Gravity localization & AdS4 / CFT3

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STRINGS 2012 (Munich)

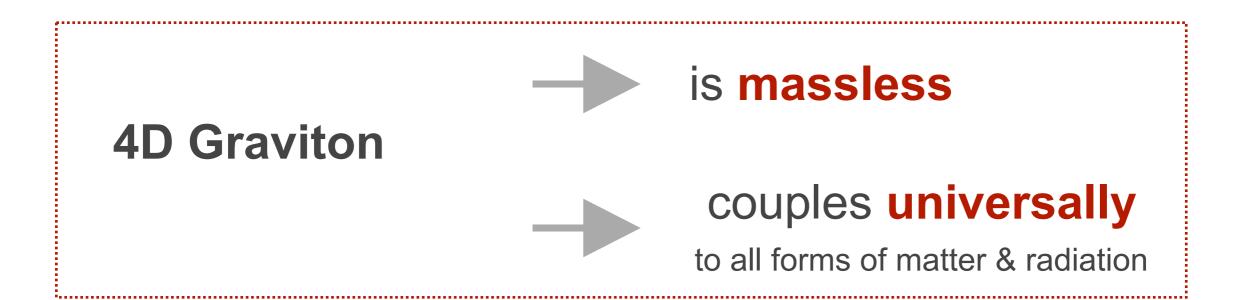
CB, J. Estes, arXiv:1103.2800 [hep-th]

B. Assel, CB, J. Estes, J. Gomis, 1106.4253 [hep-th]

1207.xxxx [hep-th]

based on:

Standard Hypothesis:

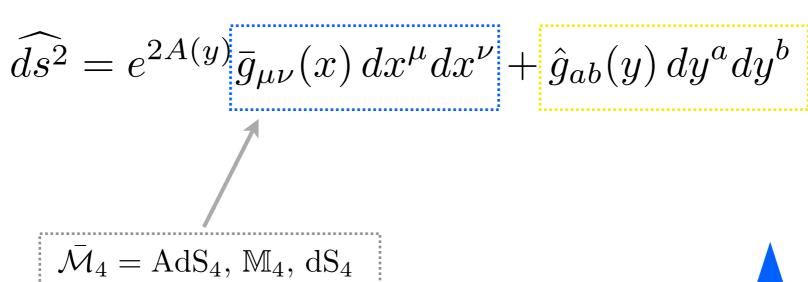


These properties are automatic in any compactification of string theory (or higher-Dim theory of gravity)

[Note: I will stick with classical 2-derivative supergravities]

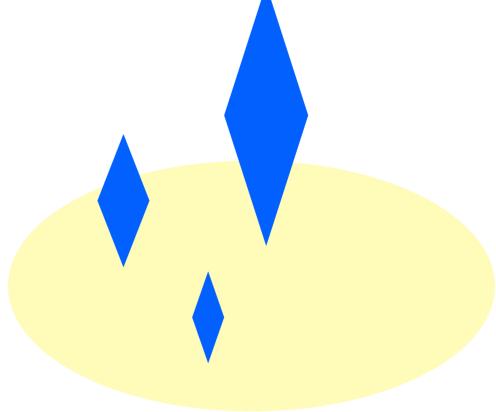
KK reduction for spin 2:

Consider warped-(A)dS vacuum,



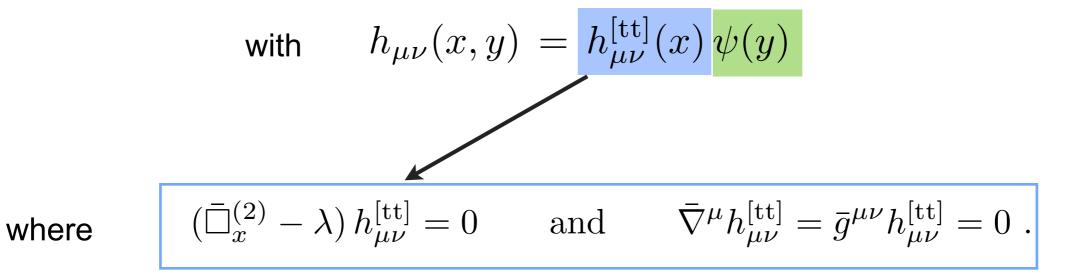
$$\bar{\mathcal{M}}_4 = \text{AdS}_4, \, \mathbb{M}_4, \, \text{dS}_4$$

$$k = -1, 0, 1$$



Consistent reduction of (spin-2) metric perturbations:

$$ds^{2} = e^{2A} \left(\bar{g}_{\mu\nu} + h_{\mu\nu} \right) dx^{\mu} dx^{\nu} + \hat{g}_{ab} dy^{a} dy^{b} ,$$



Pauli-Fierz equations $(\lambda = m^2 + 2k)$

Linearizing the Einstein equations $R_{MN}-\frac{1}{2}g_{MN}R=T_{MN}$

leads to a universal Schrödinger problem in the 6D transverse space:

depends only on geometry, not on "matter" fields

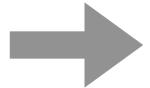
Brandhuber, Sfetsos; Csaki, Erlich, Hollowood, Shirman; CB, Estes

$$-\frac{e^{-2A}}{\sqrt{[\hat{g}]}} \left(\partial_a \sqrt{[\hat{g}]} \,\hat{g}^{ab} e^{4A} \partial_b\right) \psi = m^2 \psi$$

mass operator $\mathcal M$

Using the standard norm $\|\psi\|^2 = \int [dy] e^{2A} \psi^* \psi$ one finds :

$$\langle \psi, \mathcal{M}^2 \psi \rangle = \int [dy] e^{4A} \partial_a \psi^* \partial^a \psi$$



$$\mathcal{M}^2 \geq 0$$
 and $\mathcal{M}^2 = 0 \longrightarrow \psi_0 = constant$



A massless 4D graviton requires
$$\int [dy] \, e^{2A} < \infty \qquad \text{(automatic for smooth compactification)}$$



Its couplings are universal:
$$\int [dy] e^{2A} \psi_0 \phi_i^* \phi_j = \psi_0 \, \delta_{ij}$$

Can it be otherwise?

- 4D graviton must be massive
- transverse space non-compact

A model for this was proposed ten years ago by Karch & Randall:

Thin AdS₄ brane in AdS₅ bulk

Randall, Sundrum; Dvali, Gabadadze, Porrati;

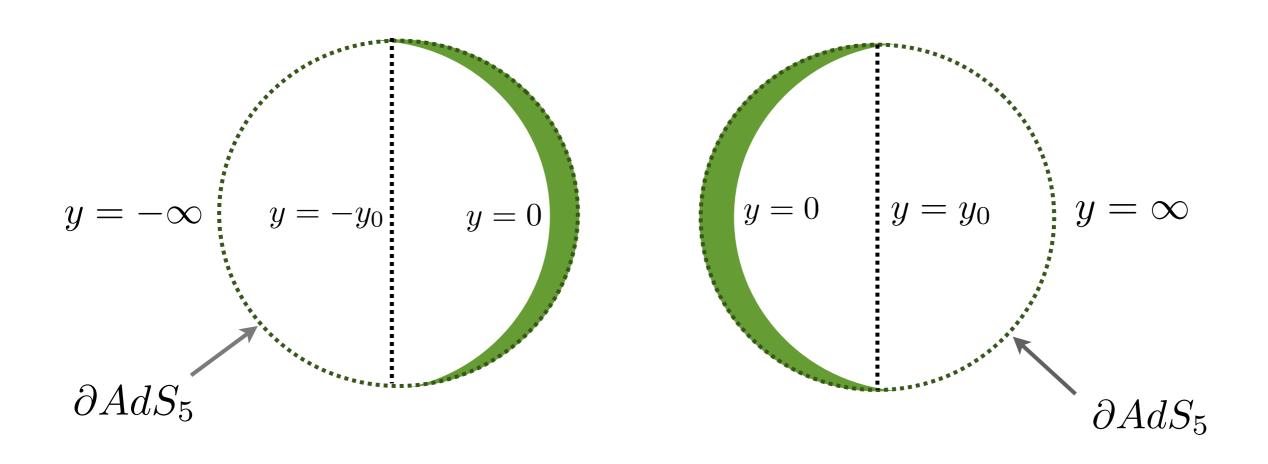
$$I_{\rm KR}\,=\,-\frac{1}{2\kappa_5^2}\int d^4x\,dy\,\sqrt{g}\left(R+\frac{12}{L^2}\right) + \lambda\int d^4x\,\sqrt{[g]_4}\,\,,$$
 AdS5 radius brane tension

$$ds^2 = L^2 \mathrm{cosh}^2 \left(\frac{y_0 - |y|}{L}\right) \; \bar{g}_{\mu\nu} dx^\mu dx^\nu + dy^2 \label{eq:ds2}$$
 AdS4 foliation

Brane radius of curvature:
$$\ell = L \cosh(y_0/L) = \frac{L}{\sqrt{1-\kappa_5^2 \lambda L/6}} \gg L$$

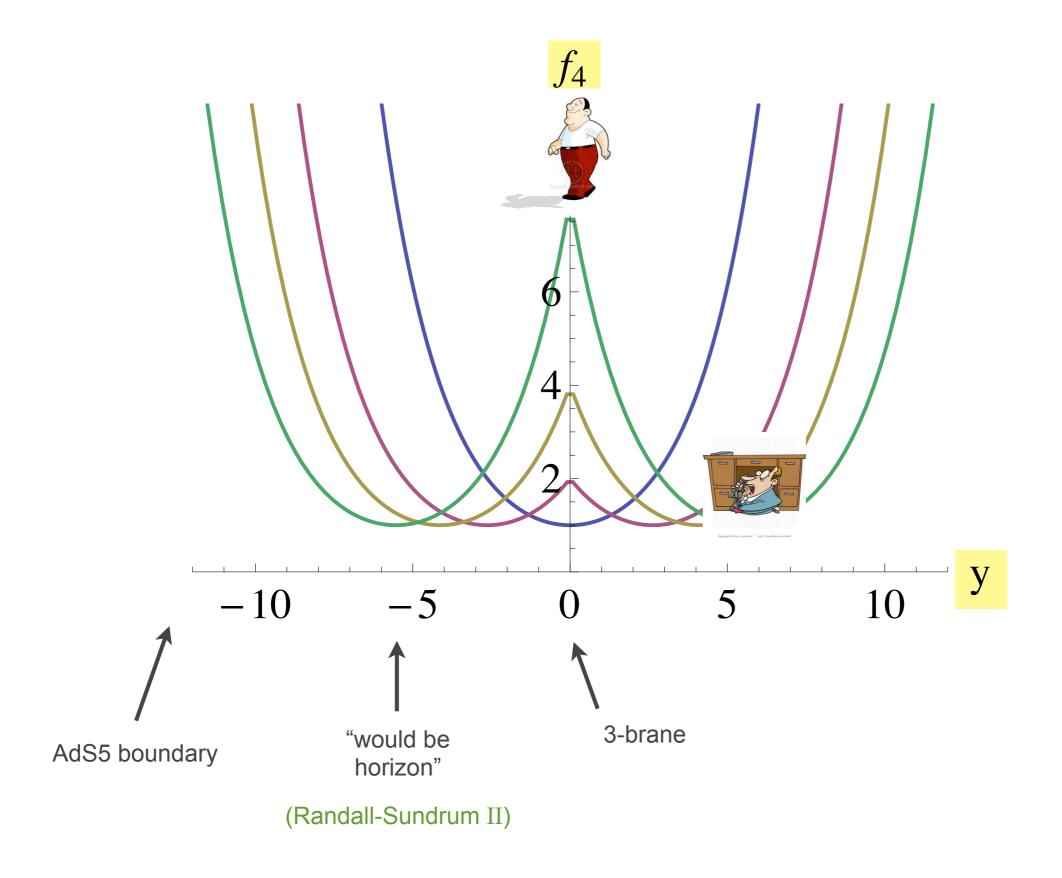
can be tuned so that

Two slices of AdS5 glued along a AdS4 brane:

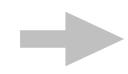


(cut away green slices and glue)

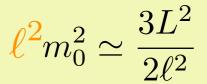
Warp factor as function of transverse coordinate:



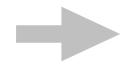
Spectrum:



one nearly-constant, nearly massless mode



vanishes near bottom of warp factor wells



two infinite towers of nearly AdS5 modes



$$\ell^2 m^2 \simeq (2n+1)(2n+4)$$
 $n = 0, 1, \cdots$

$$n=0,1,\cdots$$

supported at bottom of wells

Their wavefunctions are exponentially suppressed at the brane position, so they are hidden from 4D gravity:

$$\int [dy]e^{2A}\,\psi_0\phi_h^*\phi_h \ll \psi_0$$

No conventional effective 4D theory, but physics does look 4D at

$$L \ll r \ll \ell$$

Newton's law:

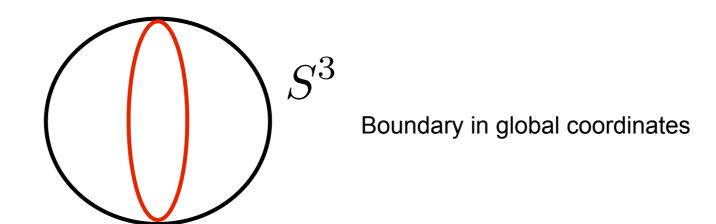
<u>Warning</u>: this is not a model for the Universe, since $\Lambda_{ ext{eff}} < 0$

$$\Lambda_{\rm eff} < 0$$

But any long-distance modification of Einstein gravity is (potentially) interesting

Can the KR model be embedded in string theory?

(holographic duals to conformal domain walls of CFT₄)



Two beautiful developments got us started:

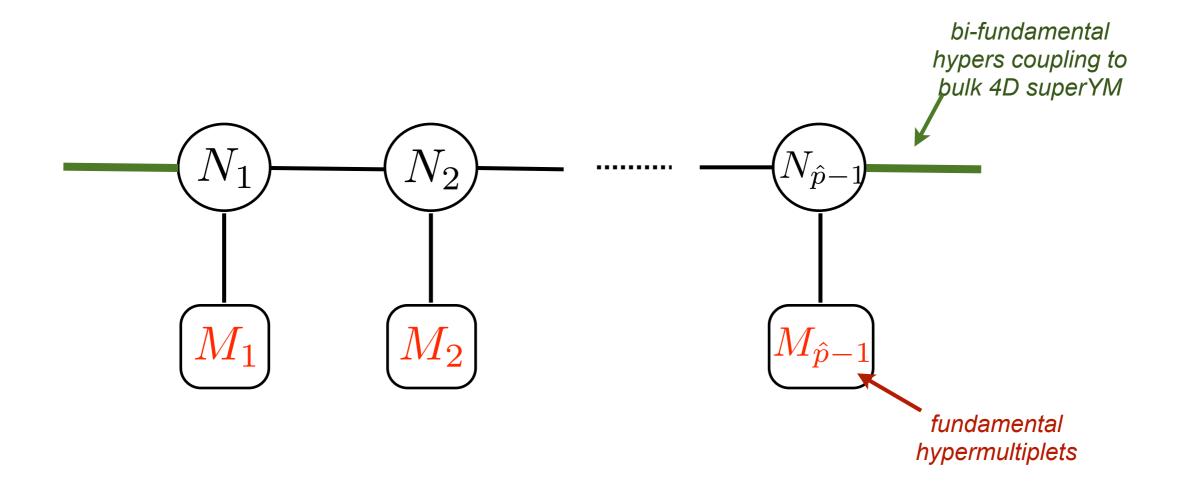
D'Hoker, Estes & Gutperle '07 found the general local form of type-IIB supergravity solutions with $OSp(2,2|4)\supset SO(2,3)\times SO(3)\times SO(3)$ supersymmetry.

This is the symmetry of conformal 1/2-BPS domain walls in N=4 D=4 super-Yang-Mills theory, which were classified by Gaiotto & Witten '08

DeWolfe, Freedman, Ooguri;

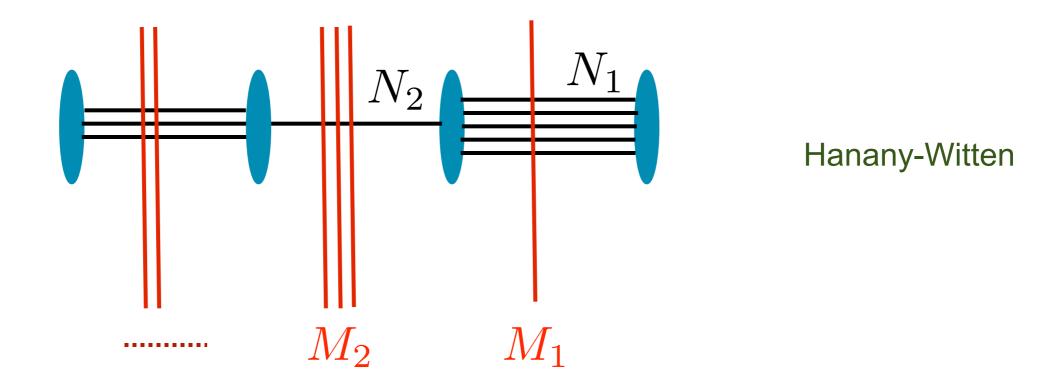
To address our question, we established the precise dictionary between these two works. Bonus: some new lessons for AdS4/CFT3.

The Gaiotto-Witten domain walls support linear-quiver N=4 CFT3s



$$U(N_1) imes U(N_2) imes \cdots imes U(N_{\hat p-1})$$
 gauge group $U(M_1) imes U(M_2) imes \cdots imes U(M_{\hat p-1})$ manifest global symmetry $U(\hat M_1) imes U(\hat M_2) imes \cdots imes U(\hat M_{p-1})$ hidden global symmetry

Engineered on D3-branes suspended/intersecting NS5- and D5- branes



Mirror symmetry exchanges NS5 with D5 (Intriligator, Seiberg; de Boer, Hori, Ooguri, Oz;)

(Almost) invariant description in terms of **linking numbers**:

```
i th D5 \ell_i = (net # of D3-branes ending on it from left) + (# of NS5 to its right)
\hat{l}_j = \text{(net # of D3-branes ending on it from right)} + \text{(# of D5 to its left)}
```

Focus on 3D gauge theories (coupling to 4D YM: slight complication)

Gaiotto-Witten conjectured that these **flow to non-trivial SCFT** whenever

 $\# \ hypers \geq 2N_a$ for each gauge group

(complete Higgsing)

These conditions translate to

$$\hat{\ell}_1 \ge \hat{\ell}_2 \cdots \ge \hat{\ell}_{\hat{p}} > 0$$

(labelled from right to left)

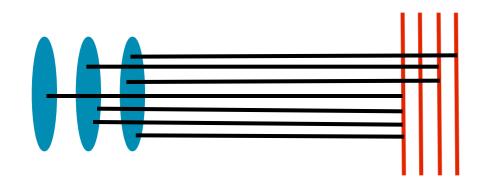
while also

$$\ell_1 \ge \ell_2 \cdots \ge \ell_p > 0$$

(labelled from left to right)

Move all NS5 to the left of all D5

$$\sum \ell_i = \sum \hat{\ell}_j = N$$

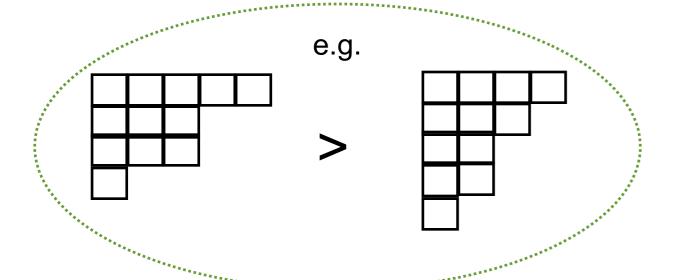


two partitions
$$\rho$$
 & $\hat{\rho}$

Supersymmetry (s-rule) implies the (partial ordering) constraints

$$\rho^T > \hat{\rho}$$

(# of boxes in first n columns of ho > # boxes in first n rows of $\hat{\rho}$ \forall η)



arise in many contexts related to Nahm's equations

Nakajima; Kronheimer; CB, Hoppe, Pioline;

Summarize: non-trivial N=4 SCFT3 $T_{
ho}^{\hat{
ho}}(SU(N))$

conjectured for any pair of partitions obeying these constraints

The IIB solutions are AdS4 x S2 x S2 fibrations over a Riemann surface basis Σ . Local solutions are determined by two harmonic functions h_1 and h_2 . Modulo an SL(2,R) rotation:

$$\underline{\text{metric}}: \quad ds^2 = f_4^2 ds_{\text{AdS}_4}^2 + f_1^2 ds_{\text{S}_1^2}^2 + f_2^2 ds_{\text{S}_2^2}^2 + 4\rho^2 dz d\bar{z} \; ,$$

$$f_4^8 = 16 \, \frac{N_1 N_2}{W^2} \, , \quad f_1^8 = 16 \, h_1^8 \frac{N_2 W^2}{N_1^3} \, , \quad f_2^8 = 16 \, h_2^8 \frac{N_1 W^2}{N_2^3}$$

$$\rho^8 = \frac{N_1 N_2 W^2}{h_1^4 h_2^4}$$

$$\underline{\text{dilaton}}: \quad e^{4\phi} = \frac{N_2}{N_1}$$

$$W = \partial h_1 \bar{\partial} h_2 + \bar{\partial} h_1 \partial h_2 = \partial \bar{\partial} (h_1 h_2) \; ,$$
 where
$$N_1 = 2h_1 h_2 |\partial h_1|^2 - h_1^2 W \; , \qquad N_2 = 2h_1 h_2 |\partial h_2|^2 - h_2^2 W \; .$$

There are also p-form backgrounds: F_5 , H_3 , F_3 expressed in terms of the dual harmonic functions:

$$h_1 = -i(\mathcal{A}_1 - \bar{\mathcal{A}}_1)$$
 \rightarrow $h_1^D = \mathcal{A}_1 + \bar{\mathcal{A}}_1$
 $h_2 = \mathcal{A}_2 + \bar{\mathcal{A}}_2$ \rightarrow $h_2^D = i(\mathcal{A}_2 - \bar{\mathcal{A}}_2)$

and the (normalized) volume forms of the (pseudo)spheres:

$$\omega^{0123}$$
 AdS_4 ω^{45} S_1^2 ω^{67} S_2^2

Note: the constant parts of $Re(A_1)$ and $Im(A_2)$ are pure gauge.

3-forms:
$$H_{(3)} + iF_{(3)} = \omega^{45} \wedge db_1 + i\omega^{67} \wedge db_2$$

$$b_1 = 2ih_1 \frac{h_1 h_2 (\partial h_1 \bar{\partial} h_2 - \bar{\partial} h_1 \partial h_2)}{N_1} + 2h_2^D$$

where:

$$b_2 = 2ih_2 \frac{h_1 h_2 (\partial h_1 \bar{\partial} h_2 - \bar{\partial} h_1 \partial h_2)}{N_2} - 2h_1^D$$

5-form:
$$F_{(5)} = -4 f_4^4 \omega^{0123} \wedge \mathcal{F} + 4 f_1^2 f_2^2 \omega^{45} \wedge \omega^{67} \wedge (*_2 \mathcal{F})$$
,

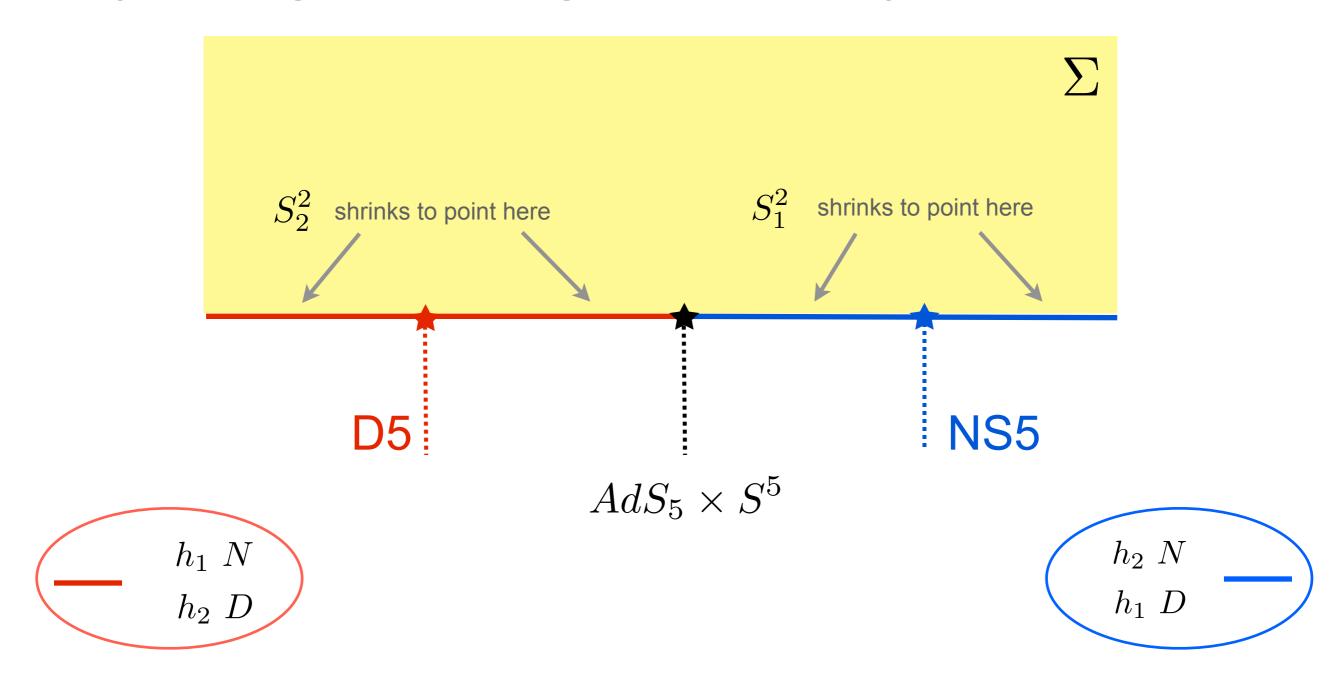
where

$$f_4^4 \mathcal{F} = dj_1$$
 with $j_1 = 3\mathcal{C} + 3\bar{\mathcal{C}} - 3\mathcal{D} + i\frac{h_1 h_2}{W} \left(\partial h_1 \bar{\partial} h_2 - \bar{\partial} h_1 \partial h_2\right)$

$$\partial \mathcal{C} = \mathcal{A}_1 \partial \mathcal{A}_2 - \mathcal{A}_2 \partial \mathcal{A}_1$$
 $\mathcal{D} = \bar{\mathcal{A}}_1 \mathcal{A}_2 + \mathcal{A}_1 \bar{\mathcal{A}}_2$

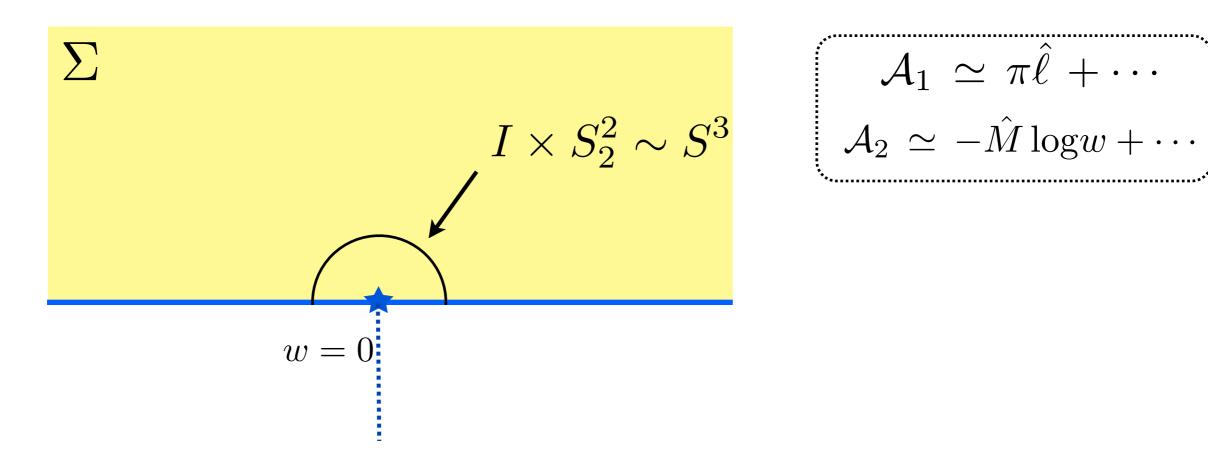
The surface Σ can have **boundaries** which are **interior points** of the 10D geometry. For this, require either h_1 to have a **Dirichlet** condition and h_2 a **Neumann** condition, or vice versa. One 2-sphere then shrinks to zero, and the geometry is (locally) $\sim AdS_4 \times S^2 \times \mathcal{D}^2$

Asymptotic regions arise as singularities on boundary:



Local form of **NS5 singularity**:

(For D5-brane: exchange roles of $\,{\cal A}_1\,$ and $\,{\cal A}_2\,$)



$$\mathcal{A}_1 \simeq \pi \hat{\ell} + \cdots$$

$$\mathcal{A}_2 \simeq -\hat{M} \log w + \cdots$$

$$\hat{M} = \frac{1}{16\pi^2} \int_{S^3} H_3 \qquad \longleftarrow \quad \text{\# NS5 branes}$$

Local form of $AdS_5 \times S^5$:

$$\sum I \times S_1^2 \times S_2^2 \sim S^5$$

$$w = 0$$

$$\mathcal{A}_1 = \frac{1}{\sqrt{w}}(a_1 + b_1 w + \cdots)$$

$$\mathcal{A}_2 = \frac{1}{\sqrt{w}}(a_2 + b_2 w + \cdots)$$

$$n = \frac{1}{(4\pi)^4} \int_{S^5} F_5 = \frac{(a_2b_1 - a_1b_2)}{2\pi}$$

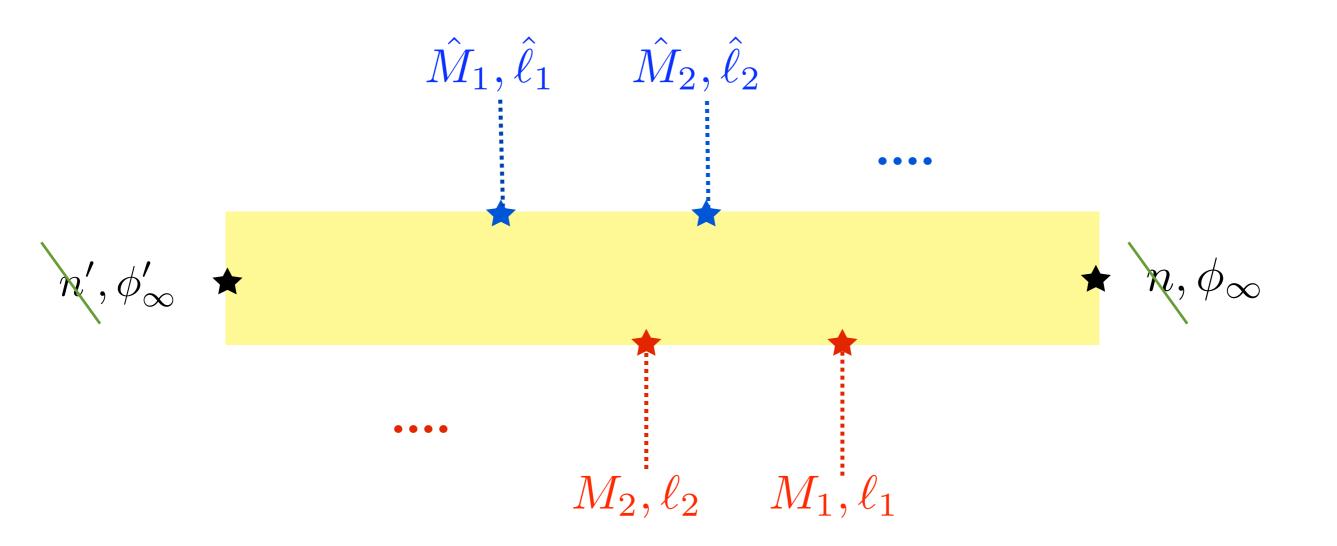
← D3 charge

Limit $a_1, a_2 \rightarrow 0$ gives a smooth capping-off of geometry

Solution for **linear quivers**

$$h_1 = \left[-i\alpha \sinh(z - \beta) - \sum_{a=1}^q \gamma_a \ln\left(\tanh\left(\frac{i\pi}{4} - \frac{z - \delta_a}{2}\right)\right) \right] + c.c.$$

$$h_2 = \left[\hat{\alpha} \cosh(z - \hat{\beta}) - \sum_{b=1}^{\hat{q}} \hat{\gamma}_b \ln\left(\tanh\left(\frac{z - \hat{\delta}_b}{2}\right)\right) \right] + c.c.$$



Parameter count:

$$\{M_a, \ell_a, \hat{M}_b, \hat{\ell}_b\} \longleftrightarrow \{\gamma_a, \delta_a, \hat{\gamma}_b, \hat{\delta}_b\}$$

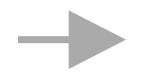
May have **zero**, **one** or **two** asymptotic $AdS_5 \times S^5$ regions, more leads to conical singularities.

All parameters in first case are discrete: no N=4 moduli.



Global symmetries of CFT3 realized as gauge symmetries on 5-branes





Can show that linking numbers define two partitions obeying the **Young-tableau inequalities**



One-to-one correspondence with Gaiotto-Witten SCFT3s

Solution for circular quivers

 $\frac{i\pi}{2}$ C $q^{\frac{1}{4}}$

$$\mathcal{A}_{1} = -i \sum_{a=1}^{p} \gamma_{a} \ln \left(\frac{\vartheta_{1} (q | \nu_{a})}{\vartheta_{2} (q | \nu_{a})} \right) + \varphi_{1} \qquad i \nu_{a} = -\frac{z - \delta_{a}}{2\pi} + \frac{i}{4}$$

$$\mathcal{A}_{2} = -\sum_{b=1}^{\hat{p}} \hat{\gamma}_{b} \ln \left(\frac{\vartheta_{1} (q | \hat{\nu}_{b})}{\vartheta_{2} (q | \hat{\nu}_{b})} \right) + i \varphi_{2} \qquad i \hat{\nu}_{b} = \frac{z - \hat{\delta}_{b}}{2\pi}$$

Annulus modulus

$$\longleftrightarrow$$

K = # of winding D3 branes

Cannot couple to 4D sYM, i.e. add asymptotic $\,AdS_5 imes S^5\,\,\,\,\,$ regions

Several calculations/observations concerning this rich set of holographic dualities, that I have no time to discuss here. Just mention some:

Free energy on S3 from sugra = from CFT:

checked that for T(SU(N)) and some related theories

$$F = \frac{1}{2}N^2 \ln N + \mathcal{O}(N^2)$$

non-trivial

Assel, Estes, Yamazaki

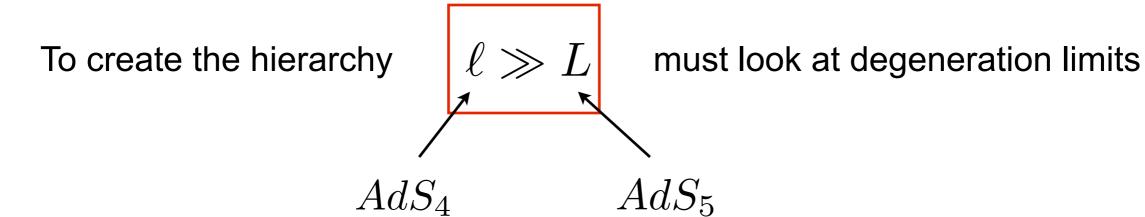
using results of Benvenuti & Pasquetti;
... Nishioka, Tachikawa, Yamazaki

Large K corresponds to $t \to 0$; T-dual to ABJM-like theories.

SL(2, Q) transformations generalize mirror symmetry in classical limit

.....

Let's return now to the original question



Factorize 5-brane singularities $\}$ "Worm-branes" (Janus throats)

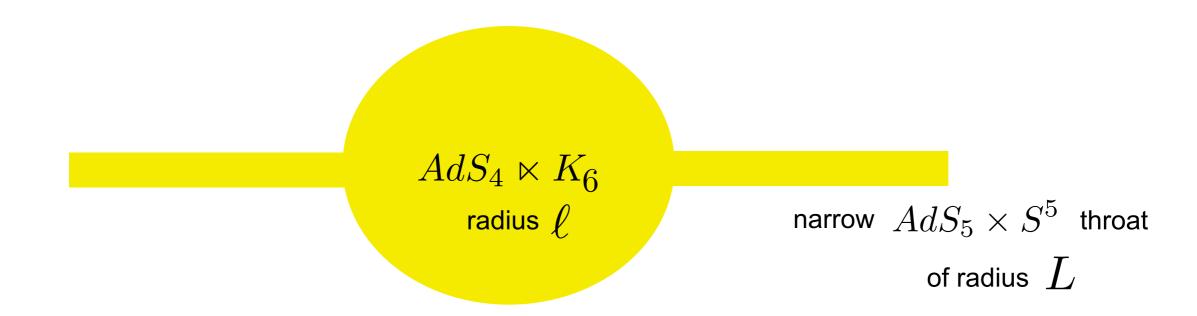
Cap-off $AdS_5 \times S^5$ (small rank for bulk D=4 SYM group)

$$\hat{\alpha}, \hat{\alpha} \rightarrow 0$$

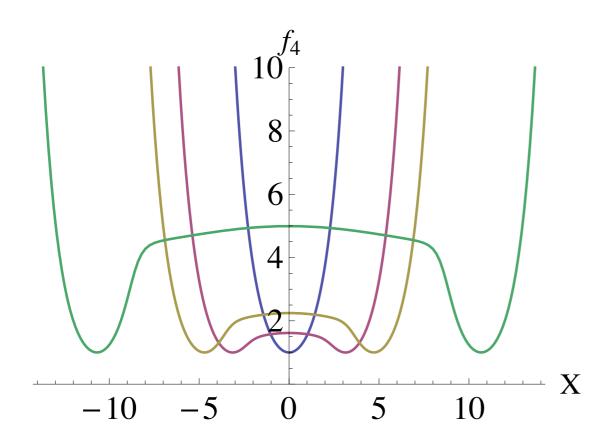
This is possible, provided there are both NS5s and D5s (to stabilize dilaton);
The lowest AdS4 spin-2 mode has a (tiny) mass and non-universal couplings.

(Assel, CB, unpublished)

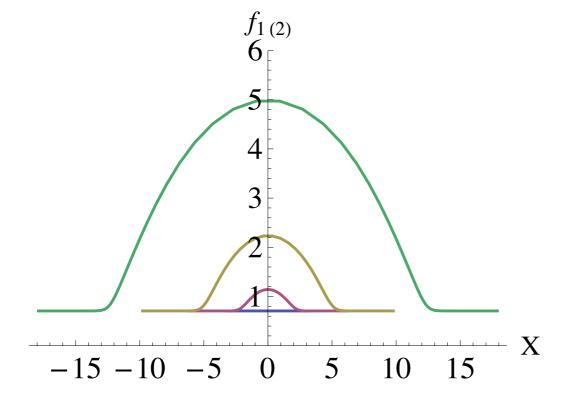
But, as in other efforts to create classical hierarchies, spacetime decompactifies



Gravity is never 4-dimensional!







sphere radii

What we need:

An AdS4/CFT3 pair for which the compact space stays small ($T_{\mu
u}$ and few other operators, then large gap)

interesting, and hard problem by itself

Couple CFT3 to SYM4 without destabilizing fixed point

need understanding of geometric singularities

Summary

Found a large class of type-IIB backgrounds, that are duals of the Gaiotto-Witten N=4 SCFT3s, and of their circular-quiver extensions.

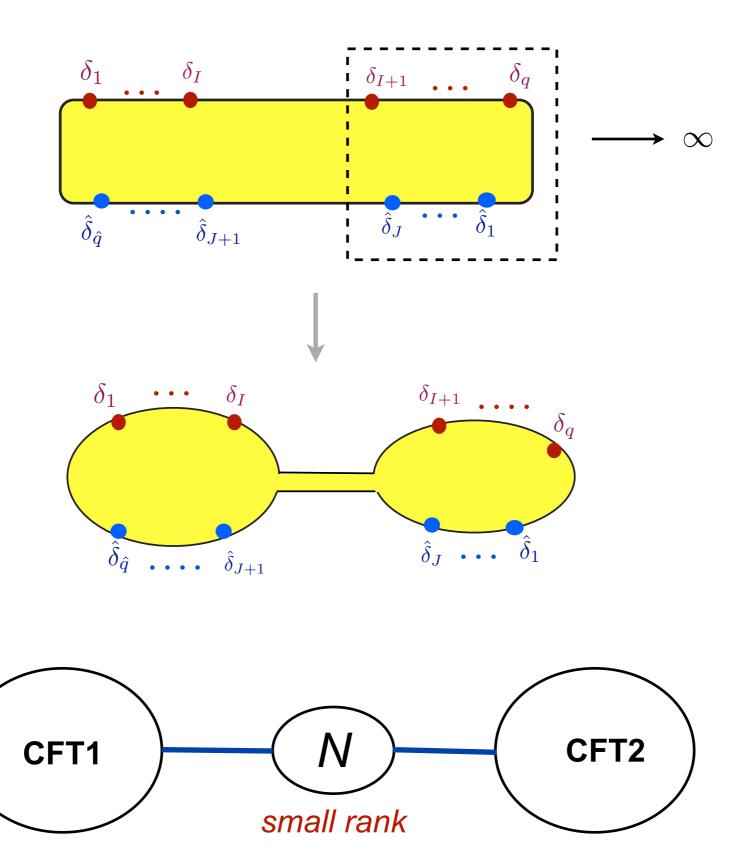
[These are AdS4 compactifications with localized 5-branes]

Degeneration limits correspond to couplings via small-rank gauge groups.

Failed to localize 4D gravity, because the brane fattens to 10D spacetime.

Given the difficulty in modifying gravity in IR, worth pursuing.....

as opposed to local multitrace couplings, e.g. Kiritsis, Niarchos; Aharony, Clark, Karch



$AKTI\Sigma \ AE\Lambda IOY \ KA\Lambda\Lambda I\Sigma TON \ \Phi AO\Sigma$

A ray of sun is the best of lights

(chorus in Antigone of Sophocles)