JT Gravity And Random Matrix Ensembles

Edward Witten

Strings 2019, Brussels

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For given X, we sum over all of the Y's, compatible with general principles.

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$$I = \int_{Y} \mathrm{d}^{2}x \sqrt{g} \phi(R+2) + \mathrm{topological}.$$

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Apparently a true quantum mechanical system is dual to something complicated, while a simple bulk system is dual to an average of quantum systems, not to a specific, bona fide quantum system.

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In the work of S^3 , it is important that JT gravity on an orientable manifold is not just 1-loop exact but tree-level exact, in the sense that the 1-loop correction is trivial. That is why everything reduces to computing volumes of moduli spaces, where the volumes are defined by classical formulas with no 1-loop correction.

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Time-reversal symmetry T is important in random matrix theory because it cannot be treated that way.

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- (3) H is an antisymmetric tensor of Sp(L).

In each case, we consider ${\cal H}$ as a random matrix with a distribution given by

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for a suitable function f. Each ensemble can be related to 2d geometry by expanding in perturbation theory and using 't Hooft's "ribbon graph" construction. The difference is that for a hermitian matrix, one gets orientable 2-manifolds only, but in the other cases, perturbation theory generates unorientable two-manifolds. That is because for the T-invariant ensembles, the propagator has a "twisted" term:

$$\left\langle H_{ij}H^{kl}\right\rangle = \delta_i^k\delta_j^l \pm \delta_j^k\delta_i^l.$$

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So we would like to compare the T-invariant Dyson ensembles to JT gravity on possibly unorientable surfaces, without or with the $(-1)^{n_c}$.

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Let us first describe what happens in the orientable case.

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Let us first describe what happens in the orientable case. Solutions of JT gravity are "hyperbolic surfaces," that is two-manifolds with constant negative curvature R = -2. Such a surface can be built by gluing together three-holed spheres:

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Let us first describe what happens in the orientable case. Solutions of JT gravity are "hyperbolic surfaces," that is two-manifolds with constant negative curvature R = -2. Such a surface can be built by gluing together three-holed spheres:



For each circle on which one glues, there are two moduli, namely a "length" parameter a and a "twist" parameter ρ .

In the orientable case, the measure on moduli space that comes from JT gravity is the classical expression

$$\mu = \prod_i \mathrm{d} \mathbf{a}_i \mathrm{d} \varrho_i$$

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that one can read off from the classical action at tree-level.

In the unorientable case, one can still make a hyperbolic two-manifold by gluing of simple building blocks, but one needs to allow a new kind of building block with one or two boundaries closed off by a cross-cap:

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Such a boundary still has a length parameter a, but it has no gluing parameter ρ .

$$\mu = \prod_{i} \mathrm{d}\boldsymbol{a}_{i} \mathrm{d}\varrho_{i} \prod_{\alpha} \frac{1}{2} \operatorname{coth} \frac{\boldsymbol{a}_{\alpha}}{4} \mathrm{d}\boldsymbol{a}_{\alpha},$$

where a_{α} are the length parameters of the cross-caps.

$$\mu = \prod_{i} \mathrm{d}\boldsymbol{a}_{i} \mathrm{d}\varrho_{i} \prod_{\alpha} \frac{1}{2} \operatorname{coth} \frac{\boldsymbol{a}_{\alpha}}{4} \mathrm{d}\boldsymbol{a}_{\alpha},$$

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where a_{α} are the length parameters of the cross-caps. The factor $\prod_{\alpha} \frac{1}{2} \coth \frac{a_{\alpha}}{4}$ is the 1-loop correction. (This measure was first obtained in another way by P. Norbury.) Note that there is a divergence in the volume (obtained by integrating the measure μ over moduli space) because of a da_{α}/a_{α} singularity for $a_{\alpha} \to 0$.

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$$\rho(E) = e^{S_0} \sinh \sqrt{E}$$

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that comes from the Schwarzian path integral. The JT path integral on such a manifold is supposed to compute a contribution to a correlator of matrix partition functions $\operatorname{Tr} e^{-\beta H}$ or (equivalently) matrix resolvents $\operatorname{Tr} \frac{1}{x-H}$.

To check whether this still works with T-invariance, we want to compare the 1/L or e^{-S_0} expansion of a matrix integral for a T-invariant Dyson ensemble to the JT path integral on an unorientable two-manifold.

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By expanding the "loop equations" of the matrix model to the first nontrivial order, i.e. the first order beyond the leading saddle point), one formally recovers precisely the same formula. This is a formal comparison because the integral diverges. (In the matrix model, one can regularize the divergence by cutting off the density of states $e^{S_0} \sinh \sqrt{E}$ at large *E*. A nice cutoff in JT gravity is not so obvious.)

A next step is to include fermions, with or without time-reversal, but without supersymmetry.

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as $(-1)^{\mathsf{F}}$, but we will not really have much time for that today.

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Instead of just including fermions (with or without T) it is more interesting to consider a supersymmetric version of the model. Thus now we hope to match super JT gravity to a supersymmetric version of random matrix theory. One preliminary question here is "what is JT supergravity?" A quick answer is this: JT gravity can be defined as *BF* theory of the group $SL(2, \mathbb{R})$, and analogously JT supergravity is *BF* theory of the supergroup OSp(1|2).

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The canonical form of a bifundamental of $U(L') \times U(L')$ is $C = \text{diag}(\lambda_1, \lambda_2, \cdots)$ where the λ_i can be assumed to be all *positive*, in contrast to Dyson ensembles where the random matrix has positive or negative eigenvalues.

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$$\mu = \prod_{i} \lambda_{i} \prod_{j < k} |\lambda_{j}^{2} - \lambda_{k}^{2}| \prod_{m=1}^{L'} \mathrm{d}\lambda_{m}.$$

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$$\mu = \prod_{i} \lambda_{i} \prod_{j < k} |\lambda_{j}^{2} - \lambda_{k}^{2}| \prod_{m=1}^{L'} \mathrm{d}\lambda_{m}.$$

To completely specify the model, we need to know the matrix potential $exp(-LTr f(C^{\dagger}C))$ or (more usefully) the corresponding saddle point which describes the limiting distribution of eigenvalues for large L'.

$$\rho(\lambda) \sim \frac{1}{\sqrt{\lambda}} \quad (\lambda \text{ small}).$$

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$$\rho(\lambda) = e^{S_0} \frac{\cosh \sqrt{\lambda}}{\sqrt{\lambda}}.$$

So this completes the specification of the model. (The same spectral curve, i.e. a model with the same density of states, has been studied by P. Norbury, with a different starting point than ours.)

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vanish in genus 0 (but not in higher orders). It turns out that this has a nice explanation in JT supergravity.

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Mirzakhani's basic idea was to build a surface with boundary Y by gluing a three-holed sphere Σ onto a simpler surface Y':

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Mirzakhani's basic idea was to build a surface with boundary Y by gluing a three-holed sphere Σ onto a simpler surface Y':



There are infinitely many choices of Σ , but there is a sum rule

$$1=\sum_{\Sigma}f(\Sigma),$$

where $f(\Sigma)$ is a certain function of the moduli of Σ . (The earliest version of this formula is due to McShane.)

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By inserting the identity $1 = \sum_{\Sigma} f(\Sigma)$ in the integral that defines the volume of the moduli space, Mirzakhani was able to get a recursion relation expressing the volume of the moduli space of hyperbolic structures on Y in terms of the analogous volume for a simpler surface Y'.

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All this has a superanalog: by imitating the purely bosonic proof, one can get an identity $1 = \sum_{\Sigma} f(\Sigma)$ in the supersymmetric context.

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We can also consider JT supergravity with time-reversal symmetry.

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We can also consider JT supergravity with time-reversal symmetry. It turns out that there are actually 8 variants of this theory because of the possibility to include a bulk topological field theory. These 8 models can be matched with 8 different random matrix ensembles (always with the same density of states that comes from the super-Schwarzian).

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In summary, I have tried to give an overview of the fact that the relation of JT gravity to a matrix model can be generalized to include time-reversal, fermions, and supersymmetry.