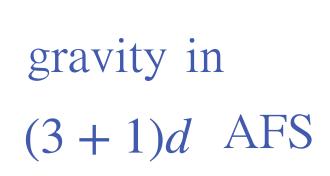
An infrared boundary action in asymptotically flat spacetimes

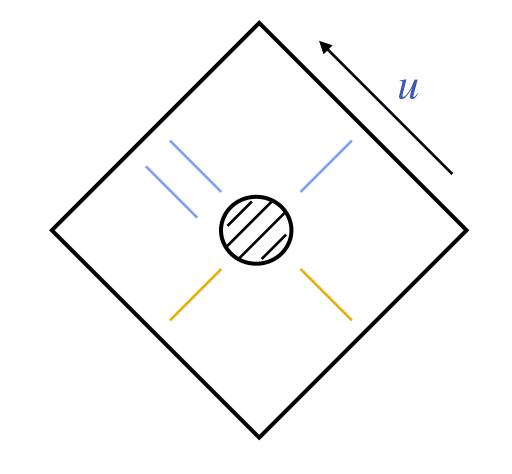
Ana-Maria Raclariu King's College London

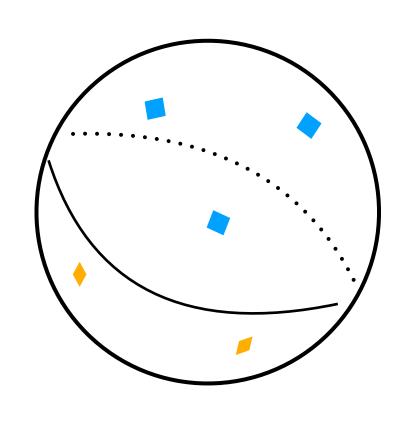
Strings 2025, NYU Abu Dhabi

based on 2408.01485 with T. He and K. Zurek [see also 2305.14411]

Flat space holography from the bottom up?







2d (C)CFT (or 3d Carrollian FT)

• Soft graviton modes

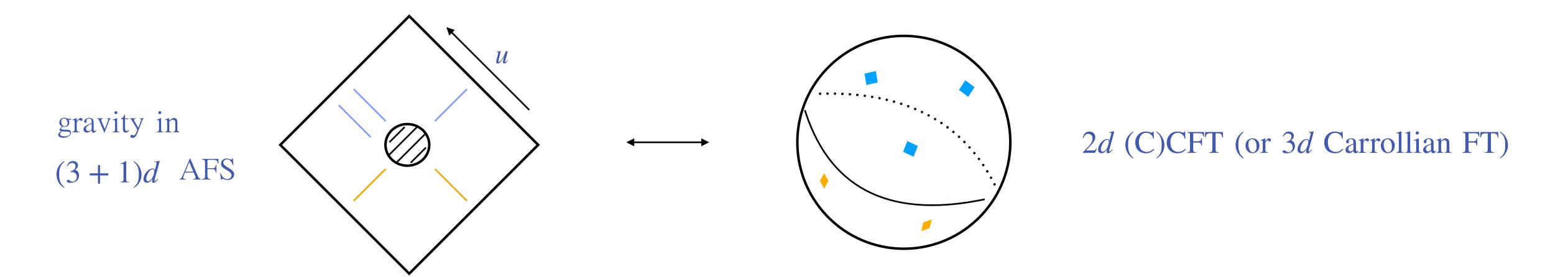
w-algebra generators

[..., Guevara, Himwich, Pate, Strominger '21]

$$\int_{-\infty}^{\infty} du u^n N_{zz}(u, z, \bar{z})$$

$$\longrightarrow w^n(z,\bar{z}) = \mathcal{L}_{2d} \left[\int_{-\infty}^{\infty} du u^n N_{zz}(u,z,\bar{z}) \right]$$

Flat space holography from the bottom up?



What about other entries in the holographic dictionary?

1. In AdS/CFT, the AdS partition function with specified boundary conditions acts as the generating function for correlators of primary operators in CFT:

$$Z_{\text{AdS}}[\phi_0, \cdots] = \langle \exp\left(\int \phi_0 \mathcal{O} + \cdots\right) \rangle_{\text{CFT}}$$

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 non-normalizable mode = source for CFT primary \mathcal{O}

classical (on-shell) gravity action
$$\longrightarrow$$
 quantum CFT correlator: $\langle \mathcal{O}(x_1) \cdots \rangle = \frac{\delta}{\delta \phi_0(x_1)} \cdots Z_{\text{AdS}}[\phi_0, \cdots]|_{\phi_0 = 0}$

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$$= \text{classical (on-shell) gravity action} \longrightarrow \text{quantum CFT correlator: } \langle \mathcal{O}(x_1) \cdots \rangle = \frac{\delta}{\delta \phi_0(x_1)} \cdots Z_{\text{AdS}}[\phi_0, \cdots] \big|_{\phi_0 = 0}$$

Q1: Is there an analogous relation in asymptotically flat spacetimes?

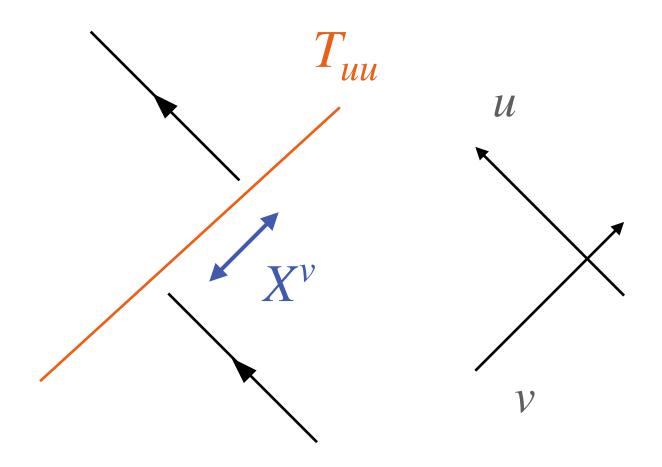
2. Proposed observability of vacuum spacetime fluctuations:

[Verlinde, Zurek '19, '22]

$$\hat{K}(z) = \int du X^{u}(z) T_{uu}(u, z) = X^{u}(z) \square X^{v}(z)$$

$$\hat{K} \equiv \int_{S^2} d^2z \hat{K}(z)$$
 (bulk) area operator bilinear in X^u, X^v

[...,Jafferis, Lewkowycz, Maldacena, Suh '15]



$$ds^{2} = h_{uu}du^{2} + ds_{\text{flat}}^{2}$$

$$\Box h_{uu} \propto T_{uu}, \quad X^{v} \propto \int du h_{uu}$$

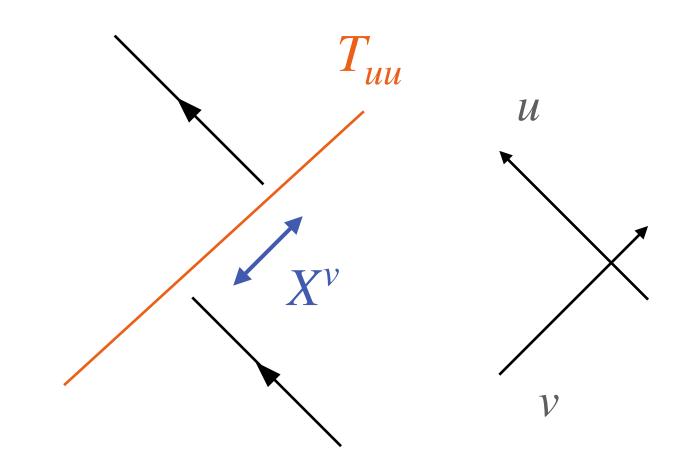
[Dray, 't Hooft '85]

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Assume that:

$$[X^{u}(z), X^{v}(z')] \propto G_N \delta^{(2)}(z, z') \implies$$

['tHooft '96]

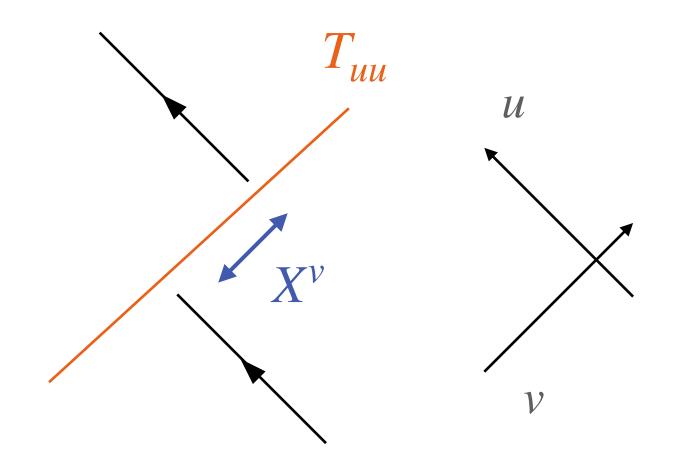
$$\langle \Delta \hat{K}^2 \rangle = \langle \hat{K}^2 \rangle - \langle \hat{K} \rangle^2 \sim A \delta^{(2)}(0) \rightarrow \frac{A}{\ell_p^2} \implies \frac{\Delta A^2}{\ell_P^4} \sim \frac{A}{\ell_P^2} \iff \Delta L^2 \sim L \ell_P$$

2. Proposed observability of vacuum spacetime fluctuations:

[Verlinde, Zurek '19, '22]

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In contrast, graviton fluctuations $\langle hh \rangle \propto \ell_P^2$... [..., Carney, Karydas, Sivaramakrishnan'24]

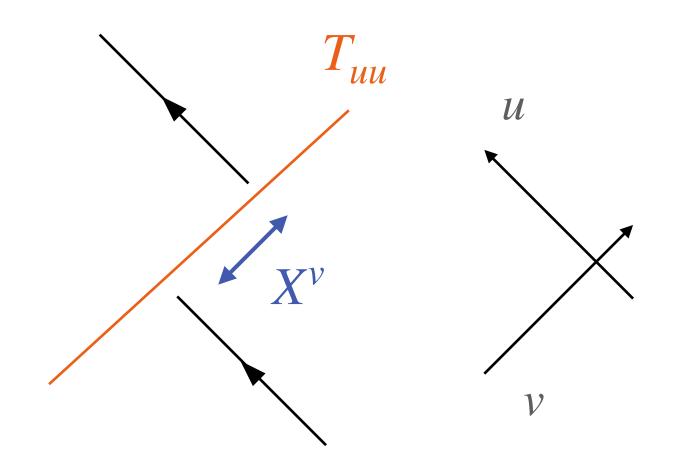
Q2: Can we define \hat{K} in 4d AFS and compute its fluctuations?

2. Proposed observability of vacuum spacetime fluctuations:

[Verlinde, Zurek '19, '22]

$$\hat{K}(z) = \int du X^{u}(z) T_{uu}(u, z) = X^{u}(z) \square X^{v}(z)$$

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In contrast, graviton fluctuations $\langle hh \rangle \propto \ell_P^2 \dots$ [..., Carney, Karydas, Sivaramakrishnan'24]

Hint: Shockwave solutions to EE related to Minkowski + gravitational memory by gauge transformation

Main results & outline

1. On-shell action in the IR sector of 4d AFS as generating function for CFT₂ current correlators

2. A relation between \hat{K} and the soft supertranslation charge

Main results & outline

- 1. On-shell action in the IR sector of 4d AFS as generating function for CFT₂ current correlators
- 2. A relation between \hat{K} and the soft supertranslation charge

- Soft modes in 4d AFS
- Boundary (on-shell) action in the (leading) soft sector of 4d AFS
- Soft charge fluctuations

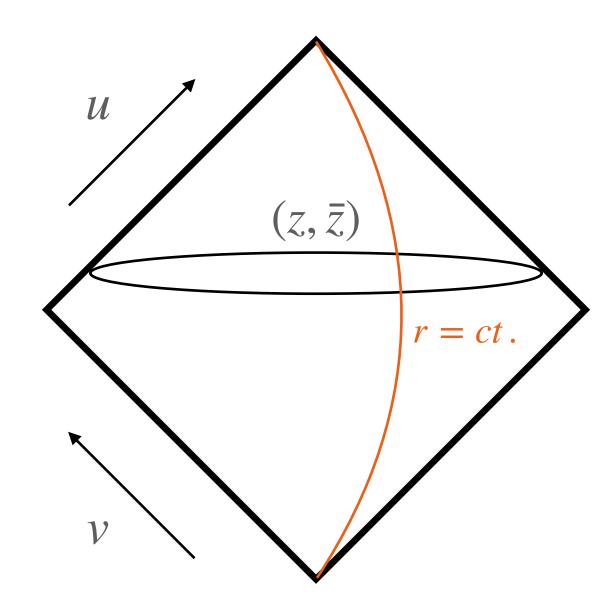
Asymptotically Flat Spacetimes (AFS)

Solutions to the Einstein eq. at large r:

$$ds^{2} = ds_{\text{Mink}}^{2} + rC_{zz}dz^{2} + rC_{\bar{z}\bar{z}}d\bar{z}^{2} + \frac{2m_{B}}{r}du^{2} + \cdots$$

 $C_{zz}(u, z, \bar{z})$: shear, free data

 $N_{zz}(u, z, \bar{z}) = \partial_u C_{zz}(u, z, \bar{z})$: news (gravitational radiation)



Equations of motion determine the subleading components in terms of C_{zz} , eg.:

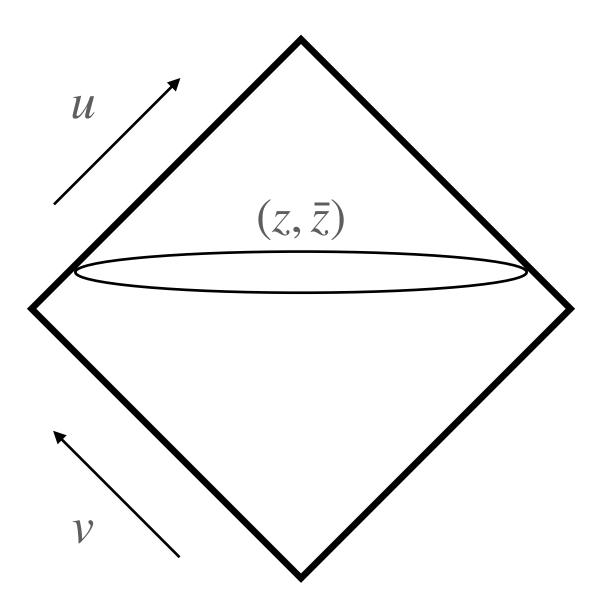
$$G_{uu} = 0 \implies \partial_u m_B = \frac{1}{4} \left(D_z^2 N_{\bar{z}\bar{z}} + D_{\bar{z}}^2 N_{zz} \right) - \frac{1}{4} N_{zz} N_{\bar{z}\bar{z}}$$

Restrict to shear profiles of the form:

$$C_{zz} = -2D_z^2 C(z, \bar{z}) + D_z^2 N(z, \bar{z})\Theta(u - u_0)$$

- $C(z, \bar{z})$ labels the vacuum degeneracy
- $N(z, \bar{z})$ is related to a soft graviton:

$$\int du N_{zz} = D_z^2 N = \Delta C_{zz}, \quad N_{zz} \equiv \partial_u C_{zz}$$

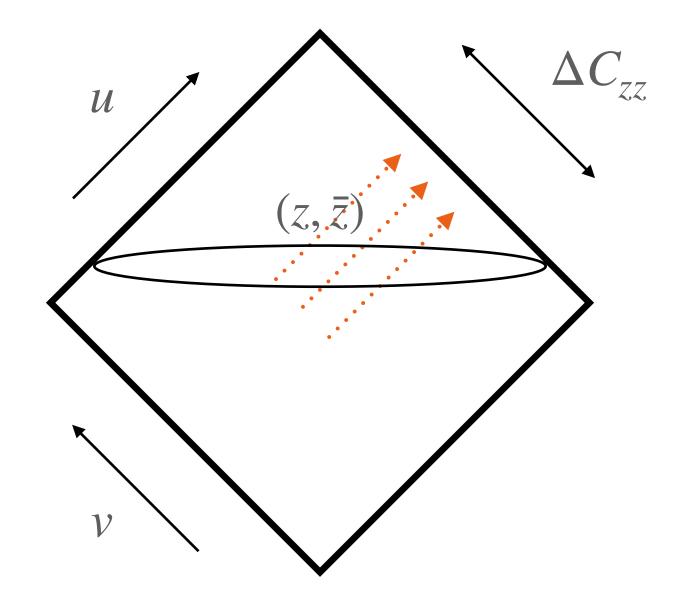


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 \leftrightarrow gravitational memory



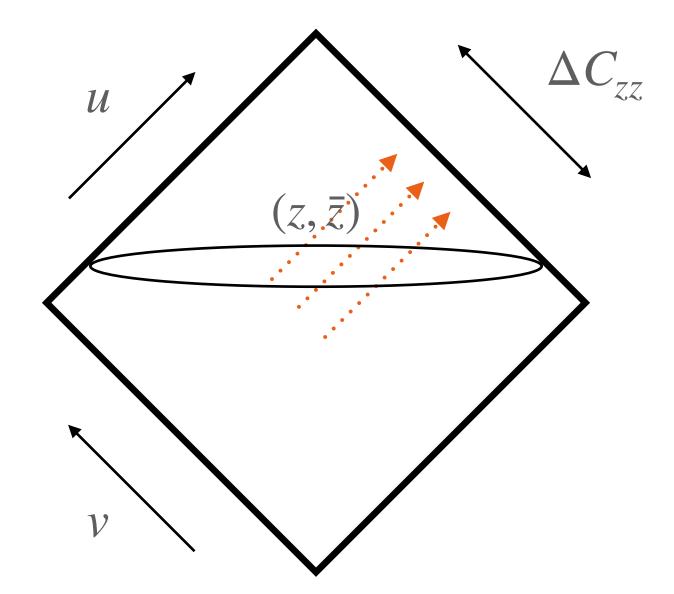
[Strominger, Zhiboedov '14]

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 \leftrightarrow gravitational memory [Strominger, Zhiboedov '14]



C and N are canonically paired variables of the leading IR sector of 4d AFS

Variation of Lagrangian form:

$$\delta L = EOM \cdot \delta g + d\theta(g, \delta g)$$

[lyer, Wald '94]

Einstein-Hilbert action evaluated on asymptotically flat solution labeled by the soft shear:

$$\Theta(g,\delta g) = \frac{1}{32\pi G} \int du \int d^2z \left[4\delta m_B + \frac{1}{4} \delta \left(\partial_z^2 N \partial_{\bar{z}}^2 N \right) \delta(u - u_0) + \frac{1}{2} \delta \left(\partial_z^2 C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \delta(u - u_0) \right]$$
$$-\frac{1}{16\pi G} \int du \int d^2z \left(\partial_z^2 \delta C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \delta(u - u_0)$$

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[Iyer, Wald '94]

Einstein-Hilbert action evaluated on asymptotically flat solution labeled by the soft shear:

$$\delta(\cdots) \implies \delta^2(\cdots) = 0$$

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$$-\frac{1}{16\pi G} \int du \int d^2z \left(\partial_z^2 \delta C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \delta(u - u_0) \implies \delta\Theta(g, \delta g) = -\frac{1}{16\pi G} \int d^2z \left(\partial_z^2 \delta C \wedge \partial_{\bar{z}}^2 \delta N + h \cdot c \cdot \right)$$

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symplectic form

⇒ C and N are canonically paired with bracket:

$$[N(z,\bar{z}),C(w,\bar{w})] = 4iG|z-w|^2\log|z-w|^2$$
 [He, Lysov, Mitra, Strominger '14]

Einstein-Hilbert action evaluated on asymptotically flat solution labeled by the soft shear:

$$\begin{split} \Theta(g,\delta g) &= \frac{1}{32\pi G} \int du \int d^2z \left[4\delta m_B + \frac{1}{4} \delta \left(\partial_z^2 N \partial_{\bar{z}}^2 N \right) \delta(u - u_0) + \frac{1}{2} \delta \left(\partial_z^2 C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \delta(u - u_0) \right] \\ &- \frac{1}{16\pi G} \int du \int d^2z \left(\partial_z^2 \delta C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \delta(u - u_0) \quad \Longrightarrow \quad \delta\Theta(g,\delta g) = -\frac{1}{16\pi G} \int d^2z \left(\partial_z^2 \delta C \wedge \partial_{\bar{z}}^2 \delta N + h \cdot c \cdot \right) \delta(u - u_0) \end{split}$$

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$$- \frac{1}{16\pi G} \int du \int d^2z \left(\partial_z^2 \delta C \partial_{\bar{z}}^2 N + h.c. \right) \delta(u - u_0) + \text{(contribution from } \mathcal{I}^-)$$

In order for variational principle to be well defined:

1. Add boundary action to cancel the first line

$$S^{\rm IR} = S_{\rm EH} + S_{\rm bdry.}, \quad S_{\rm bdry.} = -\frac{1}{32\pi G} \int d^2z \left[4 \int du m_B + \frac{1}{4} \partial_z^2 N \partial_{\bar{z}}^2 N + \frac{1}{2} \left(\partial_z^2 C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \right]$$

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since N = 0 would otherwise set the gravitational memory to 0

$$S_{\text{on-shell}}^{\text{IR}} = -\frac{1}{32\pi G} \int d^2z \left[4 \int du m_B + \frac{1}{4} \partial_z^2 N \partial_{\bar{z}}^2 N + \frac{1}{2} \left(\partial_z^2 C \partial_{\bar{z}}^2 N + h \cdot c \cdot \right) \right]$$

- m_B dependence can be traded for $\mathcal{O}(N^2)$ term by constraint, coefficient is IR divergent (related to Weinberg soft S-matrix?) [Weinberg '65]
- Linearize in N (keeping the C background fixed)
- $S_{\text{on-shell}}^{\text{IR}}$ becomes gauge invariant after adding a corner term $\propto \int d^2z \partial_{\bar{z}}^2 N \partial_z C$

• Also including the contribution from \mathcal{F}^- : $S^{\rm IR} = -\frac{1}{8\pi G} \int d^2z \partial_{\bar{z}}^2 N \partial_z^2 C$

[Donnelly, Freidel '16]

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- Also including the contribution from \mathcal{F}^- : $S^{IR} = -\frac{1}{8\pi G} \int d^2z \partial_{\bar{z}}^2 N \partial_z^2 C$ a) generating function for Kac-Moody current correlators
 - b) soft supertranslation charge

a) Current correlators in 2d CFT

Associate with $S_{\text{on-shell}}^{\text{IR}}[N, C]$ the partition function:

$$Z[C] \equiv \int [dN] e^{-S_{\text{on-shell}}^{\text{IR}}[N,C]}$$

- Z[C] is the generating function for 2d correlators of Kac-Moody (supertranslation) currents (these are dual to soft gravitons in 4d) [He, Lysov, Mitra, Strominger '14]
- The on-shell action agrees with the action derived from a 2d effective action of Goldstone modes after integrating in the memory/soft mode [Himwich, Narayanan, Pate, Paul, Strominger '20; Kapec, Mitra '21]
- Potentially related to partition function for edge modes

[Donnelly, Wall '14, ...Chen, A.R., Myers '24]

b) Soft supertranslation charge

Supertranslation charge in 4d AFS:

$$Q_{f=C} = \frac{1}{4\pi G} \int d^2z C(z,\bar{z}) m_B(z,\bar{z})$$

$$= -\frac{1}{16\pi G} \int_{\mathcal{J}^+} du d^2z C(z,\bar{z}) \left(\partial_z^2 N_{\bar{z}\bar{z}} + h.c.\right) + Q_H$$

$$Q_S$$

$$N_{zz} = \partial_z^2 N \delta(u - u_0)$$

$$\Longrightarrow \qquad Q_S = -\frac{1}{8\pi G} \int d^2 z \partial_z^2 C \partial_{\bar{z}}^2 N = S^{IR} + \mathcal{O}(N^2)$$

• C is a function

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- C is a function
- Promote C to an operator (recall $\{C, N\} \neq 0$):

$$\hat{Q}_S = -\frac{1}{8\pi G} \int d^2z \partial_z^2 \hat{C} \partial_{\bar{z}}^2 \hat{N}$$

Soft charge fluctuations?

When $m_B = 0$ (but in the presence of matter) the supertranslation charge vanishes and $Q_S + Q_H = 0$

This is condition is equivalent to the $G_{uu} = T_{uu}^{M}$ constraint:

$$\delta(u) \square h(z,\bar{z}) = T_{uu}^{M}(u,z,\bar{z}) \implies \underbrace{\int d\Omega f(\Omega) \square h}_{-Q_{S}} = \underbrace{\int du \int d\Omega f(\Omega) T_{uu}^{M}}_{Q_{H}=K}$$

• In vacua with $m_B = 0$, $\hat{Q}_S \propto \hat{K} \implies \langle \Delta \hat{Q}_S^2 \rangle \propto \langle \Delta \hat{K}^2 \rangle \sim \frac{A}{\ell_p^2}$ [provided that $f \to \hat{C}$]

Summary

- The on-shell action for the (leading) soft sector of gravity in (3+1)d AFS is non-trivial
- Fixing the zero mode of the shear (Goldstone), we recover the generating function for 2d correlators of Kac-Moody currents dual to 4d soft gravitons
- The boundary action is related to the soft component of the supertranslation charge
- Allowing for the Goldstone to fluctuate leads to a new soft charge operator whose fluctuations are enhanced by an IR scale (related to the area of the celestial sphere)

Outlook

- Why should C be promoted to an operator? Finite size causal diamonds?
- Relation to edge modes and entanglement in gravitational theories in AFS?
- How could vacuum spacetime fluctuations be experimentally probed? Can we measure \hat{Q}_S ?

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Thank you!

Extra slides

Current correlators in 2d CFT

Weinberg soft S-matrix $S = e^{-\Gamma}$, $\Gamma \propto \log \frac{\Lambda_{\rm UV}}{\Lambda_{\rm IR}} \int d^2z J_{ab} J^{ab}$ reproduced from 2d effective action of Goldstones:

[Kapec, Mitra '21]

$$e^{-\Gamma} = \int [dC] e^{-S_{\rm soft}[C] - S_{\rm int}[C,j]} \implies S_{\rm soft}[C] \propto \int d^2z \widetilde{C}_{ab}(z,\bar{z}) \widetilde{C}^{ab}(z,\bar{z}), \quad C_{zz} = -2\partial_z^2 C \quad \text{pure gauge}$$

$$S_{\rm int} \propto \int d^2z J_{ab} \widetilde{C}^{ab}$$

- In the presence of external soft gravitons, integrate in soft graviton N and further demand that $\int [dN]e^{-S'_{\rm soft}[N,C]} = e^{-S_{\rm soft}[C]} \equiv Z[C]$
- $S'_{\text{soft}}[N, C]$ agrees (at least to linear order in N) with the on-shell action we found!