

A PFS Cosmology Survey thought experiment:

Goals:

- I. Enable breakthrough galaxy science.
- 2. Maximize unique potential for cosmology.



The galaxy formation fantasy is the SDSS at z~I

Spec-z survey Wish List

- I. Mass complete to $z\sim I... > log M = 10.5$
- 2. SDSS like volume. V > 0.2 Gpc3
- 3. High sampling/success rate.. > 70%
- 4. Strong S/N.. > 10
- 5. Wide wavelength range
- 6. Blue sensitivity

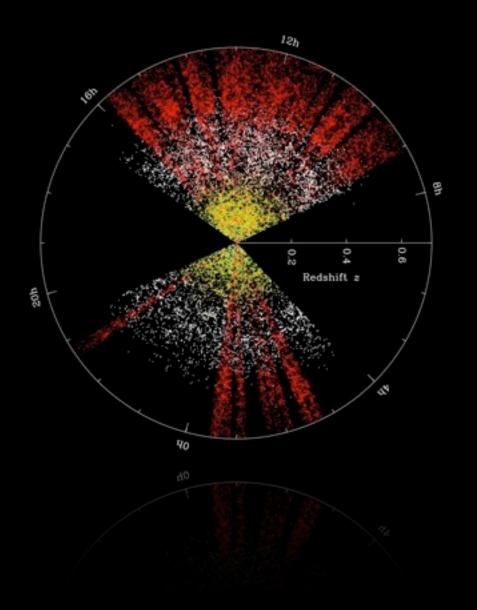


Be realistic... DEEP2 took over 100 nights on Keck

Cosmologist's desires are BAO at z~1

BAO requirements

- I. Volume! $V > 5 \text{ Gpc}^3 \text{ for } 5\% \sigma_w$
- 2. Sampling $2e10^{-4} < n < 4e10^{-4}$
- 3. $S/N \sim 1-3$



A problem marriage for BAO and galaxy formation

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BAO requirements ... for galaxy formation

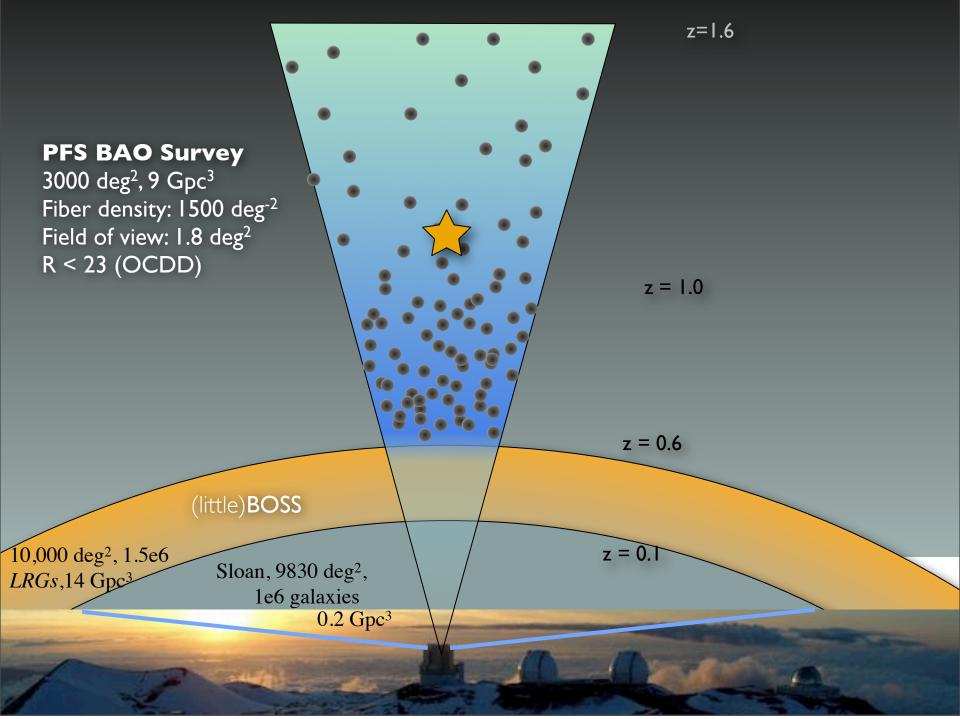
1. Volume! V > 5 Gpc<sup>3</sup> for 5% \sigma_w Way overkill (only want 0.2)

2. Sampling 2e \cdot 10^{-4} < n < 4e \cdot 10^{-4} Log M=11.3 (not 10.5), or not complete

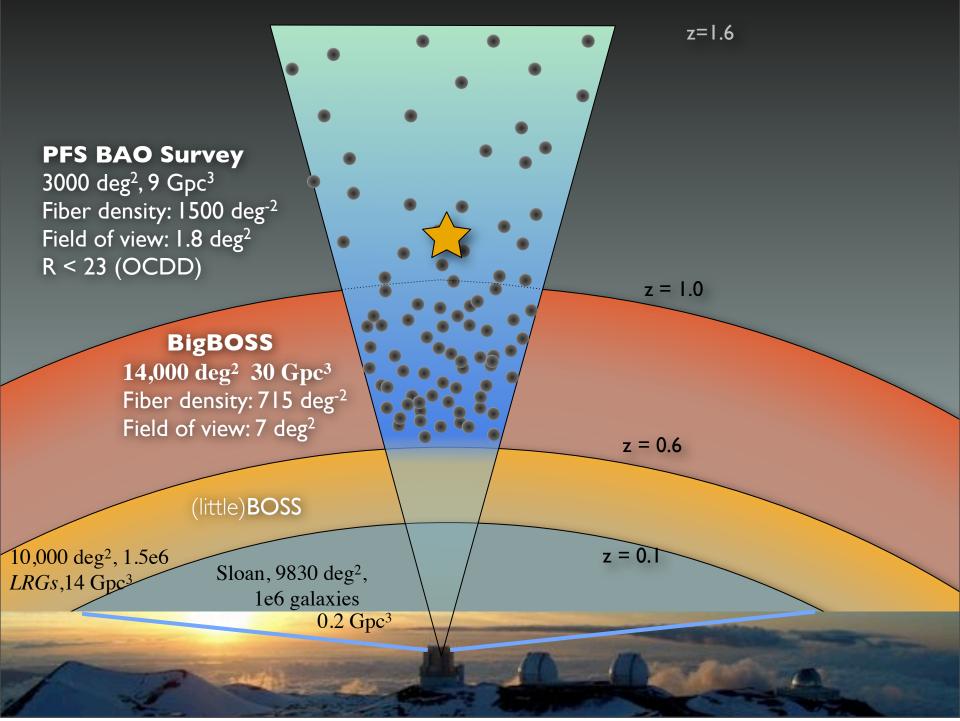
3. S/N ~ 1-3
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Let's choose both!

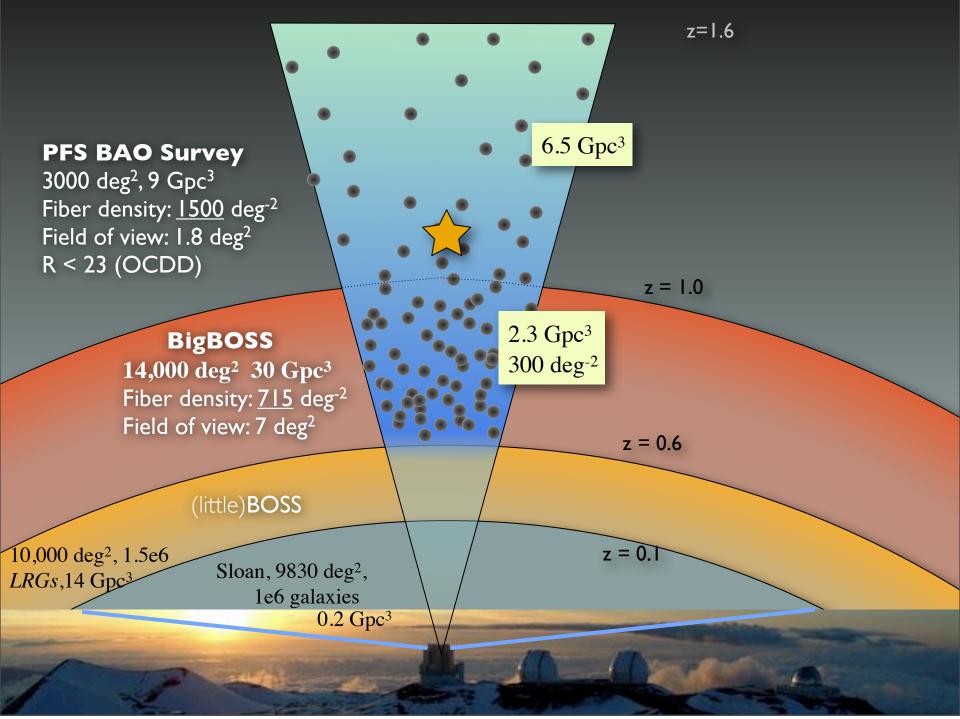
(need 10)



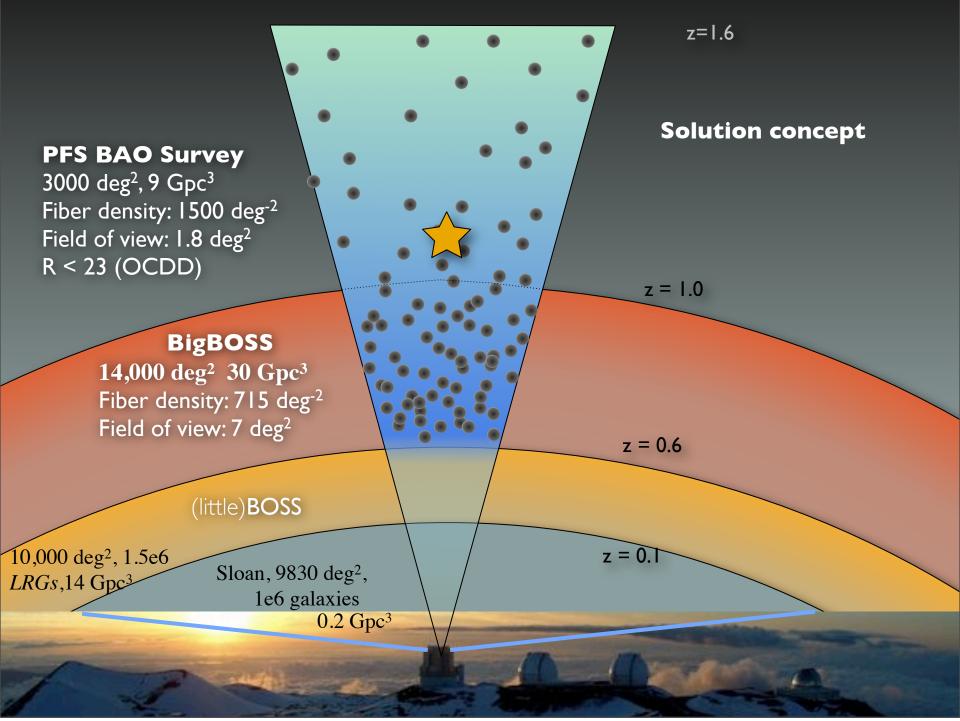
Friday, December 10, 2010



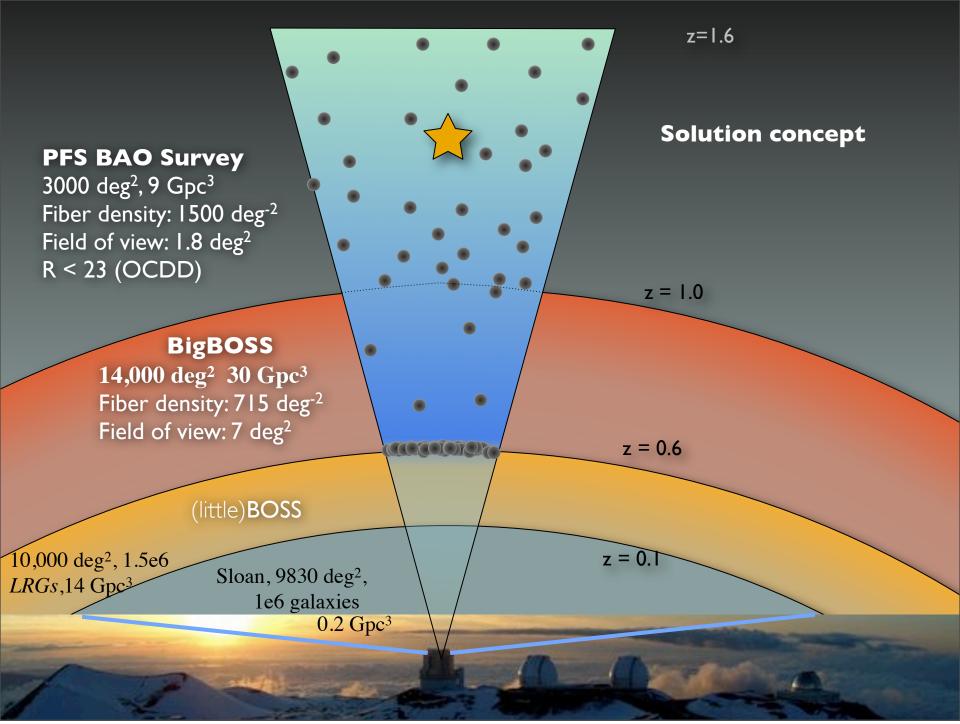
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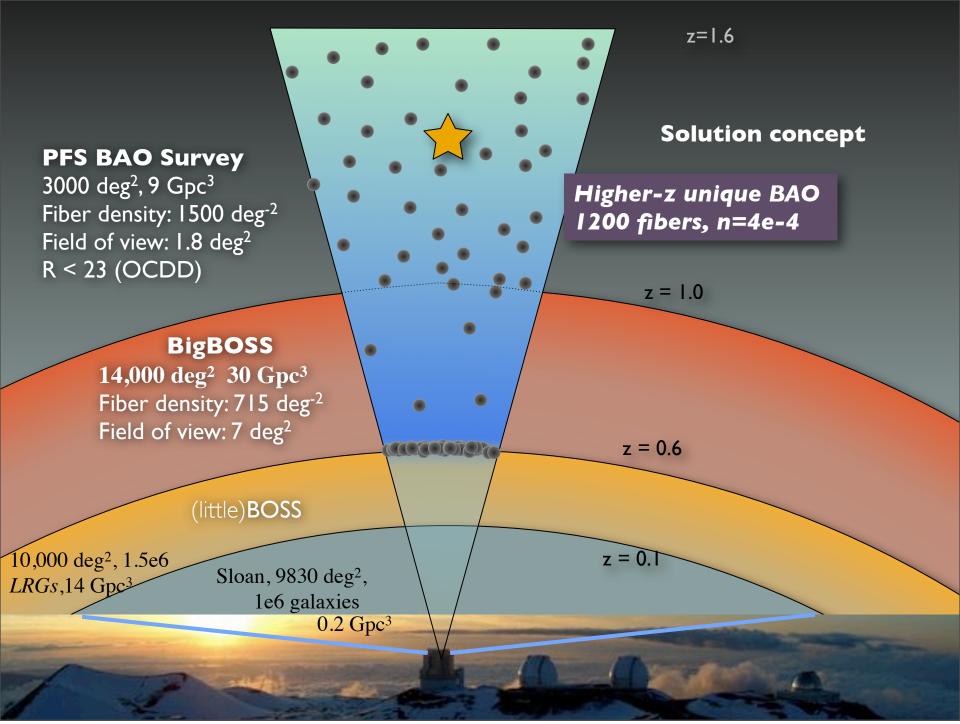
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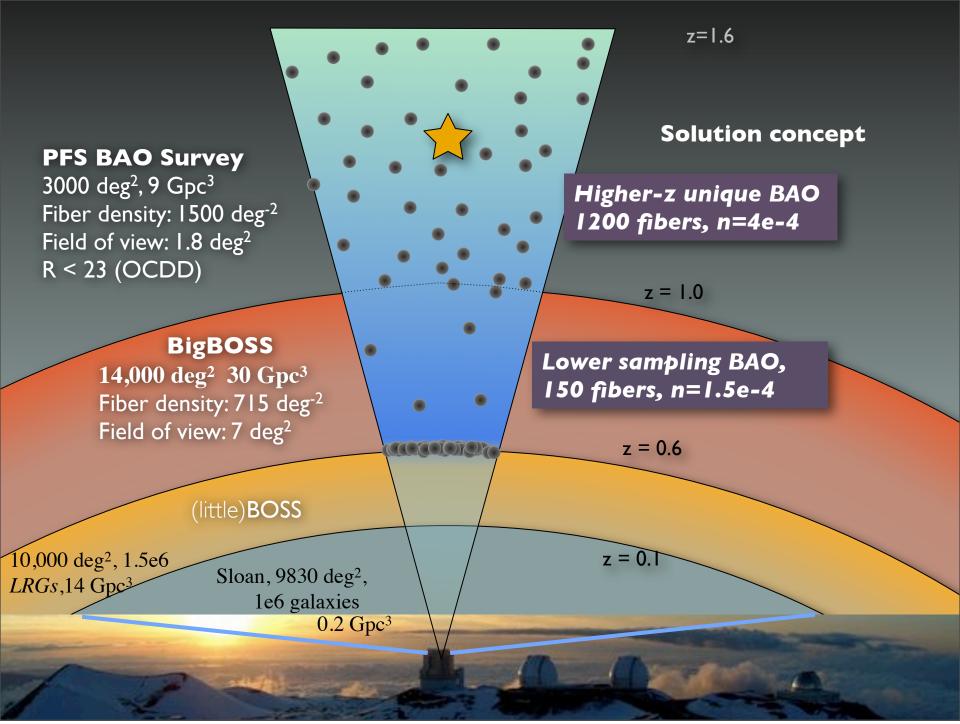
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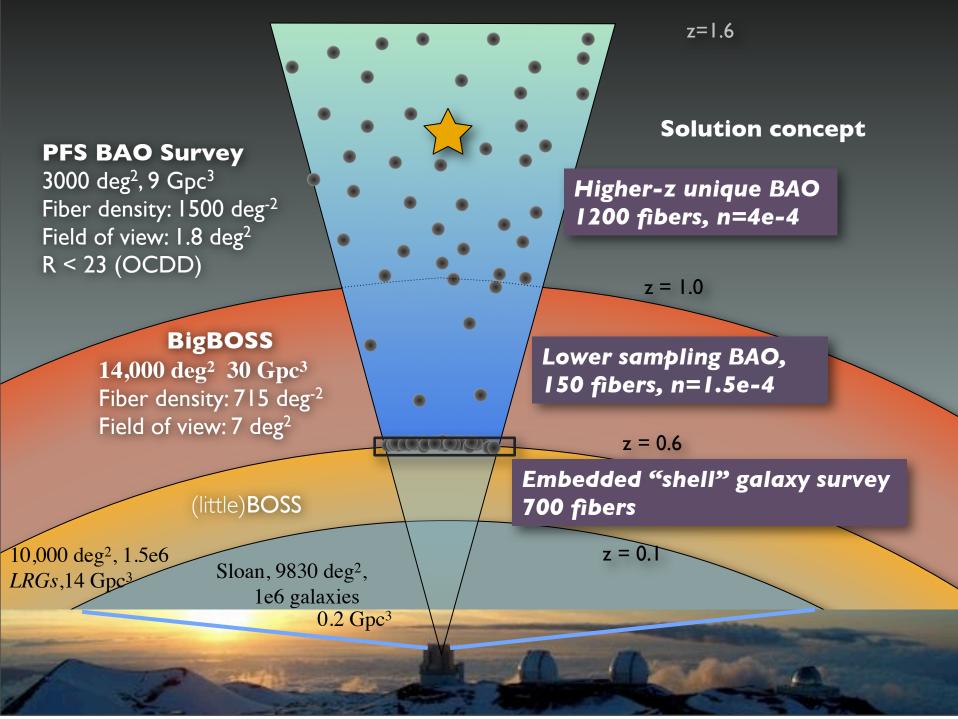
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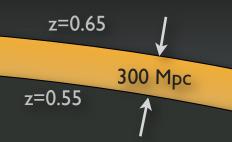


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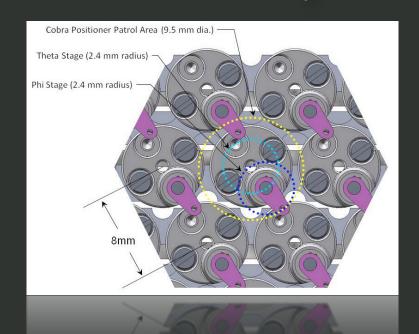
A mass-complete galaxy shell survey

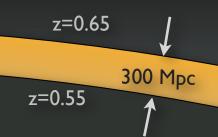
- Log M > 10.5, 0.2 Gpc 3
- 90% sampling rates (DEEP2 was 30%)
- 1% clustering at 100 kpc → unique territory and science
- Bright galaxies (i<22) S/N ~ 8-10 → velocity dispersions, chemistry
- Covers all key lines: Ha, NII, OIII, Hb, OII, etc. → AGN, SFH, balmer decrement
- 0.2 $Gpc^3 \rightarrow Vanishing cosmic variance, statistics on rare samples$

This is the galaxy survey Wish List!

How to do it? Survey Strategy...

- Aggressive fiber re-deployment, I minute bright targets: i<21, 5 min, S/N 10 fainter targets: 21< i < 22, 30 min, S/N 10
- 2500 shell galaxies per PFS pointing
- 4.5e6 shell galaxies over 3200 deg²
- Robust photo-z selection from HSC
 Synergy with HSC medium-band surveys!

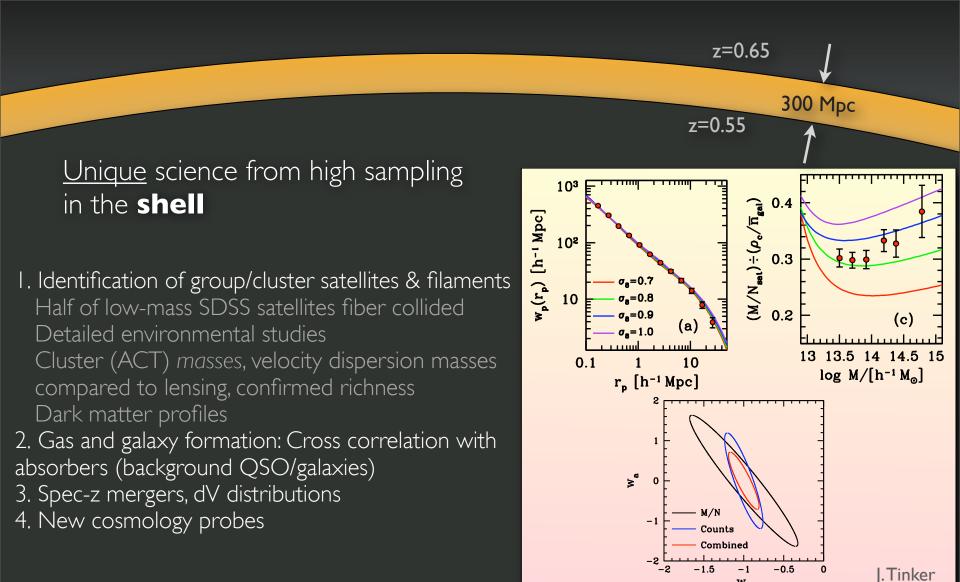


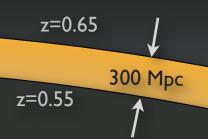


<u>Unique</u> science from high sampling in the **shell**

- I. Identification of group/cluster satellites & filaments
 Half of low-mass SDSS satellites fiber collided
 Detailed environmental studies
 Cluster (ACT) masses, velocity dispersion masses
 compared to lensing, confirmed richness
- 2. Gas and galaxy formation: Cross correlation with absorbers (background QSO/galaxies)
- 3. Spec-z mergers, dV distributions
- 4. New cosmology probes

Dark matter profiles





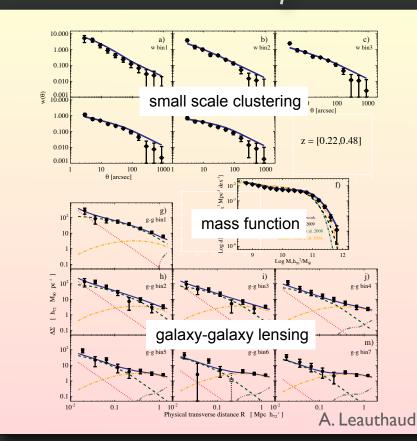
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Summary: Hybrid strategies to marry galaxy and BAO science.

- Exploit unique power of Subaru for BAO by concentrating on z > 1.0
- An embedded shell at z=0.6 enables breakthrough, highly sampled galaxy science. Area, depth, and fiber density well suited to BAO component.



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