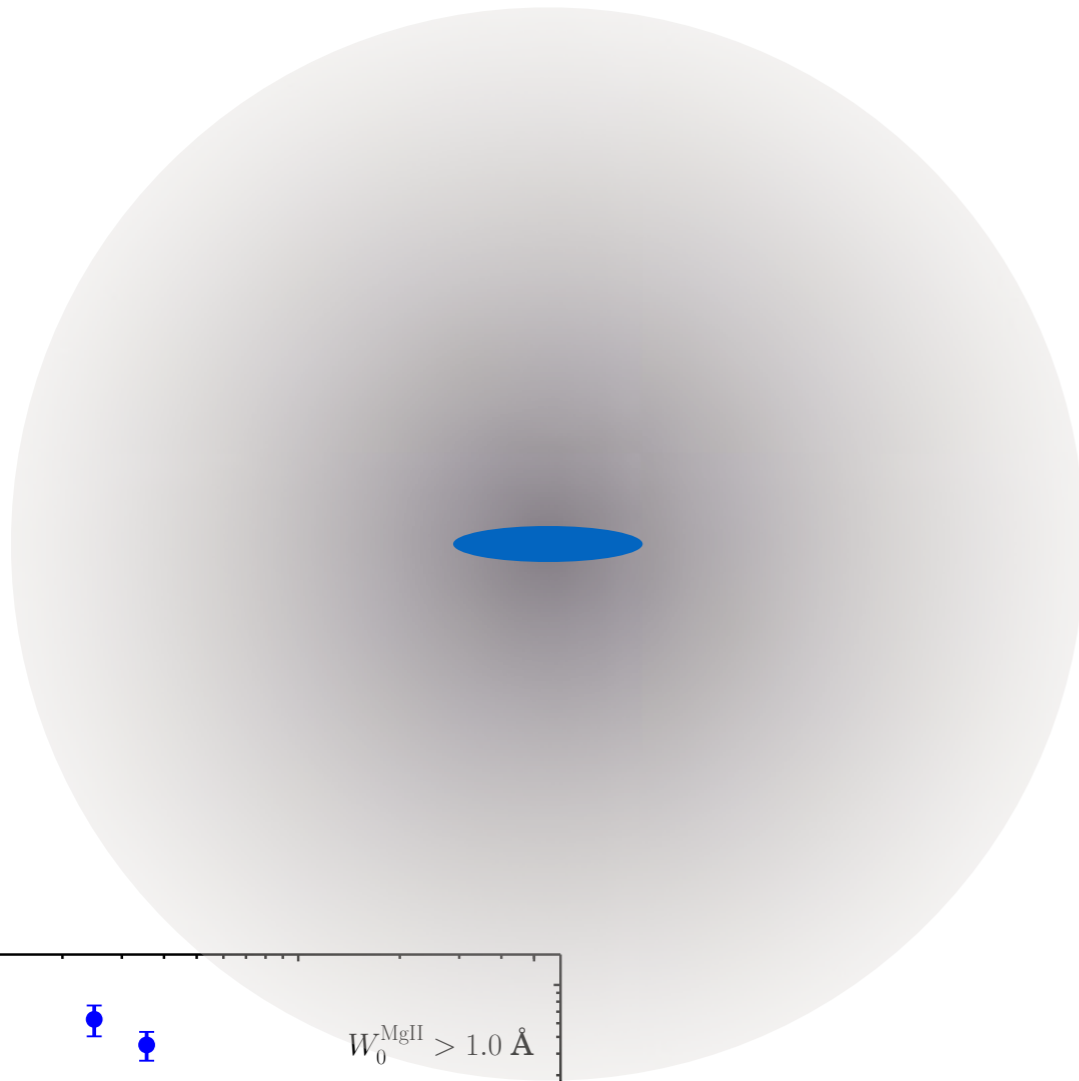
Mg II covering fraction



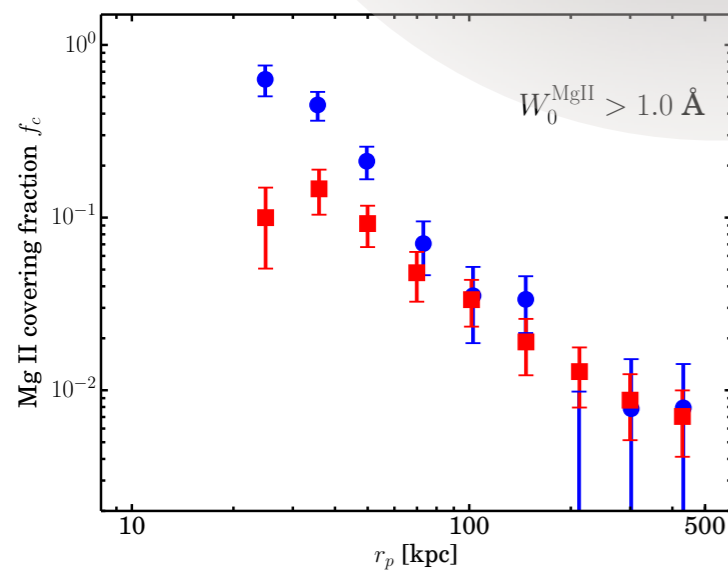
Cool gas around galaxies

Star-forming galaxies

Passive galaxies



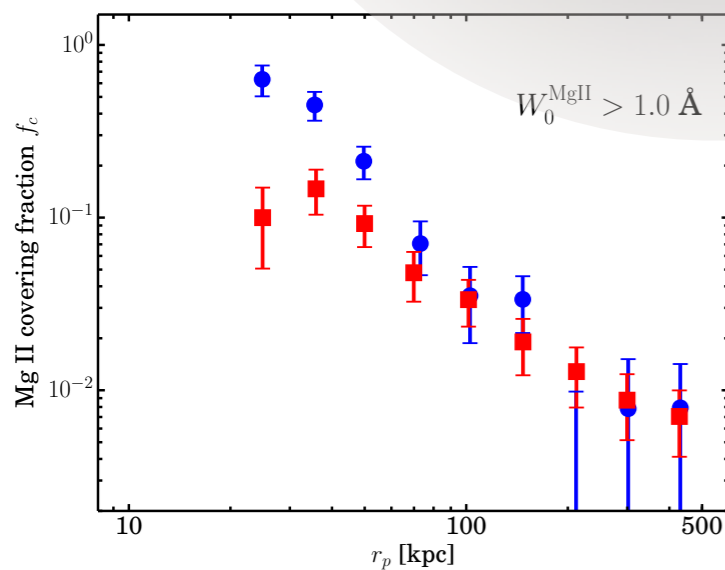
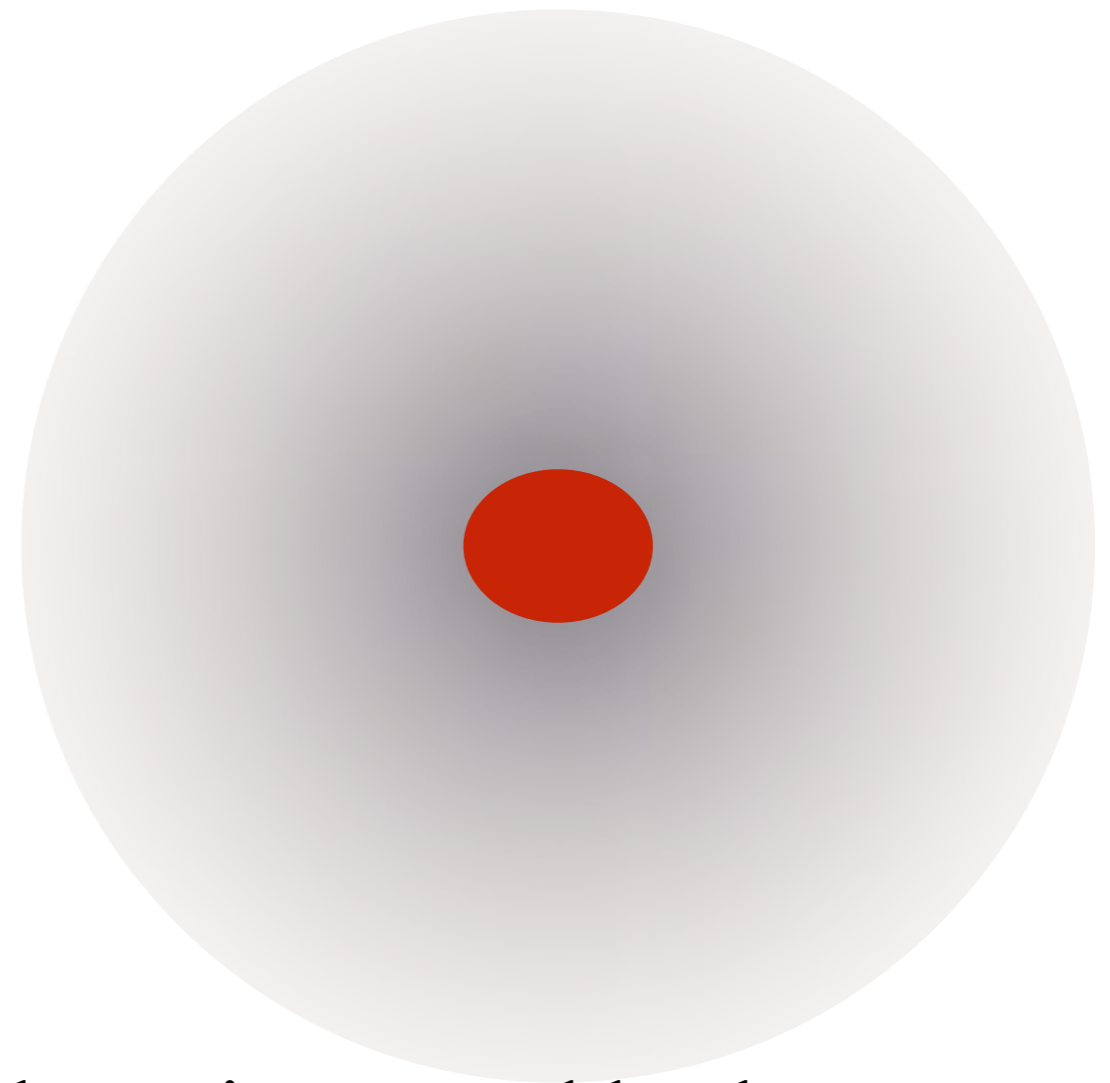
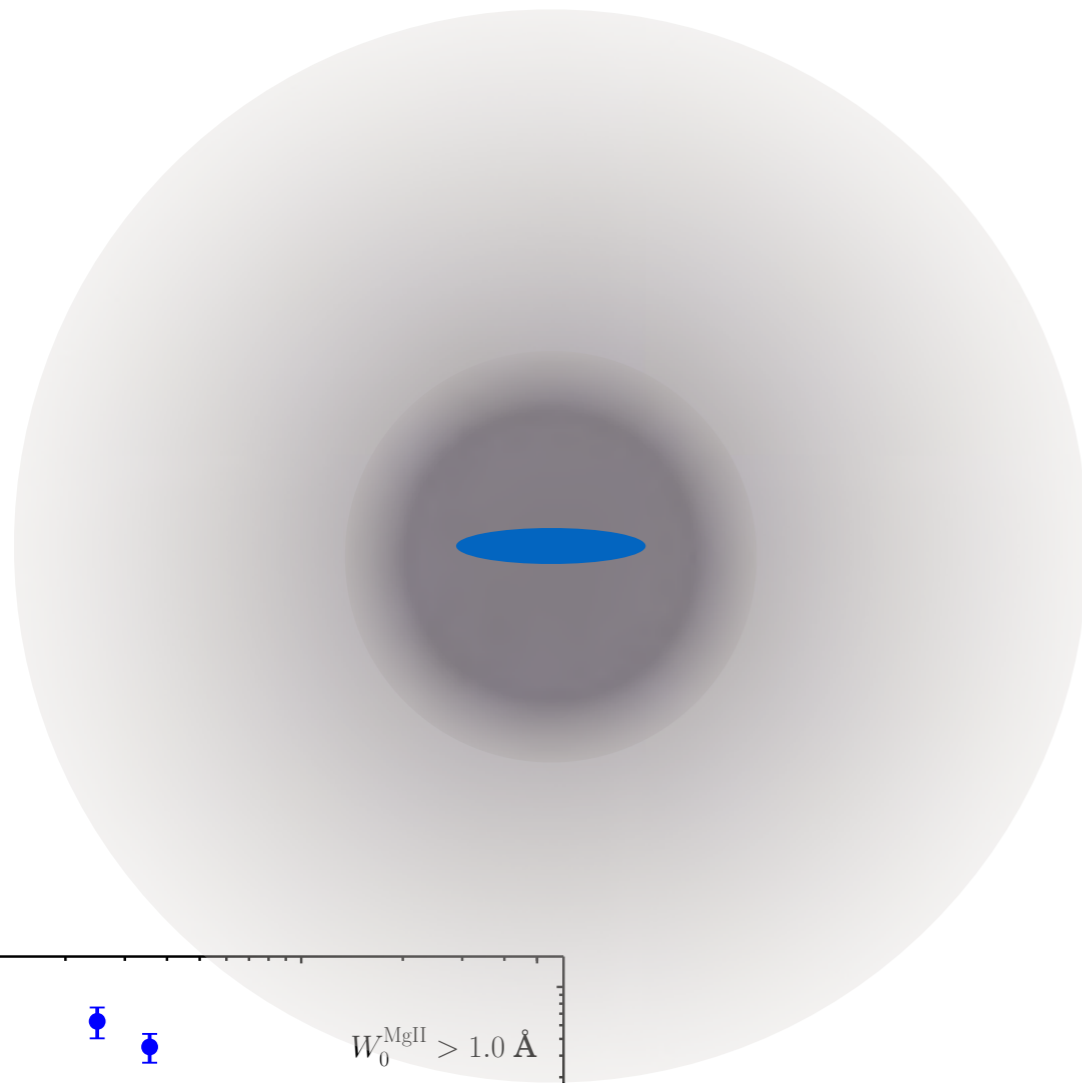
Cool gas is around both
star-forming galaxies and passive galaxies.



Cool gas around galaxies

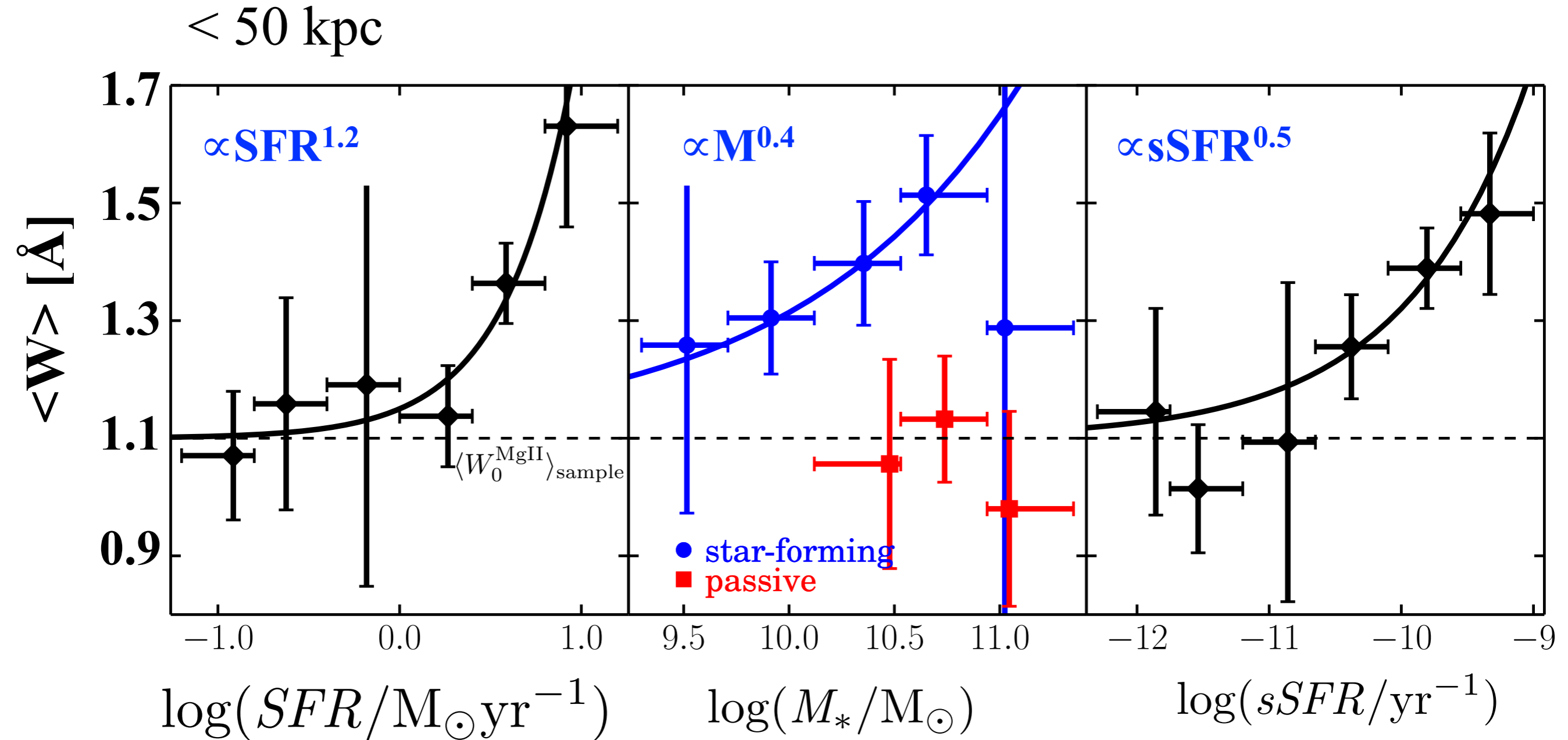
Star-forming galaxies

Passive galaxies



Cool gas is around both
star-forming galaxies and **passive** galaxies.
Excess absorption
around **star-forming** galaxies within 50 kpc.

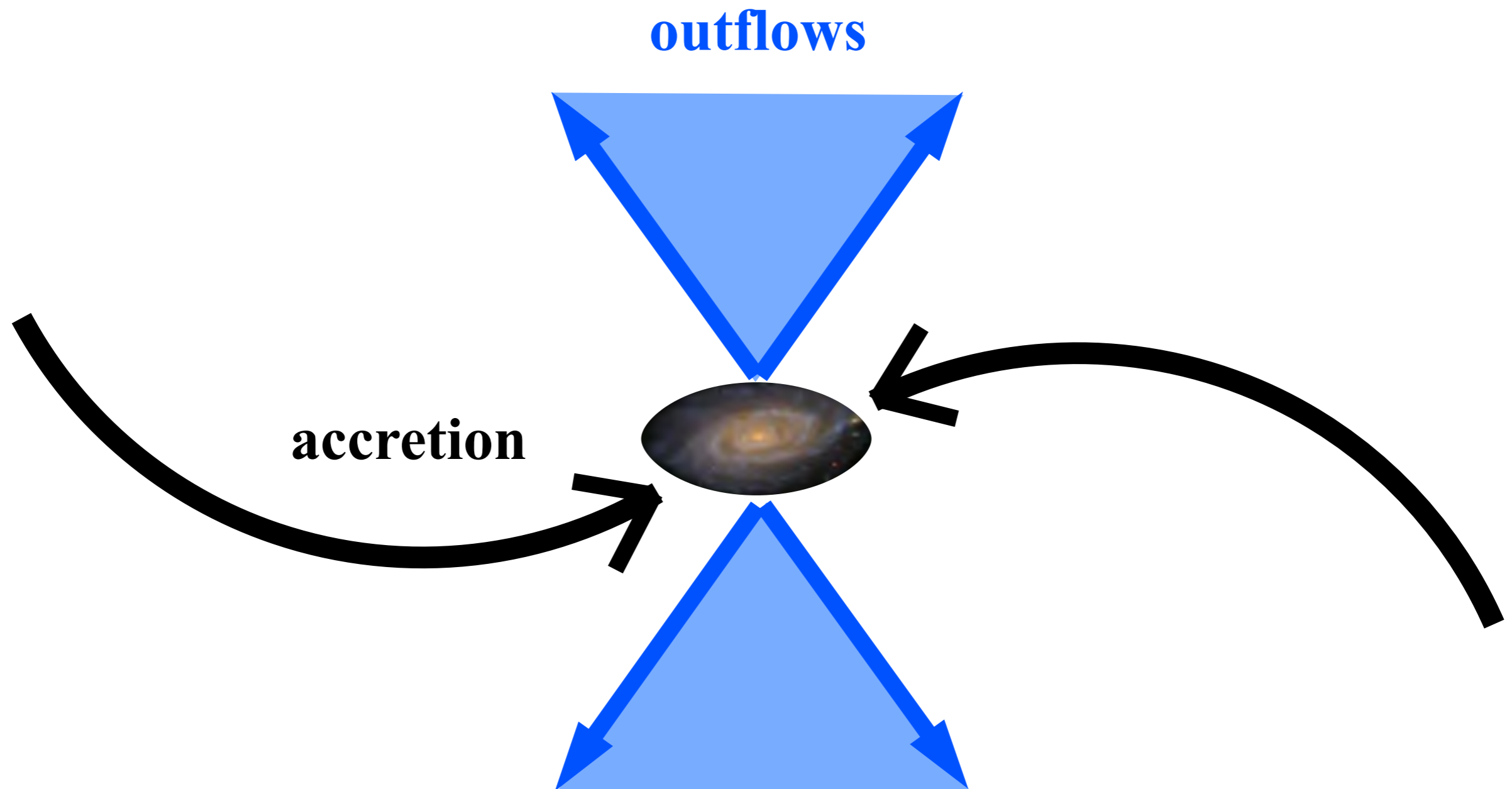
The mean absorption as a function of galaxy properties



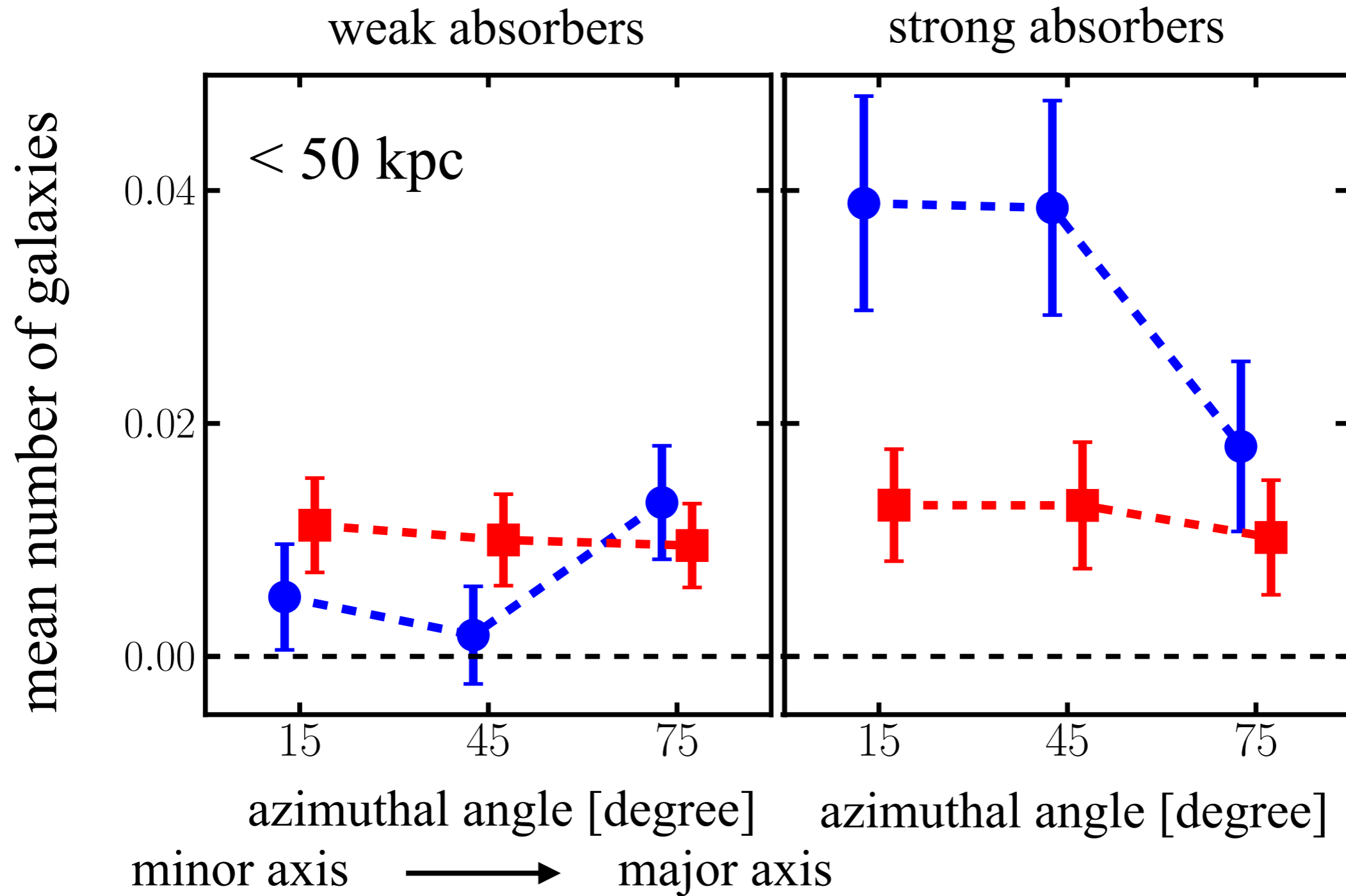
The scaling relations between absorption and galaxy properties

The cool circumgalactic gas ($r \sim 50 \text{ kpc}$) knows about galaxy properties ($r \sim 5 \text{ kpc}$).

Signatures of gas flows in the circumgalactic medium



Azimuthal Angle Dependence edge-on galaxies

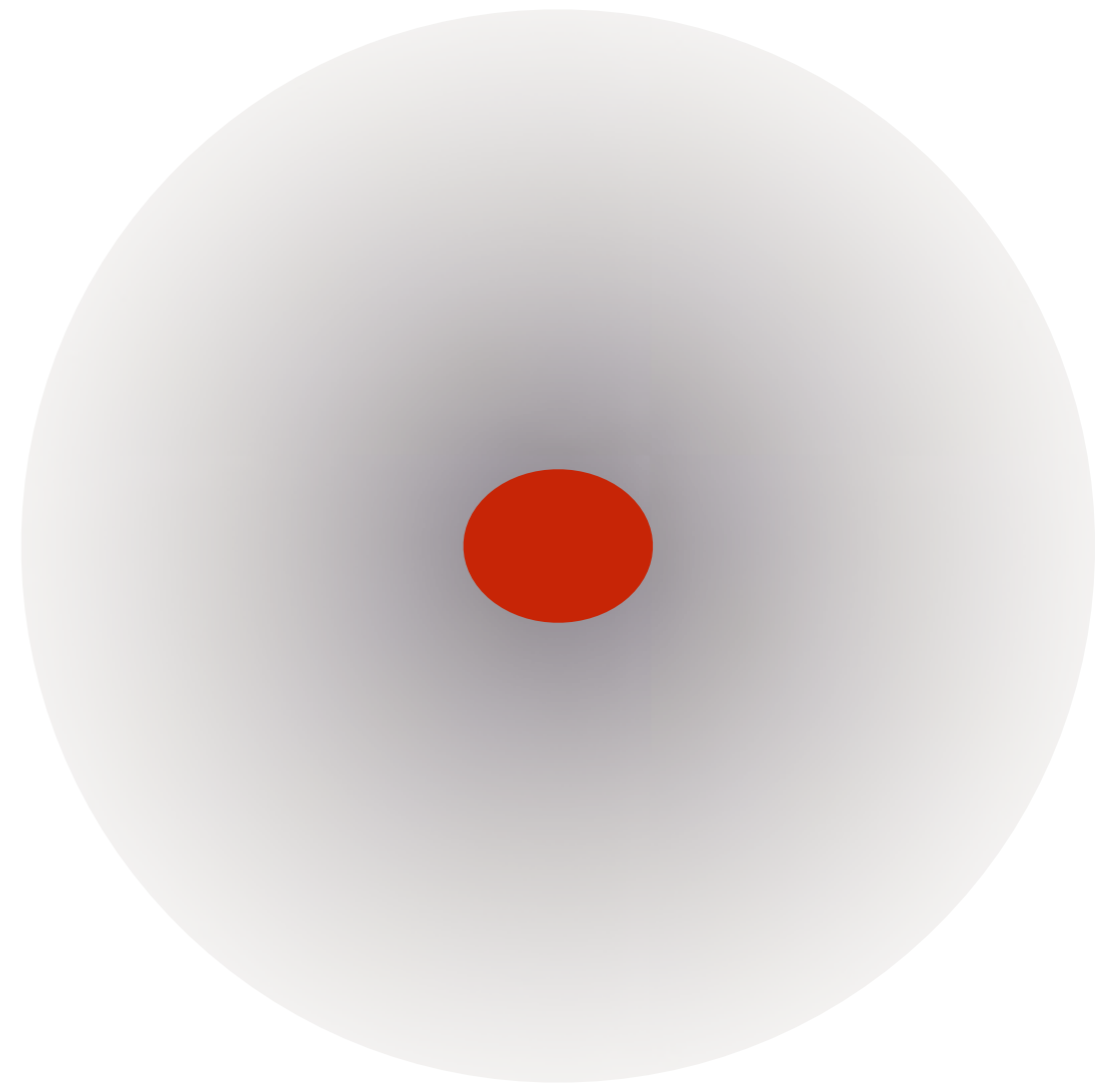
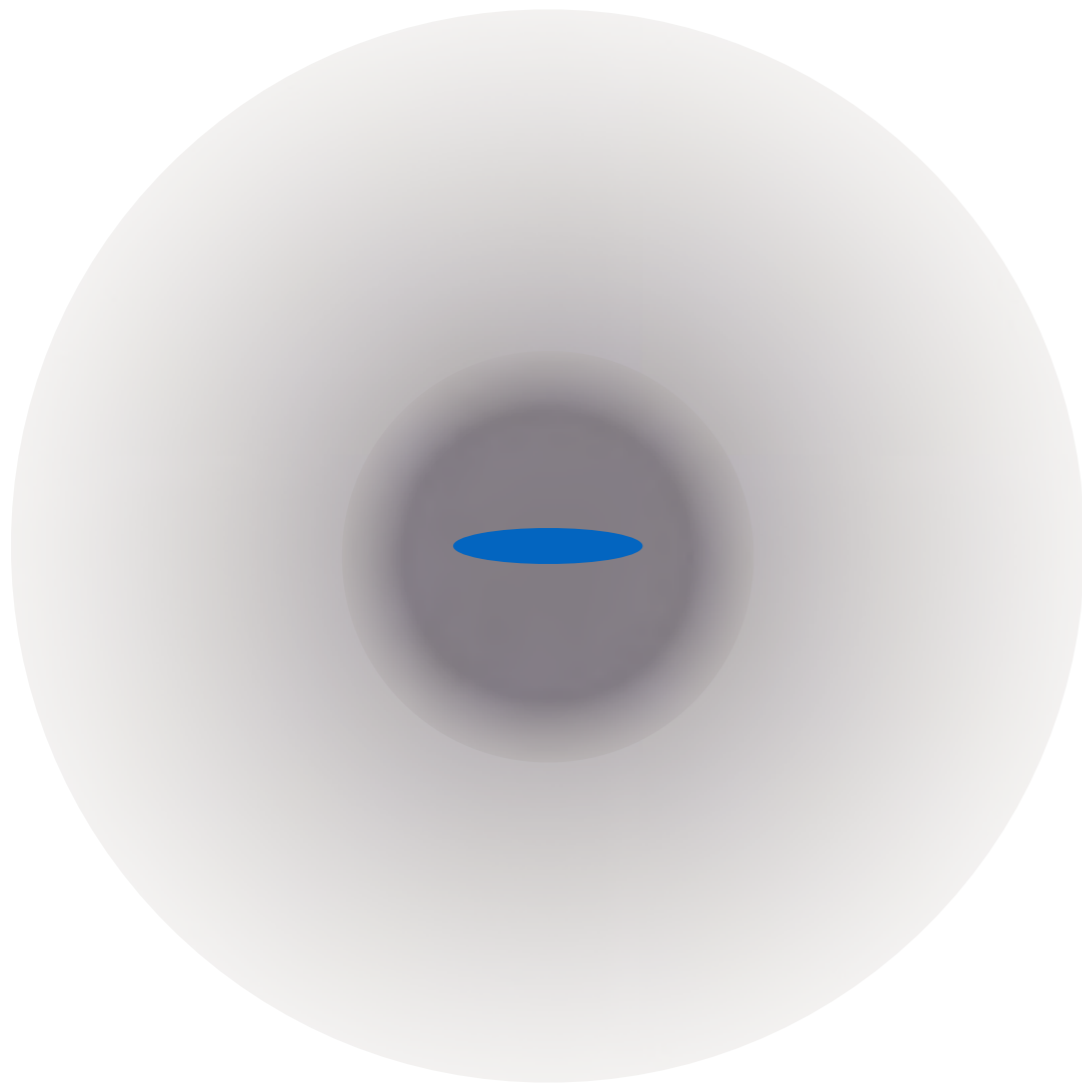


**Strong absorbers around star-forming galaxies
tend to trace outflowing gas.**

Cool gas around galaxies

Star-forming galaxies

Passive galaxies

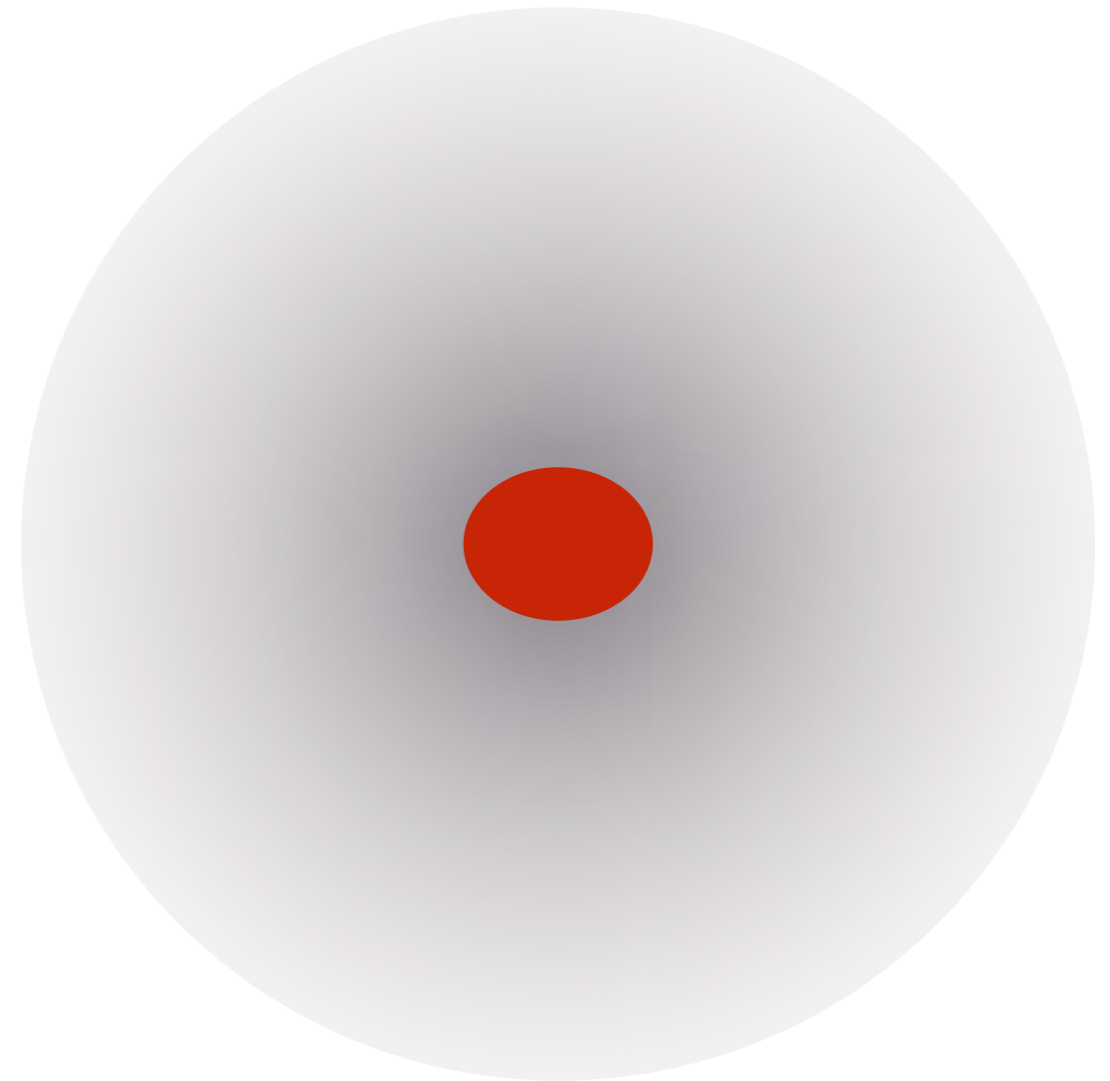
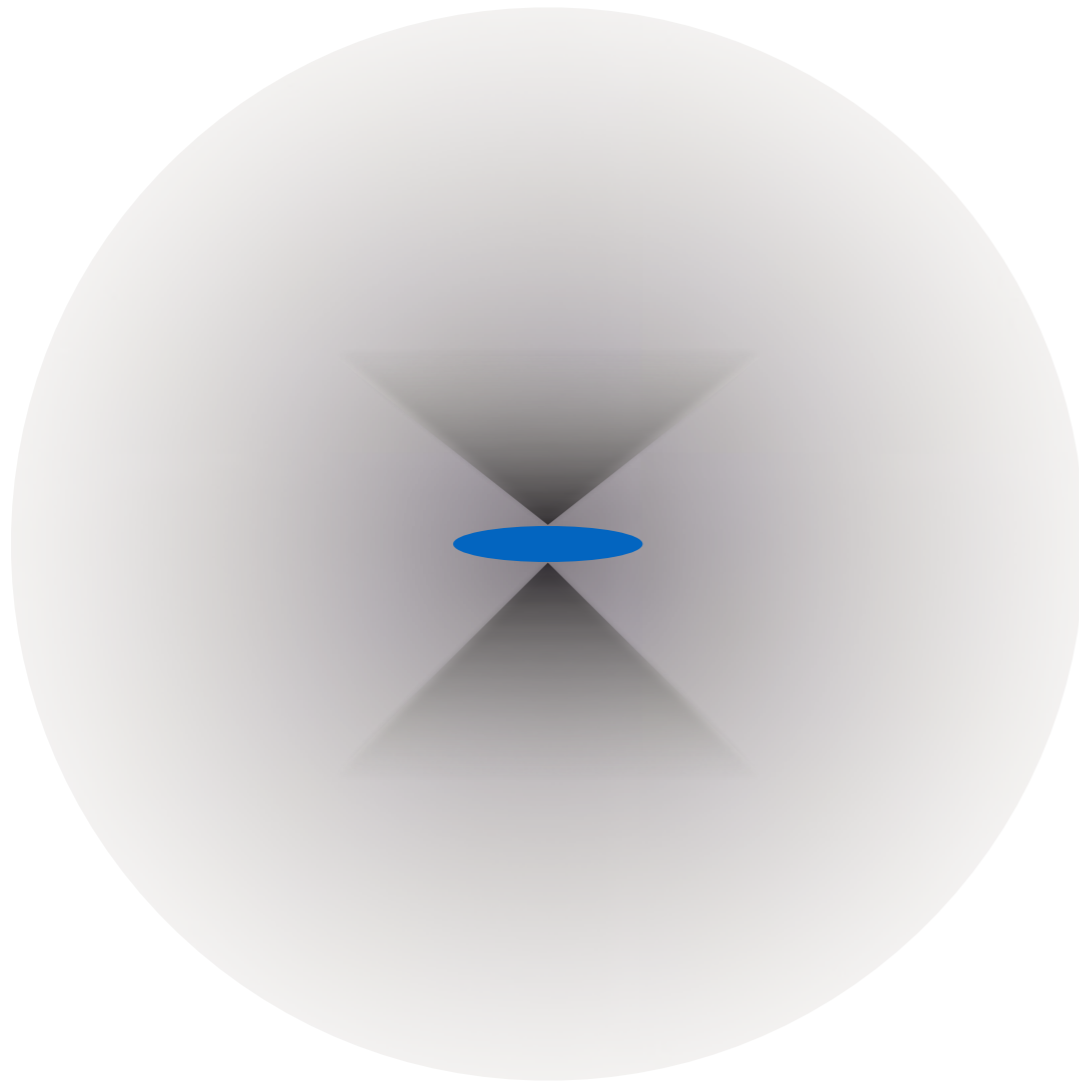


Excess absorption
around **star-forming** galaxies within 50 kpc.

Cool gas around galaxies

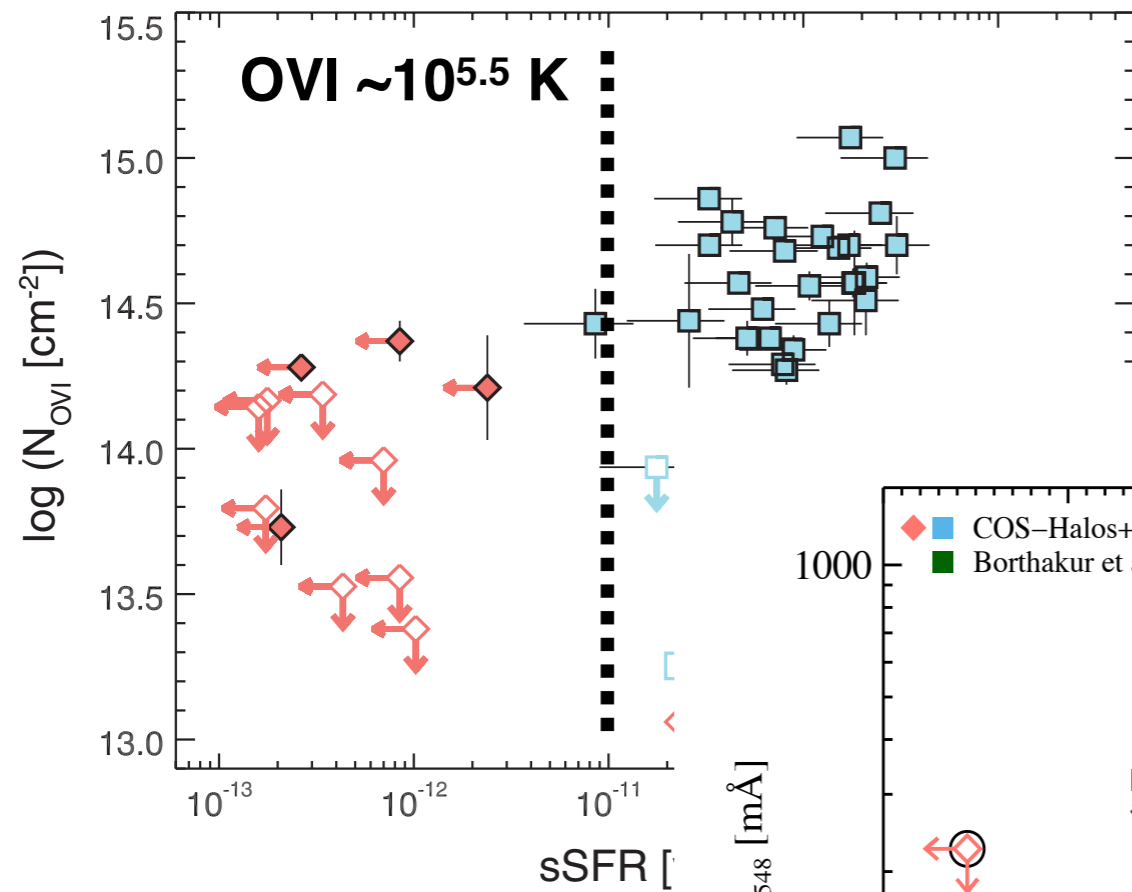
Star-forming galaxies

Passive galaxies



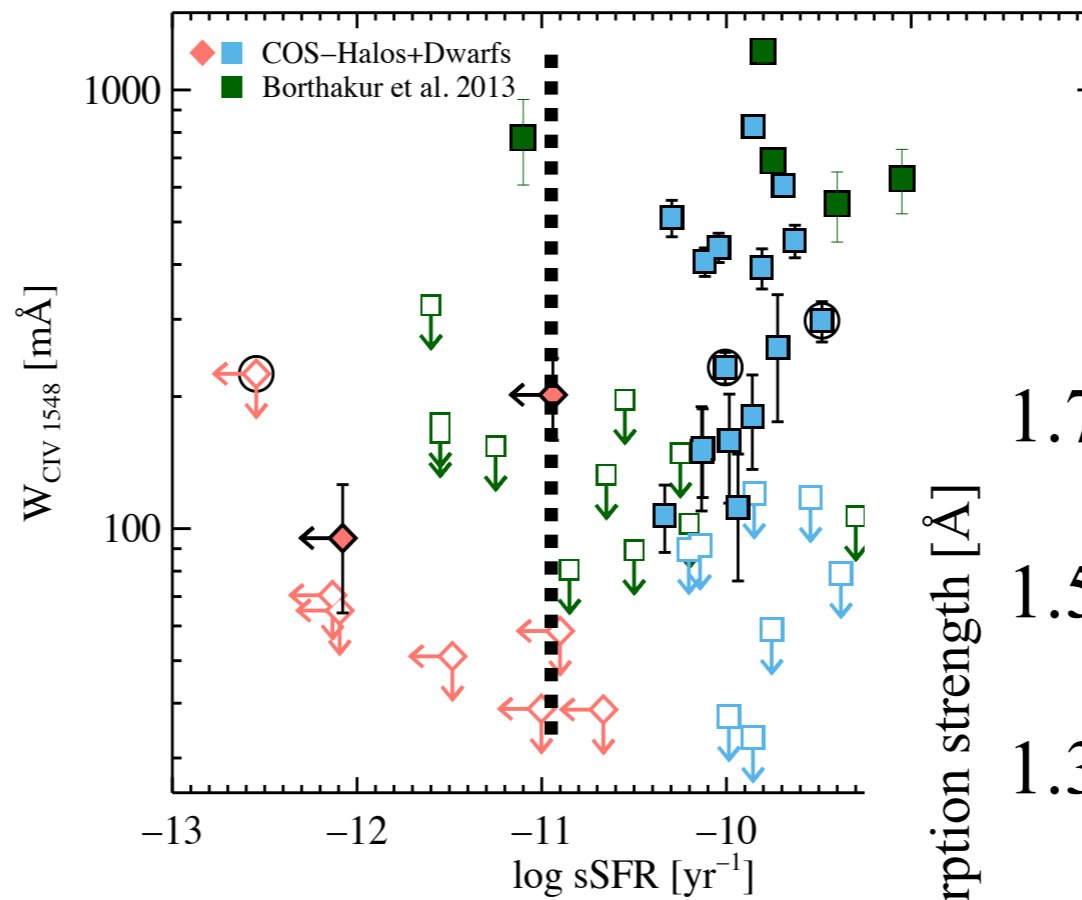
Excess absorption
around **star-forming** galaxies within 50 kpc.

The dichotomy of galaxy types in the CGM



Tumlinson & COS-Halos

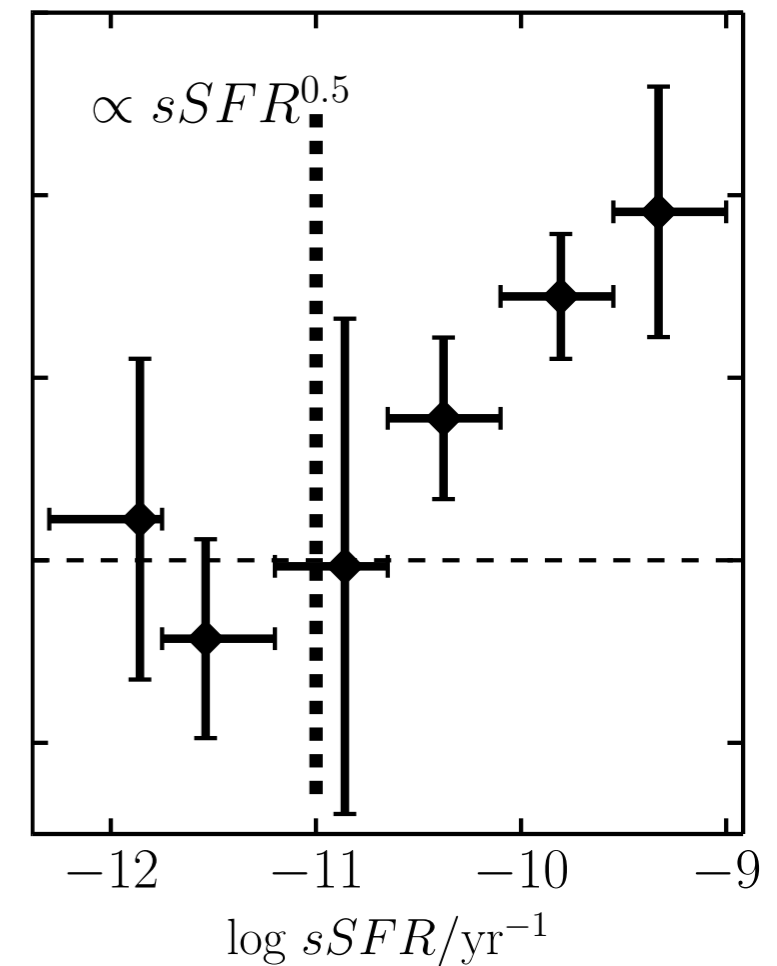
CIV $\sim 10^5 \text{ K}$



Bordoloi & COS-Halo team
Borthakur et al. 2013

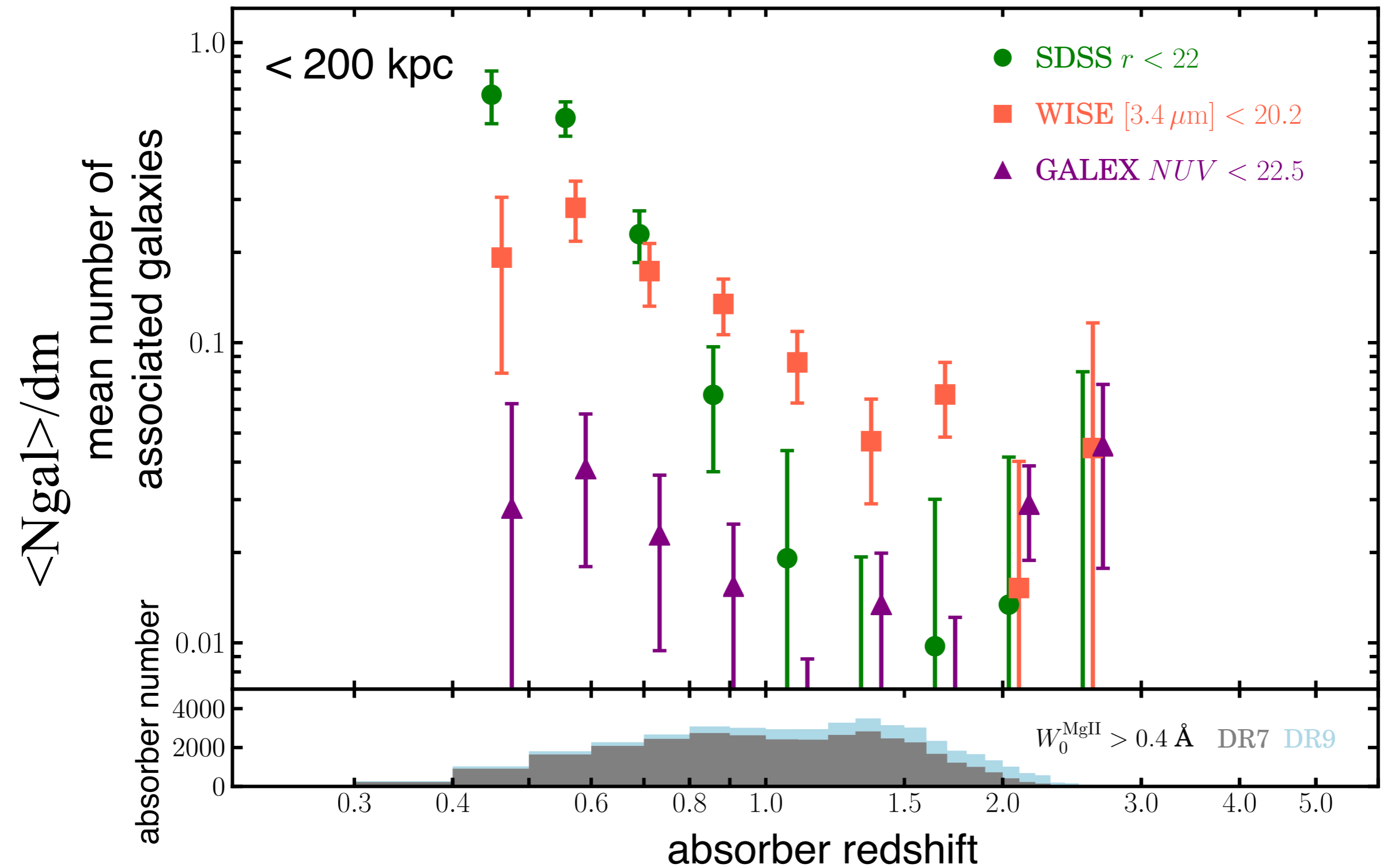
mean absorption strength [Å]

MgII $\sim 10^4 \text{ K}$

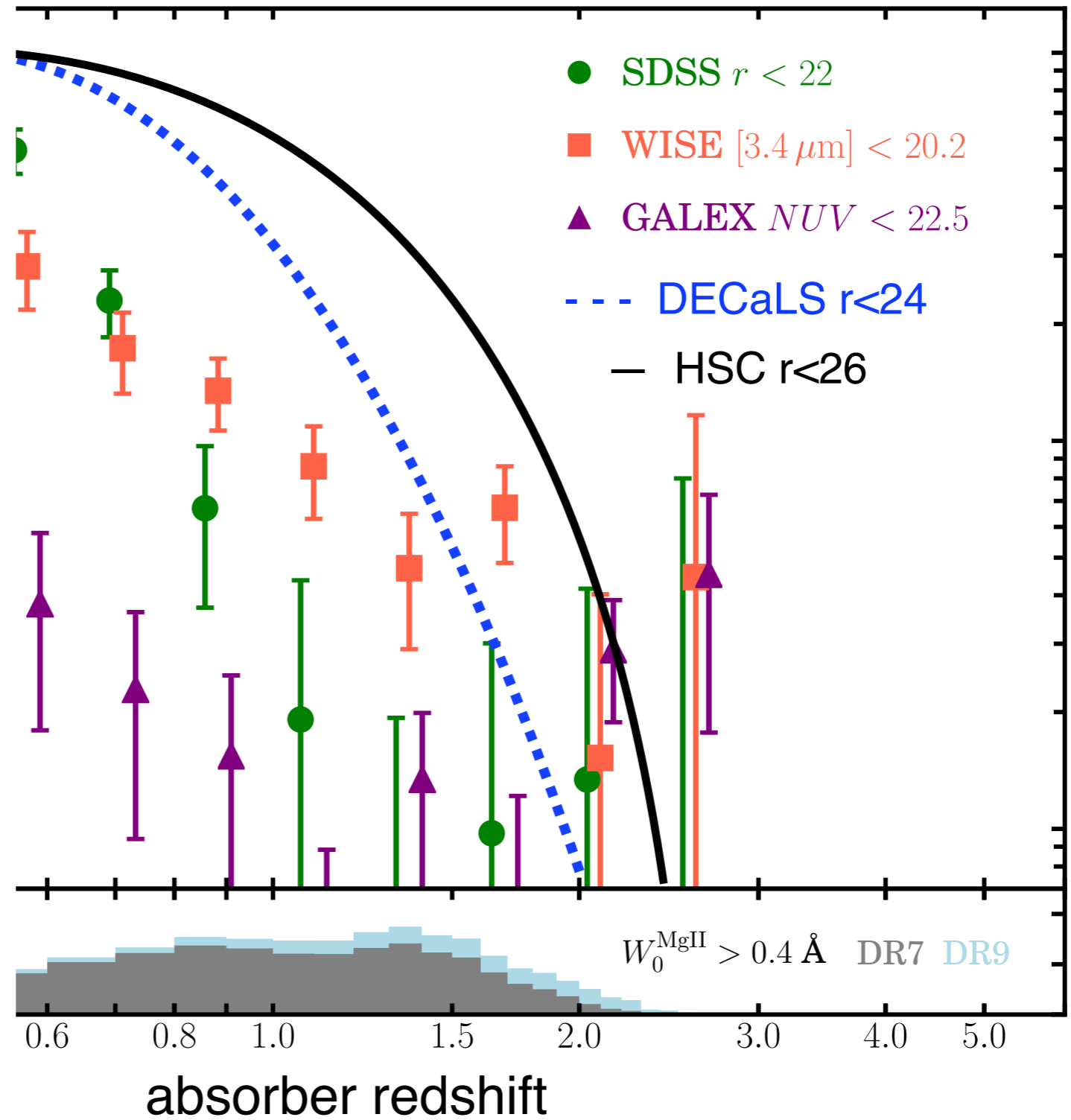
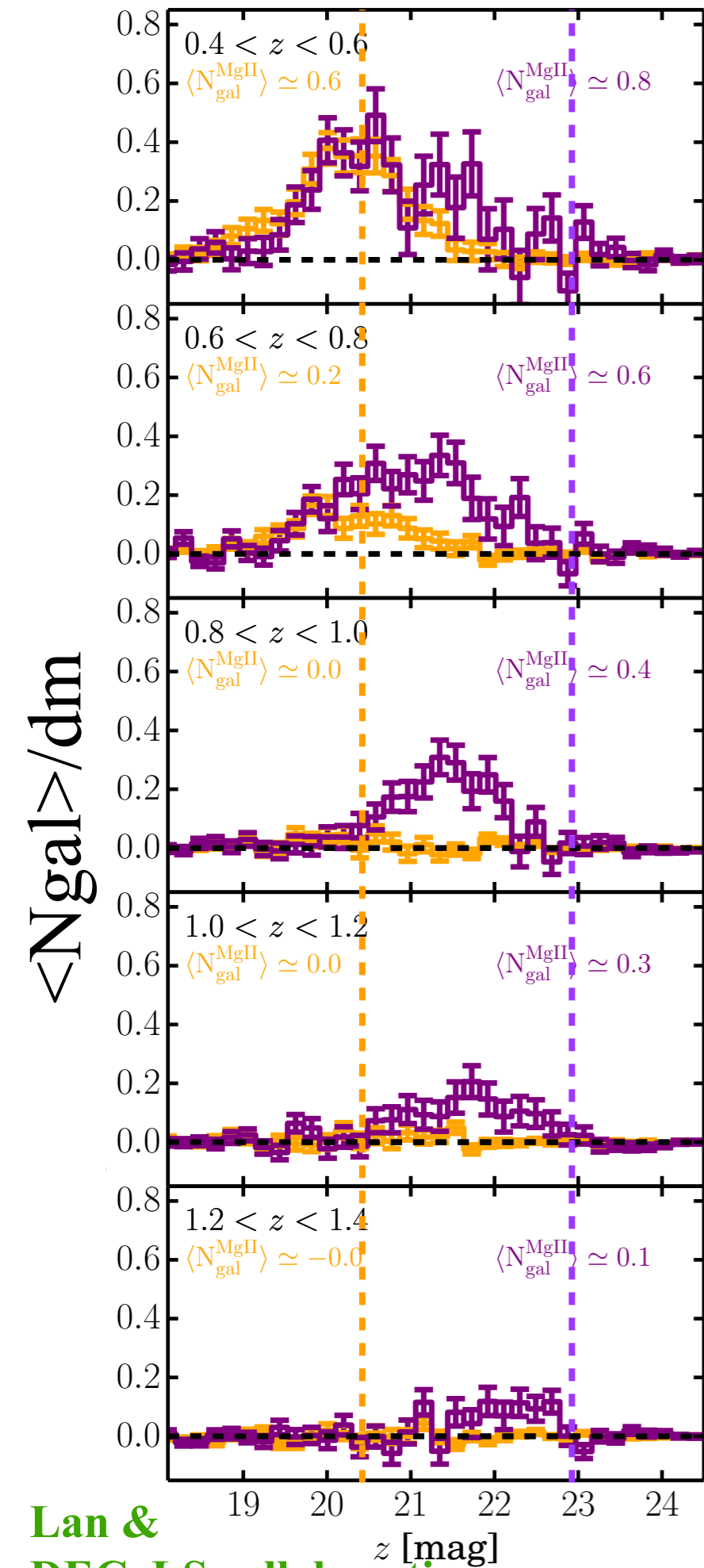


How large imaging surveys can help before PFS?

Connecting the CGM and galaxy properties with large surveys



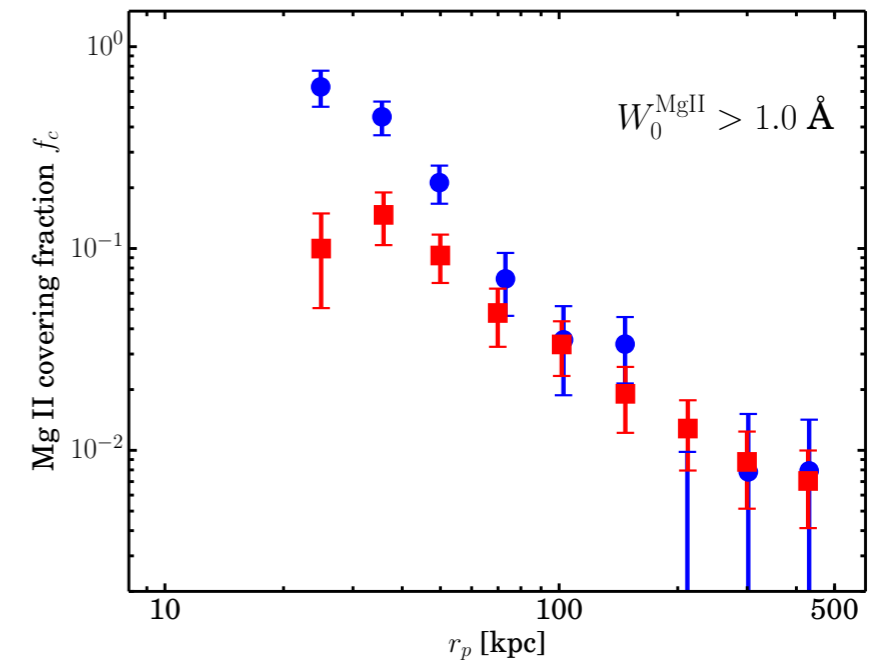
and galaxy properties with large surveys



Probing galaxies and CGM relation through cosmic time!

Is there cool gas around both types of galaxies?

Yes, there is cool gas around all galaxies.



Do the properties of the cool CGM depend on the galaxy properties?

Yes, there is more MgII absorption around star-forming galaxies.

Scaling relations between gas and galaxies.

