3d gravity as a random ensemble

Gabriel Wong

Oxford Math Institute

To appear soon with Dan Jafferis and Liza Rozenberg

Gabriel Wong (Oxford Math Institute)

3d gravity as a random ensemble

To appear soon with Dan Jafferis and Liza Rozenberg 1/27

< □ > < 同 > < 回 > < 回 >

• Semi-classical gravity averages over UV microstates. This is due to black hole solutions, whose horizon area gives the high energy density of states:

$$ho_0(E) = \exp{rac{\operatorname{Area}(E)}{4G}}$$

 This density of states is continuous because it is averaged over a high energy window. In AdS3, it corresponds to the Cardy density of the dual CFT.



Gravity and averaging II: Wormholes

• Gravity also computes ensemble averages. This is because of wormholes.

$$Z_{grav}\begin{pmatrix}\beta_1\\ 0\\ 0\\ \beta_2\end{pmatrix} = \beta_1 \bigcirc \beta_2 + \beta_1 \bigcirc \beta_2 + \cdots$$
$$Z_{CFT}\begin{pmatrix}\beta_1\\ 0\\ \beta_2\end{pmatrix} = Z_{CFT}(\beta_1)Z_{CFT}(\beta_2)$$

• The inclusion of wormhole topologies provides a coarse grained description in terms of an ensemble of quantum systems (see Monday's review talk).

$$\beta_1 \bigoplus \beta_2 + \beta_1 \bigoplus \beta_2 + \cdots = \overline{Z_{CFT}(\beta_1)Z_{CFT}(\beta_2)}$$

• These two types of coarse graining are related. For a chaotic system like a black hole, averaging over a high energy window is indistinguishable from averaging over an ensemble.

(日) (四) (日) (日) (日)

- Just like in statistical physics, the ensemble interpretation of quantum gravity provides the "best" description of a system given a set of constraints (de Boer)
- For instance, the micro-canonical ensemble maximizes entropy subject to the constraint of fixed energy.
- A random matrix model, describing a random Hamiltonian *H*, is the maximum entropy ensemble for a given average density of states.

$$\mathcal{Z} = \int dH \quad e^{-V(H)} \qquad V(H) \longleftrightarrow \overline{\rho(E)}_{\text{ensemble}}$$

・ 同 ト ・ ヨ ト ・ ヨ ト

JT gravity as a maximum ignorance ensemble

- Saad, Shenker, Stanford (SSS) showed that two dimensional JT gravity is equivalent to a double scaled matrix model.
- Inputs:

$$\overline{\rho(E)}_{disk} = E \qquad \overline{\rho(E)\rho(E')}_{cylinder} = E \qquad (1)$$

The cylinder determines the eigenvalue statistics associated to the symmetry class of the ensemble.

• This determines the gravity path integral on all 2D geometries via the machinery of topological recursion



[figure from SSS].

AdS3 gravity as a maximum ignorance ensemble

- We consider a generalization of the SSS model and study an ensemble of CFT2's dual to AdS3 gravity
- CFT2 data is given by the Dilatation operator Δ_s graded by spin, and the OPE coefficients C_{ijk} . Assuming integers spins, we get a set of random matrices and a random tensor:

$$(\Delta_s, C), \qquad s \in \mathbb{Z}$$

A (10) × A (10) × A (10)

AdS3 gravity as a maximum ignorance ensemble

- We consider a generalization of the SSS model and study an ensemble of CFT2's dual to AdS3 gravity
- CFT2 data is given by the Dilatation operator Δ_s graded by spin, and the OPE coefficients C_{ijk} . Assuming integers spins, we get a set of random matrices and a random tensor:

$$(\Delta_s, C), \quad s \in \mathbb{Z}$$

This random data satisfy locality constraints given by the modular bootstrap. This is generated by 4 point crossing equation on the sphere

$$\sum_{p} C_{12p} C_{p34} \left| \right\rangle_{1}^{2} \xrightarrow{p}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{q}^{2} - \sum_{q} C_{23q} C_{q41} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} = 0,$$

and torus modular invariance.

AdS3 gravity as a maximum ignorance ensemble

- We consider a generalization of the SSS model and study an ensemble of CFT2's dual to AdS3 gravity
- CFT2 data is given by the Dilatation operator Δ_s graded by spin, and the OPE coefficients C_{ijk} . Assuming integers spins, we get a set of random matrices and a random tensor:

$$(\Delta_s, C), \quad s \in \mathbb{Z}$$

This random data satisfy locality constraints given by the modular bootstrap. This is generated by 4 point crossing equation on the sphere

$$\sum_{p} C_{12p} C_{p34} \left| \right\rangle_{1}^{2} \xrightarrow{p}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{q}^{2} - \sum_{q} C_{23q} C_{q41} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} = 0,$$

and torus modular invariance.

• We will argue that the sum over topologies in AdS3 gravity produces the maximum ignorance ensemble consistent with the modular bootstrap.

(日) (四) (日) (日) (日)

Ensemble of approximate CFT's

2 The tensor potential and Virasoro TQFT

The matrix potential and SL(2, Z) modular invariance

Sum over all manifolds and the Schwinger Dyson equation

Ensemble of approximate CFT's

2 The tensor potential and Virasoro TQFT

3) The matrix potential and $\mathit{SL}(2,\mathbb{Z})$ modular invariance

Sum over all manifolds and the Schwinger Dyson equation

< (日) × < 三 × <

• An ensemble of exact CFT2's would have a potential V₀ defined by the Cardy density, and a delta function imposing the bootstrap constraints:

$$\mathcal{Z}_0 \equiv \sum_{s \in \mathbb{Z}} \int D[\Delta_s] D[C] e^{-V_0(\Delta_s)} \delta(ext{constraints})$$

• For irrational CFT's with only Virasoro symmetry, we don't know how to impose these constraints. Following (BdBJNS), we relax the constraints by smearing the delta function with a parameter \hbar

$$\mathcal{Z}_{\hbar} \equiv \sum_{s \in \mathbb{Z}} \int D[\Delta_s] D[C] \exp\left(-V_0(\Delta_s) - \frac{1}{\hbar}V(\Delta_s, C)\right)$$

- V(Δ_s, C) is a "constraint squared" potential that is minimized on solutions of the bootstrap.
 This defines an ensemble of *approximate* CFT's, with ħ parametrizing the violation of the bootstrap.
- Since $\hbar = 0$ localizes to exact CFT's, computing Z_{\hbar} and then taking $\hbar \to 0$ should teach us something about the space of CFT's

(日) (四) (日) (日) (日)

3d gravity from the ensemble of approximate CFT's

- We first define the ensemble as an integral over a finite tensor and matrices by truncating to N primaries. Then take a double scaling limit of Δ_s that sends $N \to \infty$
- Apriori, we want $\hbar \to 0$ at fixed central charge c. However, we will first re-organize the Feynman diagrams of the ensemble into an e^{-c} expansion, then take $\hbar \to 0$
- We will argue that the e^{-c} expansion of the ensemble reproduces the topological expansion of 3d gravity.



• Unifies many recent works on 3d gravity and random ensembles: Mertens/Turiaci,Collier/Maloney Maxfield/Tsaires, Cotler/Jensen, Belin/de Boer, Anous/Belin/de Boer/ Liska, Chandra/Collier/Hartman/ Maloney,Jafferis/Kolchmeyer/Mukhametzhanov/Sonner,Yan...



To appear soon with Dan Jafferis and Liza Rozenberg 11/2

э

イロト イヨト イヨト

Ensemble of approximate CFT's

2 The tensor potential and Virasoro TQFT

3) The matrix potential and $\mathit{SL}(2,\mathbb{Z})$ modular invariance

Sum over all manifolds and the Schwinger Dyson equation

▲ 同 ▶ → 三 ▶

• 4 point crossing = the vanishing of a vector.

$$\sum_{p} C_{12p} C_{p34} \left| \right\rangle_{1}^{2} \xrightarrow{p}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left| \right\rangle_{q}^{2} - \sum_{q} C_{23q} C_{q41} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left| \right\rangle_{1}^{2} = 0$$

We want an inner product to take its square.

• 4 point crossing = the vanishing of a vector.

$$\sum_{p} C_{12p} C_{p34} \left| \right\rangle_{1}^{2} \xrightarrow{p}{4} \left\langle \right\rangle \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left\langle \right\rangle - \sum_{q} C_{23q} C_{q41} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left\langle \right\rangle \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{\bar{q}} \left\langle \right\rangle = 0$$

We want an inner product to take its square.

• The space of Virasoro conformal blocks forms a Hilbert space equipped with the Verlinde inner product and a representation of crossing transformations.

• 4 point crossing = the vanishing of a vector.

$$\sum_{p} C_{12p} C_{p34} \left| \right\rangle_{1}^{2} \xrightarrow{p}{4} \left\langle \right\rangle \left| \right\rangle_{1}^{2} \xrightarrow{\bar{p}}{4} \left\langle \right\rangle - \sum_{q} C_{23q} C_{q41} \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{4} \left\langle \right\rangle \left| \right\rangle_{1}^{2} \xrightarrow{\bar{q}}{\bar{q}} \left\langle \right\rangle = 0$$

We want an inner product to take its square.

• The space of Virasoro conformal blocks forms a Hilbert space equipped with the Verlinde inner product and a representation of crossing transformations.



• This is the Hilbert space structure of Virasoro TQFT (Verlinde/Collier/Eberhardt/Mengyang).

Constraint squared potential

• The square of 4 point crossing defines a quartic potential

$$V_{4} = \sum_{i_{1}\cdots i_{4}} \sum_{p,q} \left(C_{i_{1}i_{2}p} C_{pi_{3}i_{4}} C_{qi_{2}i_{1}} C_{i_{4}i_{3}q} \right) \left| \left\langle \sum_{1}^{2} \frac{p}{4} \right\rangle \right|^{2} - \left(C_{i_{1}i_{2}p} C_{pi_{3}i_{4}} C_{i_{1}i_{4}q} C_{i_{3}i_{2}q} \right) \left| \left\langle \sum_{1}^{2} \frac{p}{4} \right\rangle \right|^{2} - \left(C_{i_{1}i_{2}p} C_{pi_{3}i_{4}} C_{i_{1}i_{4}q} C_{i_{3}i_{2}q} \right) \left| \left\langle \sum_{1}^{2} \frac{p}{4} \right\rangle \right|^{2}$$



< A > <

Constraint squared potential

• The square of 4 point crossing defines a quartic potential

$$V_{4} = \sum_{i_{1}\cdots i_{4}} \sum_{p,q} \left(C_{i_{1}i_{2}p} C_{pi_{3}i_{4}} C_{qi_{2}i_{1}} C_{i_{4}i_{3}q} \right) \left| \left\langle \sum_{1}^{2} \frac{p}{4} \right\rangle \right|^{2} - \left(C_{i_{1}i_{2}p} C_{pi_{3}i_{4}} C_{i_{1}i_{4}q} C_{i_{3}i_{2}q} \right) \left| \left\langle \sum_{1}^{2} \frac{p}{4} \right\rangle \right|^{2} \right|^{2}$$





• The kinetic term comes from a special case of the 6J:

$$C_{ii\mathbb{1}} = 1 \longrightarrow V_2 = -\sum_{ijk} C_{ijk} C_{jik}$$



Triple line Feynman diagrams and theirgravity interpretation

• The triple line Feynman rules are defined by removing the vertices from the graphs

$$\frac{i}{j} \underbrace{\frac{1}{j}}_{k} \frac{i}{k} = \hbar \delta_{il} \delta_{jm} \delta_{kp}, \qquad \frac{1}{j} \underbrace{\frac{1}{k}}_{i} = \frac{1}{\hbar} \frac{\delta(P_{p} - P_{q})}{S_{1P_{q}}}, \qquad \frac{1}{j} \underbrace{\frac{1}{k}}_{i} = \frac{1}{\hbar} \begin{cases} q & 4 & 3 \\ p & 2 & 2 \end{cases}$$

• We interpret these diagrams as multiboundary wormholes with Wilson lines inserted:



- They come from Wilson line networks in the three sphere, with solid balls removed around the junctions.
- These Feynman rules generate a sum over 3- manifolds. On a fixed manifold they agree with 3d gravity as defined by 2 copies of Virasoro TQFT .

$$Z_{\text{tensor}}(M) = |Z_{\text{Vir}}(M)|^2$$

Ensemble of approximate CFT's

2 The tensor potential and Virasoro TQFT

The matrix potential and SL(2, Z) modular invariance

Sum over all manifolds and the Schwinger Dyson equation

< (日) × < 三 × <



æ

< □ > < □ > < □ > < □ > < □ >

• The random Hamiltonians Δ_s belong to the GOE ensemble (Yan). The GOE measure exponentiates to the Vandermonde potential

$$K(\Delta_1, s_1; \Delta_2, s_2) = \delta(s_1 - s_2) \mathsf{Log} |\Delta_1 - \Delta_2|^1$$

• The "cylinder" contribution to $\overline{\rho\rho}$ is the inverse of the Vandermonde kernel. It is the propagator for ρ .

$$\overline{\rho(\Delta_1, s_1)\rho(\Delta_2, s_2)}_{cylinder} = \mathcal{K}^{-1}(\Delta_1, s_1; \Delta_2, s_2) =$$

It agrees with the 3d gravity path integral on the torus wormhole computed in (Cotler/Jensen 20), up to an extra factor of 2 due to extra wormholes needed to get agreement with the GOE ensemble (Yan, Jensen) .

(日) (四) (日) (日) (日)

Cardy density and the BTZ black hole

• The disk density for Δ_s is given by the Cardy formula. This is a product of Virasoro S matrix elements in the left and right moving sector.

$$\overline{\overline{h}})_{disk} = \mathbb{S}_{1h} \mathbb{S}_{1\overline{h}} =$$

$$egin{aligned} \Delta &= h + ar{h}, \quad s = h - ar{h} \in \mathbb{Z} \ \Delta &\geq s + rac{c-1}{12} \end{aligned}$$

 $\rho(h,$

< □ > < 同 > < 回 > < 回 >

Cardy density and the BTZ black hole

• The disk density for Δ_s is given by the Cardy formula. This is a product of Virasoro S matrix elements in the left and right moving sector.

$$\overline{\rho(h,\bar{h})}_{\text{disk}} = \mathbb{S}_{\mathbb{1}h} \mathbb{S}_{\mathbb{1}\bar{h}} =$$

$$Z_{\text{BTZ}}(\tau,\bar{\tau}) = \chi_{\mathbb{1}} \left(-\frac{1}{\tau}\right) \chi_{\mathbb{1}} \left(-\frac{1}{\bar{\tau}}\right) = \int_{\frac{c-1}{24}}^{\infty} \int_{\frac{c-1}{24}}^{\infty} dh d\bar{h} \ \mathbb{S}_{\mathbb{1}h} \mathbb{S}_{\mathbb{1}\bar{h}} \ \chi_{h}(\tau) \chi_{\bar{h}}(\bar{\tau})$$

(日) (四) (日) (日) (日)

Cardy density and the BTZ black hole

 The disk density for Δ_s is given by the Cardy formula. This is a product of Virasoro S matrix elements in the left and right moving sector.

$$Z_{\mathsf{BTZ}}(\tau,\bar{\tau}) = \chi_{\mathbb{1}}\left(-\frac{1}{\tau}\right)\chi_{\mathbb{1}}\left(-\frac{1}{\bar{\tau}}\right) = \int_{\frac{c-1}{24}}^{\infty} \int_{\frac{c-1}{24}}^{\infty} dh d\bar{h} \,\,\mathbb{S}_{\mathbb{1}h}\mathbb{S}_{\mathbb{1}\bar{h}}\chi_{h}(\tau)\chi_{\bar{h}}(\bar{\tau})$$

 $\bullet\,$ To get all spins, write $\sum_{s\in\mathbb{Z}}=\int ds\sum_{n=-\infty}^{\infty}e^{2\pi ins}$

(日) (四) (日) (日) (日)

Constraint squared potential for S-modular invariance I:

• The (approximate) CFT partition function is a vector on the Hilbert space ${\cal H}_{T^2}\otimes {\cal H}_{T^2}$ of Virasoro characters

$$\ket{Z(au,ar{ au})} = \sum_i \ket{h_i} \ket{ar{h}_i}$$

• Given an inner product, we can define a double trace potential for modular invariance

$$V_{\mathcal{S}} \equiv |(\mathbb{1} - \mathbb{S})|Z(au, ar{ au})
angle|^2 = \sum_{i,j} \langle h_i, ar{h}_i|\hat{V}_{\mathcal{S}}|h_j, ar{h}_j
angle$$

• We then expand in powers of the potential $\frac{1}{\hbar}V_S$:

$$\mathcal{Z} \equiv \sum_{n=-\infty}^{\infty} \int_{-\infty}^{\infty} ds \ e^{2\pi i n s} \int D\Delta_s D[C] \exp(-V_0(\Delta_s) - \frac{1}{\hbar} V_S)$$



• We need to find the gravitational inner product: it is not the one defined by VTQFT.

20 / 2

Gravity inner product: Torus wormhole

• The GOE Vandermonde K is the gravity inner product on $\mathcal{H}_{T^2} \otimes \mathcal{H}_{T^2}$

• The GOE Vandermonde K is the gravity inner product on $\mathcal{H}_{T^2}\otimes\mathcal{H}_{T^2}$



• Because 3d gravity is topological, this propagator is the identity on $\mathcal{H}_{T^2} \otimes \mathcal{H}_{T^2}$. \mathcal{K}^{-1} gives the resolution of identity in an non-orthogonal basis of Virasoro characters • The GOE Vandermonde K is the gravity inner product on $\mathcal{H}_{T^2} \otimes \mathcal{H}_{T^2}$



- Because 3d gravity is topological, this propagator is the identity on $\mathcal{H}_{T^2} \otimes \mathcal{H}_{T^2}$. \mathcal{K}^{-1} gives the resolution of identity in an non-orthogonal basis of Virasoro characters
- The non orthogonality comes from the the gauging of the bulk mapping class group.



 $\bullet\,$ The VTQFT inner product \longrightarrow orthogonal Virasoro characters. VTQFT has different random matrix statistics than gravity.

$SL(2,\mathbb{Z})$ modular invariance from sum over topologies

 Before adding V_S, we have a simple set of 3 manifolds obtained from the genus expansion of the double scaled matrices Δ_s. The expansion parameter the level spacing e^{-c}.



$SL(2,\mathbb{Z})$ modular invariance from sum over topologies

 Before adding V_S, we have a simple set of 3 manifolds obtained from the genus expansion of the double scaled matrices Δ_s. The expansion parameter the level spacing e^{-c}.



• Introducing V_S using the gravity inner product inserts S transforms into the propagation of a bulk toriodal slice

In the $\hbar \to 0$ limit this produces a projector onto the $\mathbb{S} = 1$ states \Rightarrow S modular invariance.

$SL(2,\mathbb{Z})$ modular invariance from sum over topologies

 Before adding V_S, we have a simple set of 3 manifolds obtained from the genus expansion of the double scaled matrices Δ_s. The expansion parameter the level spacing e^{-c}.



• Introducing V_S using the gravity inner product inserts S transforms into the propagation of a bulk toriodal slice



In the $\hbar \to 0$ limit this produces a projector onto the $\mathbb{S} = 1$ states \Rightarrow S modular invariance.

- Full $SL(2,\mathbb{Z})$ modular invariance is achieved when combined with the sum over T transforms.
- These new wormholes produces Seifert manifolds that are needed to cure the negative density of states that would arise from the $SL(2,\mathbb{Z})$ sum over BTZ black holes (Maxfield-Turiaci).

22 / 2

Ensemble of approximate CFT's

2 The tensor potential and Virasoro TQFT

) The matrix potential and $\mathit{SL}(2,\mathbb{Z})$ modular invariance

Sum over all manifolds and the Schwinger Dyson equation

< (日) × < 三 × <

• Loops in the tensor model diagrams \leftrightarrow insertion of the Cardy density



• This implements toroidal surgery, which is equivalent to Moore-Seiberg identities



· Gluing 6J's with a relative rotation creates arbitrary braids



This allows us to build all 3 manifolds (Lickorish)

Schwinger Dyson and the sum over topologies

• Due to surgery relations, we obtain many tensor model diagrams with the same topology, weighted by different powers of \hbar .

$$\langle C \cdots C \rangle = \sum_{M} Z_{\text{grav}}(M, c) f(M, \hbar).$$
 $Z_{\text{grav}}(M, c) \sim e^{-c \text{Vol}M}$

• To match with 3d gravity, we want $f(M,\hbar)
ightarrow 1$ as $\hbar
ightarrow 0$

・ 同 ト ・ ヨ ト ・ ヨ ト

• Due to surgery relations, we obtain many tensor model diagrams with the same topology, weighted by different powers of \hbar .

$$\langle C \cdots C \rangle = \sum_{M} Z_{\text{grav}}(M,c) f(M,\hbar).$$
 $Z_{\text{grav}}(M,c) \sim e^{-c \text{Vol}M}$

- $\bullet\,$ To match with 3d gravity, we want $f(M,\hbar) \to 1$ as $\hbar \to 0$
- We can check this conjecture using the $\hbar \to 0$ limit of the Schwinger Dyson equation, e.g.



• For a single manifold on the LHS, we have a finite counting problem



25.

- We showed how the topological expansion of 3d gravity arises from a random ensemble of approximate CFT's.
- The sum over topologies implements the bootstrap constraints order by order in the e^{-c} expansion
- A successful non-perturbative completion of the sum would solve the bootstrap equations exactly, and land us in the space of exact CFT's. We expect this is a $e^{-e^{c}}$ question.
- The matrix model sums over a simpler class of 3 manifolds: perhaps we should start by understanding the non-perturbative completion of this sum.

- 4 回 ト 4 三 ト

- A matrix model prediction: 3 boundary torus wormhole
- Some loose ends:
 - Solve the Schwinger-Dyson equation : generalization of topological recursion?
 - renormalization /cancellation of S^2 handle divergences
 - Regularization of the SL(2, Z) accumulation point = hyperbolic cusp
- Future work:
 - Random BCFT and EOW branes
 - Adding matter
 - Random ensemble for dS3?
 - Random ensemble of RCFT's

< 回 > < 三 > < 三 >