de Sitter vacua in string theory.

Liam McAllister Cornell Strings 2019, Brussels

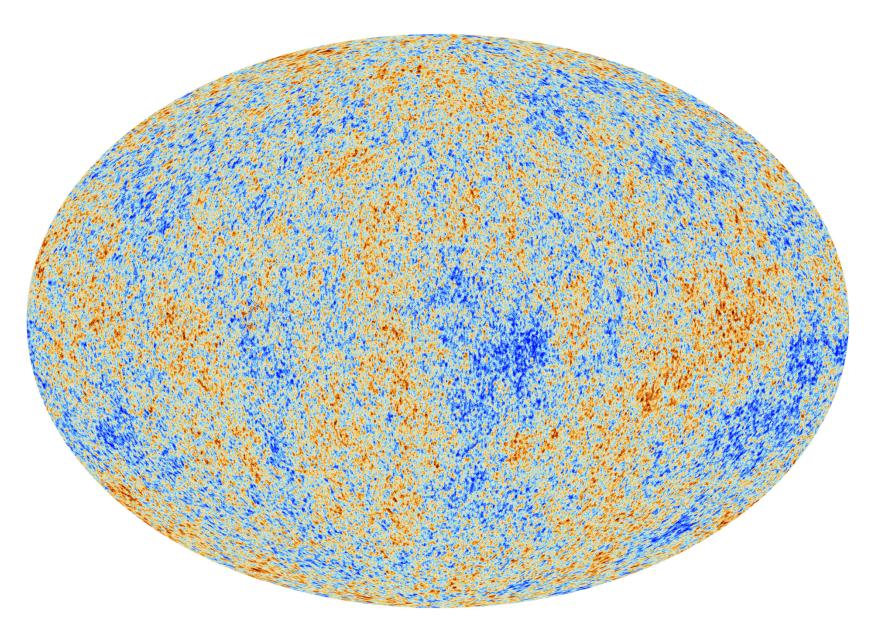
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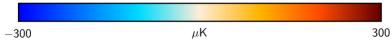
Byrne Observatory at Sedgwick Reserve and the Palomar Transient Factory | LCOGT

Why de Sitter space?

Simplest explanation for cosmological observations. Foundation of concordance model of cosmology.









Why de Sitter space?

Measurements of supernovae, the cosmic microwave background, and large-scale structure give overwhelming evidence for

two stages of accelerated expansion:

early:	$t \ll 1 s$	Inflation
late:	$t \gtrsim 10^{17} s$	Dark energy

Simplest explanation: (quasi) de Sitter space

 ΛCDM + inflation consistent with all observations.

Evidence goes far beyond homogeneous expansion.

Why in string theory?

Is string theory compatible with the concordance cosmology?

Why in string theory?

Dark energy:

- Cosmological constant problem.
- Does string theory have enough de Sitter vacua, $\gg 10^{120}$, to accommodate anthropic "solution"?
- Does it admit a better solution?

Inflation:

- CMB signatures depend on Planck-scale physics. Dramatic for primordial tensors, but holds generally. Understanding inflation requires quantum gravity.
- Powerful link from string theory to observations.



I. Overview of task II. KKLT scenario III. Other constructions

Focus: constructions of dS₄

Quantum gravity in de Sitter space is challenging.

Today: review status of constructions of dS_4 in EFTs derived from compactifications of superstrings.

Require $\mathcal{R}_4 \alpha' \ll 1$, derive and use EFT below string scale.

Possible worry: are EFTs meaningful in quantum gravity? Banks 00, 03 Banks, Dine, Corbatov 03

Will assume that vacua of an EFT derived from string theory descend from vacua of full string theory.

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10d supergravity + localized sources \Rightarrow 4d (super)gravity

$$S_{10} = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{G} \left(\mathcal{R}_{10} + \dots \right)$$

$$\Rightarrow S_4 = \int d^4x \sqrt{g} \left(\frac{M_{\rm pl}^2}{2} \mathcal{R}_4 + \mathcal{L}(\phi) \right)$$

Task: derive $\mathcal{L}(\phi)$

Strategy for finding de Sitter

Conceptually easy, impractical at present:

Find solution preserving $\mathcal{N} = 1$ SUSY, e.g. type II on CY_3 - \mathcal{O} . Directly compute 4d EFT to N^kLO, in α' and g_s . Exhibit de Sitter solutions in EFT at N^{k-1}LO, show that N^kLO negligible.

Practical: apply further approximations. Look under special lampposts.

e.g. parameter regimes where sectors decouple into 'modules' that interact weakly.

Analyze modules in isolation, then weakly couple them.

At present, many such modules well-understood. Final 'assembly' is the key challenge.

Status of constructions

Many scenarios for constructing de Sitter vacua in EFTs derived in compactifications of string theory.

No incontrovertible example to date. But, many highly nontrivial tests already passed, and remaining difficulties appear purely technical.

Challenge is not unique to de Sitter vacua.

e.g., no derivation of Standard Model + Einstein gravity (with no light moduli) to standard demanded for de Sitter.

Non-supersymmetric compactifications are hard, but this does not imply they do not exist.

KKLT de Sitter vacua

Compactification of type IIB on an orientifold X of a $\mathrm{CY}_3,$ including:

three-form flux $G_3 = F_3 - \tau H_3 \in H_3(X, \mathbb{Z})$ an $\mathcal{N} = 1$ pure SYM sector on $N_c > 1$ D7-branes a warped deformed conifold region (structure) containing one or more anti-D3-branes

Claim [KKLT]: in a suitable parameter regime, these sources can yield metastable dS_4 , and corrections to approximations are small. Large Volume Scenario (LVS): different parameter regime, crucially including $\alpha'^3 \mathcal{R}^4$ correction to \mathcal{K}



Type IIB string theory compactified on O3/O7 orientifold, X, of a CY_3 .

 $ds^{2} = G_{AB}dX^{A}dX^{B} = e^{-6u(x)+2A(y)}g_{\mu\nu}dx^{\mu}dx^{\nu} + e^{2u(x)-2A(y)}g_{ab}dy^{a}dy^{b}$ Take $h^{1,1}_{+} = 1$, $\Sigma \in H_{4}(X,\mathbb{Z})$. Choose $G_{3} = F_{3} - \tau H_{3} \in H_{3}(X,\mathbb{Z})$. Moduli: axiodilaton $\tau := C_{0} + ie^{-\phi}$ complex structure ζ_{a} , $a = 1, \dots h^{2,1}$ Kähler: $T = e^{4u} + i \int_{\Sigma} C_{4}$.

 $4d \mathcal{N} = 1$ supergravity:

$$W_{ ext{classical}} = \int_X G_3 \wedge \Omega$$
 Given, Value, Witten 99
 $\mathcal{K} = -3\log(T + \overline{T}) - \log(-i(\tau - \overline{\tau})) - \log(-i\int_X \Omega \wedge \Omega)$

For generic G_3 , solutions of $D_{\tau}W = D_{\zeta_a}W = 0$ are isolated $\Rightarrow \tau, \zeta_a$ fixed. Giddings, Kachru, Polchinski Ol Kachru, Kallosh, Linde, Trivedi (

$$\begin{aligned} & \textbf{Setup} \\ \text{Below the scale } m_{\zeta_a} \sim \frac{\alpha'}{\sqrt{\text{Vol}(X)}} \text{ of the complex structure moduli masses} \\ & \mathcal{K} = -3\log(T + \overline{T}) \qquad W \rightarrow \left\langle \int_X G_3 \wedge \Omega \right\rangle =: W_0 \\ & D_T W = K_T W_0 = -\frac{3}{T + \overline{T}} W_0 \neq 0 \\ & \text{SUSY broken, but} \\ & V_F = e^{\mathcal{K}} \left(K^{T\overline{T}} D_T W_0 \overline{D_T W_0} - 3W_0 \overline{W_0} \right) = 0 \end{aligned}$$

Consider a stack of D7-branes on Σ that support pure $SU(N_c)$ SYM.

$$W_{\rm np} = -\frac{N_c}{32\pi^2} \langle \lambda \lambda \rangle = \mathcal{A} \, e^{-\frac{2\pi}{N_c}T} \qquad W = W_0 + \mathcal{A} \, e^{-\frac{2\pi}{N_c}T} \qquad \qquad \text{[or, ED3 on 2]}$$

This supergravity theory has a SUSY AdS_4 minimum. If $W_0 \ll 1$, the minimum is at large volume, and $m_T \ll m_{\zeta_a}$. $W_0 \ll 1$ occurs for some of the $\mathcal{O}(e^{2b_3})$ choices of G_3 . Ashok, Douglas 63 Gryavets, Kachru, Tripathy, Trivedt 63 Denef, Douglas 64



Introduce p anti-D3-branes at the tip of a Klebanov-Strassler throat.

 $S_{\rm DBI+CS} \Rightarrow V_{\overline{D3}} = 2p \, T_3 e^{4A} e^{-8u}$ Kachru, Pearson, Verlinde Ol

In a noncompact throat, this gives a metastable SUSY-breaking state.

In a **compact** throat, anti-D3-branes affect EOM of T. and *can* affect D7-brane $\langle \lambda \lambda \rangle$.

To the extent that the antibranes affect the gaugino condensate only by contributing to the EOM for T,

it follows that $V_{\text{tot}} \approx V_{\overline{D3}} + V_F$.

This theory has a metastable dS_4 minimum.

KKLT

MODULI STABILIZATION

Do there exist consistent global models with:

- i. Quantized fluxes giving small classical superpotential W_0
- ii. D7-brane stack(s) supporting gaugino condensate
- iii. Klebanov-Strassler throat region

ANTIBRANE UPLIFTING

Can anti-D3-branes be described with a SUSY action? Is decompactification the only important instability from anti-D3-branes? Do anti-D3-branes and D7-brane $\langle \lambda \lambda \rangle$ interact weakly? Can the de Sitter solution be described in ten-dimensional supergravity?

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Can one exhibit an explicit and fully-controlled compactification that unifies all necessary components?

MODULI STABILIZATION

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Explicit CY_3 orientifolds:

e.g., $h^{2,1} = 3$, $h^{1,1} = 51$: resolution of $T^6/\mathbb{Z}_2 \times \mathbb{Z}_2$ 12 D7/O7 stacks, pure $\mathcal{N} = 1$ SYM, gauge group $SO(8)^{12}$ 48 exceptional divisors $D_{\alpha} \cong \mathbb{P}^1 \times \mathbb{P}^1, h^{0,i} = 0$. each unique in homology class, and $h^3(\widehat{D}_{\alpha}) = 0 \Rightarrow W_{np} \neq 0$ $W = W_0 + \sum_{i=1}^{12} e^{-2\pi T_i/6} + \sum_{\alpha=1}^{48} e^{-2\pi T_{\alpha}}$

Corrections: higher instantons in W, $\lesssim 10^{-3}$; in K, $\lesssim 10^{-2}$

 $\alpha'^3 \mathcal{R}^4, \quad \lesssim 10^{-2}$

ANTIBRANE UPLIFTING

Can anti-D3-branes be described with a SUSY action? yes

Ferrara, Kallosh, Linde 14 Kallosh, Wrase 14 Kallosh, Quevedo, Uranga 15 Bergshoeff, Dasgupta, Kallosh, Van Proeyen, Wrase 19 Bandos, Heller, Kuzenko, Martucci, Sorokin 16 Vercnocke, Wrase 16 Dall'Agata, Dudas, Farakos 16 Kallosh, Vercnocke, Wrase 16 Aalsma, Van Der Schaar, Vercnocke 17 Garcia del Moral, Parameswaran, Quiroz, Zavala 17 Aalsma, Tournoy, Van Der Schaar, Vercnocke 18

Using constrained multiplets, can write complete supergravity action.

Volkov, Akulov 72 Cribiori, Roupec, Wrase, Yamada 19 Komargodski, Seiberg 09

SUSY spontaneously broken.

ANTIBRANE UPLIFTING

Kachru, Pearson, Verlinde Ol

Is decompactification the only important instability from anti-D3-branes? yes

Bena, Grana, Halmagyi 09 Bena, Giecold, Graña, Halmagyi, Massai 11 Dymarsky 11 Bena, Graña, Kuperstein, Massai 12 Bena, Graña, Kuperstein, Massai 14 Michel, Mintun, Polchinski, Puhm, Saad 14 Blåbäck, Danielsson, Junghans, Van Riet, Vargas 14

Singularities in fluxes near anti-D3-branes?

Yes, if ansatz too restrictive and excludes puffing up into NS5. Cohen-Maldonado, Diaz, Van Riet, Vercnocke 15

Study beyond probe, and at finite temperature, matches KPV.

Armas, Nguyen, Niarchos, Obers, Van Riet 18

"We regard this as very strong evidence for the existence of the meta-stable states, since by now they have been argued in rather complementary ways.

armas, Nguyen, Niarchos, Obers, Van Riet 18

de Sitter vacua from 10d

Can the de Sitter solution be described in ten-dimensional supergravity? yes

KKLT: after dimensional reduction of D3-branes and D7-brane $\langle \lambda \lambda \rangle$,
the EOM of the 4d EFT lead to a dS4 vacuum.Possible complaint: "this is not a 10d construction"Maldacena, Maldacena, Maldacen

$$\frac{1}{4}M_{\rm pl}^2 \mathcal{R}_4[g] = \underbrace{2T_3 e^{-12u} e^{4A}(z_{\overline{D3}})}_{V_{\overline{D3}}} - \frac{1}{4}\int_X \sqrt{g_6} e^{-4A} g^{\mu\nu} T_{\mu\nu}^{\lambda\lambda}$$

Task: compute $T^{\lambda\lambda}_{\mu\nu}$ in 10d. Is it such that $\mathcal{R}_4[g] \to \mathcal{R}_4[g] \Big|_{\text{KKLT}}$?

Moritz, Retolaza, Westphal 17, 18; Hamada, Hebecker, Shiu, Soler 18, 19; Gautason, Van Hemelryck, Van Riet 18 Gautason, Van Hemelryck, Van Riet, Venken 19; Carta, Moritz, Westphal 19; Kachru, Kim, L.M., Zimet 19

Consider a stack of D7-branes supporting $SU(N_c)$ SYM.

$$S_{G\lambda\lambda}=-rac{i}{(4\pi^2lpha')^2}\int\sqrt{g}e^{-4A}e^{\phi}G_3\cdot\Omegarac{ar{\lambda}ar{\lambda}}{16\pi}\delta^{(0)}+c.c.$$
 Cánara, Ibáñez, Uranga 03

 G_3 causes D7-brane gaugino mass $\Leftrightarrow \langle \lambda \lambda \rangle$ sources G_3 .

Strong consistency check: potential for position z of a D3-brane.

Dewolfe, Giddings 02 Kachru, Kallosh, Linde, Maldacena, L.M., Trivedi 03 Berg, Haack, Körs 04 Baumann, Dymarsky, Klebanov, Maldacena, L.M., Murugan 06 Koerber, Martucci 07 Baumann, Dymarsky, Kachru, Klebanov, L.M. 10 Heidenreich, L.M, Torroba 10 Dymarsky, Martucci 10

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$$V_{\text{DBI+CS}}\Big|_{G_3} = V_F\Big[\mathcal{K}(T,\overline{T}, \boldsymbol{z}, \boldsymbol{\bar{z}}), W_{\text{np}}(T, \boldsymbol{z})\Big]$$

Baumann, Dymarsky, Kachru, Klebanov, L.M. 10 Dymarsky, Martucci 10 Kachru, Kim, L.M., Zimet

The 4d F-term potential due to $W_{\rm np}(T, z)$ matches the 10d DBI+CS potential due to G_3 sourced by $\langle \lambda \lambda \rangle$.

Compelling evidence that gaugino condensation sources flux via $S_{G\lambda\lambda}$. We can compute the associated 10d stress-energy $T^{\lambda\lambda}_{\mu\nu}$.

Stress-energy $T^{\lambda\lambda}_{\mu\nu}$ due to

$$S_{G\lambda\lambda} = -\frac{i}{(4\pi^2\alpha')^2} \int \sqrt{g} e^{-4A} e^{\phi} G_3 \cdot \Omega \frac{\bar{\lambda}\bar{\lambda}}{16\pi} \delta^{(0)} + c.c.$$

is such that

$$\frac{1}{4}M_{\rm pl}^2 \mathcal{R}_4[g] = -3e^{\mathcal{K}}|W|^2 = V_{\rm KKLT}^{\rm AdS}$$

without anti-D3-branes

$$\frac{1}{4}M_{\rm pl}^2 \mathcal{R}_4[g] = V_{\overline{D3}} + V_F = V_{\rm KKLT}^{\rm dS}$$

with anti-D3-branes

Match is exact.

Hamada, Hebecker, Shiu, Soler 19 Gautason, Van Hemelryck, Van Riet, Venken 19 Carta, Moritz, Westphal 19 Kachru, Kim, L.M., Zimet

Upshot: consistency between 10d and 4d calculations.

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Match is exact.



Manki Kim

Upshot: consistency between 10d and 4d calculations.

KKLT status summary

No evidence of obstructions or inconsistencies.

Still awaiting complete explicit compactifications.

Heterotic string

Complete moduli stabilization at weak coupling difficult.

 $W = \int H \wedge \Omega + \mathcal{A} e^{-S}$

Issue: flux quantization.

Dine, Rohm, Seiberg, Witten 85

Gukov, Kachru, Liu, L.M. 03 Anderson, Gray, Lukas, Ovrut 11 Cicoli, de Alwis, Westphal 13 Apruzzi, Gautason, Parameswaran, Zagermann 14 Anderson, Gray, Lukas, Wang

Constructions global \Rightarrow non-modular.

No module for metastable SUSY breaking in de Sitter.



Classical flux compactifications already interesting.

Grimn, Louis 04 DeWolfe, Giryavets, Kachru, Taylor 05 Villadoro, Zwirner 05

 ${
m CY}_3$ with H_3, F_p $(p=0,2,4,6), {
m D6}, {
m O6}:$ no-go. Hertzberg, Kachru, Taylor, Tegmark 07

Negative curvature \Rightarrow no no-go, dS critical points.

Flauger, Paban, Robbins, Wrase 08 Caviezel, Koerber, Kors, Lust, Wrase, Zagermann 08 Danielsson, Haque, Shiu, Van Riet 09 Wrase, Zagermann 10 Danielsson, Haque, Koerber, Shiu, Van Riet, Wrase 11 Andriot 18

Nilmanifolds with fluxes, O-planes, KK-monopoles: dS vacuum. Silverstein 07

Problematic simplification: smearing of O-planes. Do solutions with localized (i.e. not smeared) O-planes exist?

See talk by Tomasiello

Conclusion

Considerable promise for exhibiting controlled de Sitter vacua in EFTs derived in solutions of string theory.

Steady progress, increasingly precise and explicit computations. Many highly nontrivial tests already passed.

No incontrovertible example to date.

Little prospect of de Sitter at *arbitrarily* weak coupling.

Learning to control non-SUSY solutions a central challenge.

